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International Centre for Theoretical Physics**



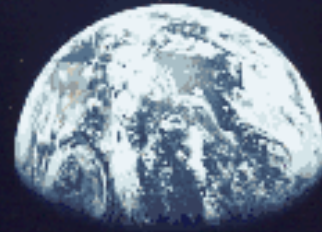
1930-1

**Joint ICTP-IAEA Advanced Workshop on Model Codes for Spallation
Reactions**

4 - 8 February 2008

**Experimental Cross Sections for the Production of Residual Nuclides at Medium
Energies.**

Rolk Michel
*University of Hannover
Department of Physics
Germany*



**Experimental Cross Sections for the Production of
Residual Nuclides at Medium Energies:
Status, Recent Progress, and Challenges for Modeling**

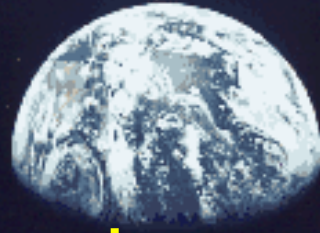
**R. Michel,
D. Hansmann, Shams A.M. Issa, M. Tutuc**

**Joint ICTP-IAEA Advanced Workshop on Model Codes for Spallation Reactions,
Trieste / February 4 – 8, 2008**

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- **Relevance of data for the production of residual nuclides at medium energies**
- **The 1997 NEA intercomparison**
- **Modeling of cosmogenic nuclides in extraterrestrial matter**
- **Candidate data for testing the performance of model codes**
 - ❖ **Thin-target data**
 - **proton-induced reactions**
 - **alpha-induced reactions**
 - **neutron-induced reactions**
 - ❖ **Thick-target data**
- **conclusions**



➤ **nuclear physics**

- systematics of nuclear reactions
- equilibrium and pre-equilibrium reactions
- spallation, fragmentation
- medium energy fission
- CN-, PE- and INC-models

➤ **accelerator technology**

- activation of detectors
- radiation protection
- on-line mass separation
- radioactive ion beams

➤ **dosimetry**

- mixed nucleon fields
- medicine

➤ **radionuclide production**

- radiation therapy

➤ **space and aviation technology**

- radiation protection
- material damage

➤ **astrophysics**

- abundance of heavy CR particles
- T-Tauri and WR stars
- p-process nucleosynthesis

➤ **geo- and environmental physics**

- cosmogenic nuclides as natural tracers in geology, archeometry, climatology, hydrology, glaciology

➤ **planetology**

- remote sensing of planetary surfaces

➤ **cosmophysics and -chemistry**

- cosmic ray exposure history of extraterrestrial matter
- terrestrial ages of meteorites
- variations of cosmic radiation with space and time

➤ **accelerator driven systems (ADS)**

- waste transmutation
- energy amplification
- spallation neutron sources

Production Rate P_i of a Residual Nuclide i at Medium Energies

depth in the target size of the target spectral fluences

cross sections

$$P_i(d, R, \vec{c}_s, \vec{c}_b) = N_L \cdot \sum_j \frac{c_{s,j}}{A_j} \cdot \sum_k \int_0^\infty \sigma_{i,j,k}(E_k) \cdot J_k(E_k, d, R, \vec{c}_b) dE_k$$

sample chemistry bulk target chemistry target elements particle types

The diagram illustrates the production rate equation for a residual nuclide i at medium energies. The equation is:

$$P_i(d, R, \vec{c}_s, \vec{c}_b) = N_L \cdot \sum_j \frac{c_{s,j}}{A_j} \cdot \sum_k \int_0^\infty \sigma_{i,j,k}(E_k) \cdot J_k(E_k, d, R, \vec{c}_b) dE_k$$
 The parameters and their meanings are:

- d : depth in the target (black arrow)
- R : size of the target (black arrow)
- \vec{c}_s : sample chemistry (black arrow)
- \vec{c}_b : bulk target chemistry (black arrow)
- N_L : target elements (red arrow)
- $c_{s,j}$: sample chemistry (black arrow)
- A_j : target elements (red arrow)
- $\sigma_{i,j,k}(E_k)$: cross sections (red arrow)
- $J_k(E_k, d, R, \vec{c}_b)$: spectral fluences (yellow arrow)
- k : particle types (yellow arrow)

Production of Residual Nuclides by Intermediate-Energy Particles

primary particles: p, ^4He

secondary particles: p, n, ^2H , ^3H , ^3He , ^4He , ...

up to 200 MeV/A: secondary particles negligible

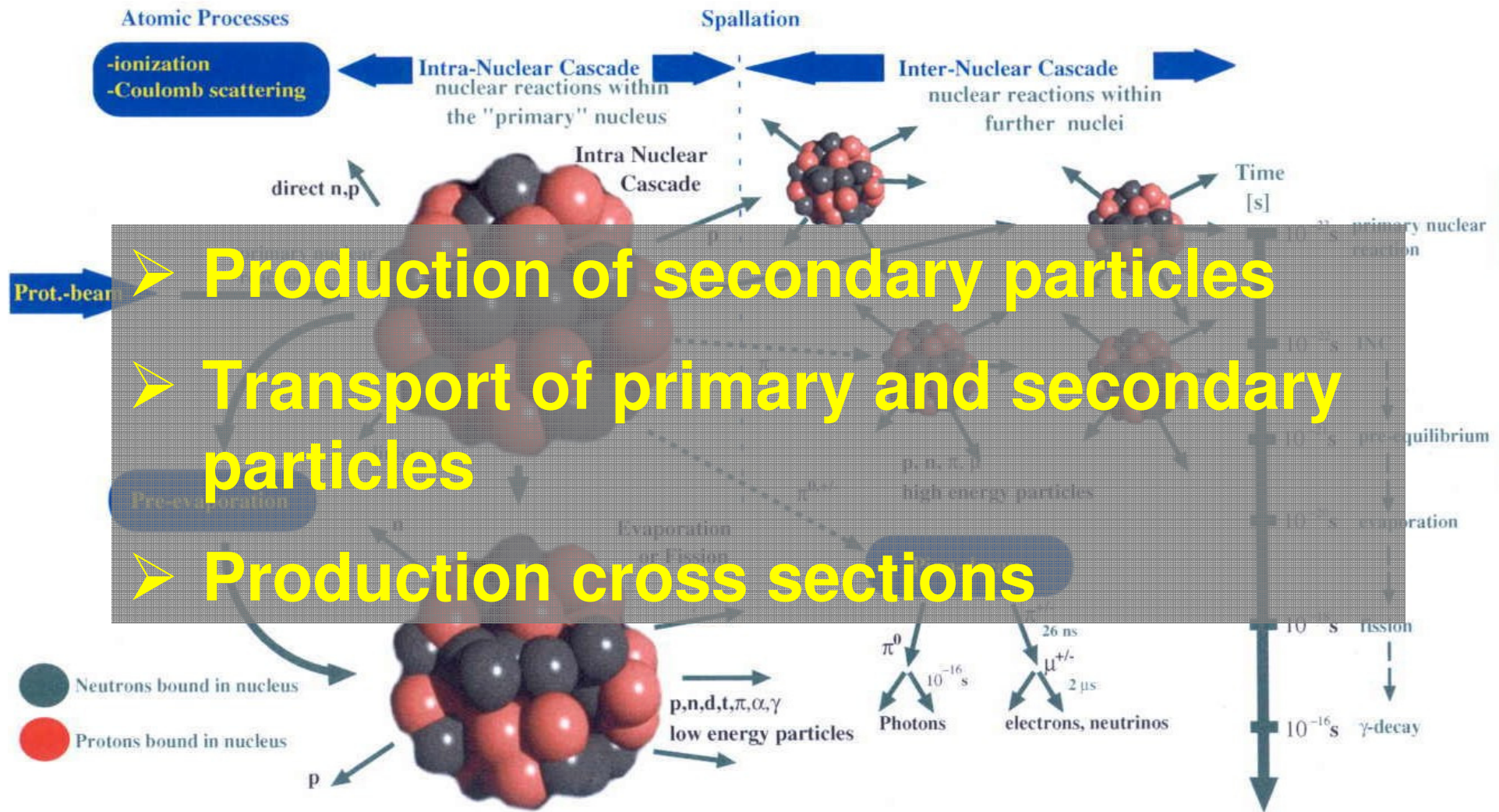
> 200 MeV up to 10 GeV/A: secondary particles dominant

Ockham's razor: Use as few particle types as possible!

relevant energies: 1 MeV - 10 GeV for protons

1 meV - 200 MeV for neutrons

Essentials of Modeling Residual Nuclide Production at Intermediate Energies



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Specification of the 1997 NEA International Codes and Model Intercomparison for Intermediate Energy Activation Yields

O-16 Be-7, Be-10, C-11, C-14

Al-27 H-3, He-3, He-4, Be-7, Be-10, Na-22 (Mg-22), Na-24 (Ne-26), (Si-26)

Fe (nat) H-3, He-3, He-4, Be-7, Be-10, Ne-20 (all mass 20 nuclides), Ne-21 (all mass 21 nuclides), Ne-22 (F-22), Na-22 (Mg-22), Na-24 (Ne-24), Mg-28 (Na-28), Al-26 (Si-26), Cl-36, Ar-36 (K-36, Ca-36), Ar-38 (all mass 38 nuclides), Sc-46, V-48 (Cr-48), Cr-51 (Mn-51, Fe-51), Mn-52m+g (Fe-52), Mn-53 (Fe-53, Co-53), Mn-54, Fe-55 (Co-55, Ni-55), Co-56

Co-59 Co-56, Co-57, Co-58, Ni-56, Ni-57

Zr (nat) Be-7, Na-22 (Mg-22), Sc-46, V-48, (Cr-48), Cr-51 (Mn-51, Fe-51), Mn-54, Co-56 (Ni-56), Co-58, Co-60, Zn-65 (Ga-65, Ge-65), Ga-67 (Ge-67, As-67), Ge-69 (As-69, Se-69), As-71 (Se-71, Br-71, Kr-71) As-74, Se-75 (Br-75, Kr-75, Rb-75), Br-77 (Kr-77, Rb-77, Sr-77), Kr-78 (Br-78, Rb-78), Kr-79 (Rb-79, Sr-79), Kr-80 (Br-80, Rb-80, Sr-80, Y-80), Kr-81 (Rb-81, Sr-81, Y-81, Zr-81), Kr-82 (Br-82, Rb-82, Sr-82, Y-82, Zr-82), Kr-83 (all mass 83 nuclides), Kr-84 (Br-84, Se-84, Rb-84), Kr-85 (Se-85, Br-85), Kr-86 (Se-86, Br-86, Rb-86), Rb-83 (Sr-83, Y-83, Zr-83), Rb-84, Rb-86, Sr-82 (Y-82, Zr-82), Sr-83 (Y-83, Zr-83), Sr-85 (Y-85, Zr-85, Nb-85), Y-86 (Zr-86, Nb-86), Y-86m, Y-87 (Zr-87, Nb-87), Y-87m, Y-88 (Zr-88, Nb-88), Zr-86 (Nb-86), Zr-88 (Nb-88), Zr-89 (Nb-89), Zr-95 (Y-95), Nb-90, Nb-92m, Nb-95, Nb-95m, Nb-96

Au-197 Be-7, Na-22 (Mg-22), Na-24 (Ne-24), Sc-46, V-48 (Cr-48), Mn-54, Fe-59 (Mn-59), Co-56 (Ni-56), Co-58, Co-60, Zn-65 (Ga-65, Ge-65), As-74, Se-75 (Br-75, Kr-75, Rb-75), Rb-83 (Sr-83, Y-83, Zr-83), Rb-84, Rb-86, Sr-85 (Y-85, Zr-85, Nb-85), Y-87 (Zr-87, Nb-87), Y-88 (Zr-88, Nb-88), Zr-88 (Nb-88), Zr-89 (Nb-89), Zr-95 (Y-95), Nb-95 (Rb-95, Sr-95, Y-95, Zr-95), Tc-96, Ru-103 (Nb-103, Mo-103, Tc-103), Rh-102, Ag-105 (Cd-105, In-105), Ag-110m, Ag-110, Sn-113 (Sb-113, Te-113, I-113, Xe-113), Te-121 (I-121, Xe-121, Cs-121, Ba-121), Te-121m, Te-121m+g, Xe-127 (Cs-127, Ba-127, La-127), Ba-131 (La-131, Ce-131), Ce-139 (Pr-139, Nd-139, Pm-139, Sm-139), Eu-145 (Gd-145), Eu-147 (Gd-147, Tb-147), Eu-148, Eu-149 (Gd-149, Tb-149, Dy-149, Ho-149), Gd-146 (Tb-146), Gd-147 (Tb-147, Dy-147), Gd-149 (Tb-149, Dy-149, Ho-149), Gd-151 (Tb-151, Dy-151, Ho-151), Gd-153 (Tb-153, Dy-153, Ho-153), Tb-149 (Dy-149, Ho-149), Tb-151 (Dy-151, Ho-151), Tb-153 (Dy-153, Ho-153), Tm-165 (Y-165, Lu-165, Hf-165), Tm-166 (Y-166, Lu-166, Hf-166, Ta-166), Tm-167 (Y-167, Hf-167, Ta-167), Tm-168, Yb-166 (Lu-166, Hf-166, Ta-166), Yb-169 (Lu-169, Hf-169, Ta-169), Lu-169 (Hf-169, Ta-169), Lu-170 (Hf-170, Ta-170), Lu-171 (Hf-171, Ta-171), Lu-172 (Hf-172, Ta-172), Lu-173 (Hf-173, Ta-173), Hf-172 (Ta-172, W-172, Re-172), Hf-173 (Ta-173, W-173), Hf-175 (Ta-175, Re-175, Os-175), Re-181 (Os-181, Ir-181), Re-182 (Os-182, Ir-182, Pt-182), Re-183 (Os-183, Ir-183, Pt-183, Au-183), Os-182 (Ir-182, Pt-182, Au-182, Hg-182), Os-185 (Ir-185, Pt-185, Au-185, Hg-185), Os-191 (Re-191), Ir-185 (Pt-185, Au-185, Hg-185), Ir-186 (Pt-186, Au-186, Hg-186), Ir-187 (Pt-187, Au-187, Hg-187), Ir-188 (Pt-188, Au-188, Hg-188), Ir-189 (Pt-189, Au-189, Hg-189), Ir-190, Ir-192, Pt-188 (Au-188, Hg-188), Pt-191 (Au-191, Hg-191), Au-193 (Hg-193), Au-194 (Hg-194), Au-195 (Hg-195), Au-196, Hg-193, Hg-194, Hg-195, Hg-195m, Hg-197, Hg-197m

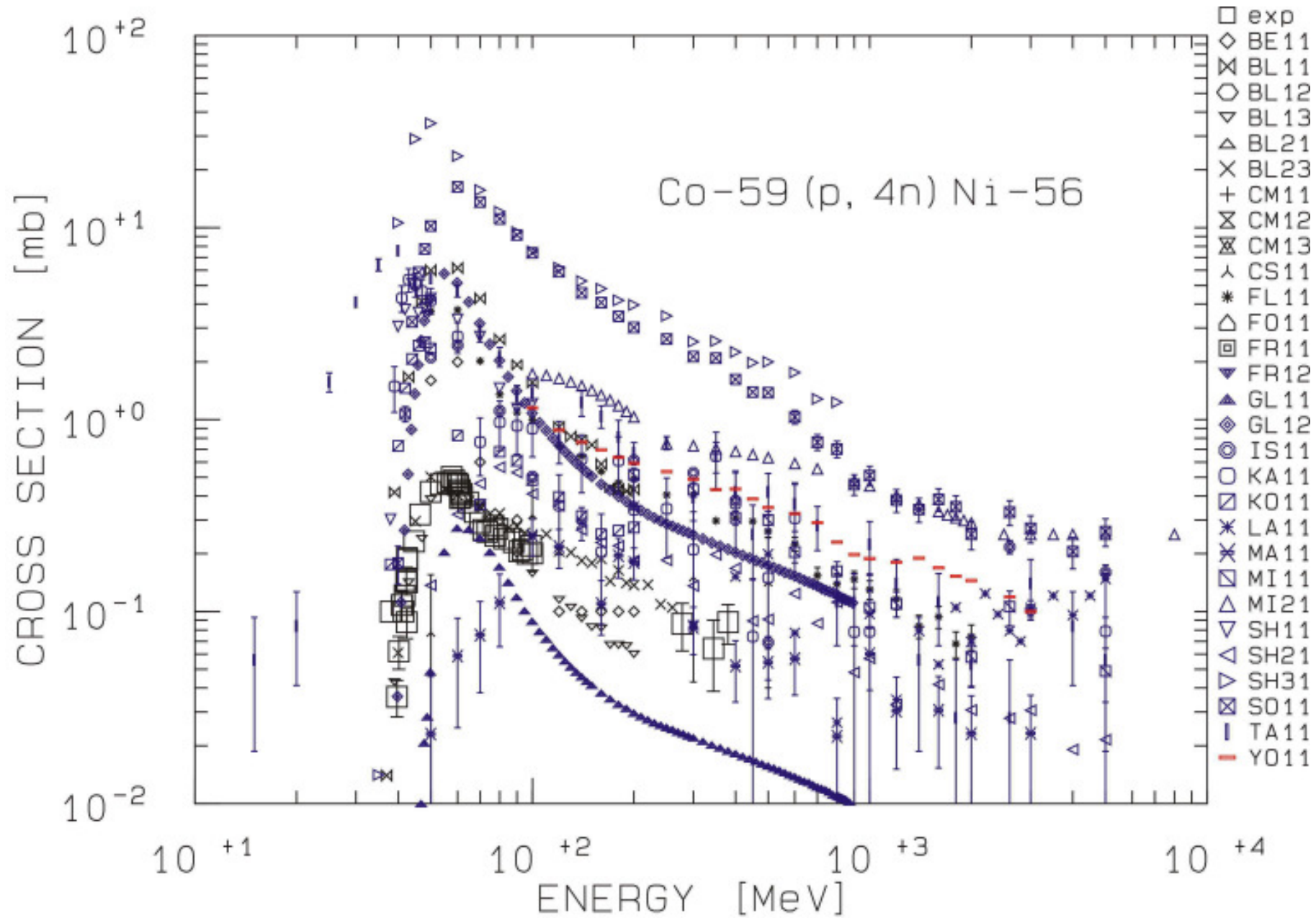
International Codes and Model Intercomparison for Intermediate Energy Activation Yields

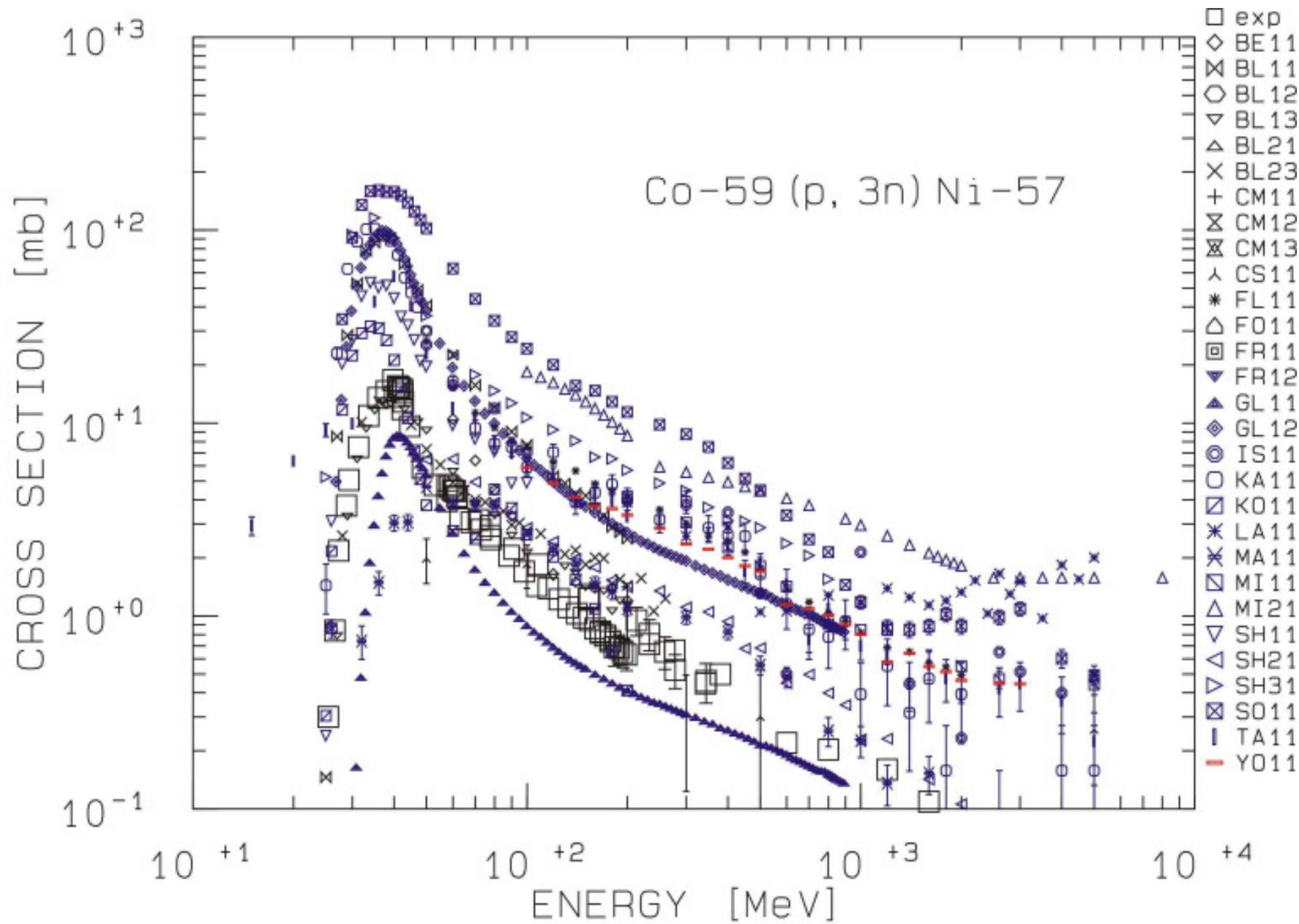
**Blind
intercomparison**

$E < 5$ GeV

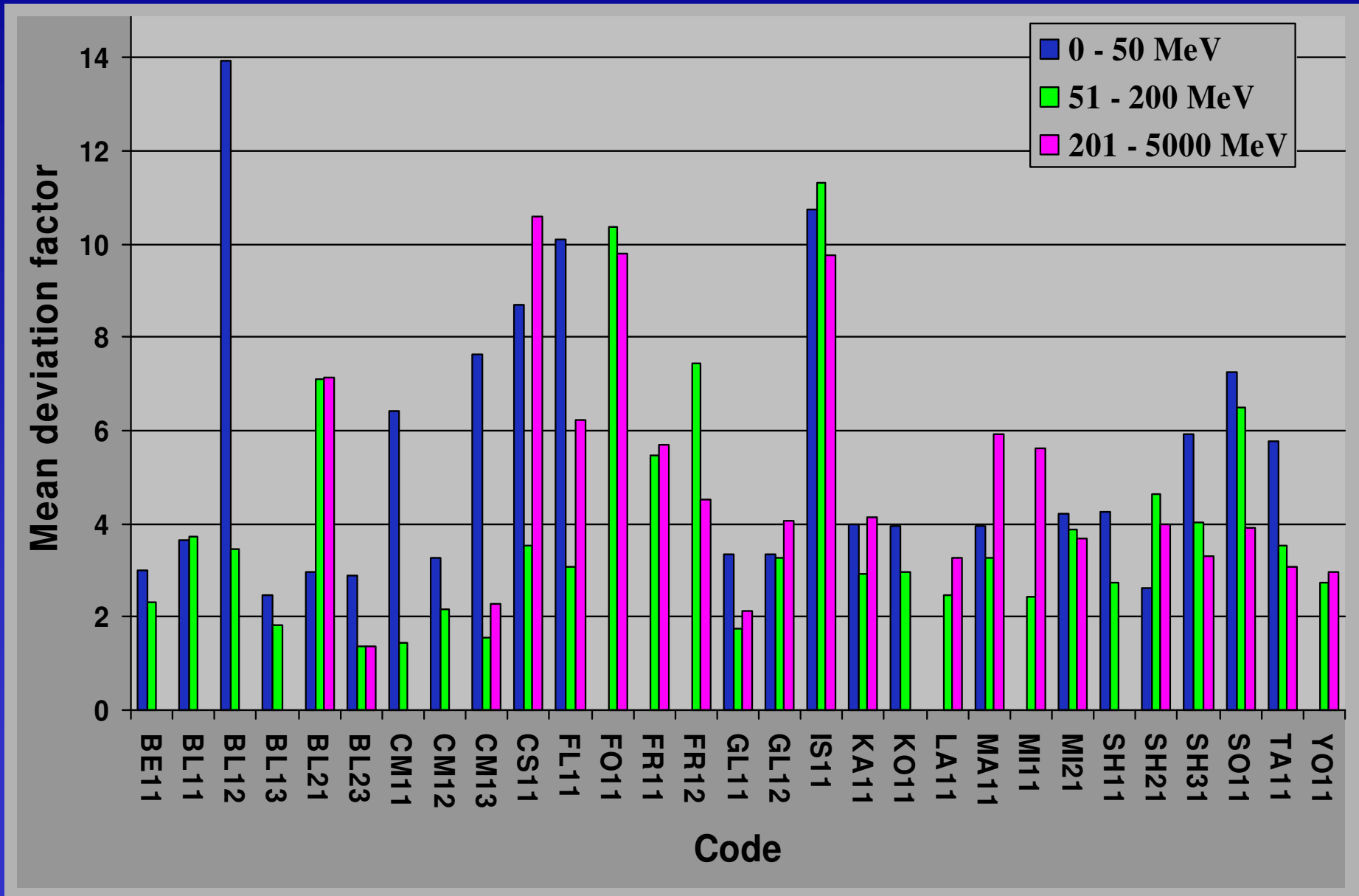
**target
elements:
O, Al, Fe,
Co, Zr, Au**

code	physical model employed	code used	options used	energy range in MeV
BE11	PE + EVAP via MASTER EQ.	PEQAG2 (extended)		50 - 200 (Fe, Co) 26 - 200 (Zr,Au)
BL11	PE + EVAP	ALICE 92	Fermi gas level densities	7 - 200
BL12	PE + EVAP	ALICE 92	Chadwick level densities	1 - 200
BL13	PE + EVAP	ALICE 92	Kataria-Ramamurty level densities	1 - 200
BL21	HMS + EVAP	HMS-ALICE	Fermi gas level densities	3 - 290
BL23	HMS + EVAP	HMS-ALICE	Kataria-Ramamurty level densities	12 - 250
CM11	FKK + EXCITON + HAUSER FESHBACH EVAP	FKK-GNASH	no evaporation of Be-7	10 - 200
CM12	FKK + EXCITON + HAUSER FESHBACH EVAP	FKK-GNASH	no evap. of H-3, He-3 and Be-7	8 - 200
CM13	FKK + EXCITON + HAUSER FESHBACH EVAP	FKK-GNASH		15 - 200
CS11	QMD + SDM	QMDRELP+SDMRELP		50 - 5000
FL11	INC + EVAP	HET/BRUYERE		50 - 200
FO11	INC + MSM	PACE + MSM		100 - 300 (Al), 800 (Fe, Co), 900 (Zr), 1000 (Au)
FR11	INC + EVAP	ISABEL-EVA	local thomas fermi density approximation for momenta	100 - 1000
FR12	INC + EVAP	ISABEL-EVA	uniform thomas fermi density approximation for momenta	100 - 1000
GL11	PE + EVAP (GDH)	AREL	Myers-Swiatecki (MS) masses	10 - 900
GL12	PE + EVAP (GDH)	AREL	exp. + MS masses + shell corr.	10 - 900
IS11	INC + PE + EVAP + FRAGMENTATION	HETC-FRG		10 - 5000
KA11	INC + EVAP	INUCL		4 - 5000
KO11	FKK + EVAP	MINGUS		8 - 200
LA11	INC + SMM + EVAP	ISABEL/SMM		200 - 5000
MA11	INC + PE + EVAP	CEM 95		10 - 5000
MI11	INC + EVAP	HET-KFA2		200 - 5000
MI21	TSAO & SILBERBERG SYSTEMATICS	SPALL (modified)/YIELD		10 - 5000
SH11	PE + EVAP	ALICE -IPPE		3 - 100
SH21	INC + EVAP	CASCADE		50 - 5000
SH31	INC + EVAP	DISCA		14 - 800
SO11	INC + PE + SMM + EVAP + FERMI BREAKUP	MSDM		1 - 5000
TA11	INC + PE + EVAP	HETC-3STEP		15 - 5000
YO11	INC + EVAP	MECC7 + EVAP_F		100 - 3000

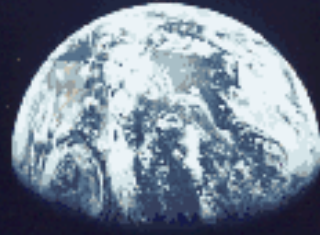




R. Michel, P. Nagel, International Codes and Model Intercomparison for Intermediate Energy Activation Yields, NSC/DOC(97)-1, NEA/OECD, Paris, 1997

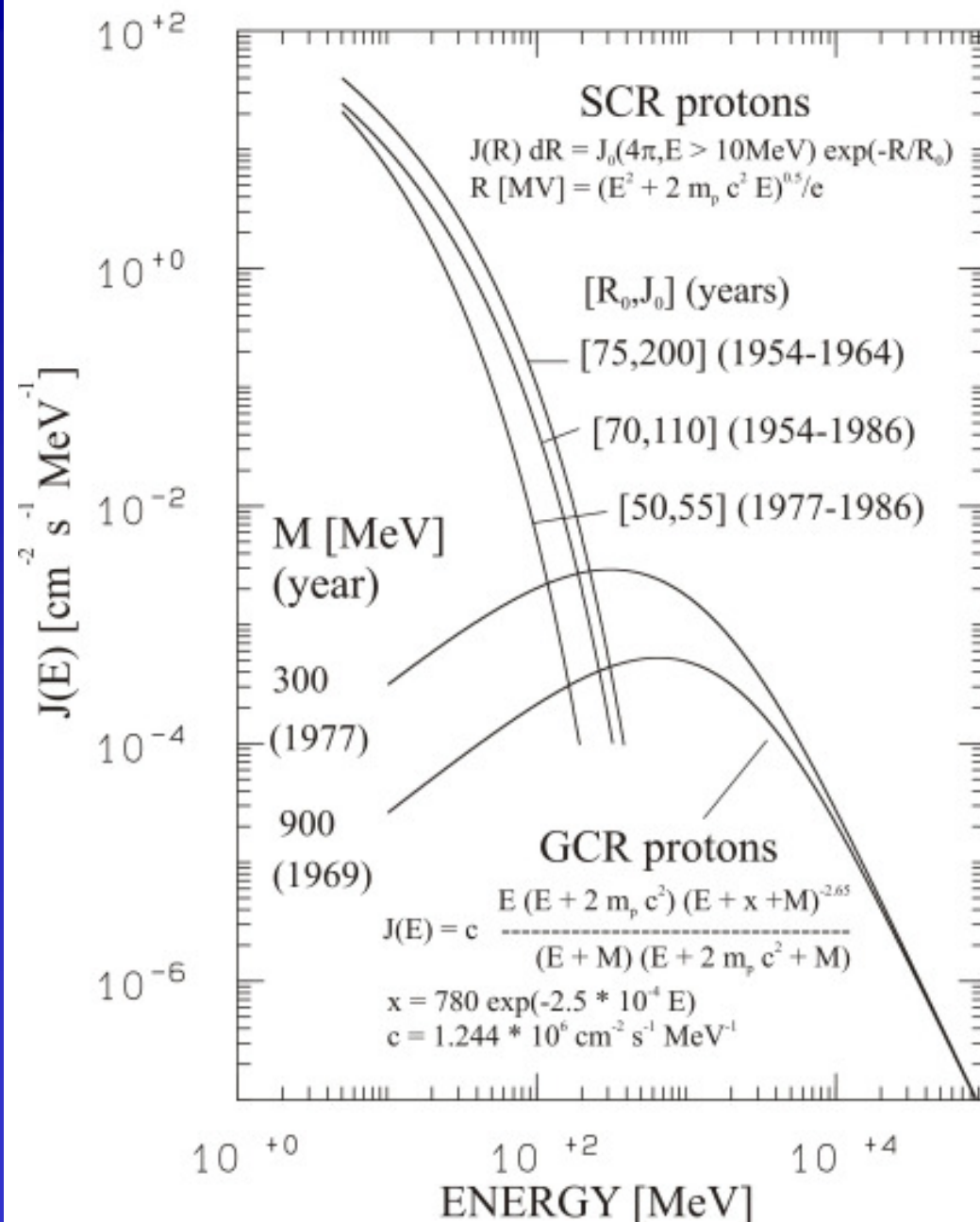


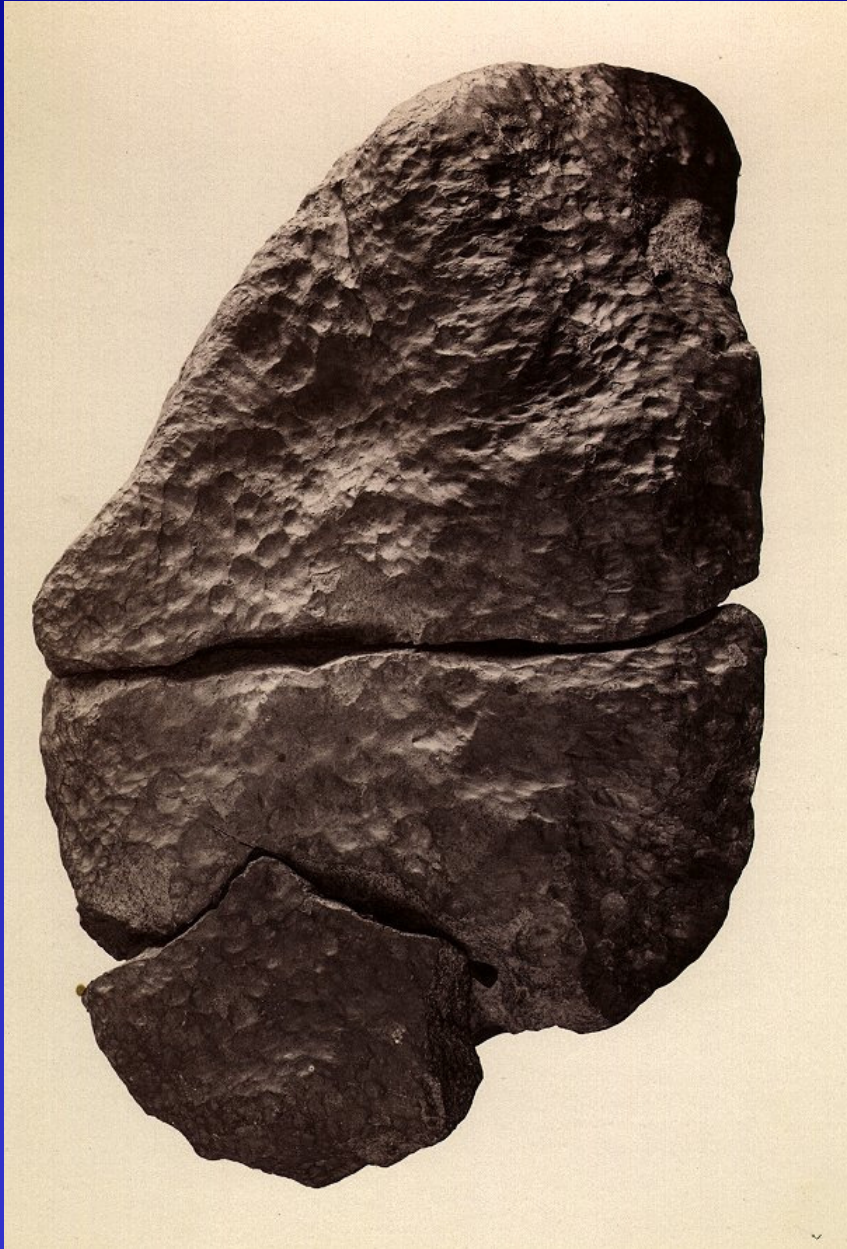
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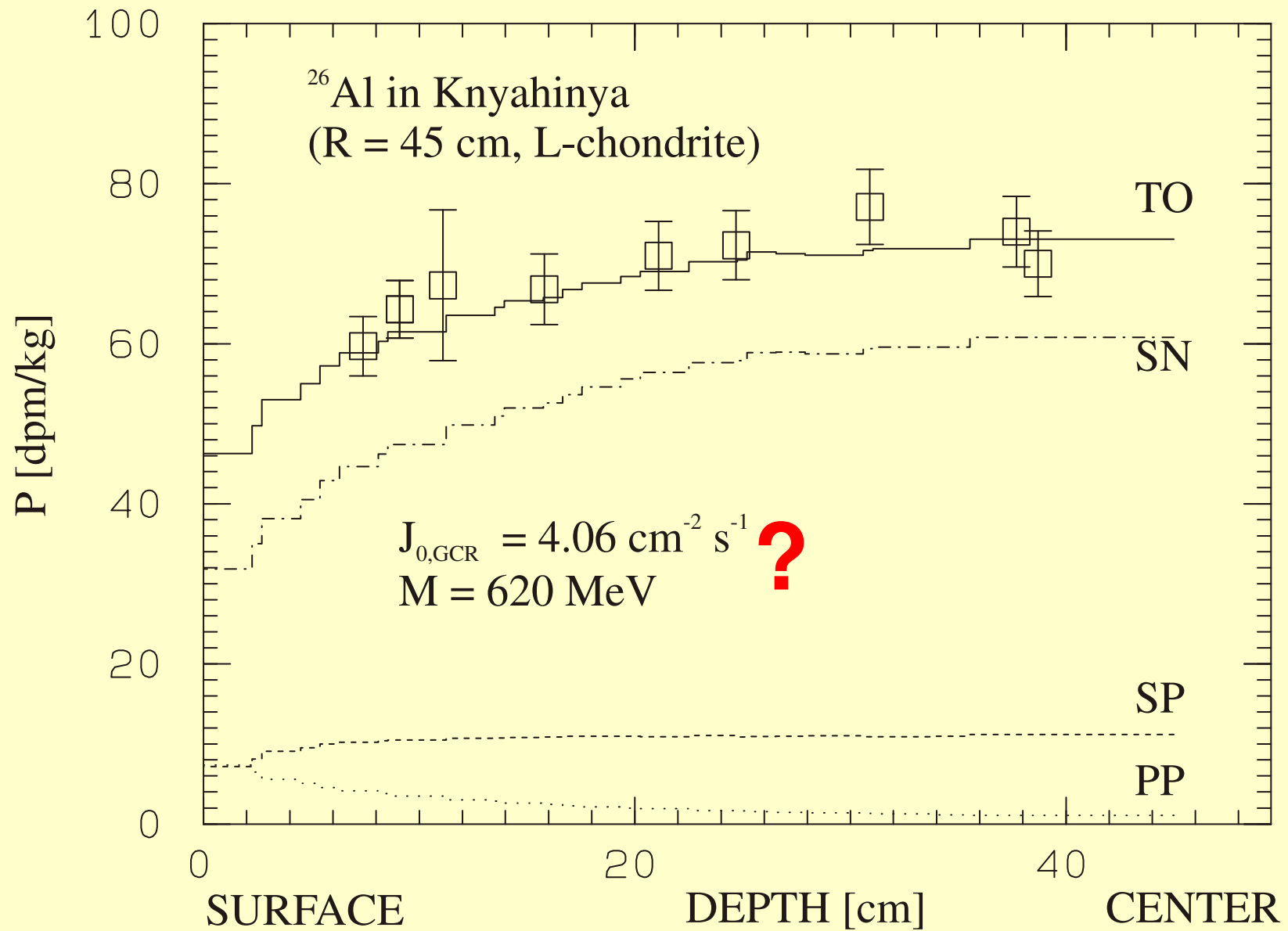
Mean
SCR-spectra
for three solar
11-years cycles
and
extremes of
GCR-spectra
observed.






Stony Meteorite Knyahinya

- **Chondrite (L5)**
- **fell on 9. Juni 1866
in Ukraine**
- **recovered mass ca. 300 kg**





Simulation of the production of cosmogenic nuclides by isotropic irradiation of an artificial stony meteoroid with 1.6 GeV protons.

- R = 25 cm
- 1400 individual targets
- 27 elements
- in 282 h: simulation of 1,6 Ma in space

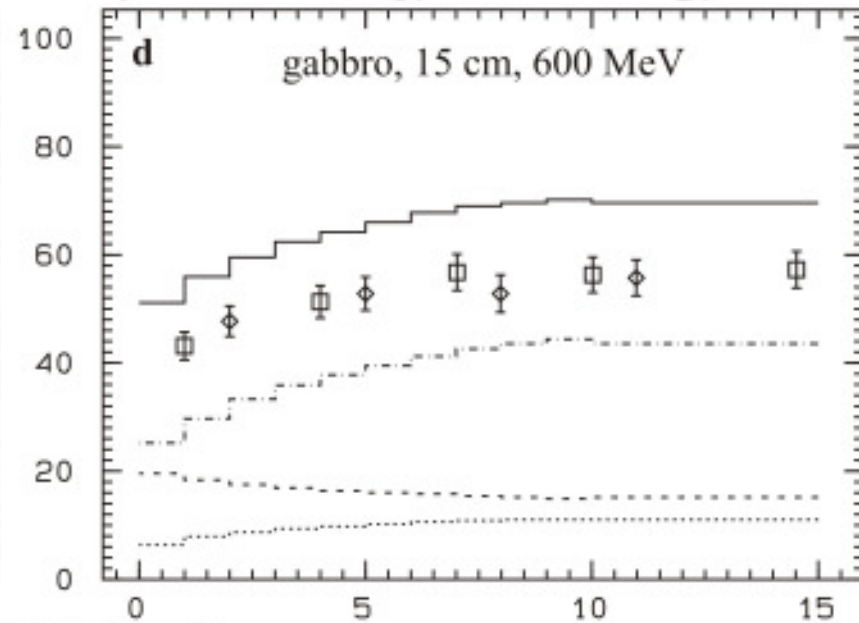
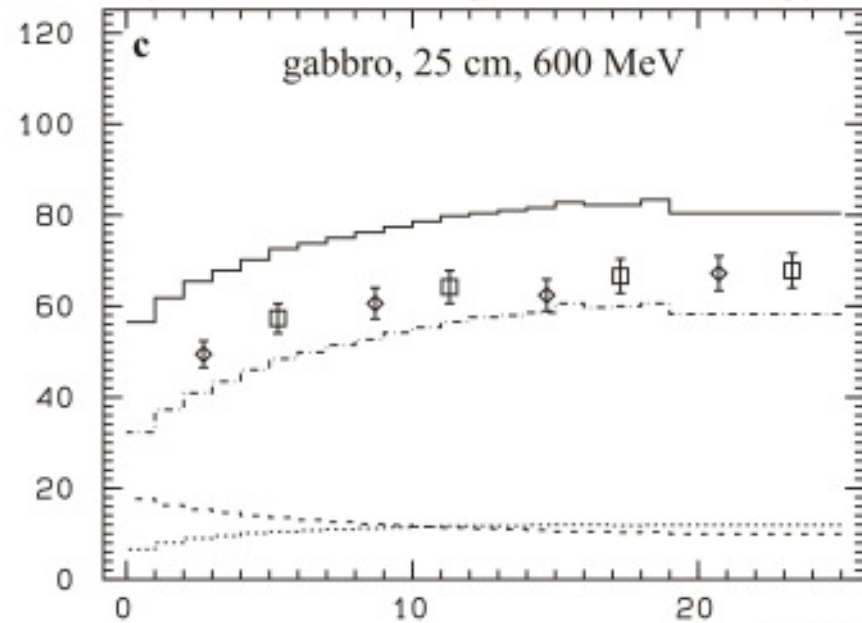
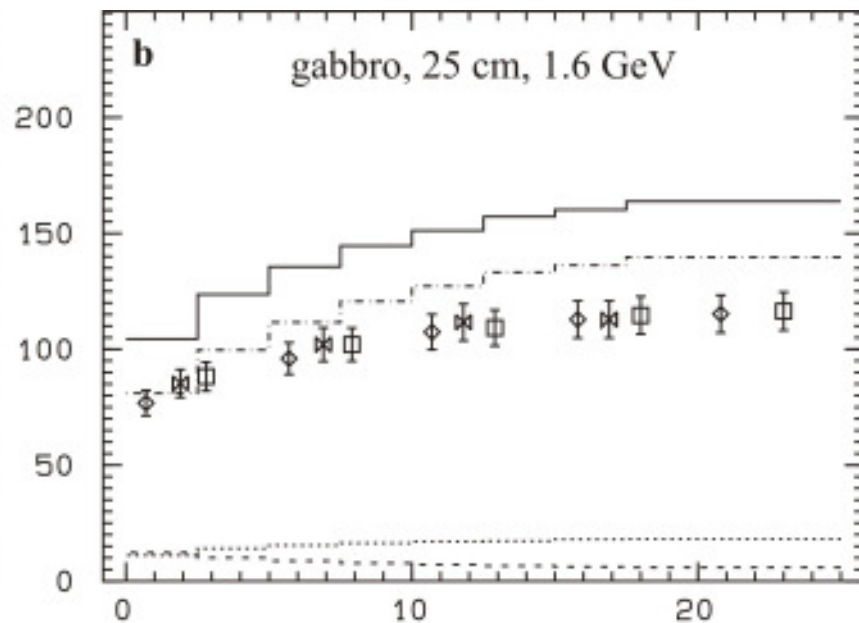
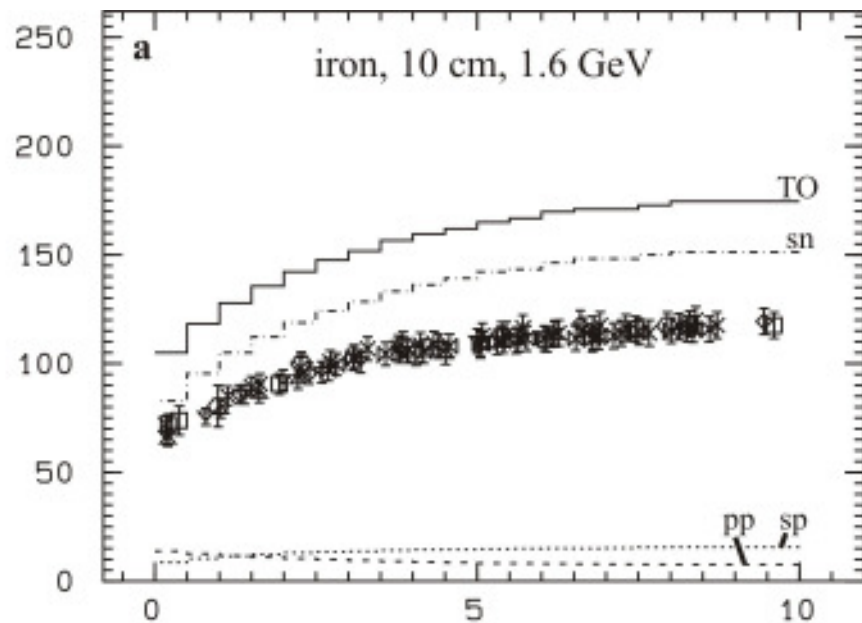


Simulation of the production of cosmogenic nuclides at LNS/Saclay by an isotropic irradiation of an artificial iron meteoroid with 1.6 GeV protons.

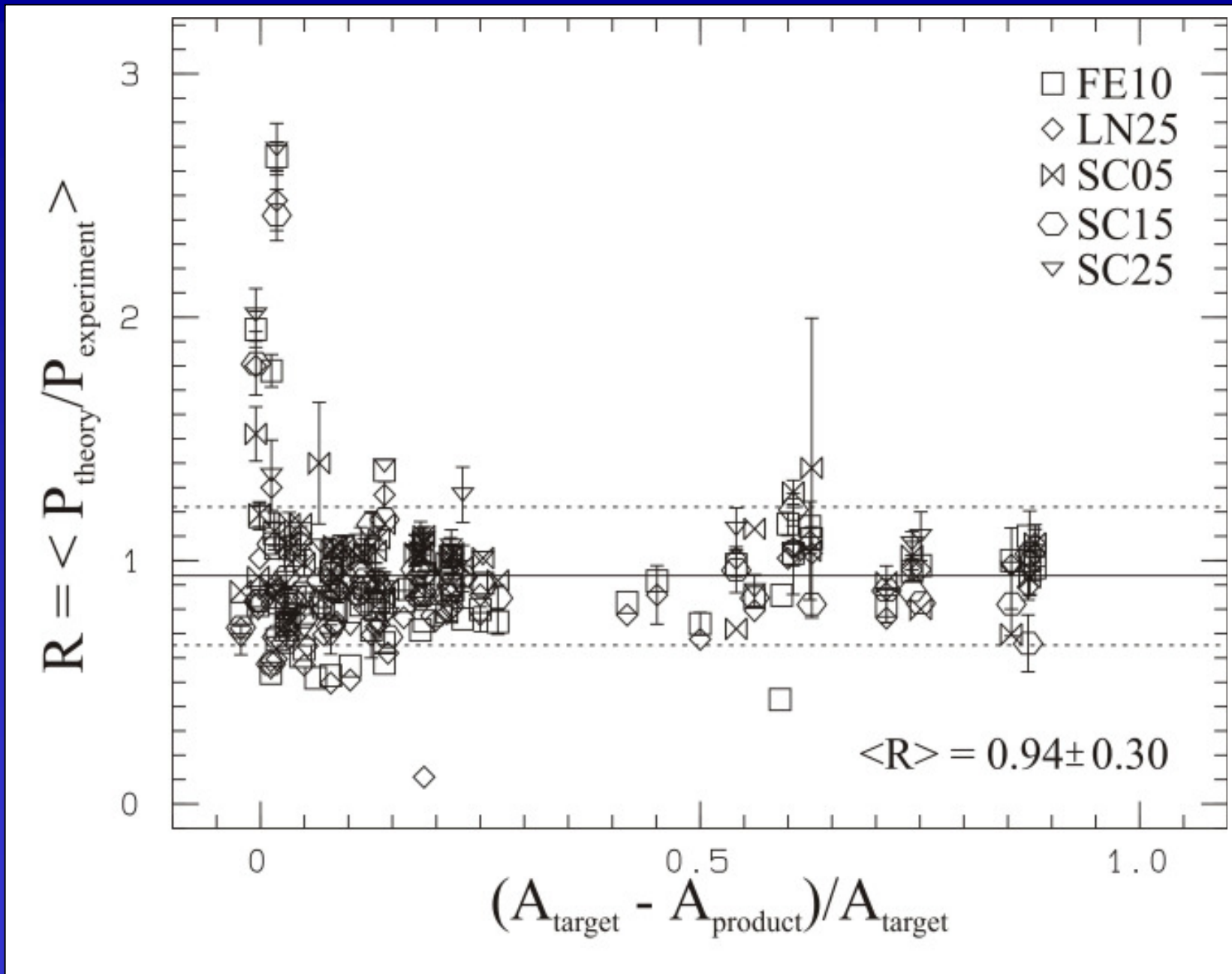
- R = 10 cm
- 870 individual targets
- 28 elements
- in 42 h: simulation of 3,0 Ma in space

R. Michel, ZSR, Leibniz Universität Hannover

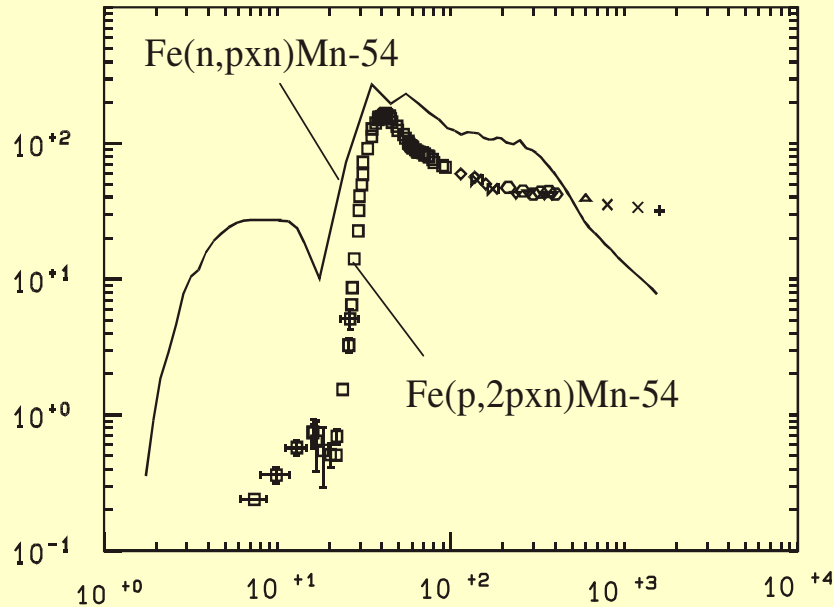
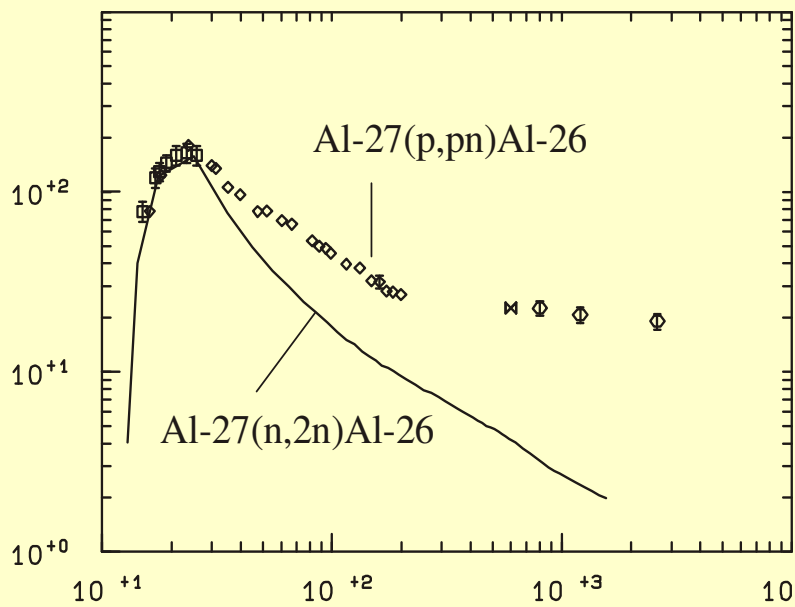
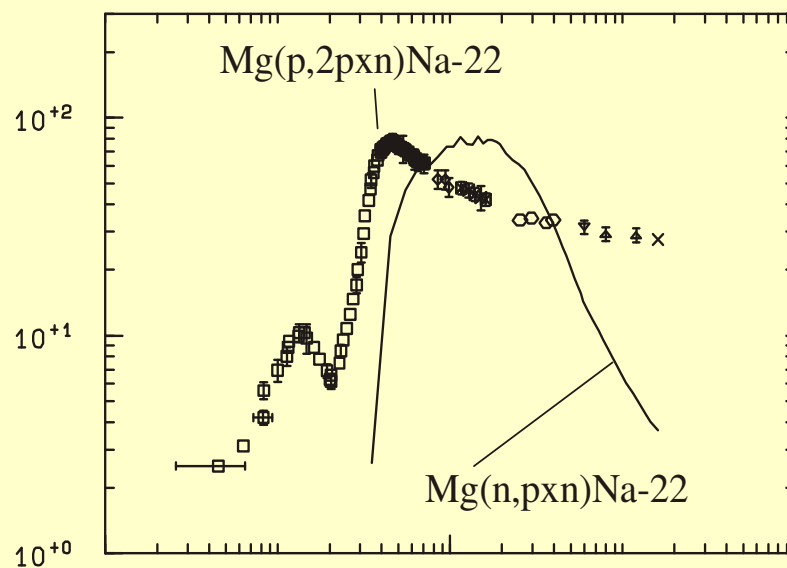
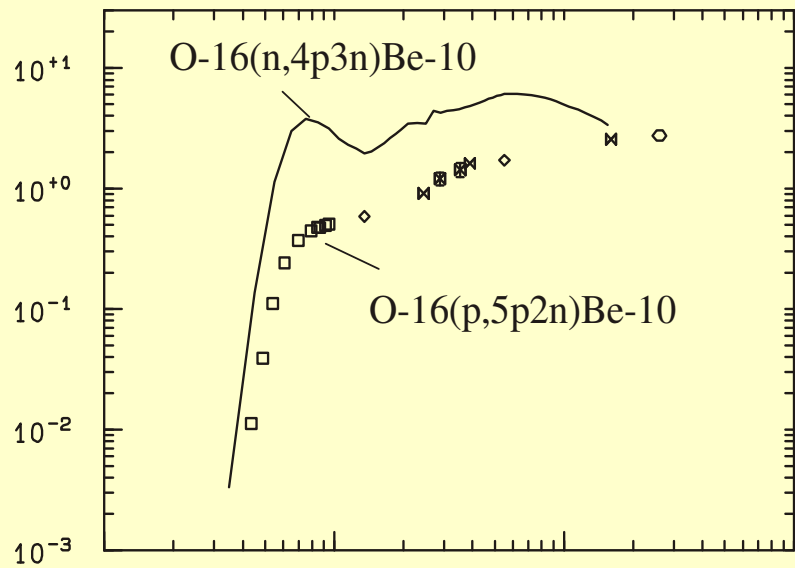
P(Mn-54 from iron) [dpm/kg]



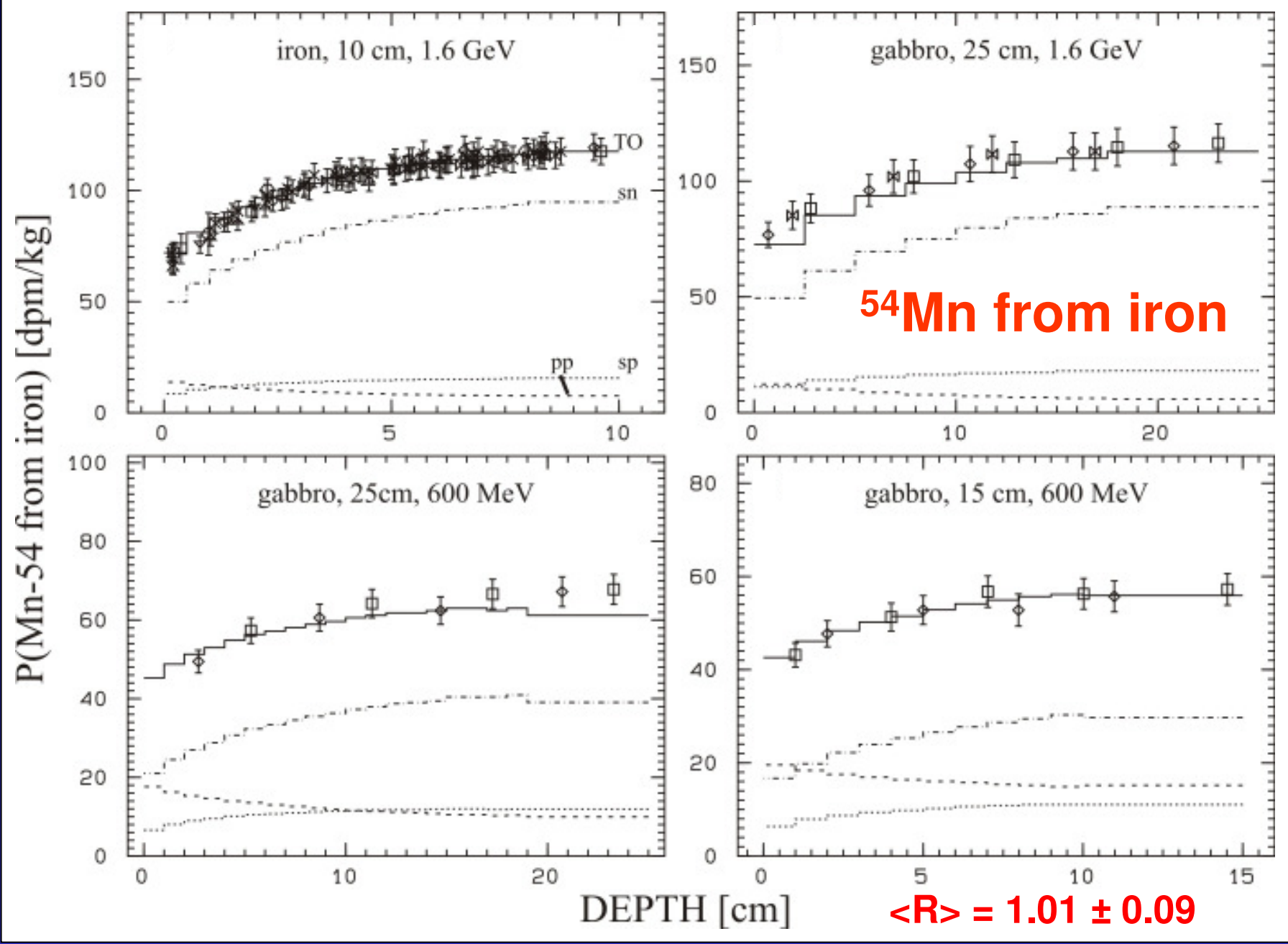
DEPTH [cm]

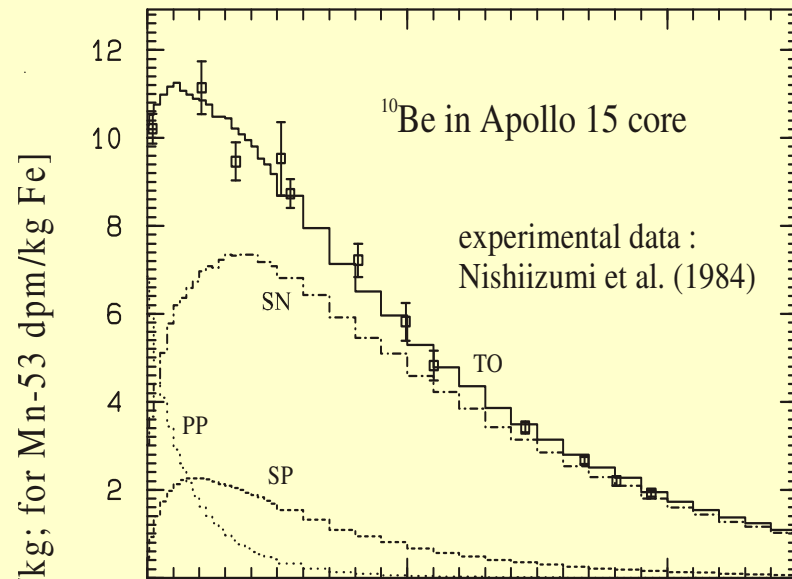


CROSS SECTION [mb]

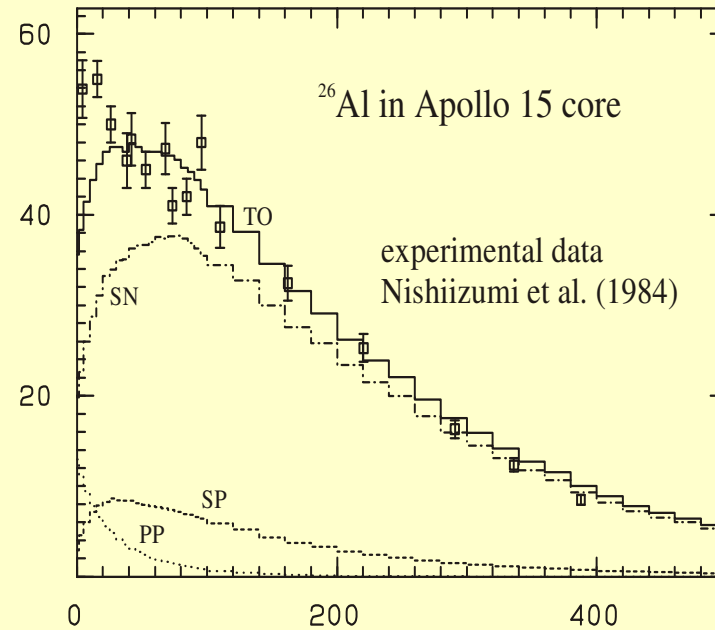
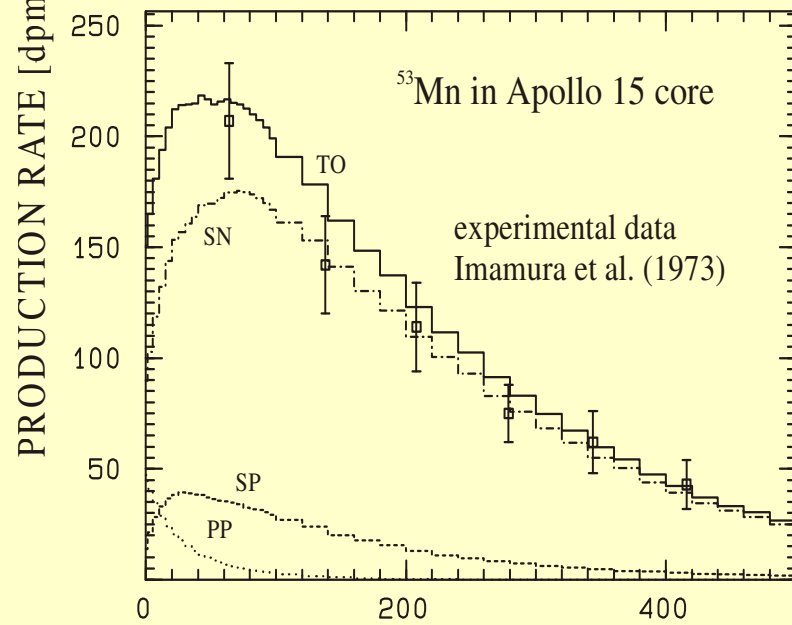


ENERGY [MeV]



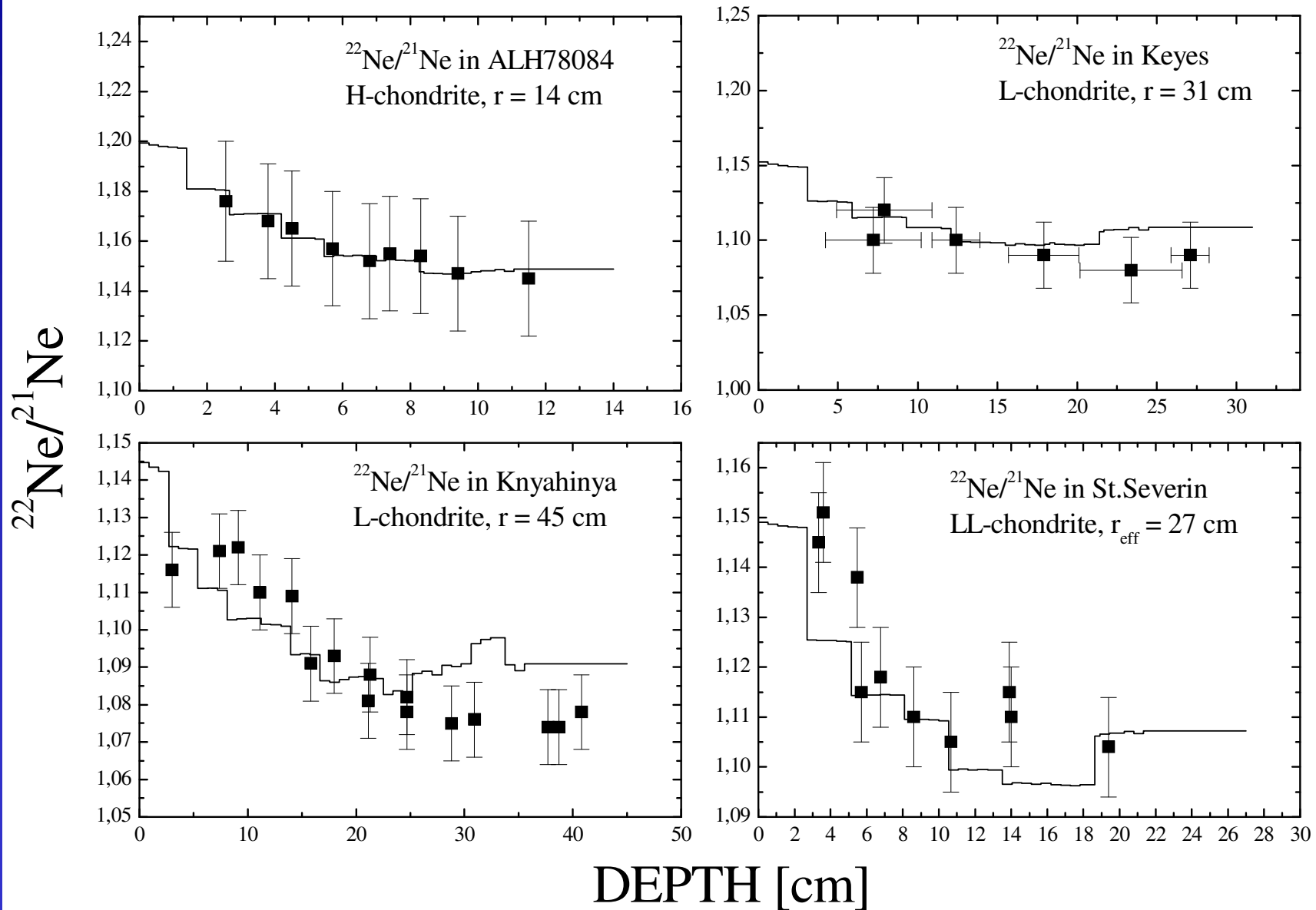


Constancy of the galactic cosmic radiation during the last 10 Ma

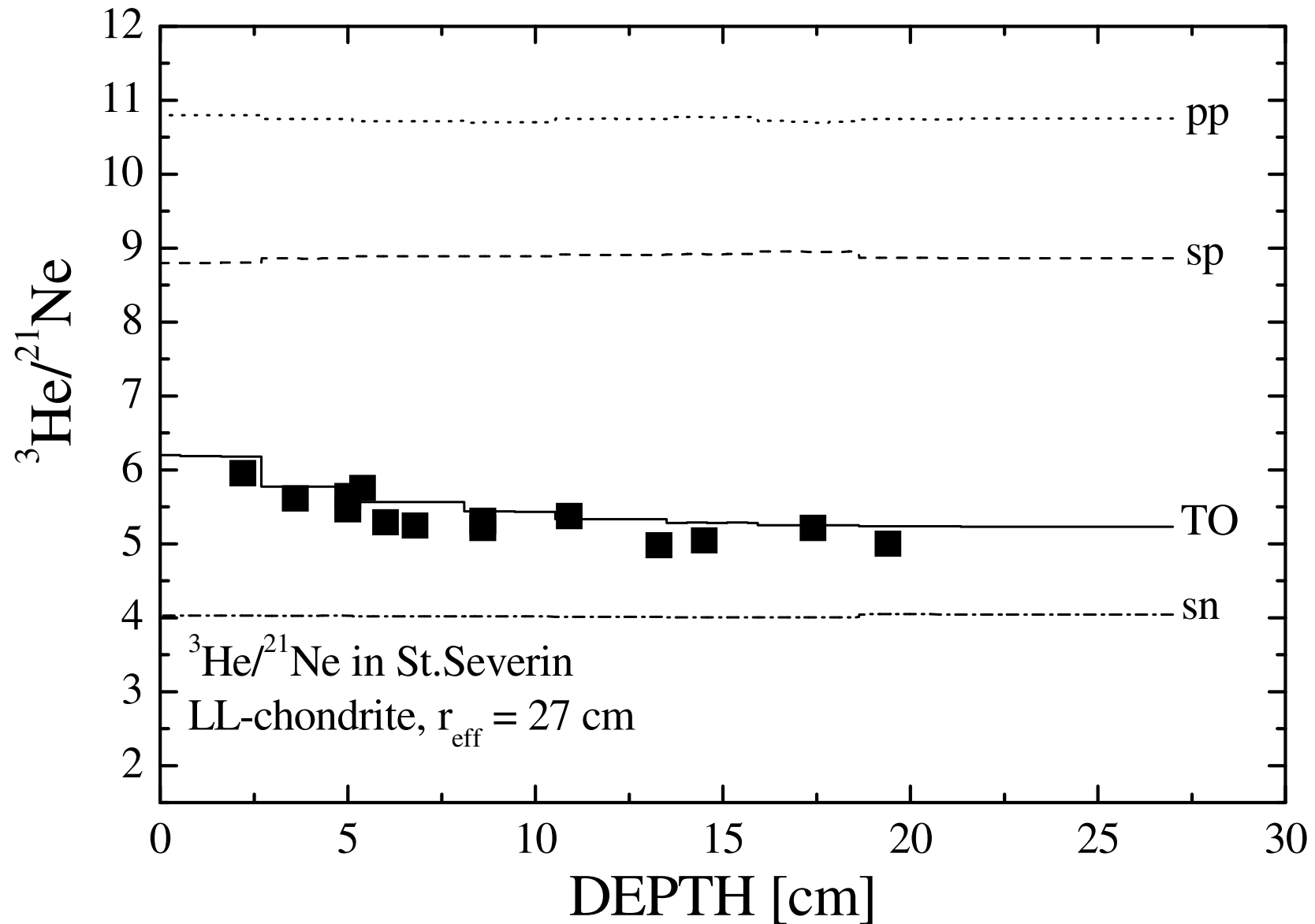


DEPTH [g/cm**2]

$^{22}\text{Ne}/^{21}\text{Ne}$ in chondrites



$^3\text{He}/^{21}\text{Ne}$ in St. Severin



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Data Needs of Accelerator-Driven Technologies

materials	target elements
air, water	N, O, Ar
concrete	C, O, Na, Mg, Al, Si, Ca, Mn, Fe
beam pipe	Mg, Al, Zn, Cu
stainless steel	Cr, Mn, Fe, Co, Ni
accelerator structural materials	Al, Cu, Fe, ...
spallation target	Na, Fe, Zr, Y, W, Ta, Hg, Pb, Bi, U, Th
fission products & actinides	Sr, Tc, I, Cs, Np, Pu, Am, Cm

**It will not be possible to measure all data needed.
One has largely to rely on calculated data.**

High- and Intermediate Energy Nuclear Data for Accelerator-Driven Systems

HINDAS

Experimental and theoretical studies on the following
elements:

1. One shielding material: Fe
2. One target element : Pb
3. One actinide: U

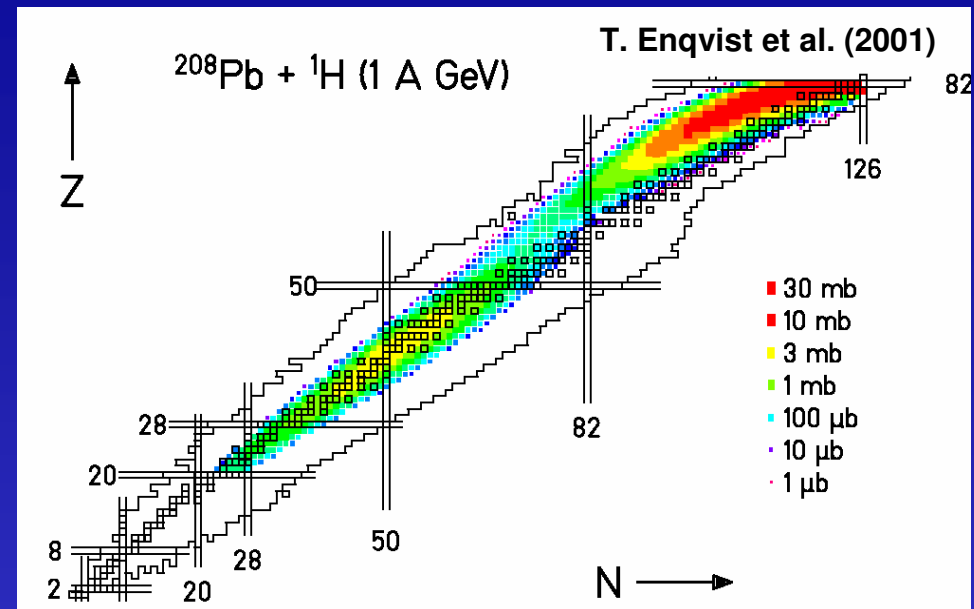
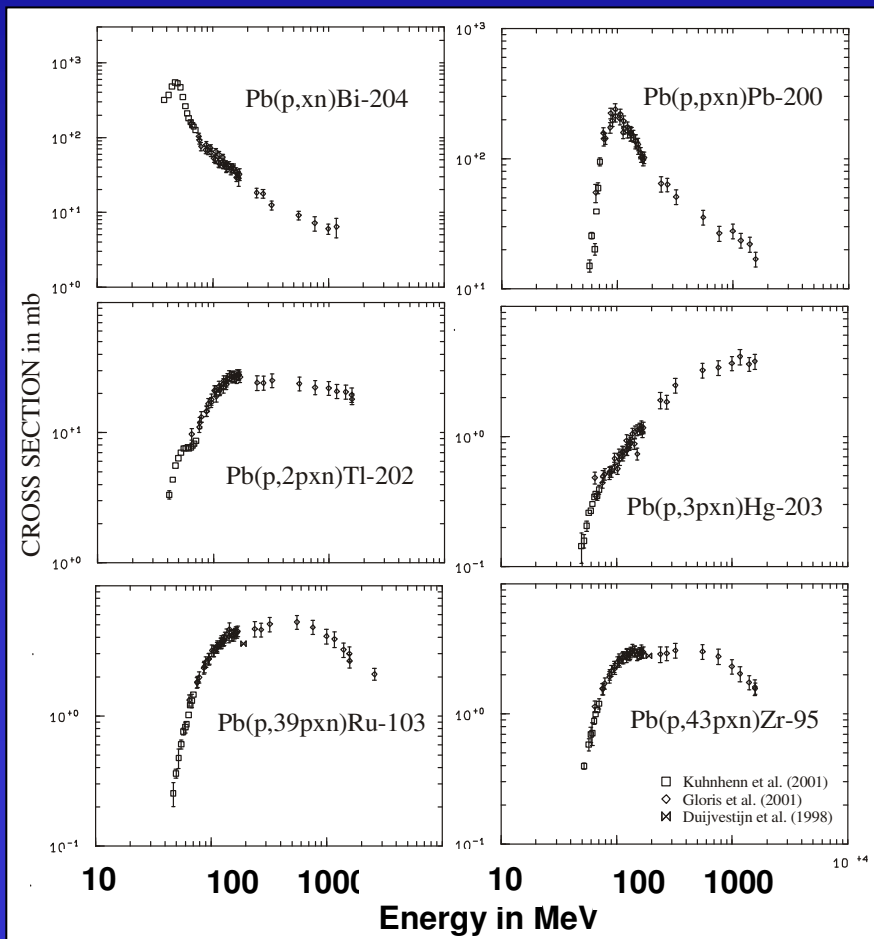
Task 1 : Experiments between 20 and 200 MeV

Task 2 : Experiments above 200 MeV

Task 3 : Theory (TALYS, INCL, ABLA) and evaluation

The two approaches: Classical versus Inverse Kinematics

Complete excitation functions for
long-lived radionuclides and
stable rare gas nuclides.

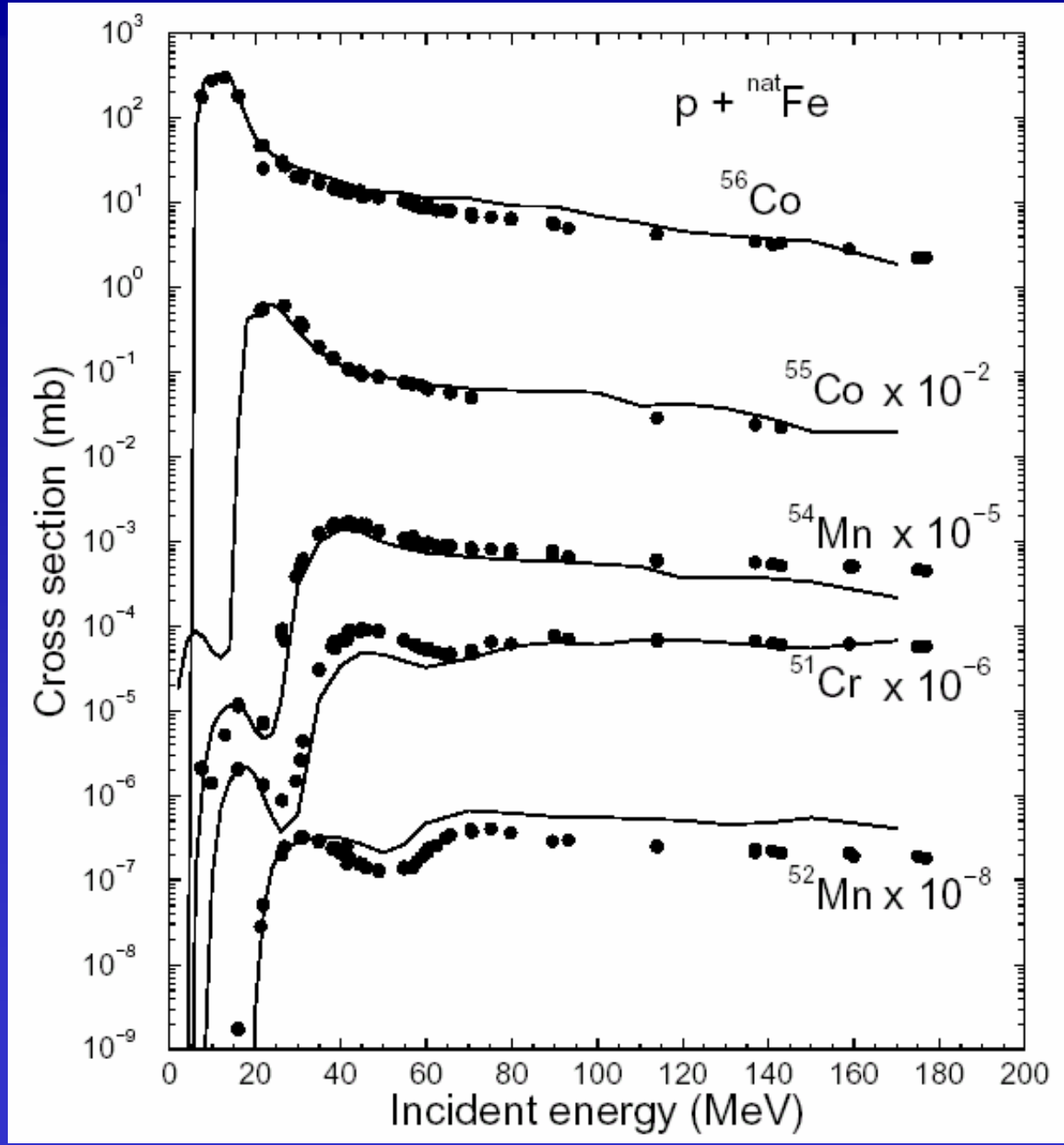


Complete product-nuclide
spectrum at one energy
(except very low masses).

Inventory of an accelerator-
driven facility

- during operation
- shortly after shut-down

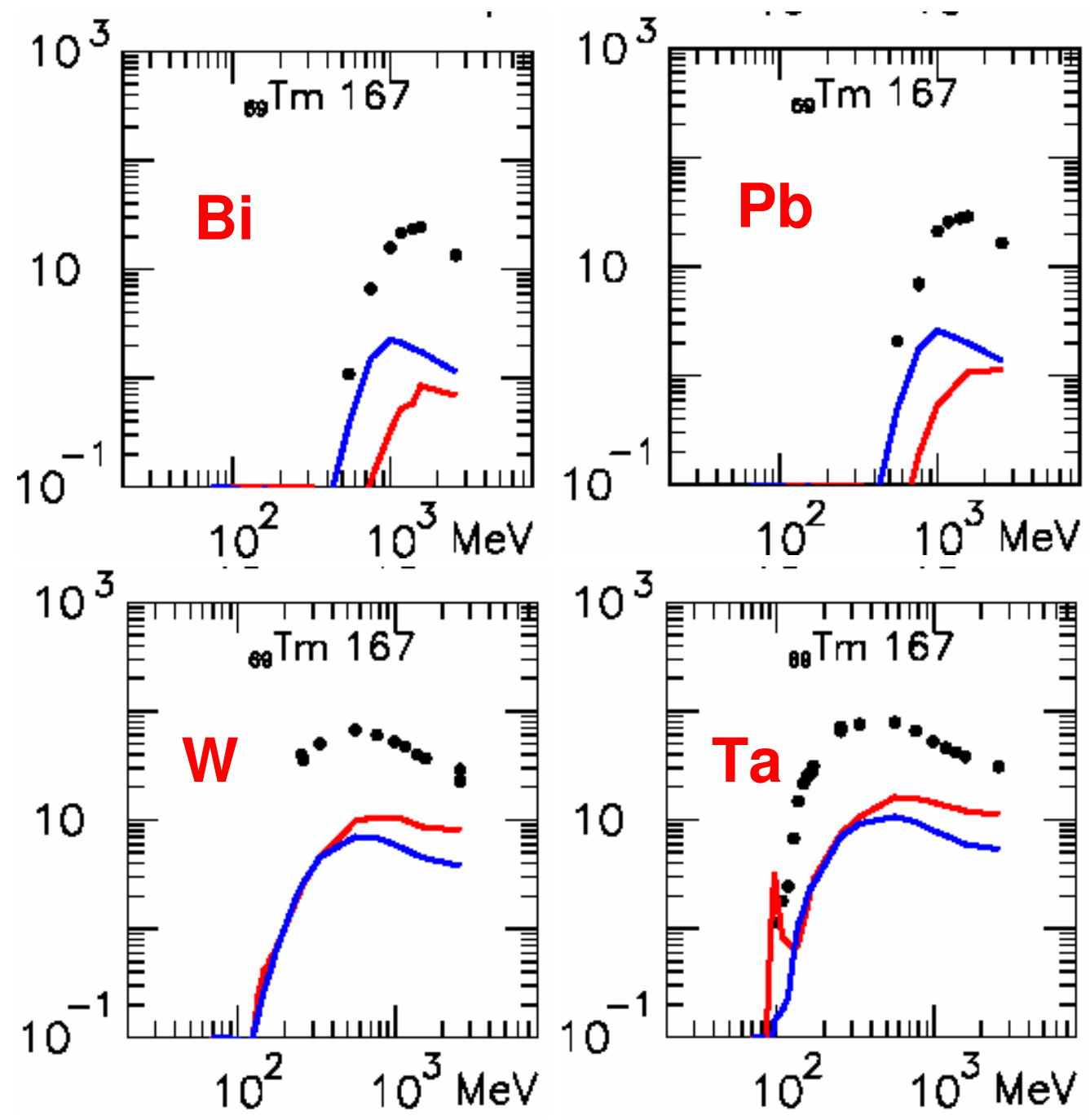
R. Michel, ZSR, Leibniz Universität Hannover



Modeling by the TALYS Code

A. Koning (2002)

Production of ^{167}Tm



— INCL4 + ABLA
— Bertini-Dresner-PE

Requirements for Testing the Performance of Model Codes

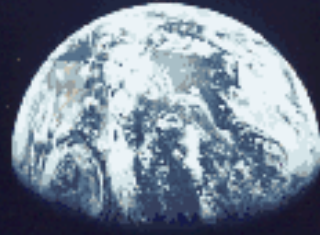
Comprehensive, systematic and representative coverage of

- particle types,
- target elements,
- product nuclides,
- energies.

Special reactions for testing extremes of nuclear reactions:

- unbalanced nucleons in the exit channel,
- competition between spallation and fission,
- competition between evaporation, PE decay, and fission,
- (multi-)fragmentation.

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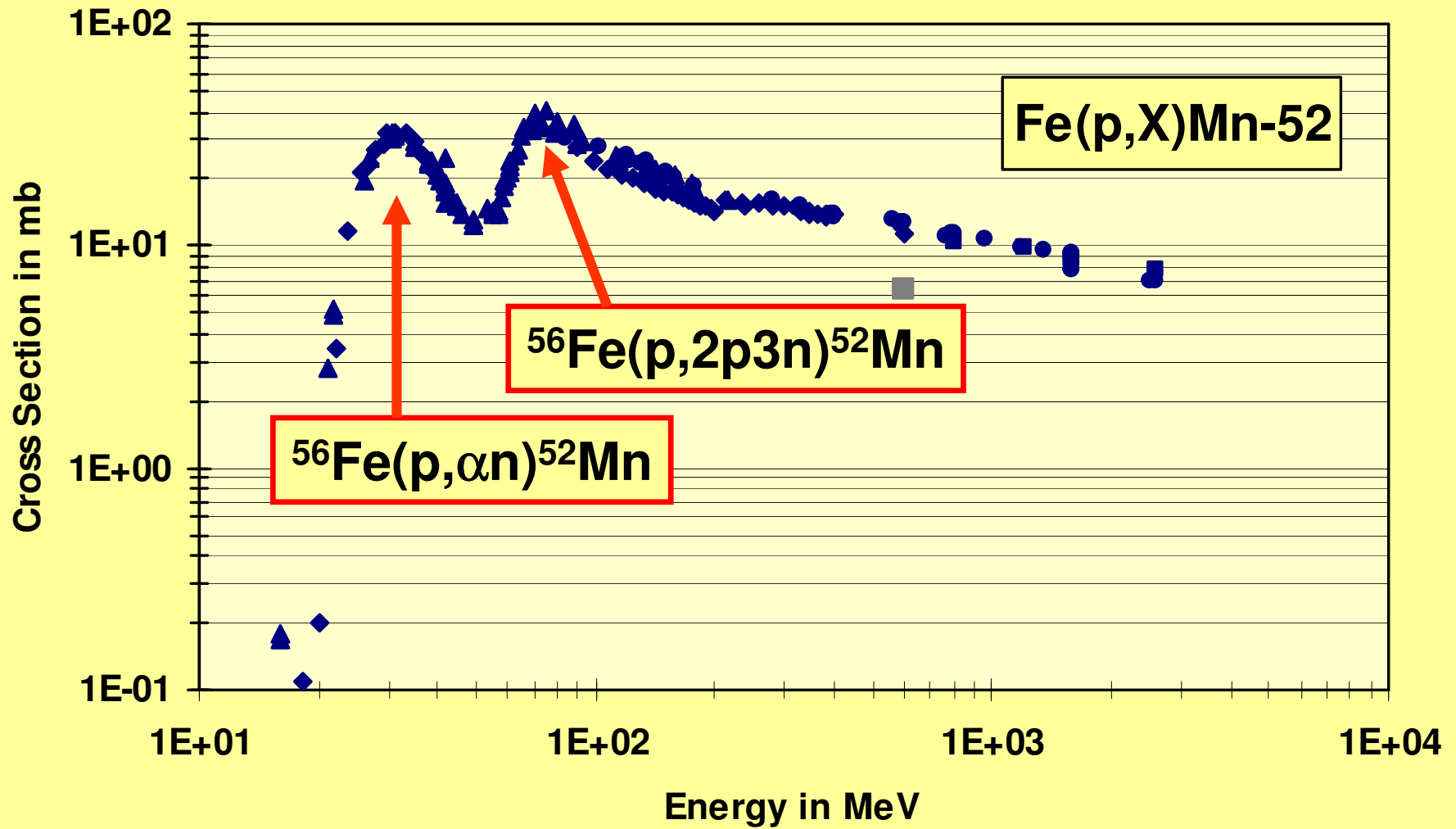


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Target Elements Investigated for the Production of Residual Nuclides by p-Induced Reactions

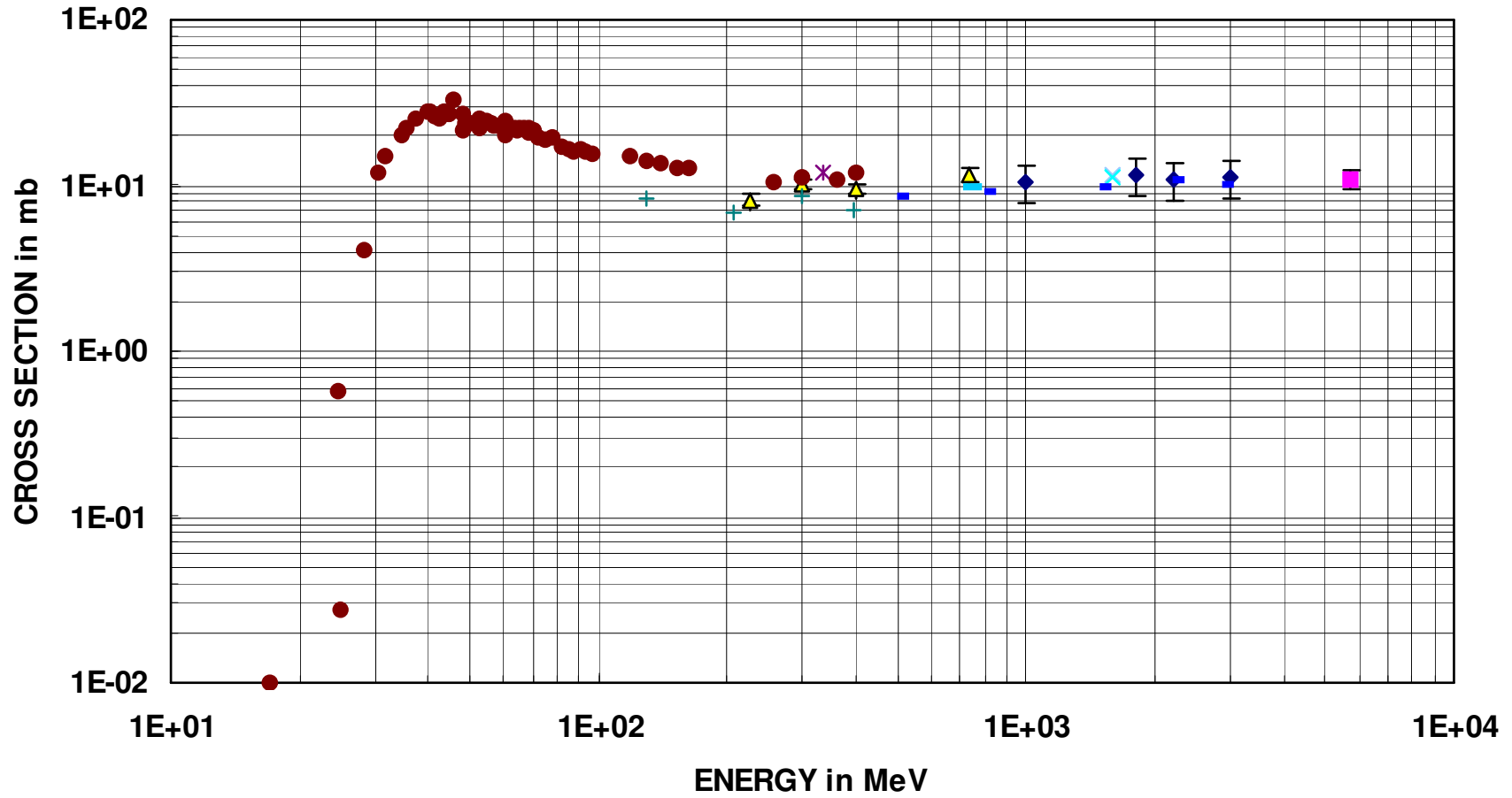
A consistent database containing
> 22 000 cross sections for more than
1000 target-product combinations.

H																	He
Li	Be											B	C	N	O	F	Ne
Na	Mg											Al	Si	P	S	Cl	Ar
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
Fr	Ra	Ac	Rf	Db	Sg	Bh	Hs	Mt	110	111	112		114		116		118
lanthanides																	
Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu				
actinides																	
Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr				



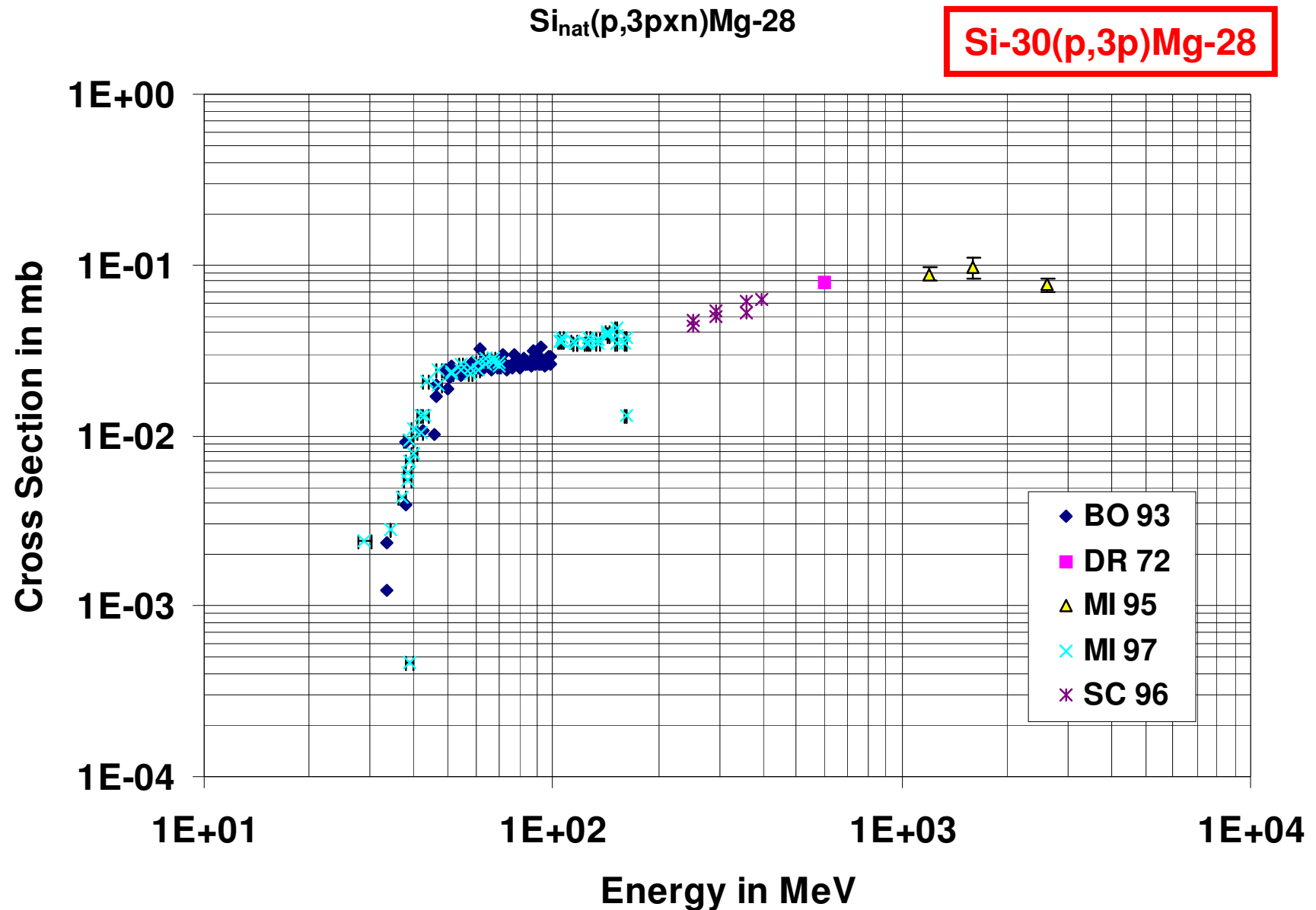
■ WE75 ◆ MI79 ◆ MI83 ▲ MI84 ◆ MI89 ■ LU93 ▲ DE96 ▲ MI96 ◆ SC96 ● this

$C_{nat}(p,3pxn)Be-7$



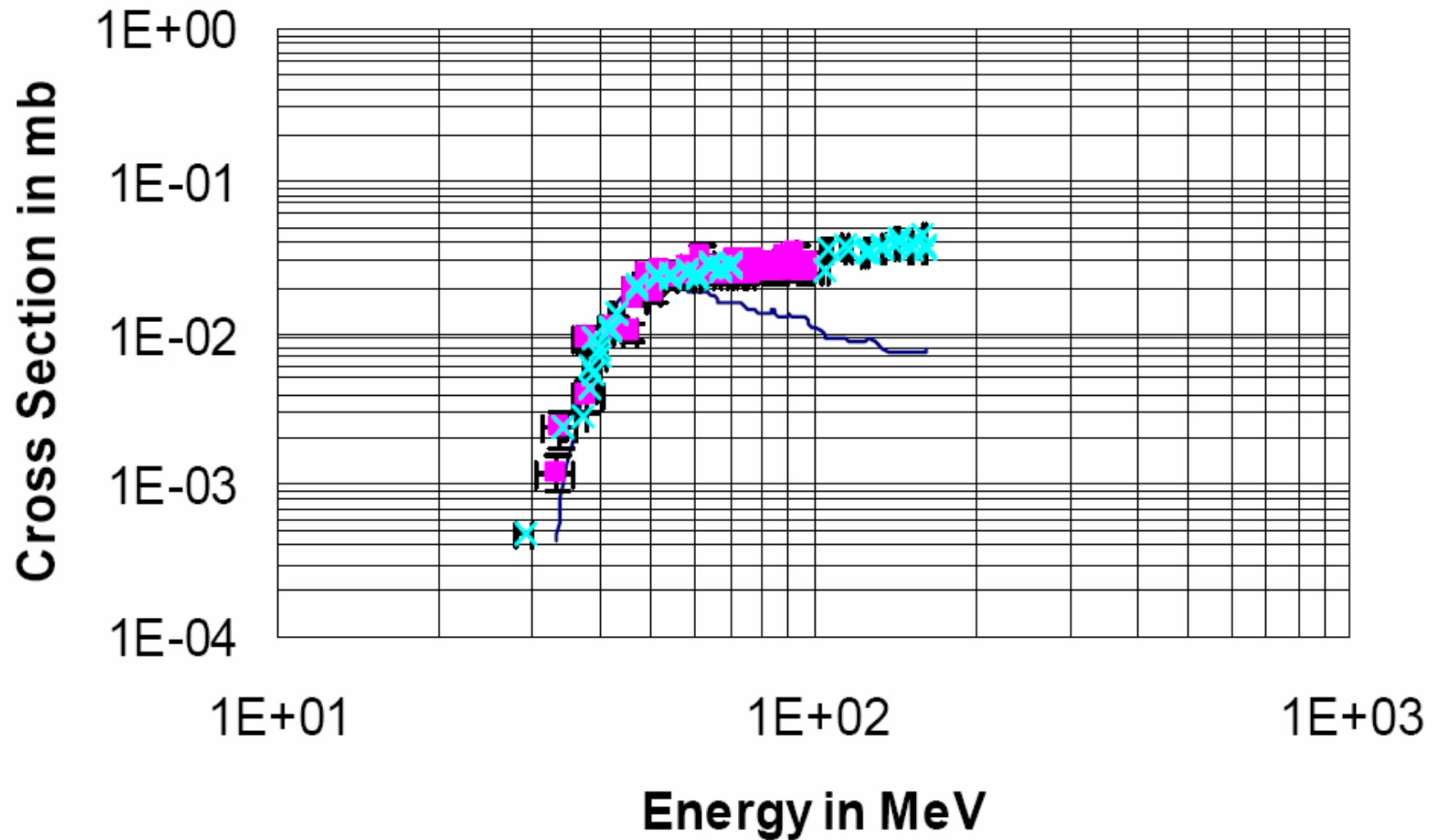
◆ BA 58 ■ BE 60 ▲ HO 60 × MI 95 × LY 97 × MA 51 ● MI 97 + RA 64 - RA 68 - RA 79 A

Special reactions



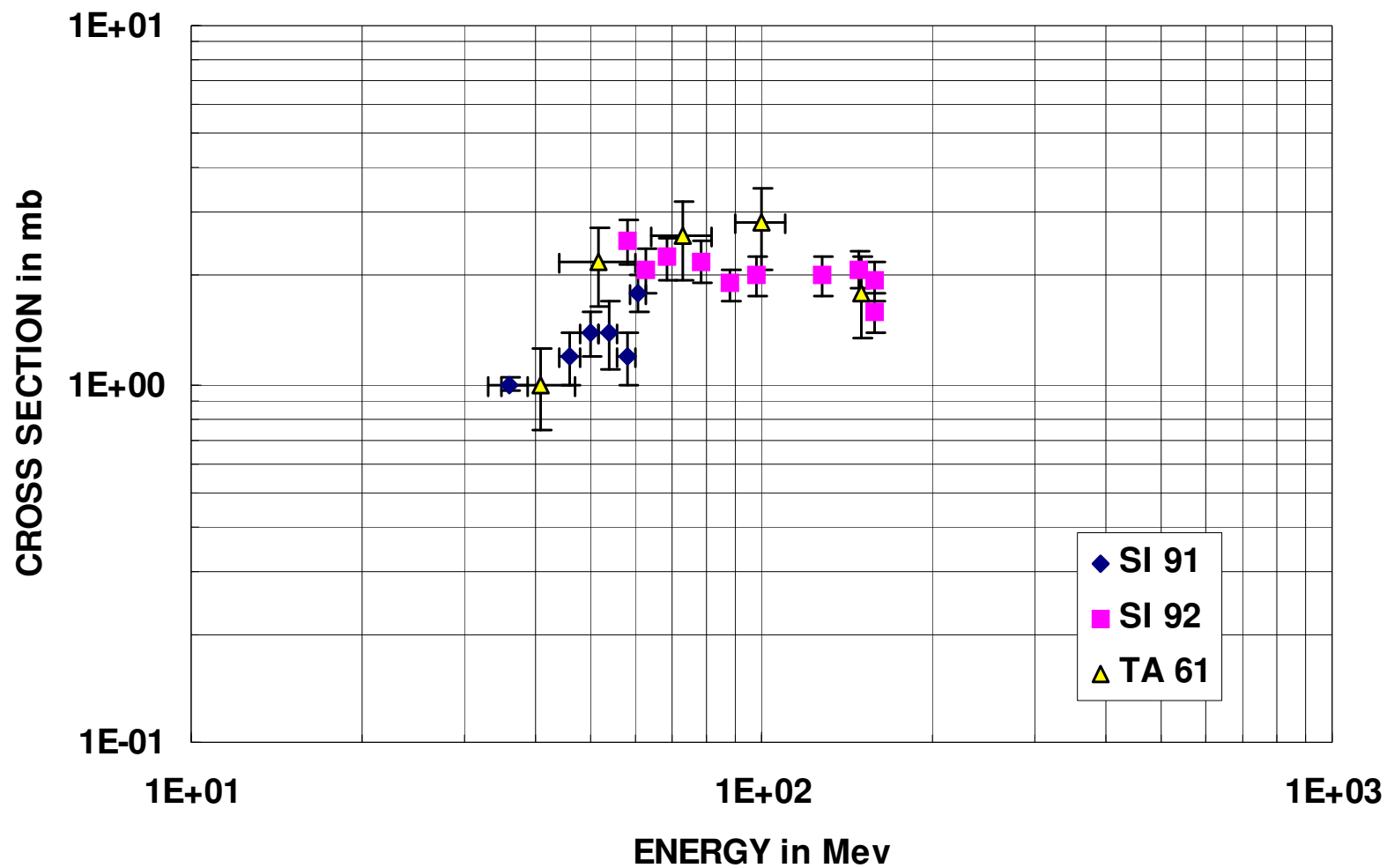
$\text{Si}_{\text{nat}}(\text{p},3\text{pxn})\text{Mg-28}$

$\text{Si-30}(\text{p},3\text{p})\text{Mg-28}$

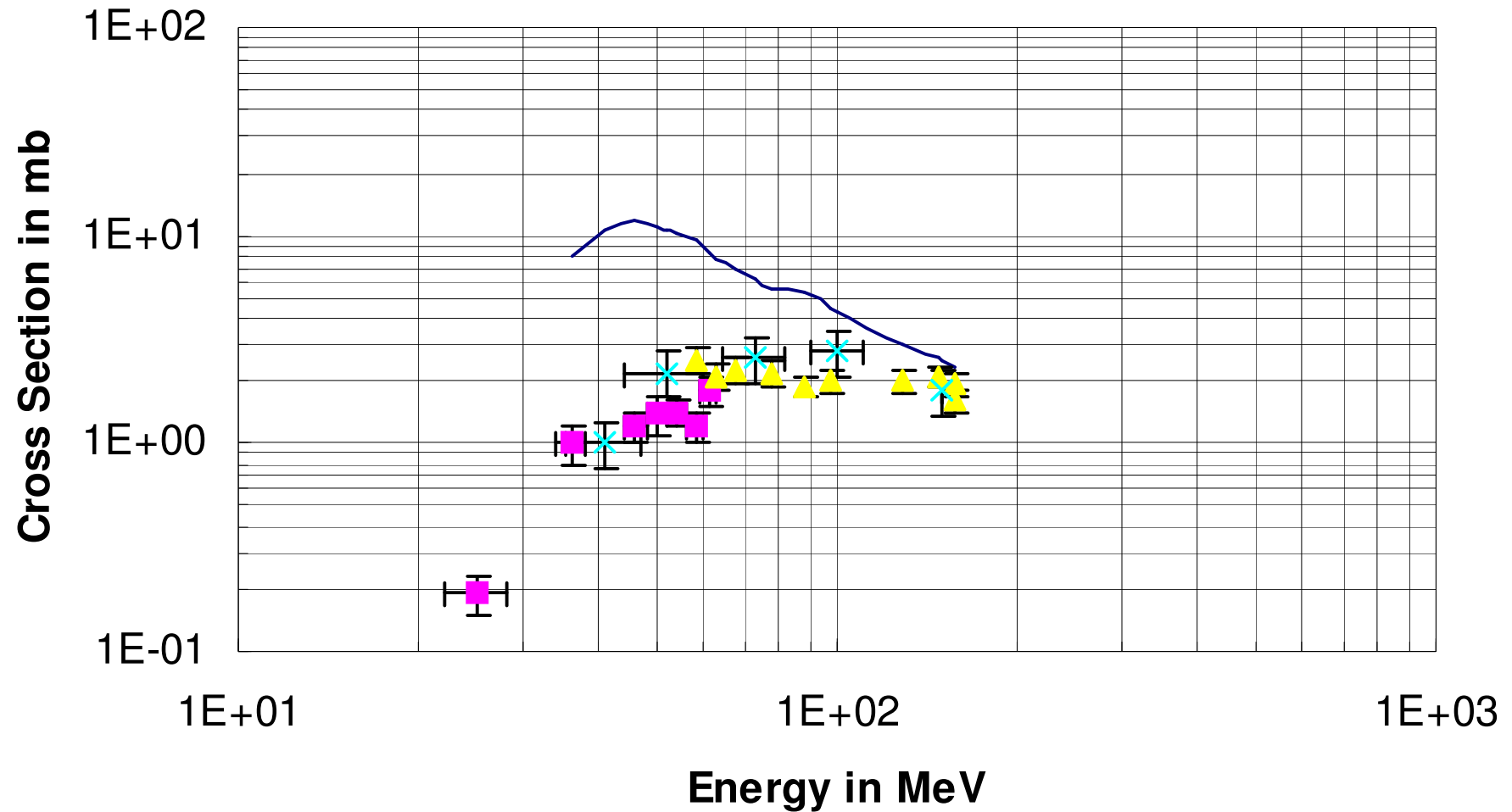


Special reactions

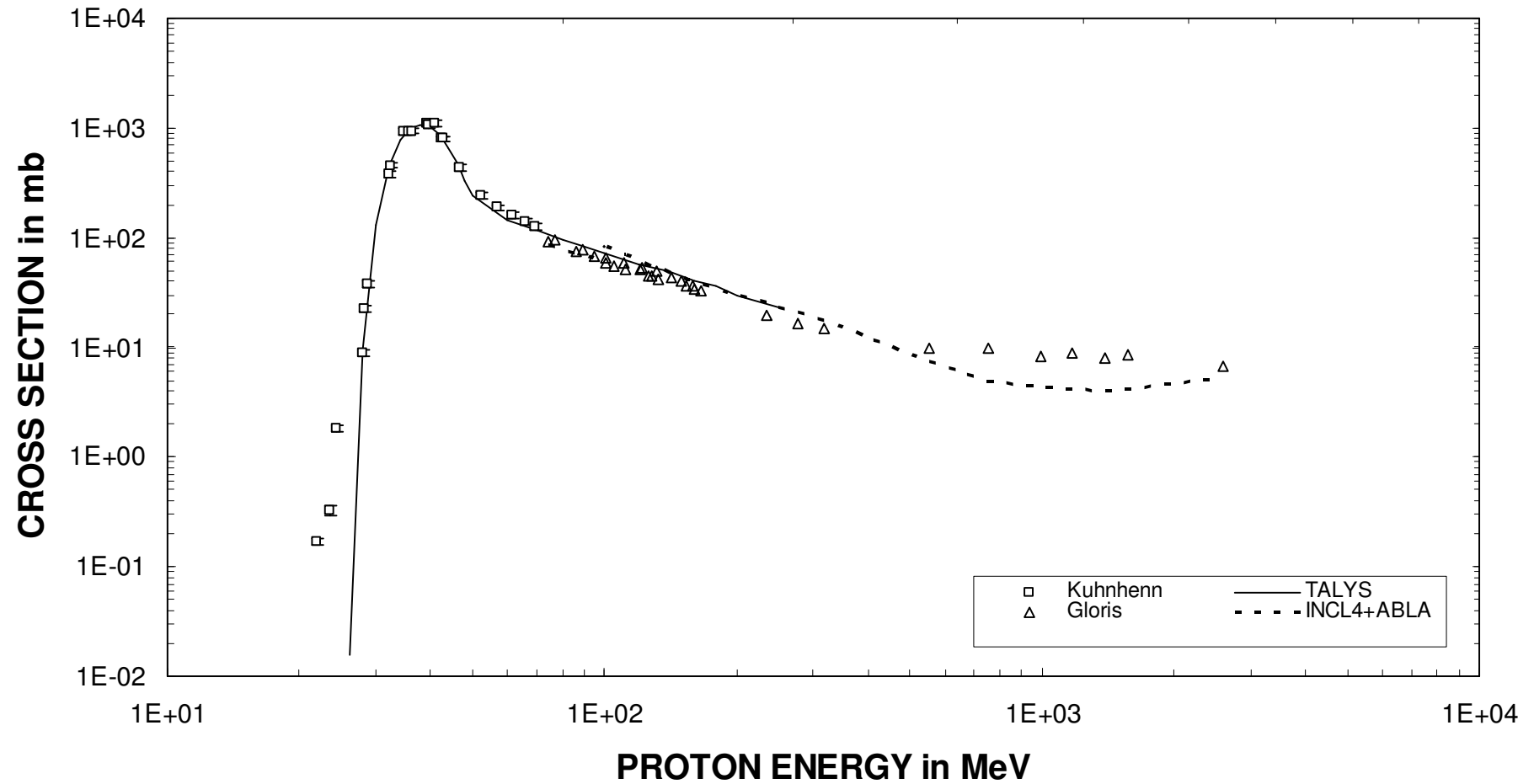
O-16(p,3p)C-14



$^{16}\text{O}(p, 3p)^{14}\text{C}$

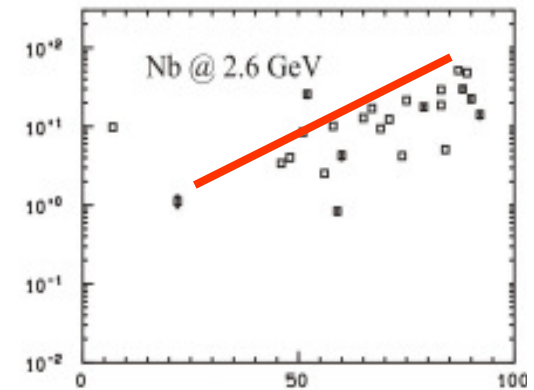
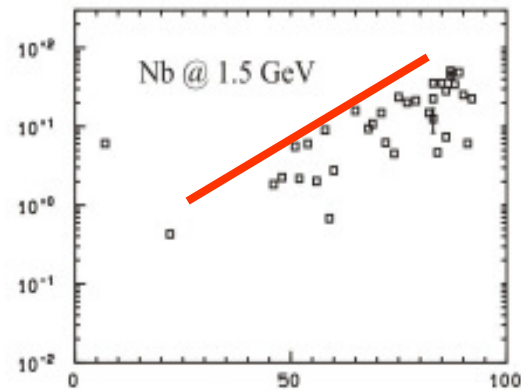
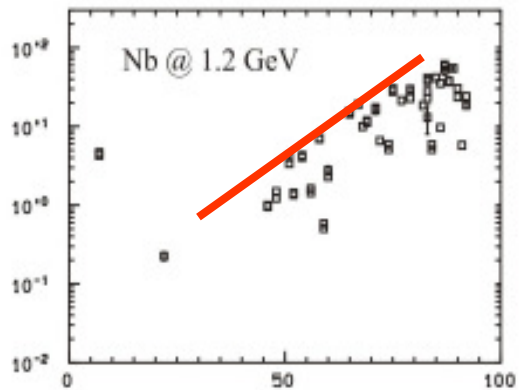
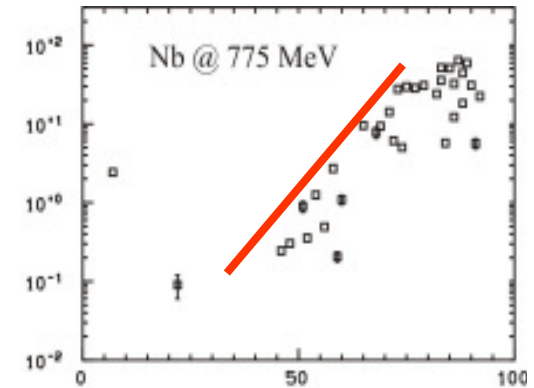
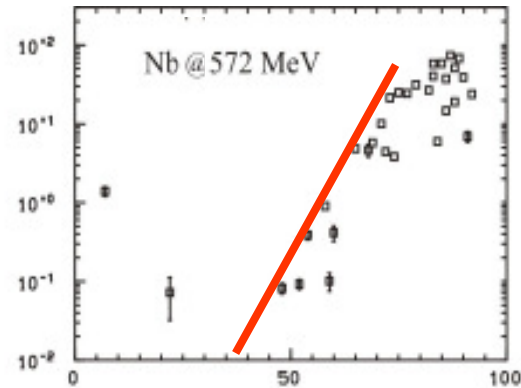
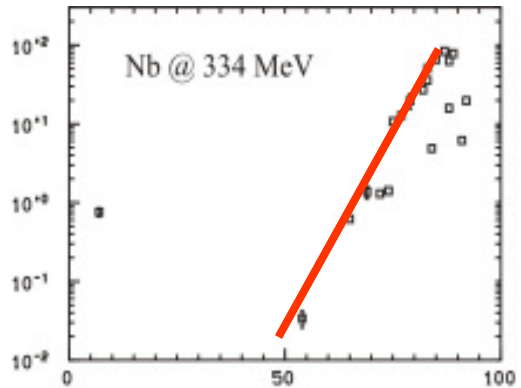


Bi(p,4n)Po-206



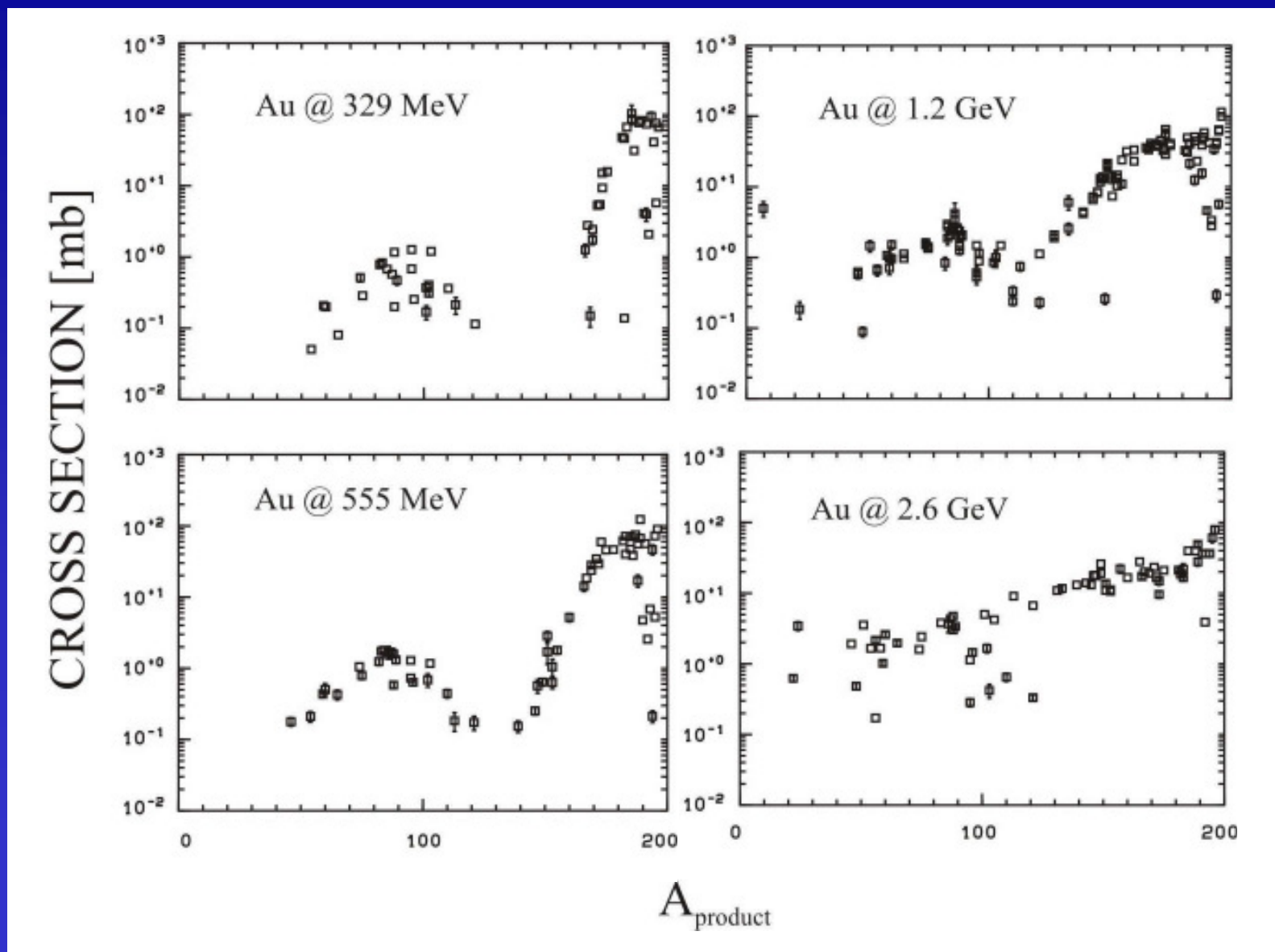
Nb(p,X) Isobaric Yields

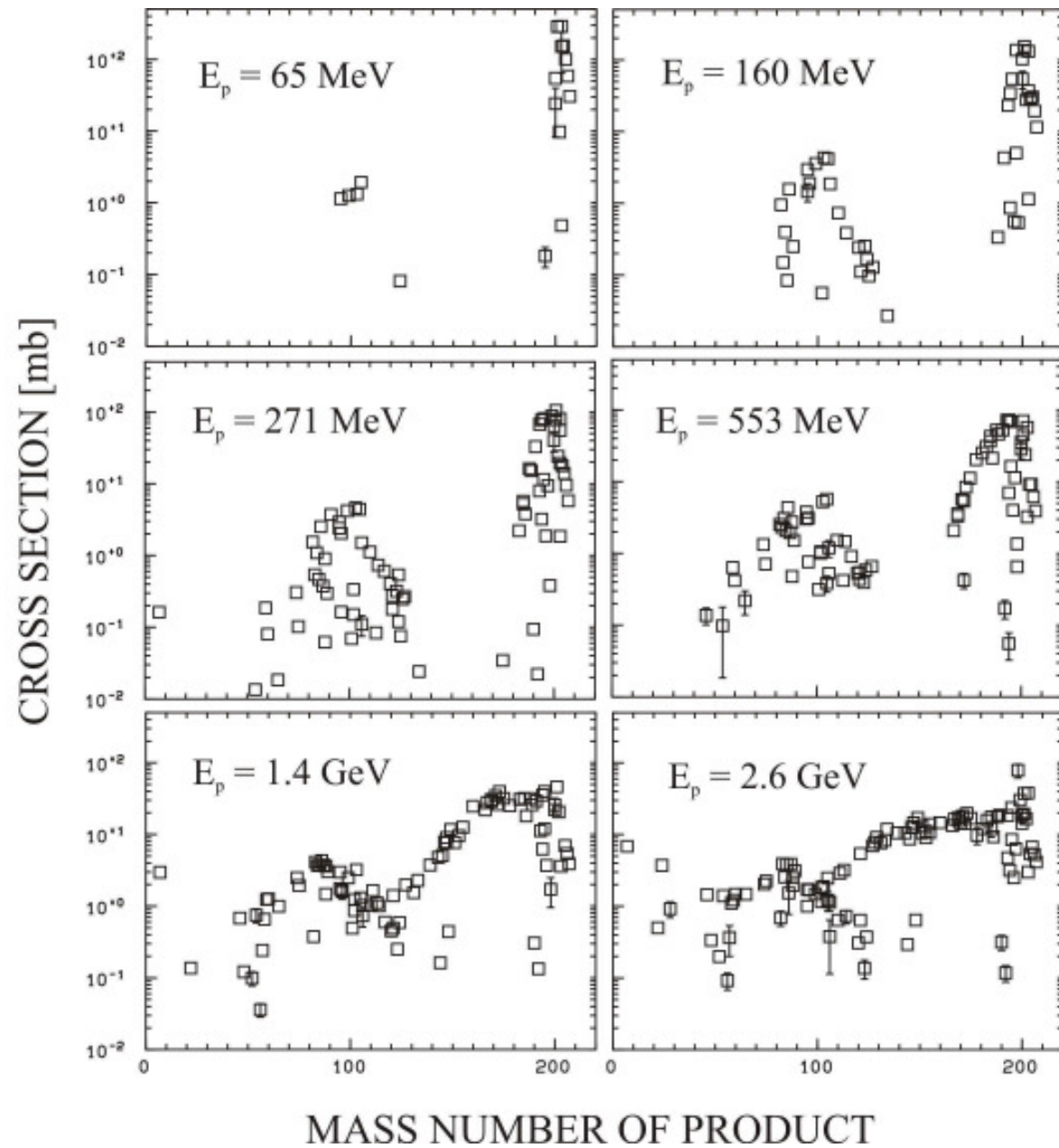
Cross section in mb



Product mass number

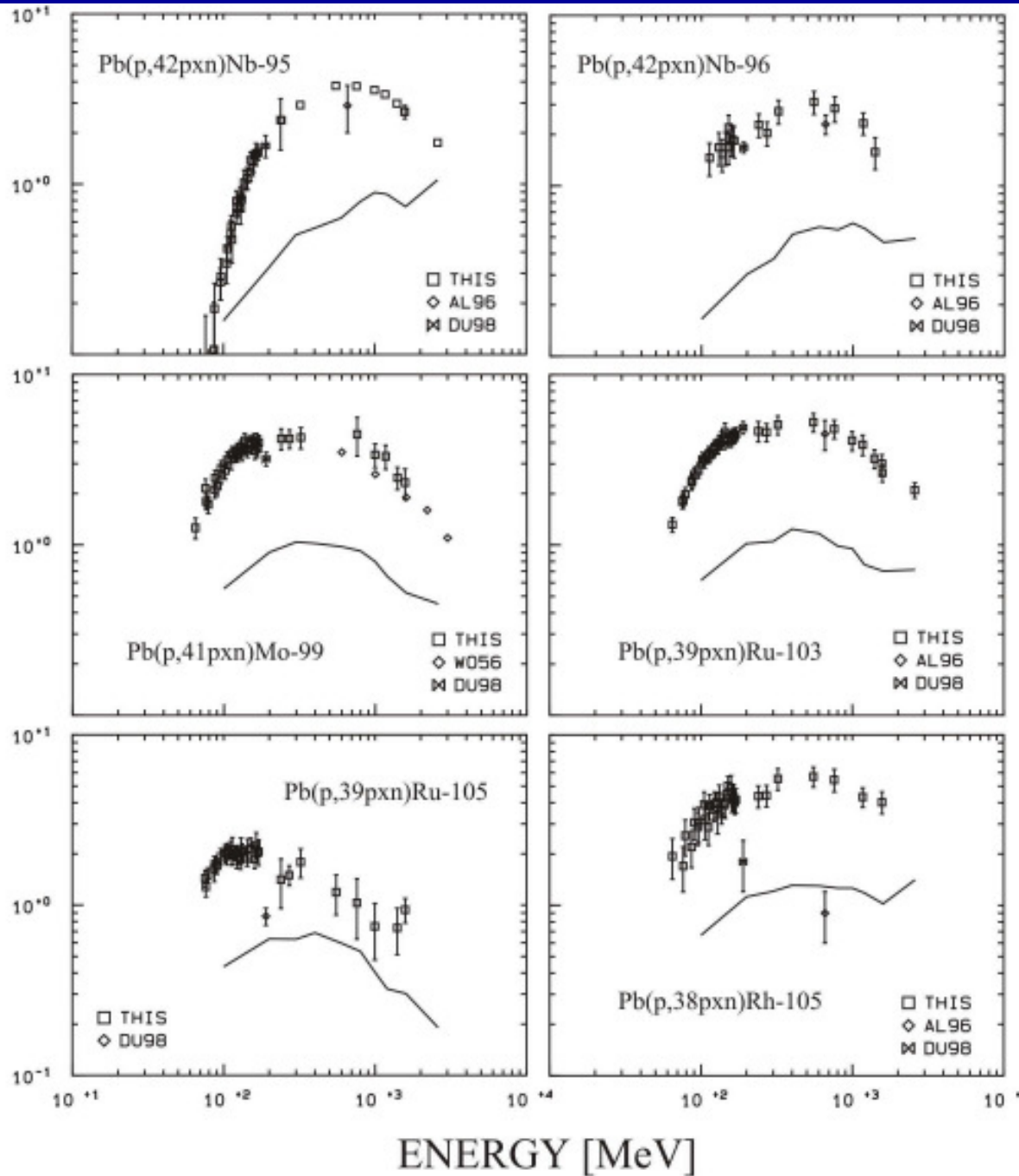
Cross Sections versus Product Mass Numbers for Au(p,X)





**Cross sections
versus product
mass numbers
for $\text{Pb}(p,X)$**

CROSS SECTION [mb]

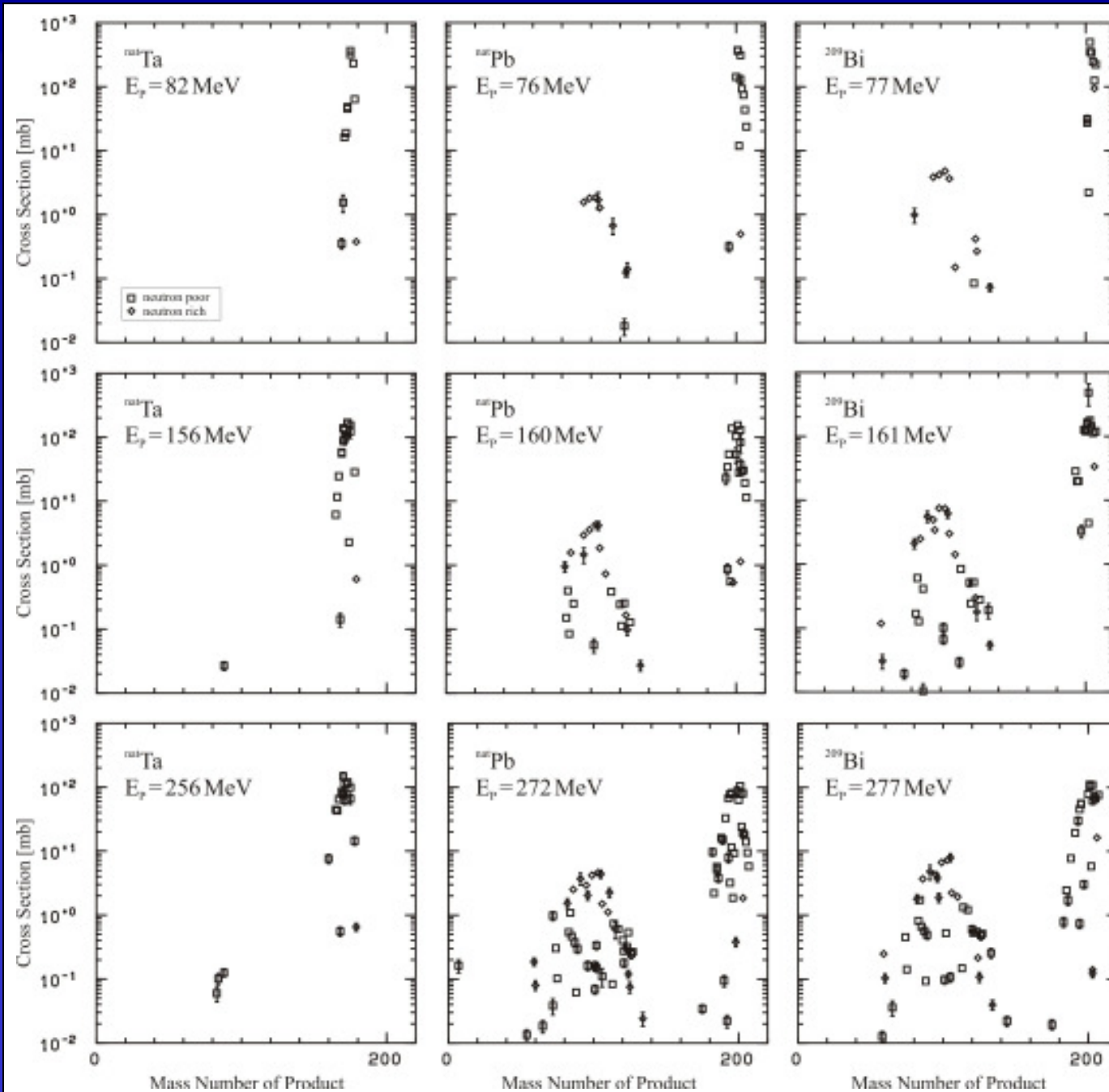


Medium
energy
fission of
lead

Theory:
LAHET code

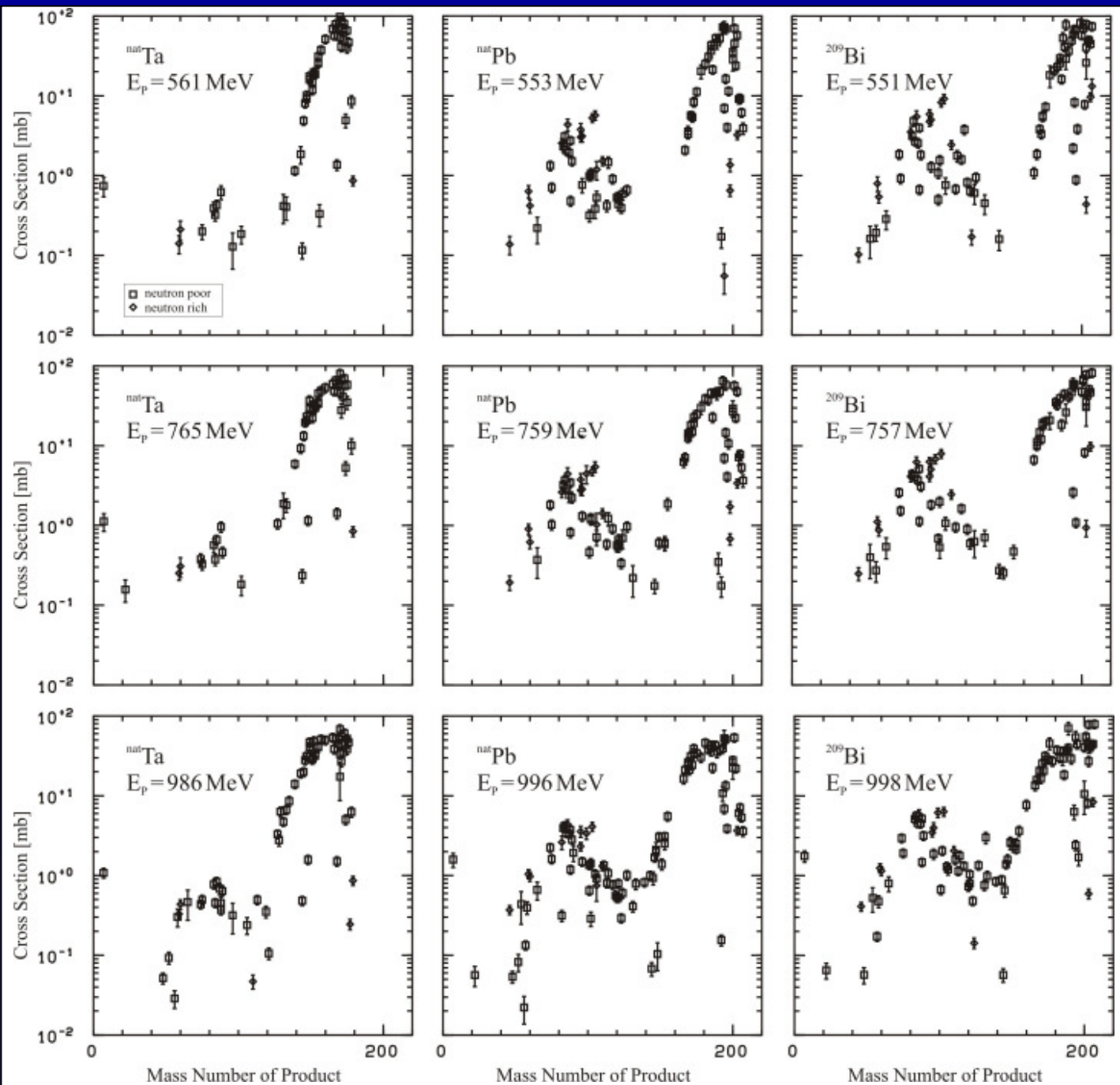
Gloris et al. (2001)

R. Micner, ZSR, Leibniz Universität Hannover



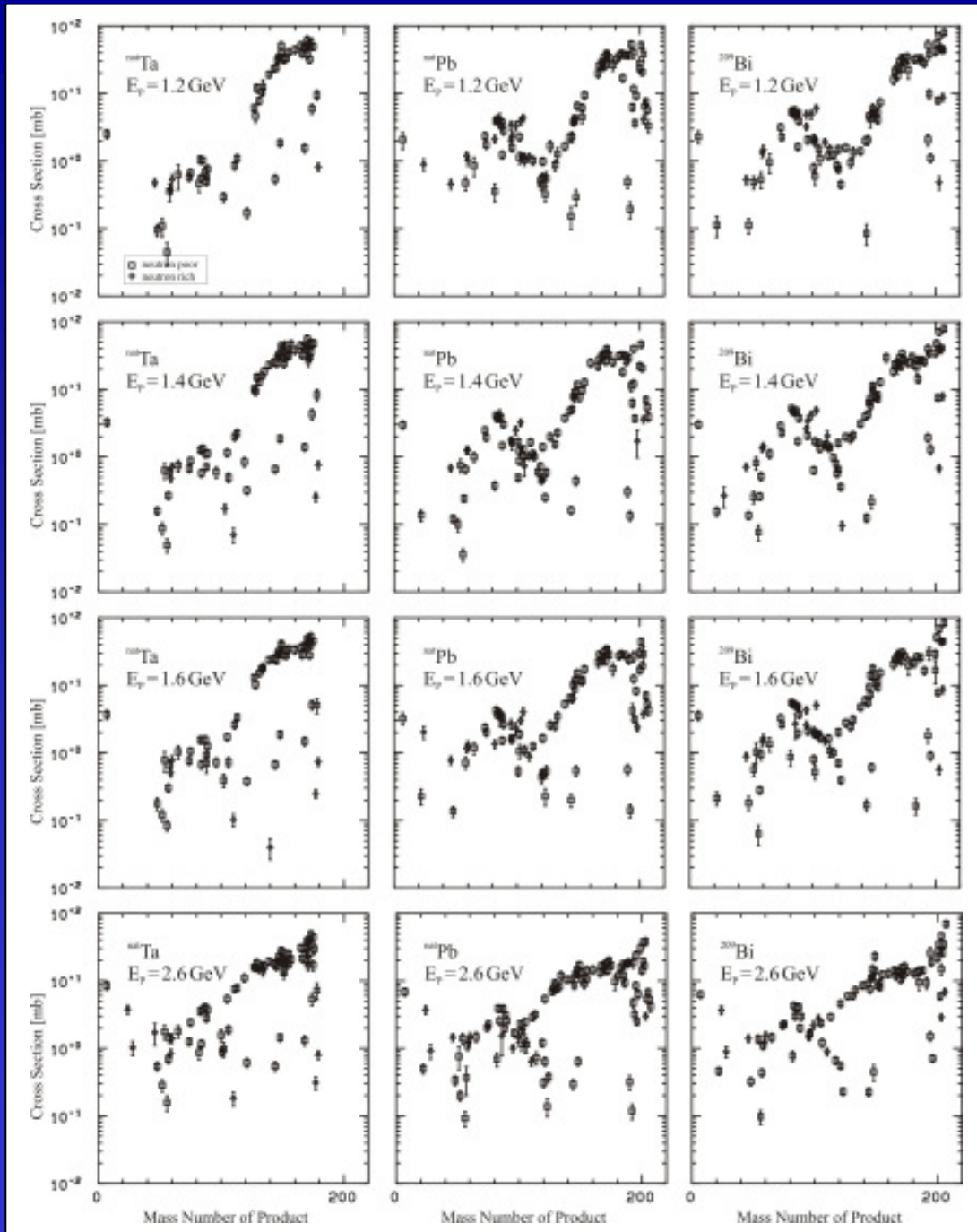
**Ta, Pb, Bi
cross
sections
versus
product
mass
numbers**

**76 - 300
MeV**



**Ta, Pb,
Bi cross
sections
versus
product
mass
numbers**

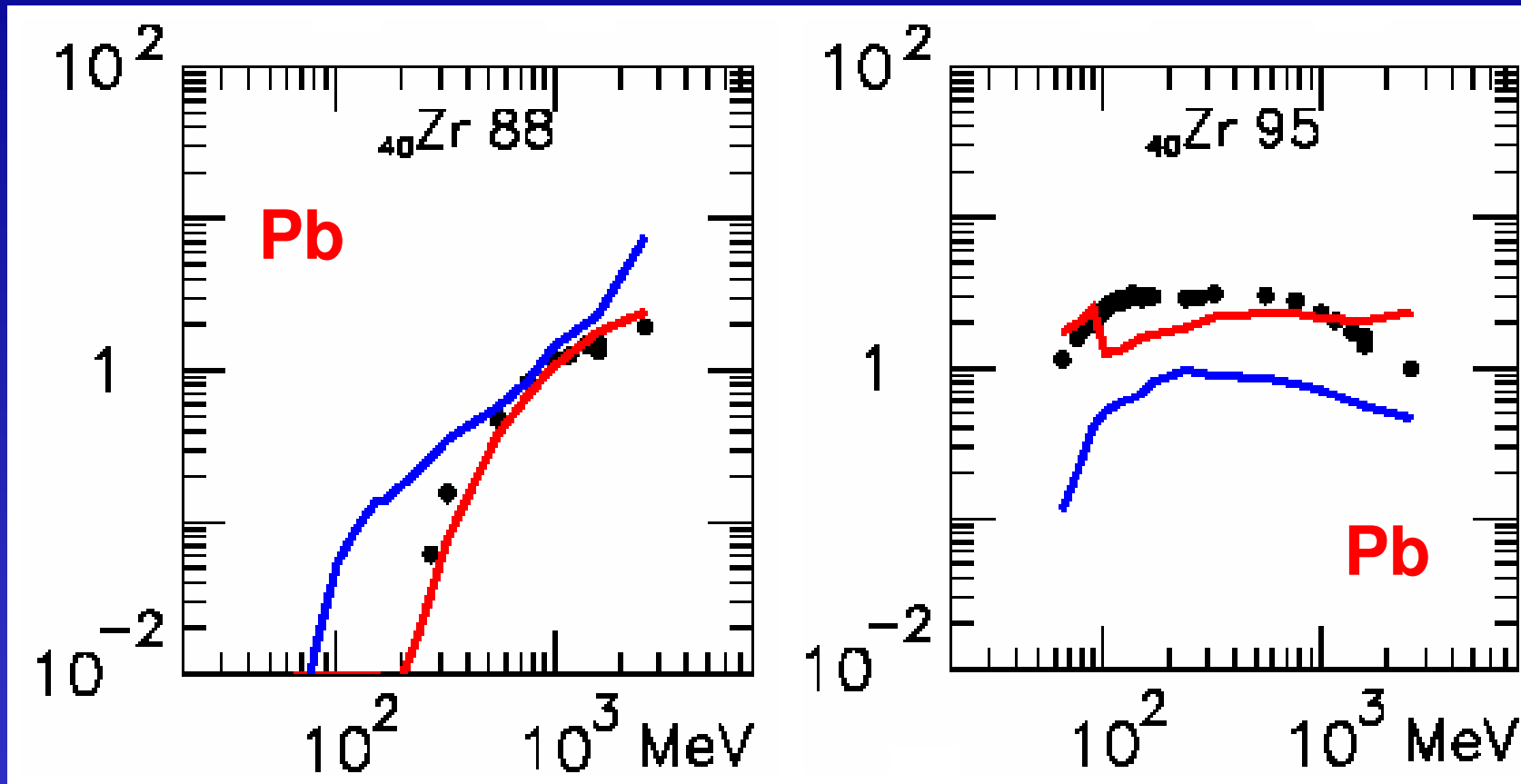
**560 - 996
MeV**



Ta, Pb, Bi
cross sections
versus product
mass numbers

1.2 - 2.6 GeV

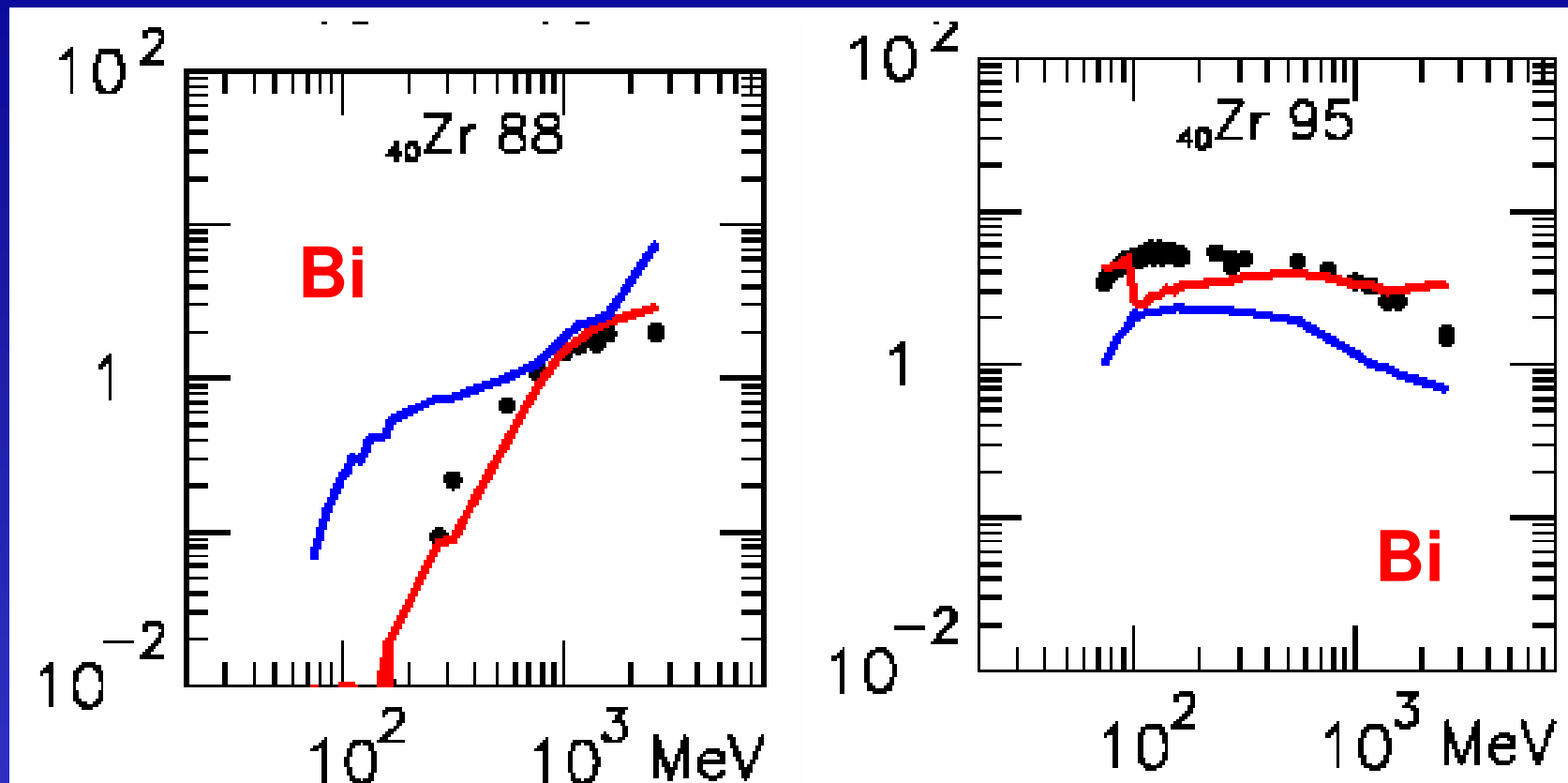
Production of ^{88}Zr and ^{95}Zr from Lead



Nb 89 66 m 2.0 h 1832... 207... 2093... 0	Nb 90 15.8 s 14.6 h β^+ 1.5... 1126... 2218... 141... 1285	Nb 91 95.9 d 690 a β^+ (105) α 2.1... 141... 1285	Nb 92 10.15 d 3.5... 10 ⁵ a 1034... 934	Nb 93 16.13 a 100 β^+ (91) α 0.1... 1.5	Nb 94 5.26 m 2 · 10 ⁴ a β^+ (41) α 0.5... 87... 703... 103... 14.4	Nb 95 86.6 h 34.97 d β^+ 238... α 0.2... 5.9... 776, 589... 1091...	Nb 96 23.4 h β^+ 1.3... 158	Nb 97 53 s 74 m β^+ 2.0... 3.9... 787... 728... 1189... 1024	Nb 98 51 m 2.9 s β^+ 4.5... 1787... 787... 1024
Zr 88 83.4 d 4.16 s 78.4 h 1568... 1173... 1187... 393	Zr 89 4.16 s 78.4 h β^+ 0.9... 1173... 1187... 393	Zr 90 51.45 α 0.014	Zr 91 11.22 α 1.2	Zr 92 17.15 α 0.2	Zr 93 1.5 · 10 ⁵ a β^+ 0.06... m α 2	Zr 94 17.3... α 0.049	Zr 95 64.0 d β^+ 0.4; 1.1... m γ 757, 724... 9	Zr 96 2.80 3.9 · 10 ¹⁹ a α 0.019	Zr 97 16.8 h β^+ 1.9... γ 506; 1148; 355... m
Y 87 80... β^+ 4... 495... m γ 1036; 898...	Y 88 106.6 d β^+ 0.057... 1.25	Y 89 16.9 s 100 α 0.057... 1.25	Y 90 3.19 h 64.1 h β^+ 2.3... γ (219)... 1218...	Y 91 49.7 m 58.5 d β^+ 1.5... 11050 m α 1.4	Y 92 3.54 h β^+ 3.6... γ 934; 1405; 561; 449...	Y 93 10.1 h β^+ 2.9... γ 919; 1139; 1918...	Y 94 10.3 m β^+ 4.9... γ 254; 2178; 3577; 1324; 2633...	Y 95 10.3 m β^+ 2.8... γ 919; 1139; 1918...	Y 96 9.8 s 5.31 s β^+ 7.1... 1766...

— INCL4 + ABLA
— Bertini-Dresner-PE

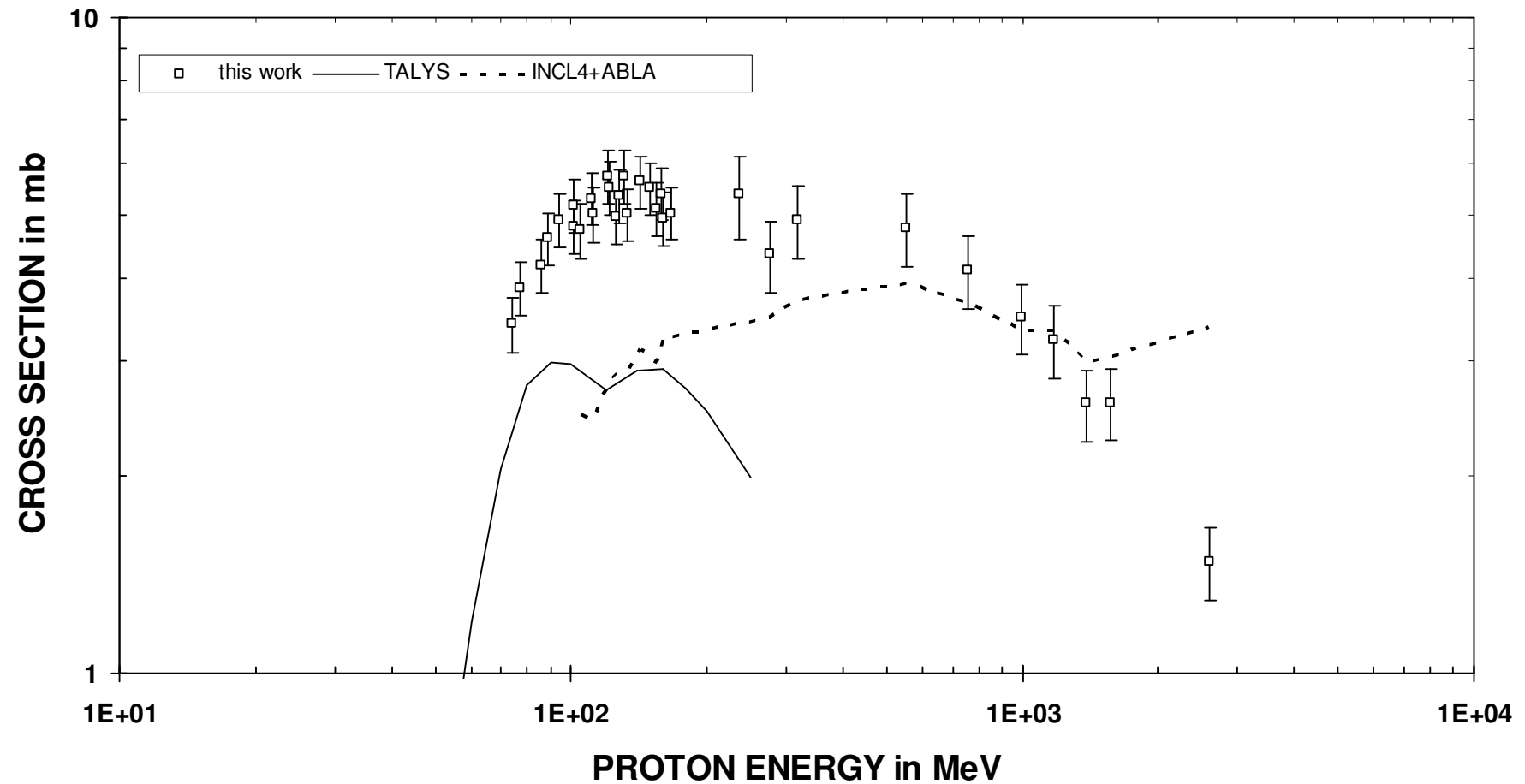
Production of ^{88}Zr and ^{95}Zr from Bismuth



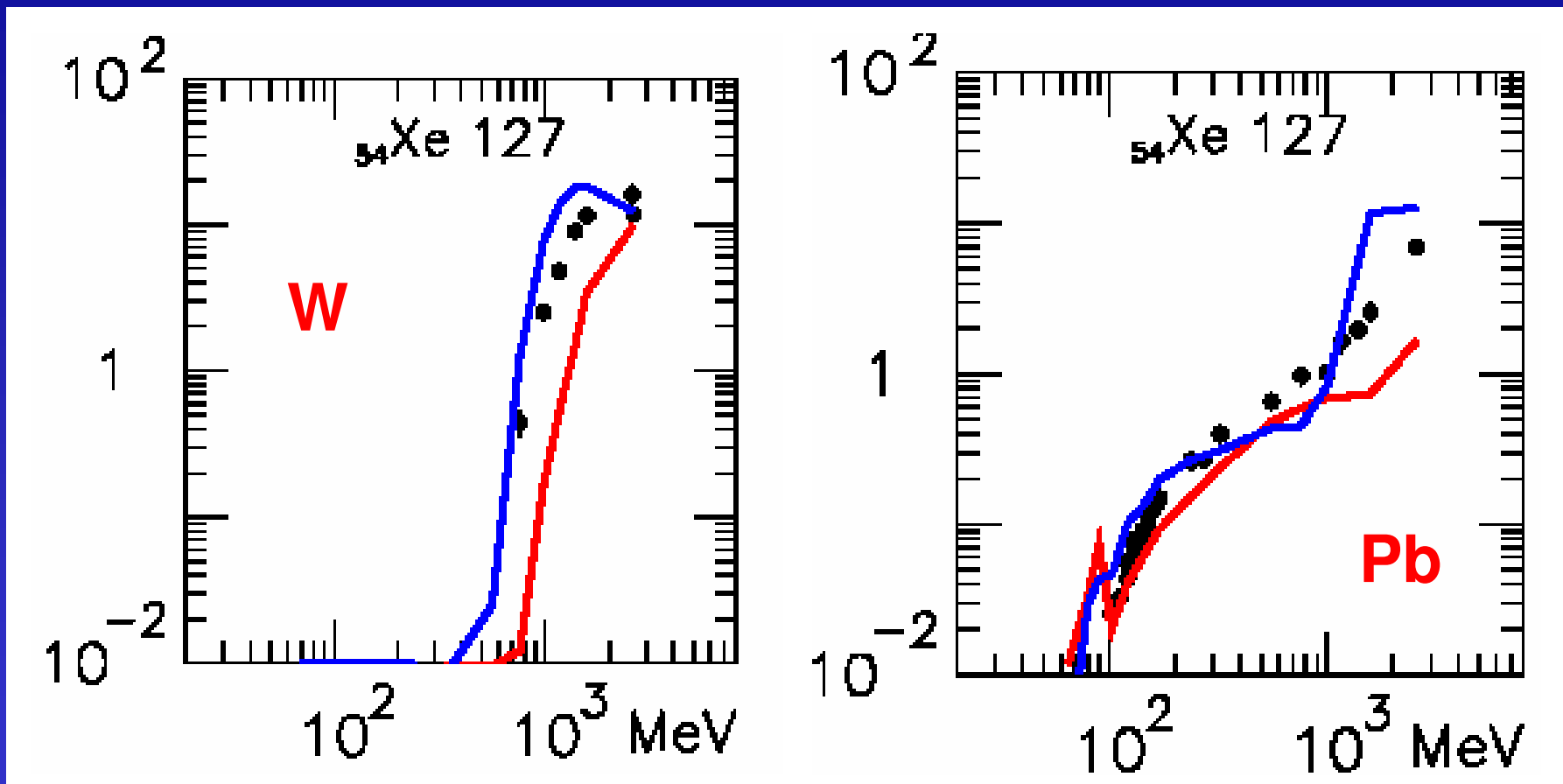
Nb 89 66 m 2.0 h 2.9 s 1832 2092 0	Nb 90 15.5 s 14.6 h $\beta^+ 1.5$ $\gamma 1125$ 2218 141	Nb 91 95.9 d 690 a $\beta^+ (105)$ $\alpha 0.1$ $\gamma 1095$	Nb 92 10.15 d 3.6 · 10 ⁵ a $\beta^+ 0.1$ $\alpha 0.1$ $\gamma 924$	Nb 93 16.13 a 100 $\beta^+ 0.1$ $\alpha 0.1$ $\gamma 91$	Nb 94 5.26 m 2 · 10 ⁴ a $\beta^+ 0.5$ $\alpha 0.1$ $\gamma 238$ 303 14.4	Nb 95 86.6 h 34.97 d $\beta^+ 0.2$ $\alpha 0.1$ $\gamma 785$ 1091	Nb 96 23.4 h $\beta^+ 1.3$ $\alpha 0.1$ $\gamma 658$	Nb 97 53 s 74 m $\beta^+ 2.6$ 2.9 787 728 1149	Nb 98 51 m 2.9 s $\beta^+ 4.5$ $\alpha 0.1$ $\gamma 787$ 1024
Zr 88 83.4 d $\beta^+ 0.9$ $\gamma 393$	Zr 89 4.16 s 78.4 h $\beta^+ 0.9$ $\alpha 0.1$ $\gamma 1173$	Zr 90 51.45 $\alpha 0.014$	Zr 91 11.22 $\alpha 1.2$	Zr 92 17.15 $\alpha 0.2$	Zr 93 1.5 · 10 ⁵ a $\beta^+ 0.06$ m $\alpha 2$	Zr 94 17.3 $\alpha 0.049$	Zr 95 64.0 d $\beta^+ 0.4; 1.1$ $\gamma 757; 724$ g	Zr 96 2.80 $\beta^+ 1.9$ $\gamma 506; 1148$ 355 m	Zr 97 16.8 h $\beta^+ 1.9$ $\gamma 506; 1148$ 355 m
Y 87 80 $\beta^+ 4.95$ m $\gamma 1036; 898$	Y 88 106.6 d $\beta^+ 0.007$ 1.25	Y 89 16.0 s 100 $\alpha 0.007$ 1.25	Y 90 3.19 h 64.1 h $\beta^+ 2.3$ $\gamma (2319)$ $\gamma (2198)$	Y 91 49.7 m 58.5 d $\beta^+ 1.5$ $\gamma (1028)$ $\alpha 1.4$	Y 92 3.54 h $\beta^+ 3.6$ $\gamma 934; 1405$ 561; 449	Y 93 10.1 h $\beta^+ 2.9$ $\gamma 257; 947$ 1918	Y 94 10.7 $\beta^+ 4.9$ $\gamma 919; 1139$ 551	Y 95 10.3 m $\beta^+ 4.4$ $\gamma 954; 2176$ 3577; 1324 2633	Y 96 9.8 s 534 s $\beta^+ 2.8$ $\gamma 1761$ 915; 617 1197

— INCL4 + ABLA
— Bertini-Dresner-PE

Bi-209(p, f)Zr-95

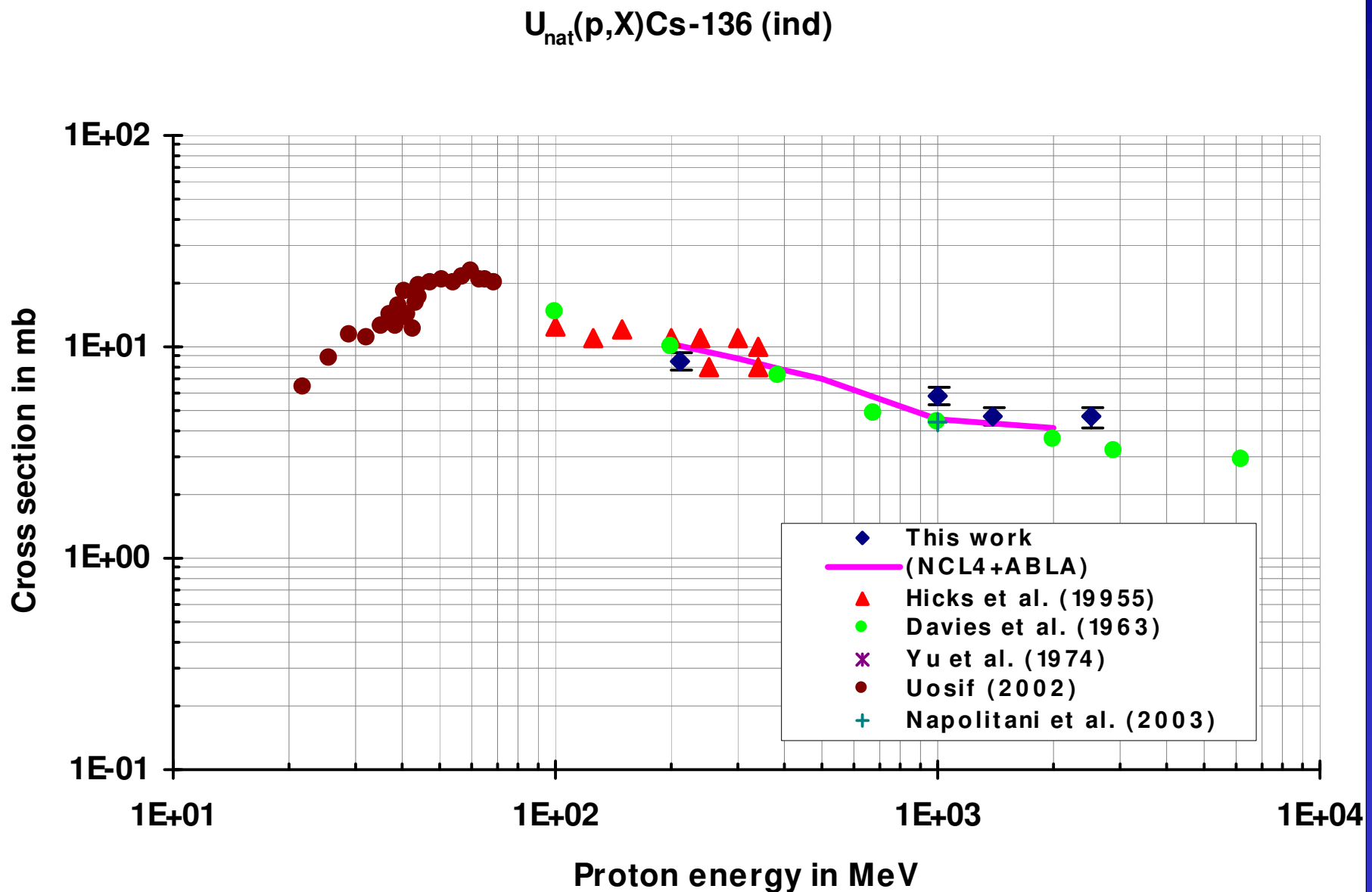


Production of ^{127}Xe

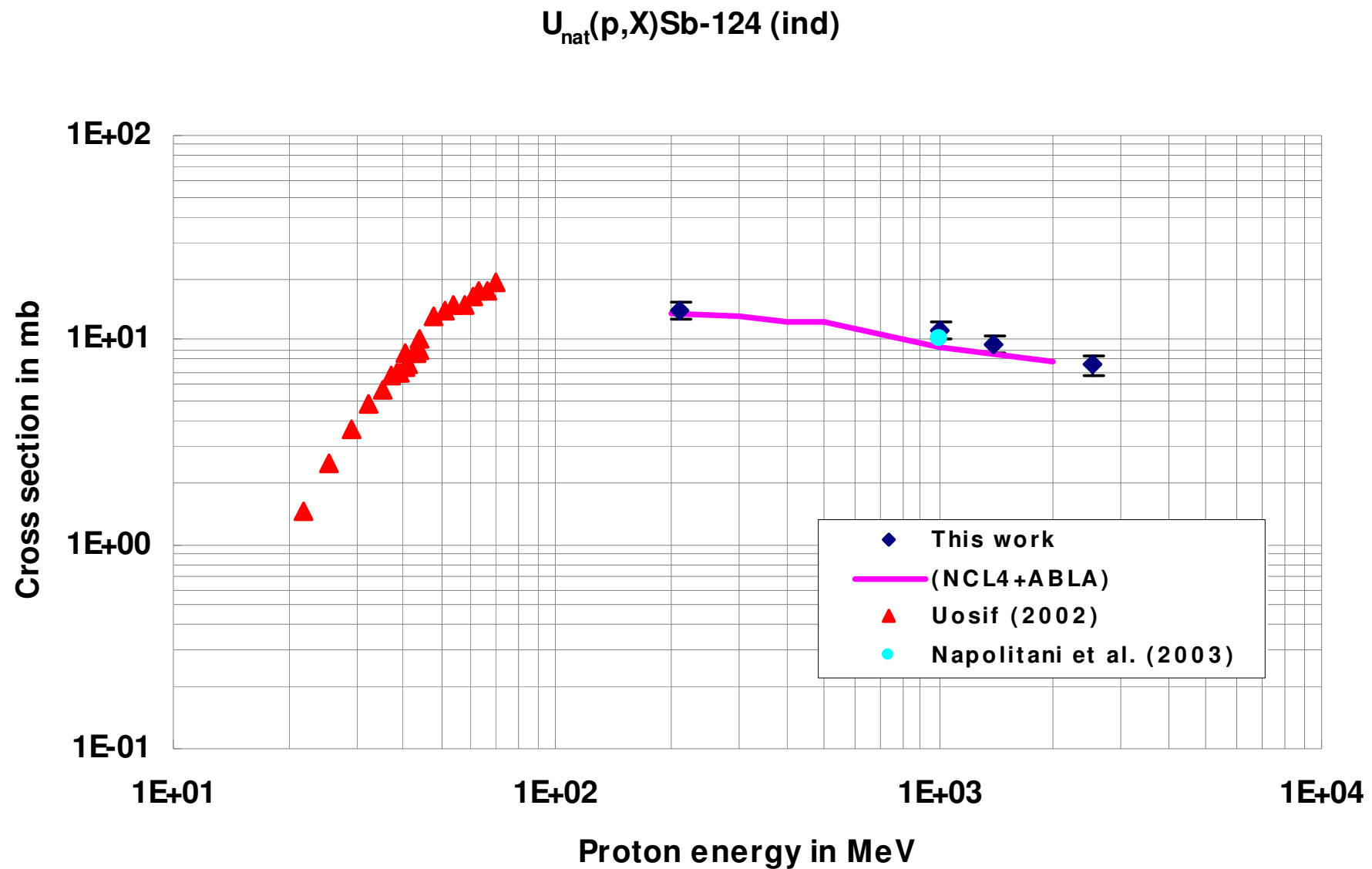


— INCL4 + ABLA
— Bertini-Dresner-PE

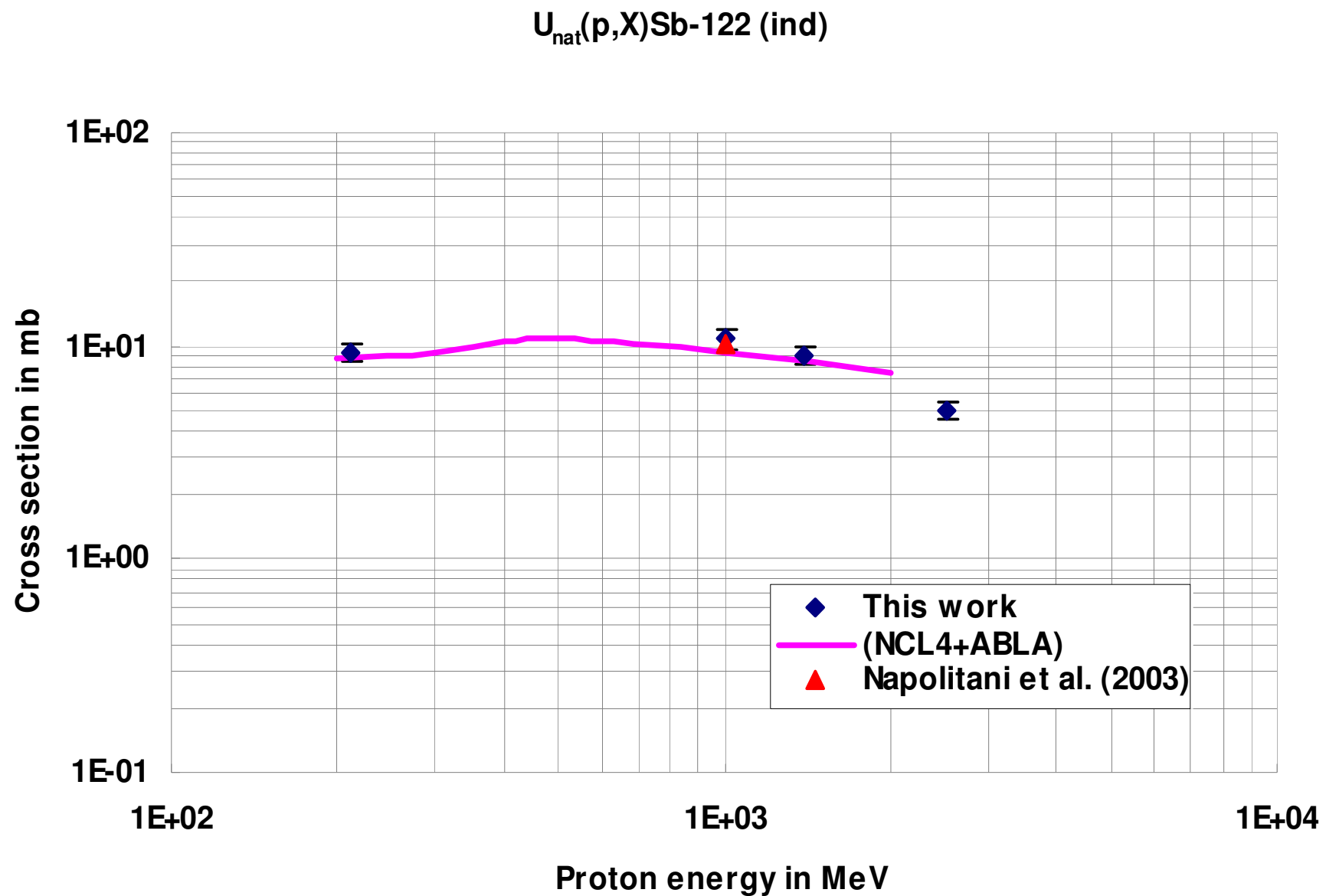
Proton-induced reactions on Uranium



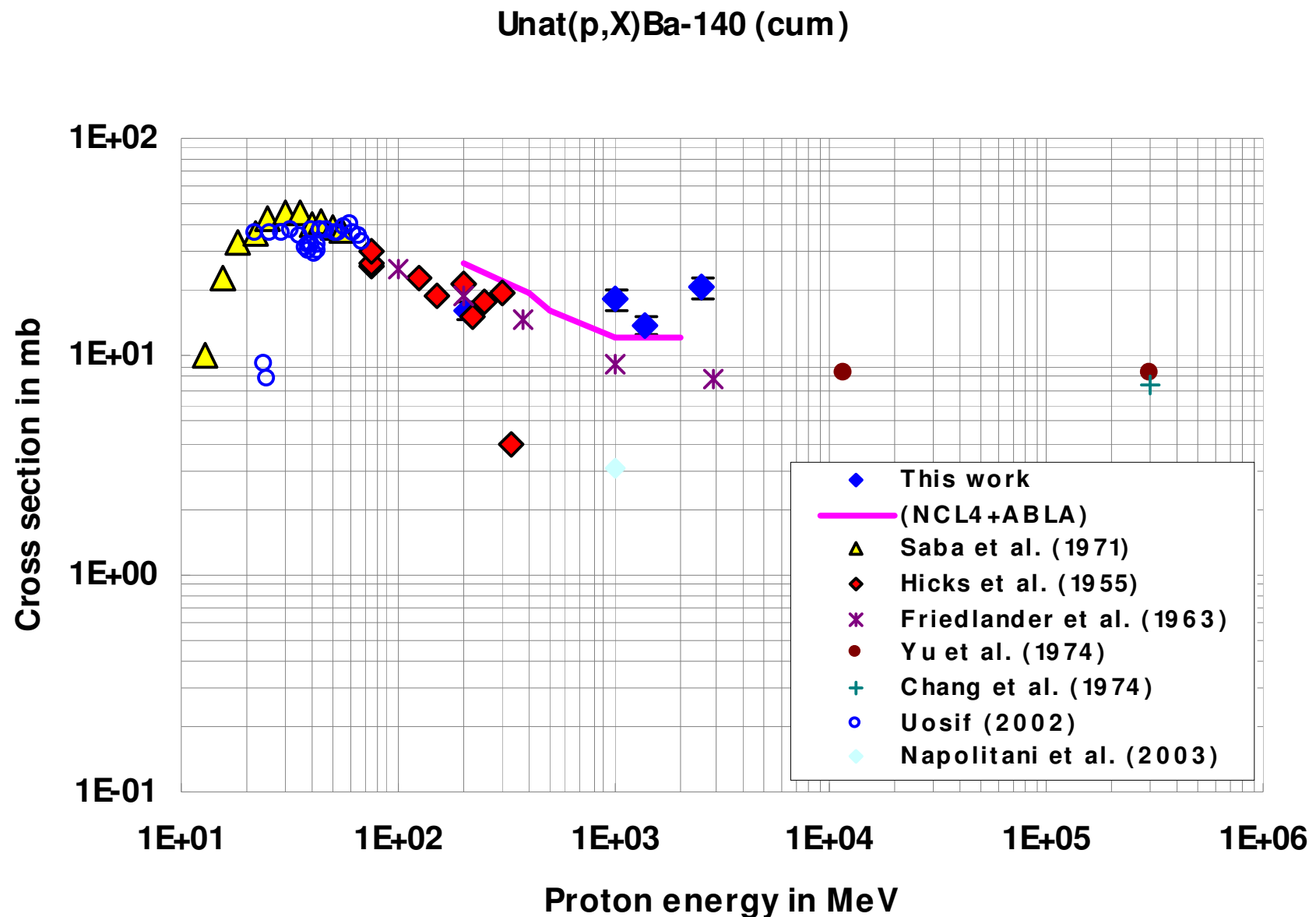
Proton-induced reactions on Uranium



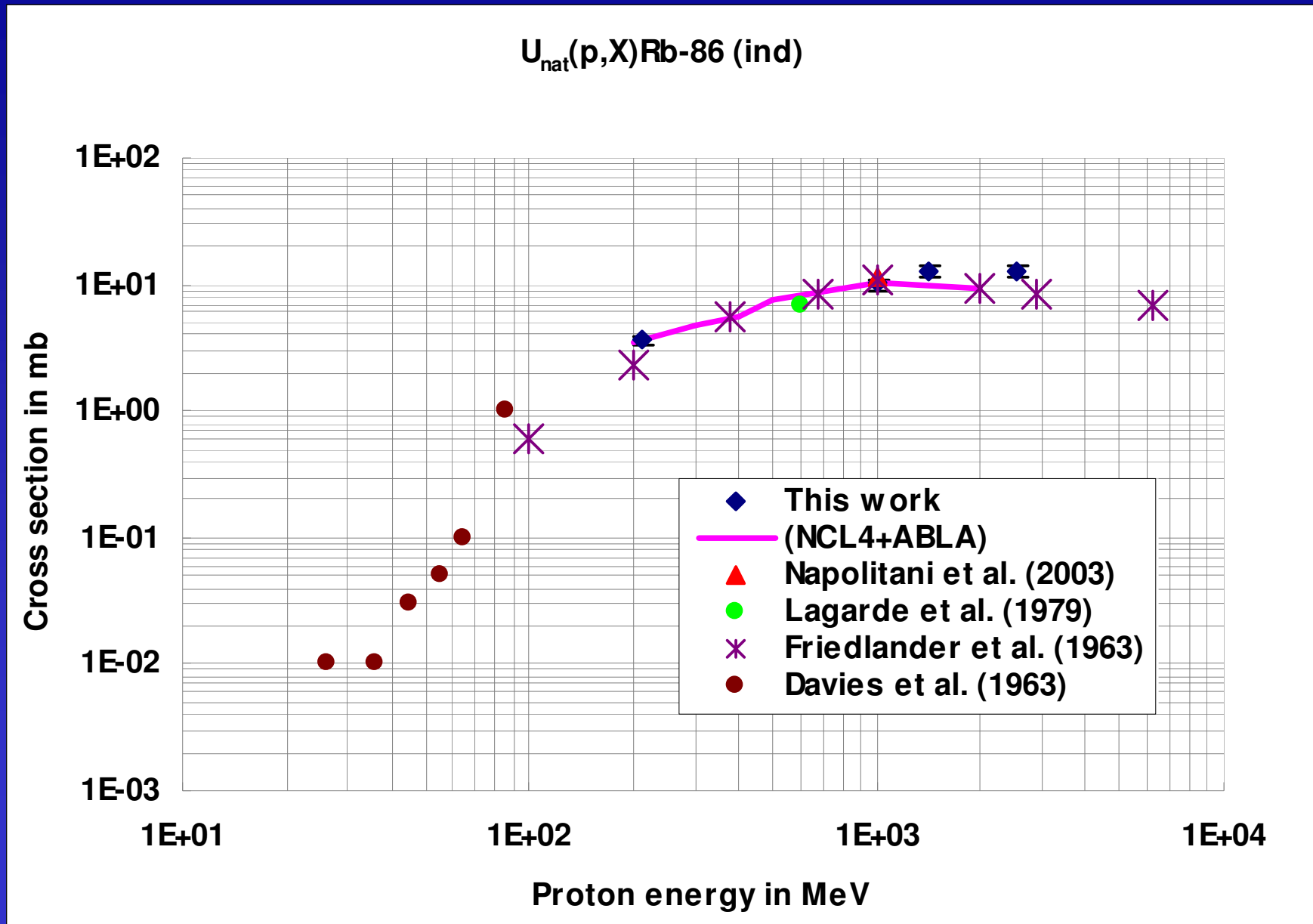
Proton-induced reactions on Uranium



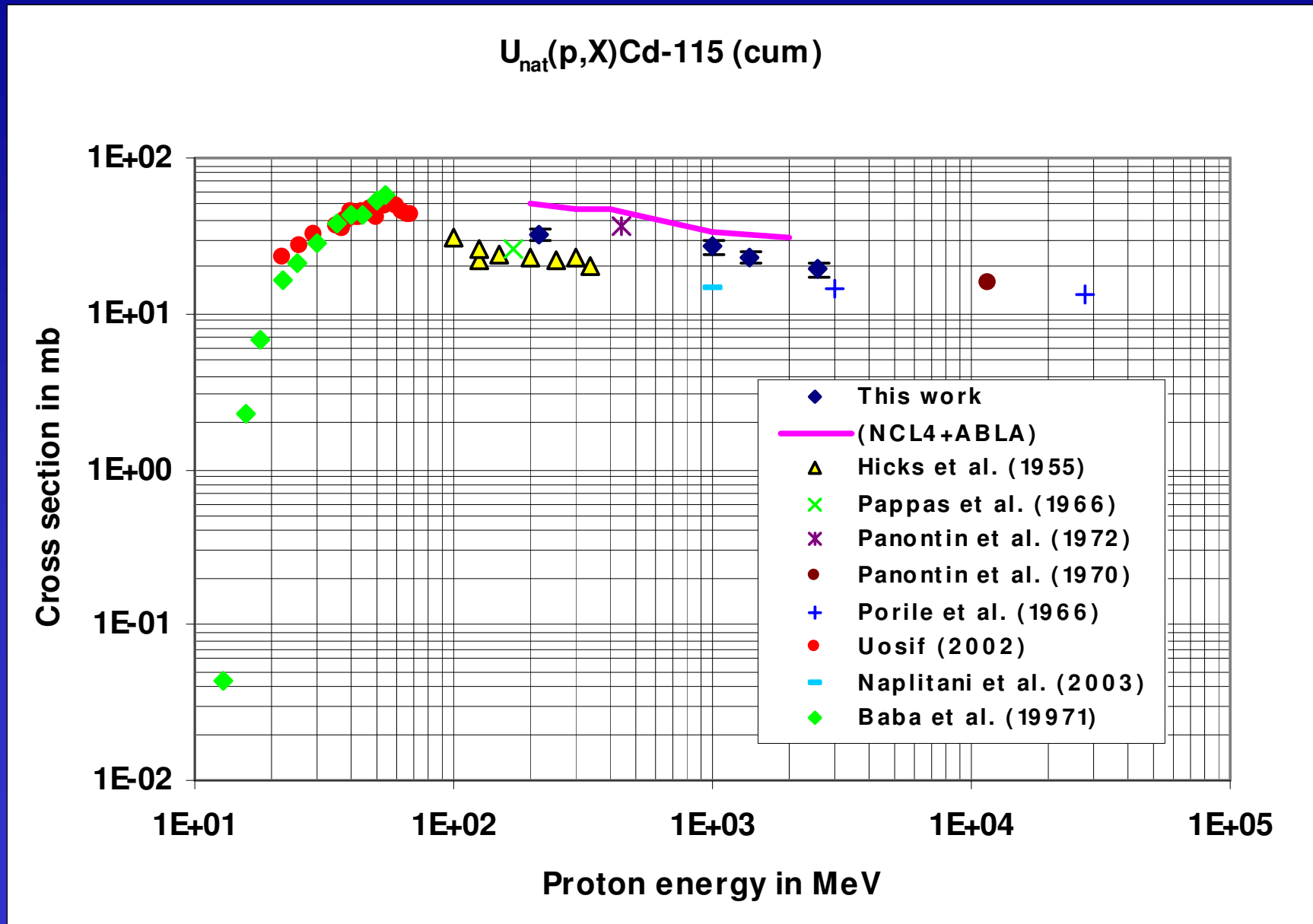
Proton-induced reactions on Uranium



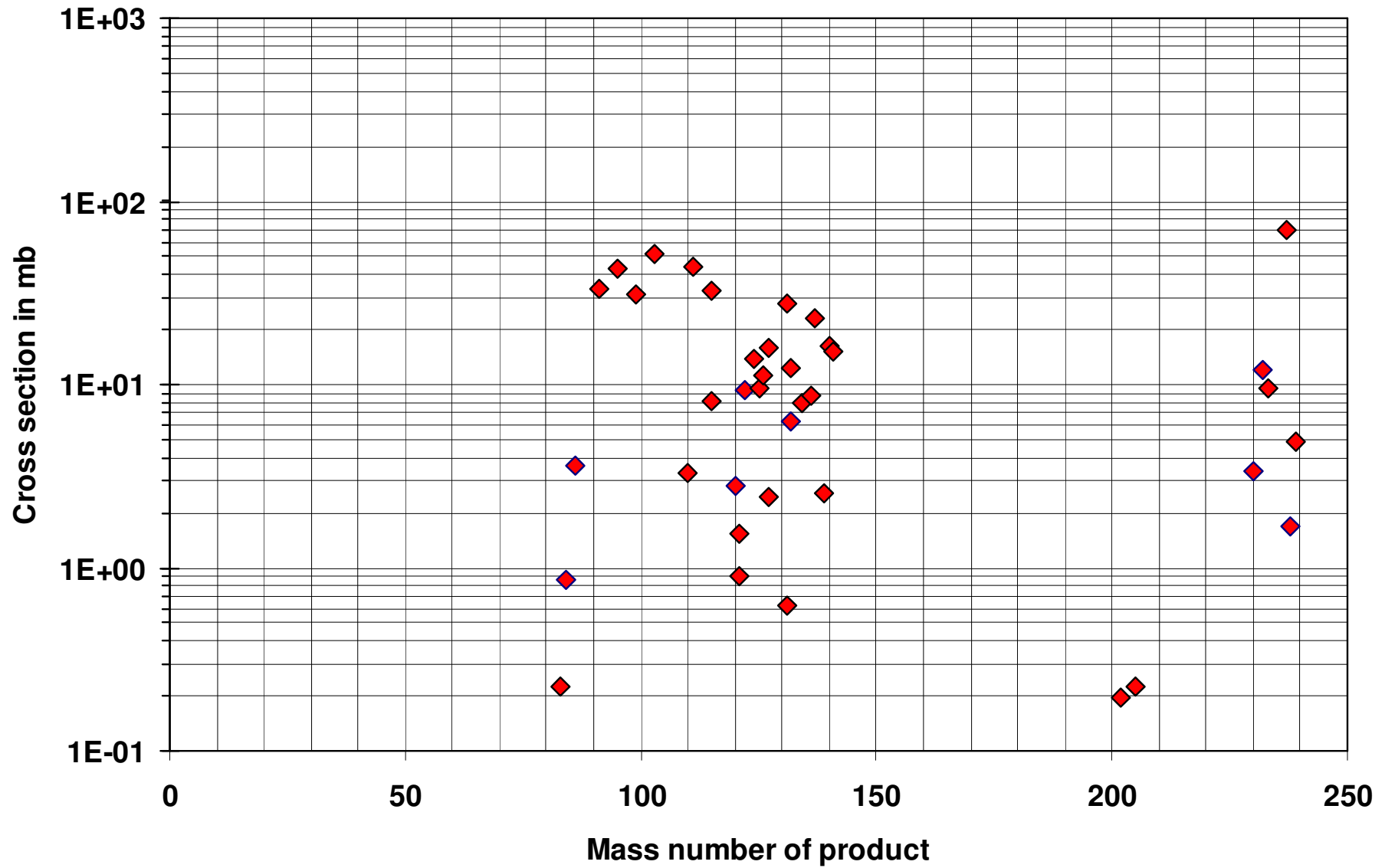
Proton-induced reactions on Uranium



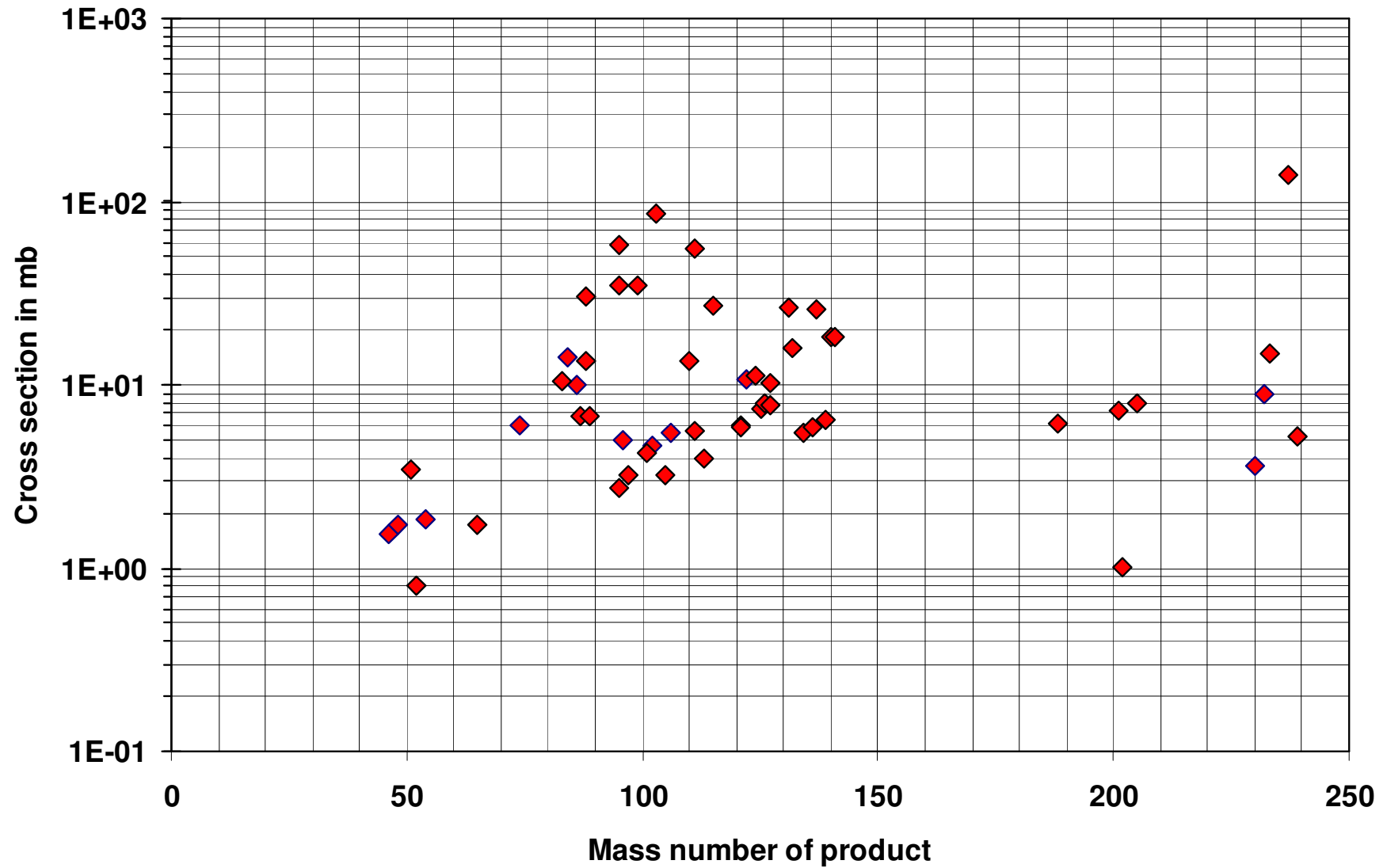
Proton-induced reactions on Uranium



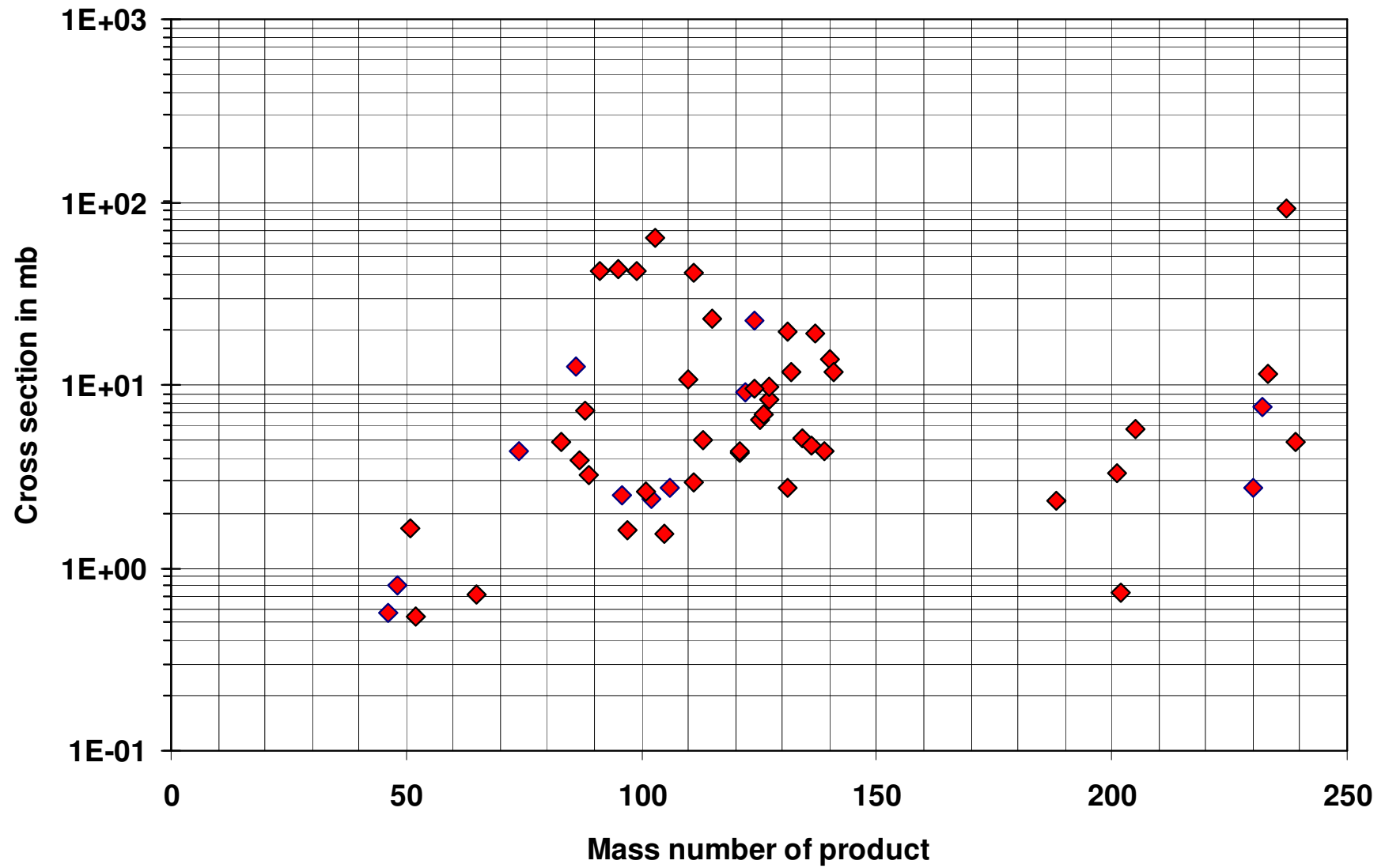
$U_{\text{nat}}(p,X) @ 211 \text{ MeV}$



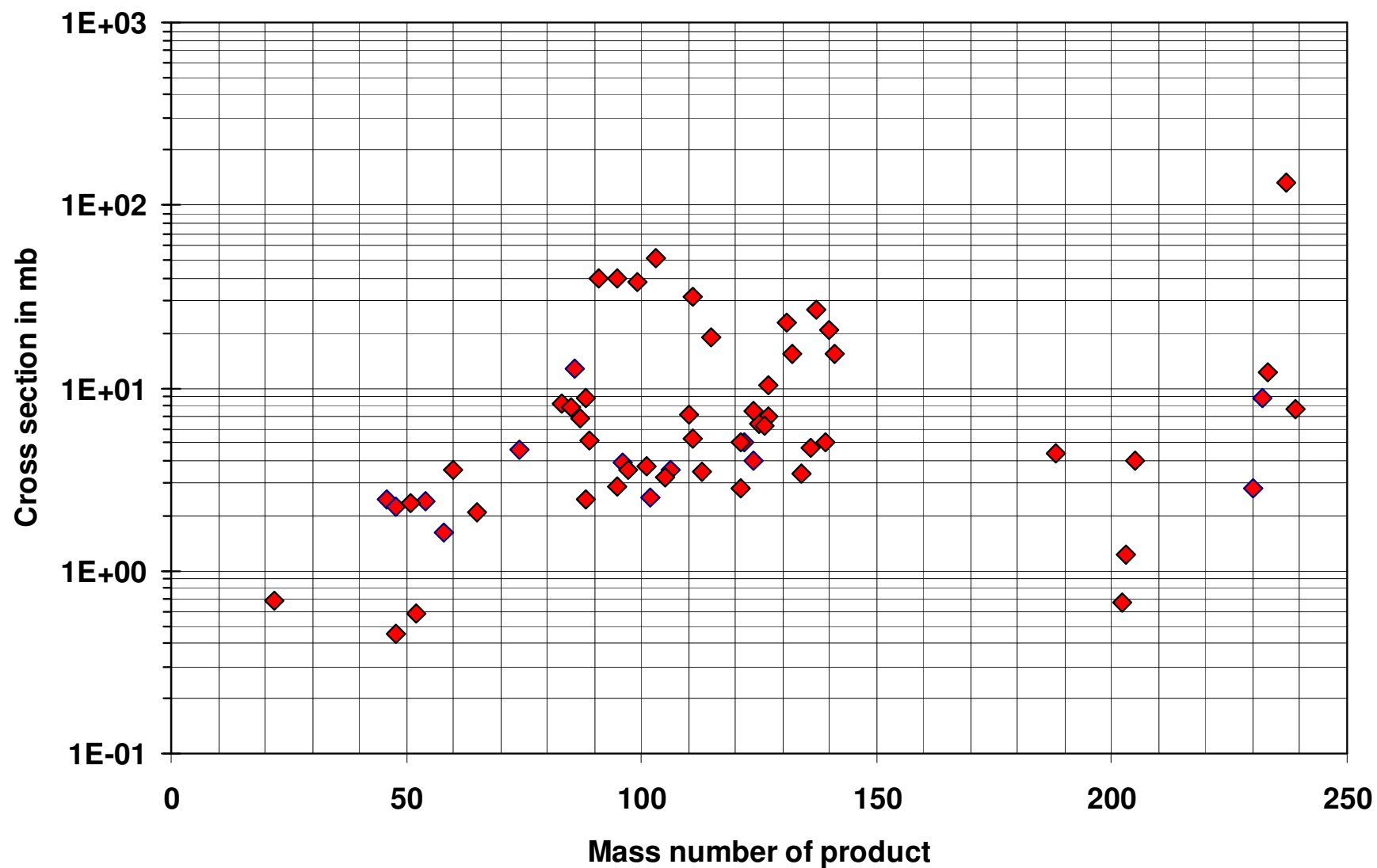
$U_{\text{nat}}(p,X) @ 1000 \text{ MeV}$



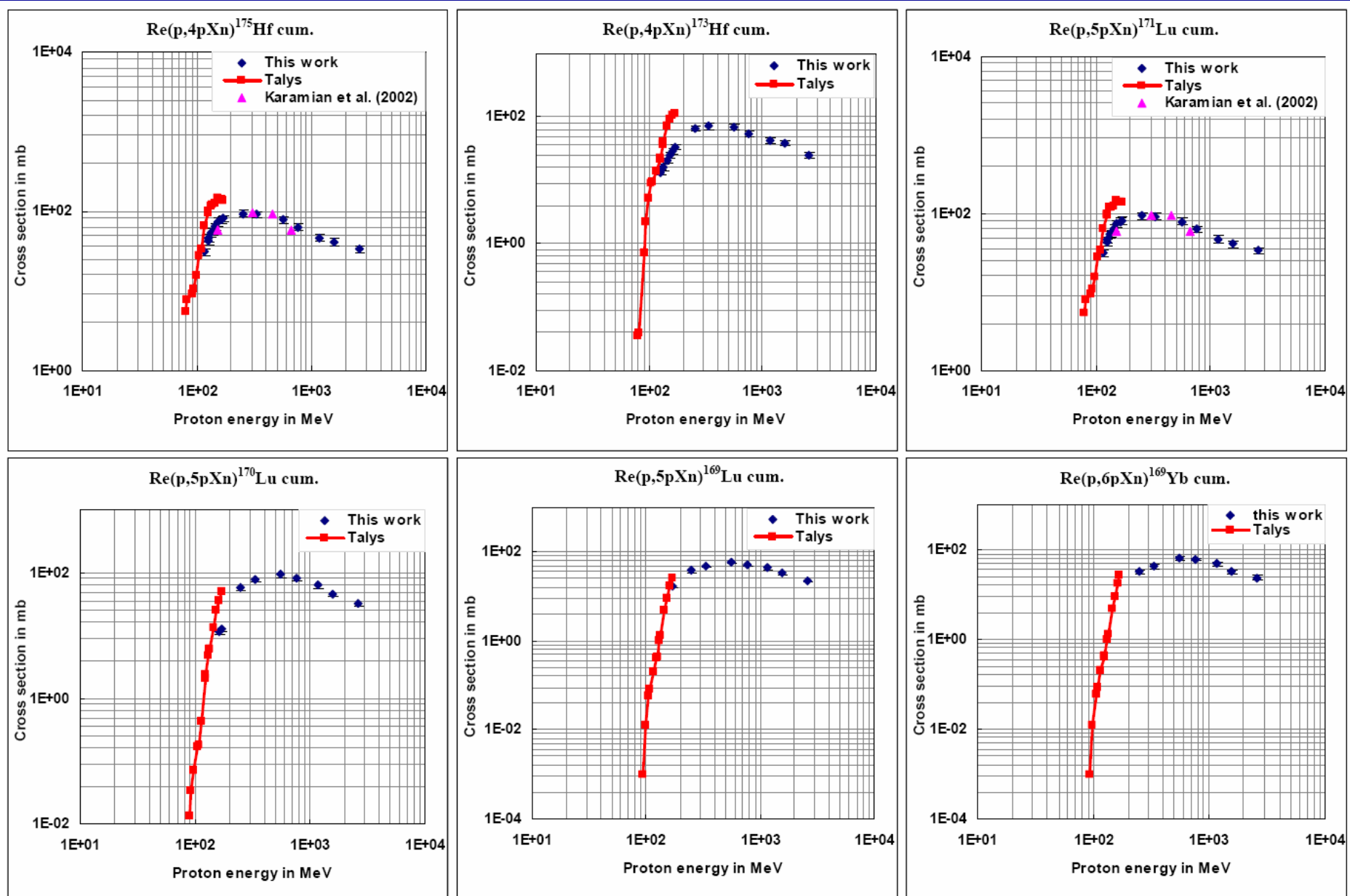
$U_{\text{nat}}(p,X) @ 1400 \text{ MeV}$



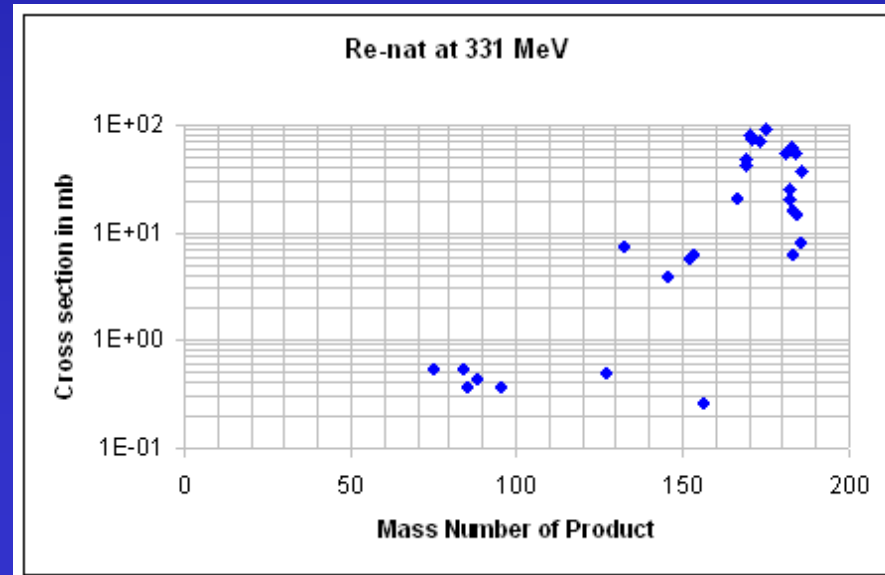
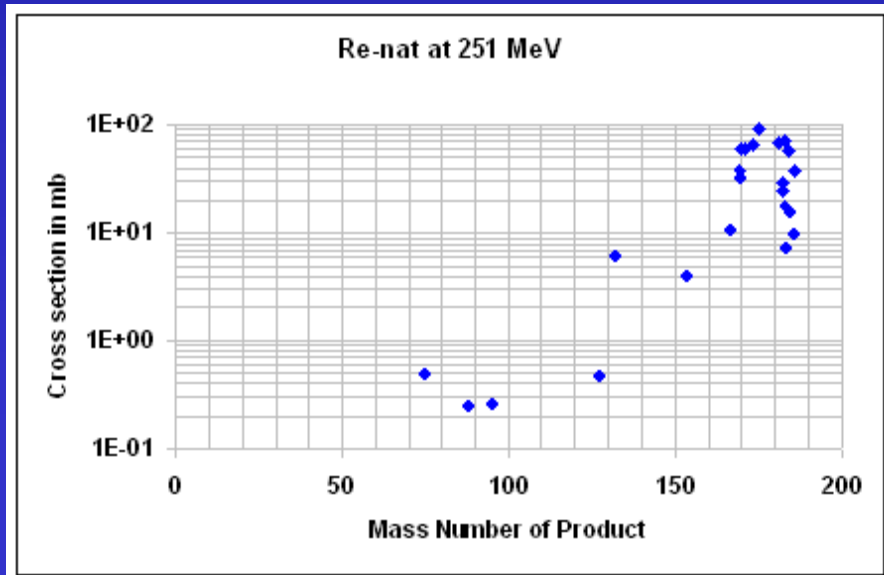
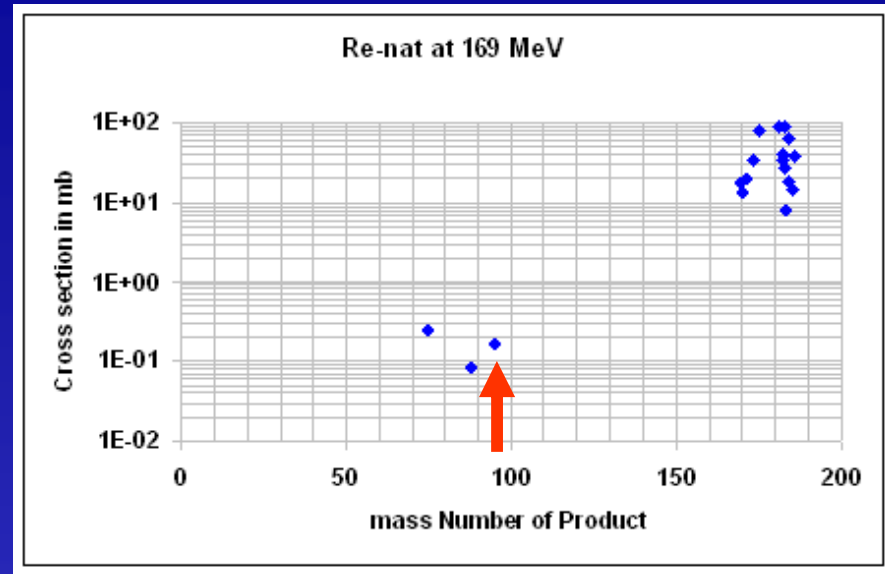
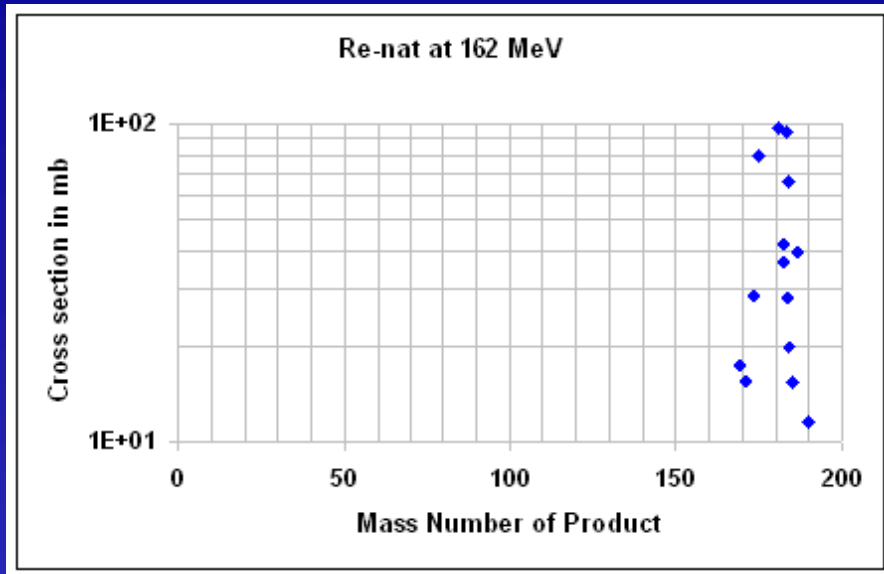
$U_{\text{nat}}(p,X) @ 2530 \text{ MeV}$



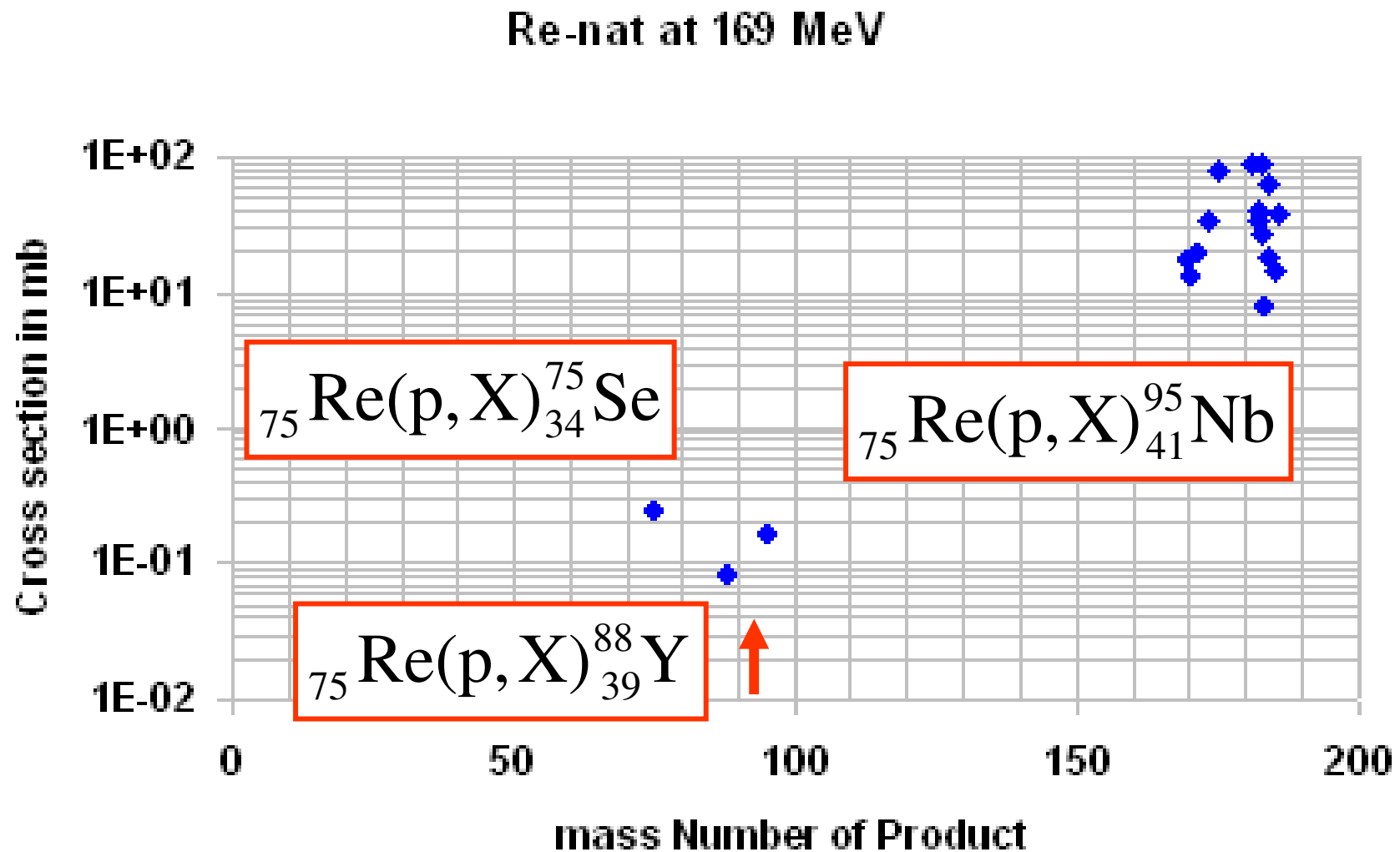
Proton-induced reactions on Rhenium



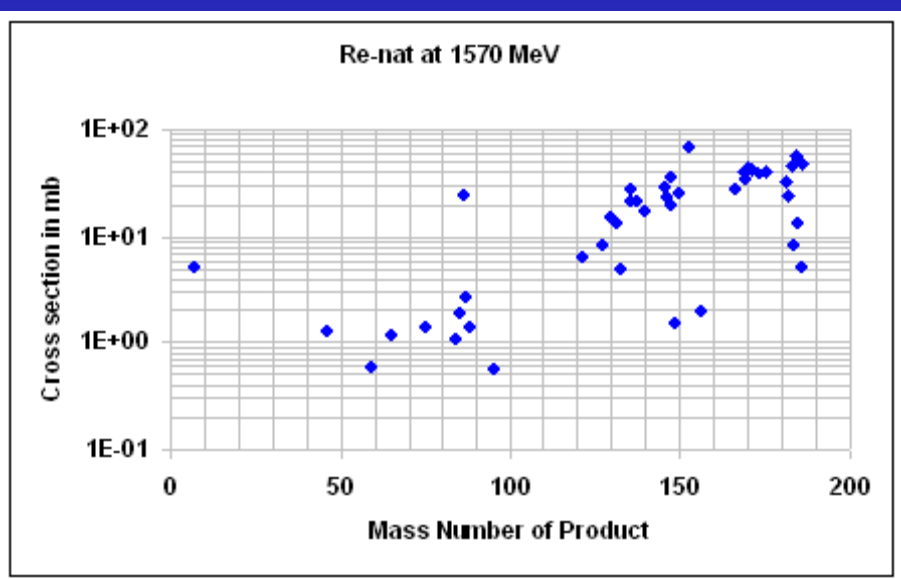
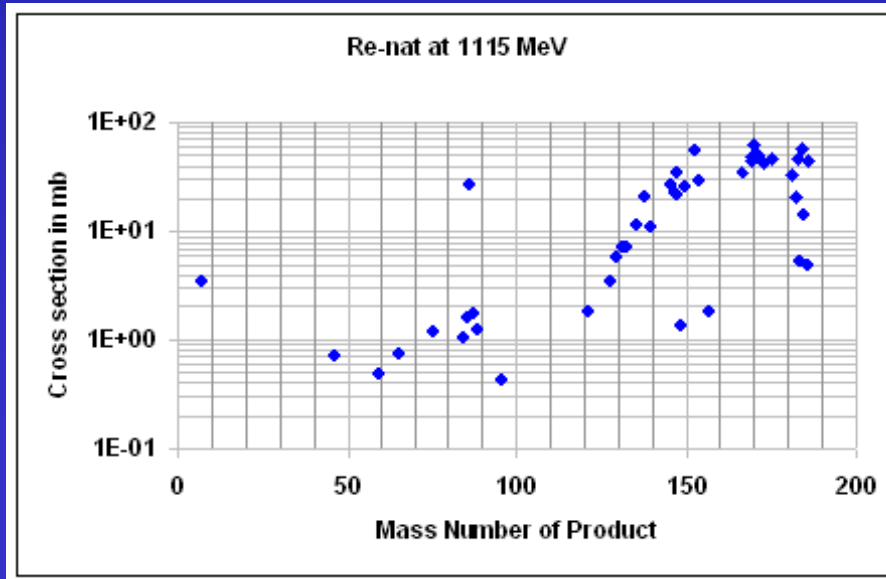
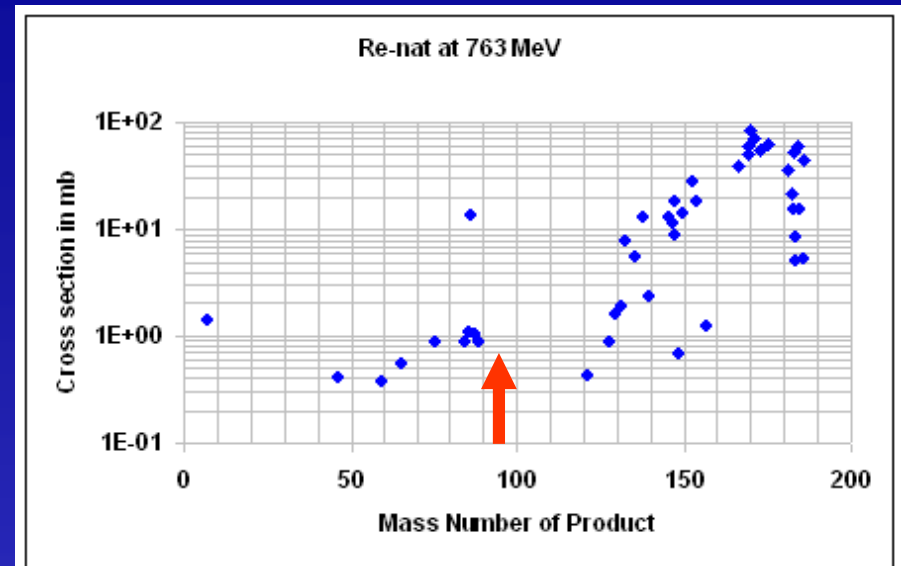
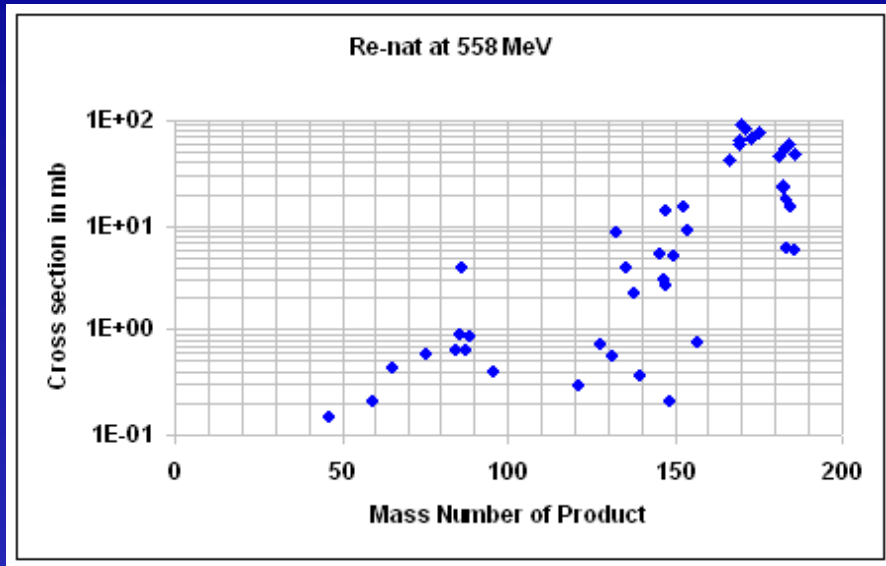
Proton-Induced Reactions on Rhenium



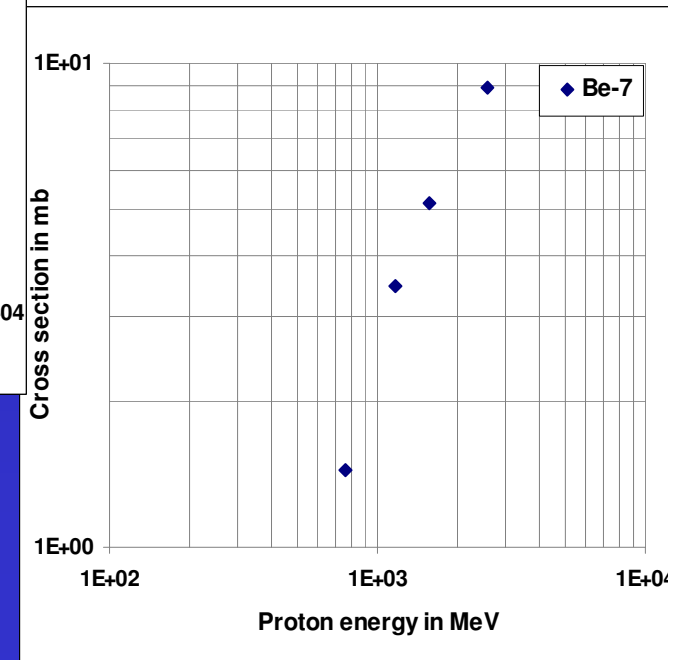
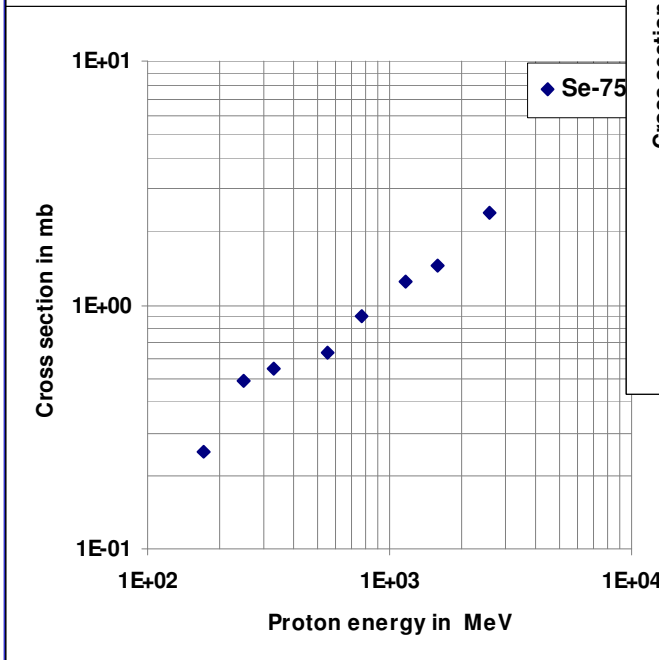
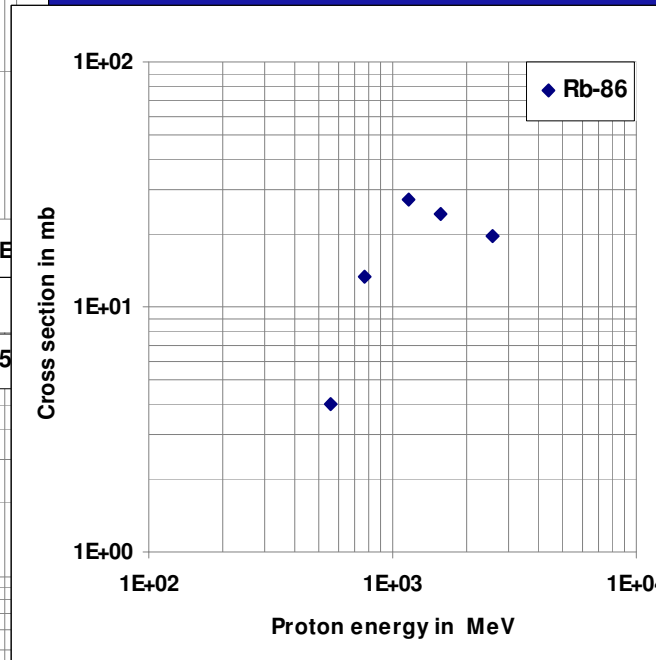
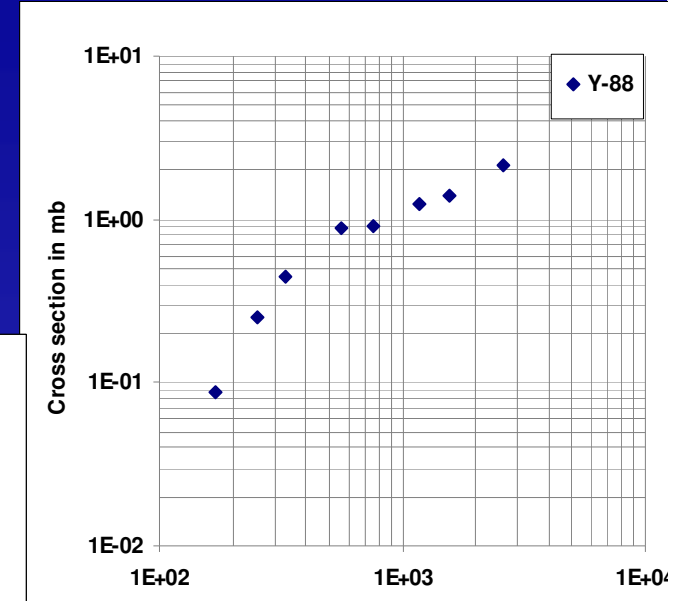
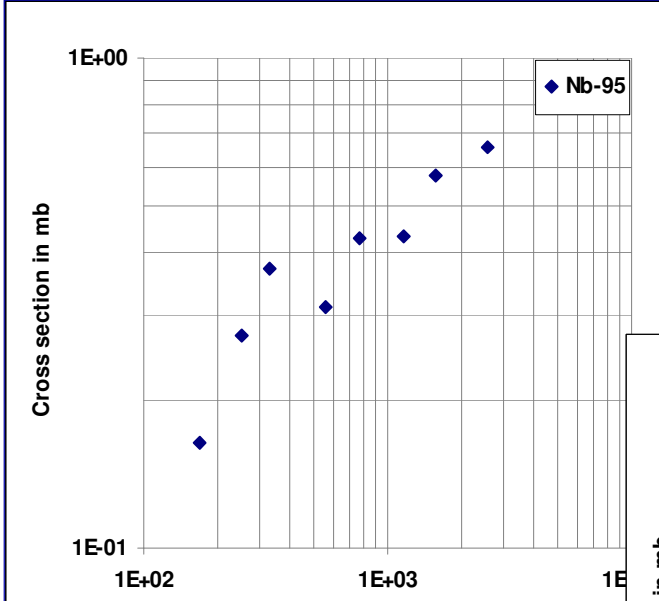
Proton-Induced Reactions on Rhenium



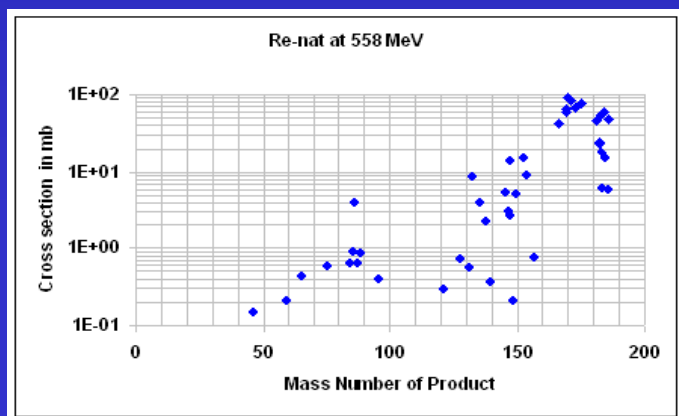
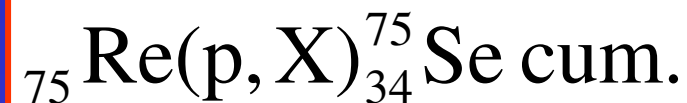
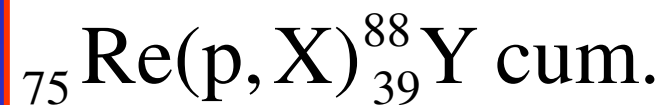
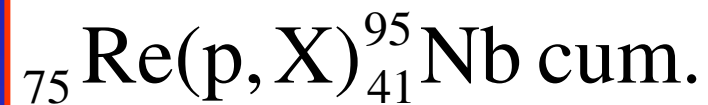
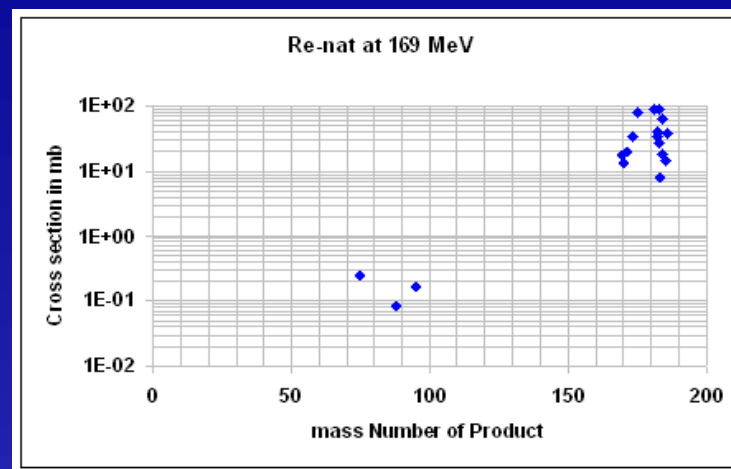
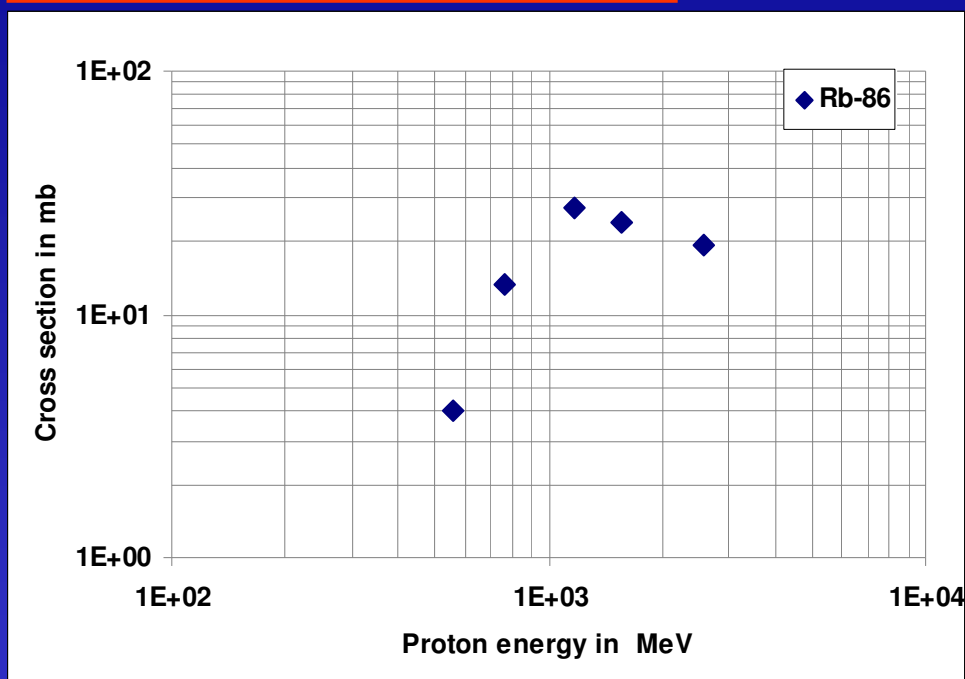
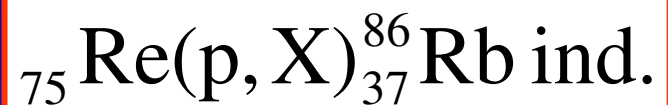
Proton-Induced Reactions on Rhenium



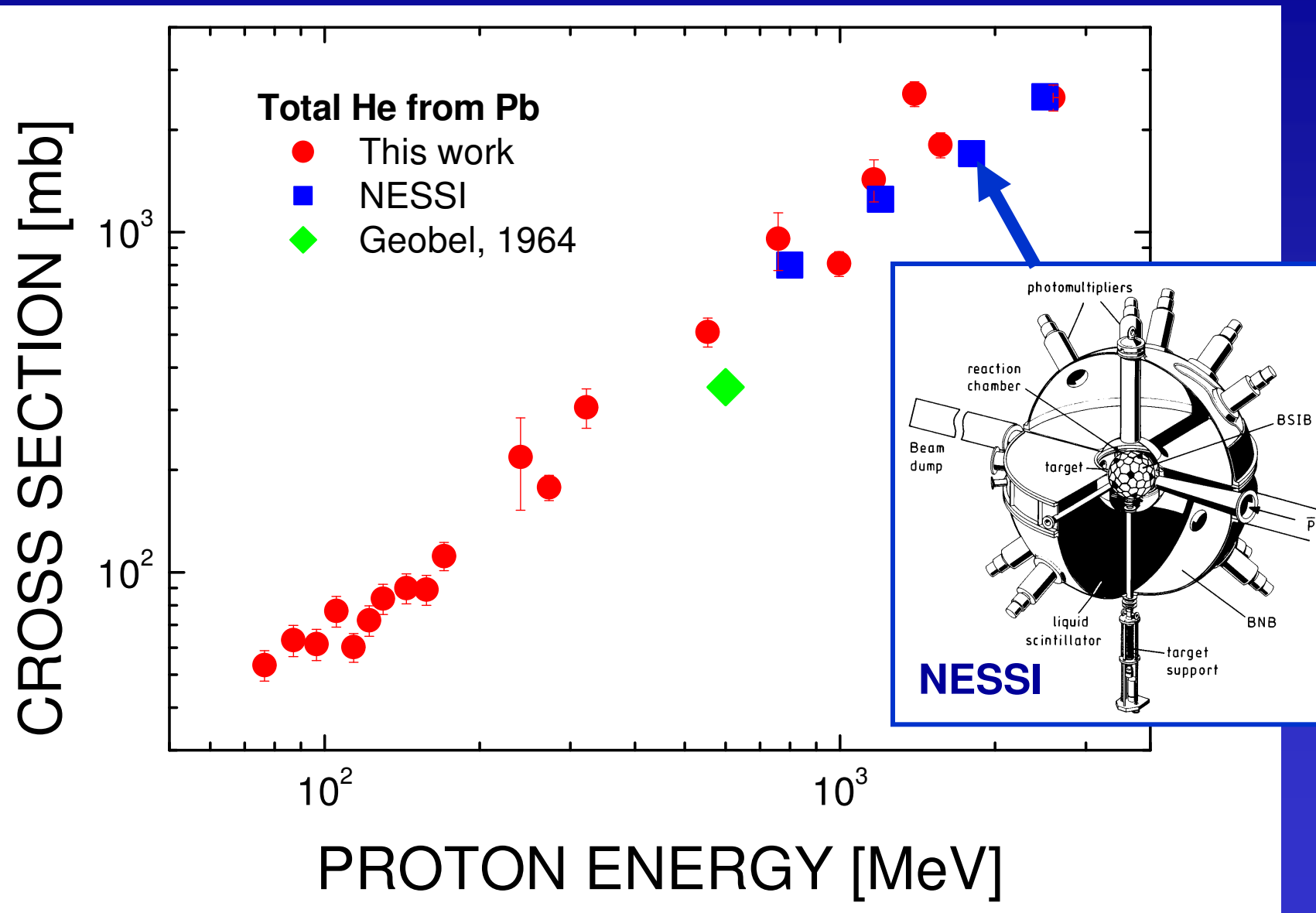
Proton-Induced Reactions on Rhenium

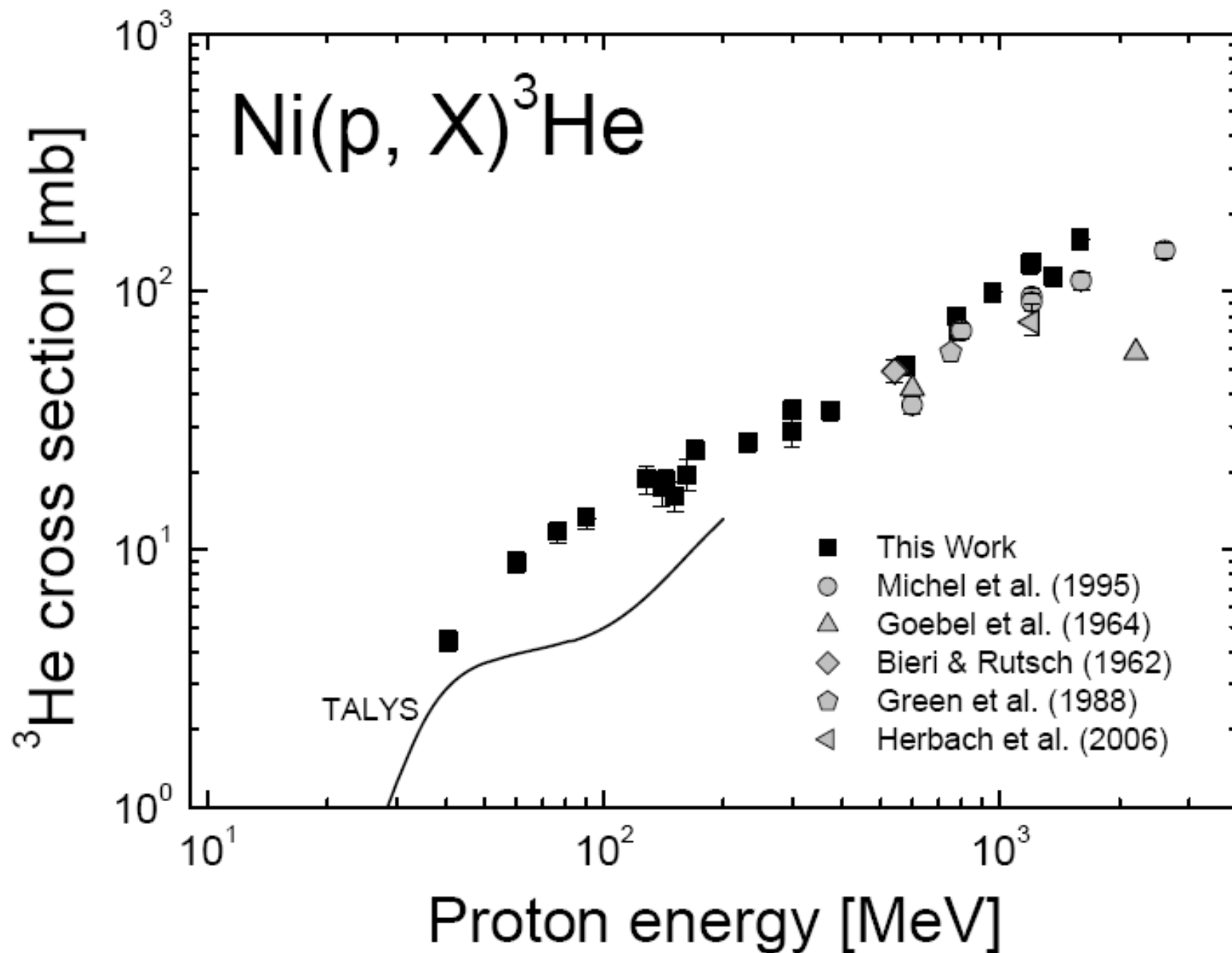


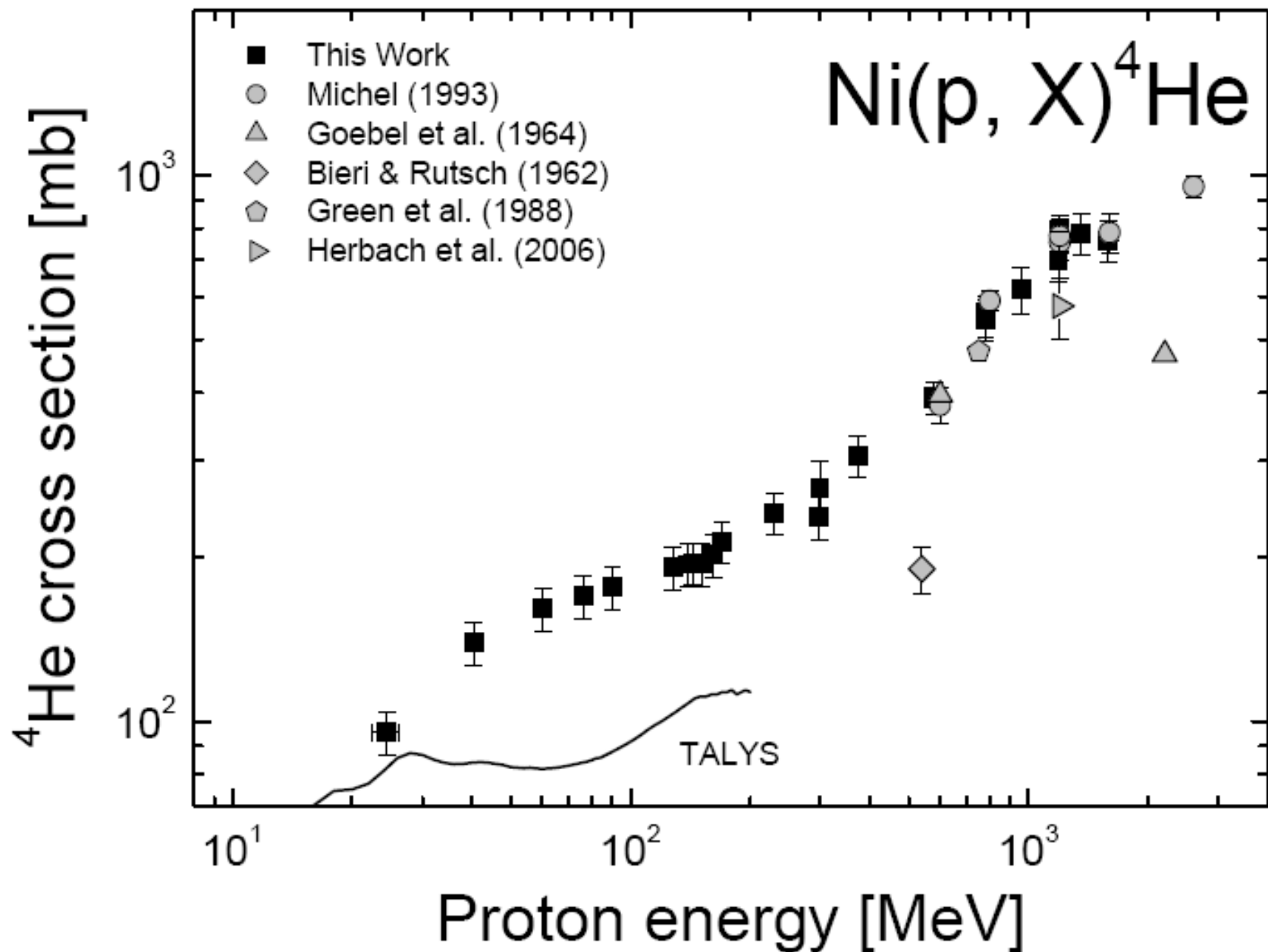
Special Reactions on Rhenium



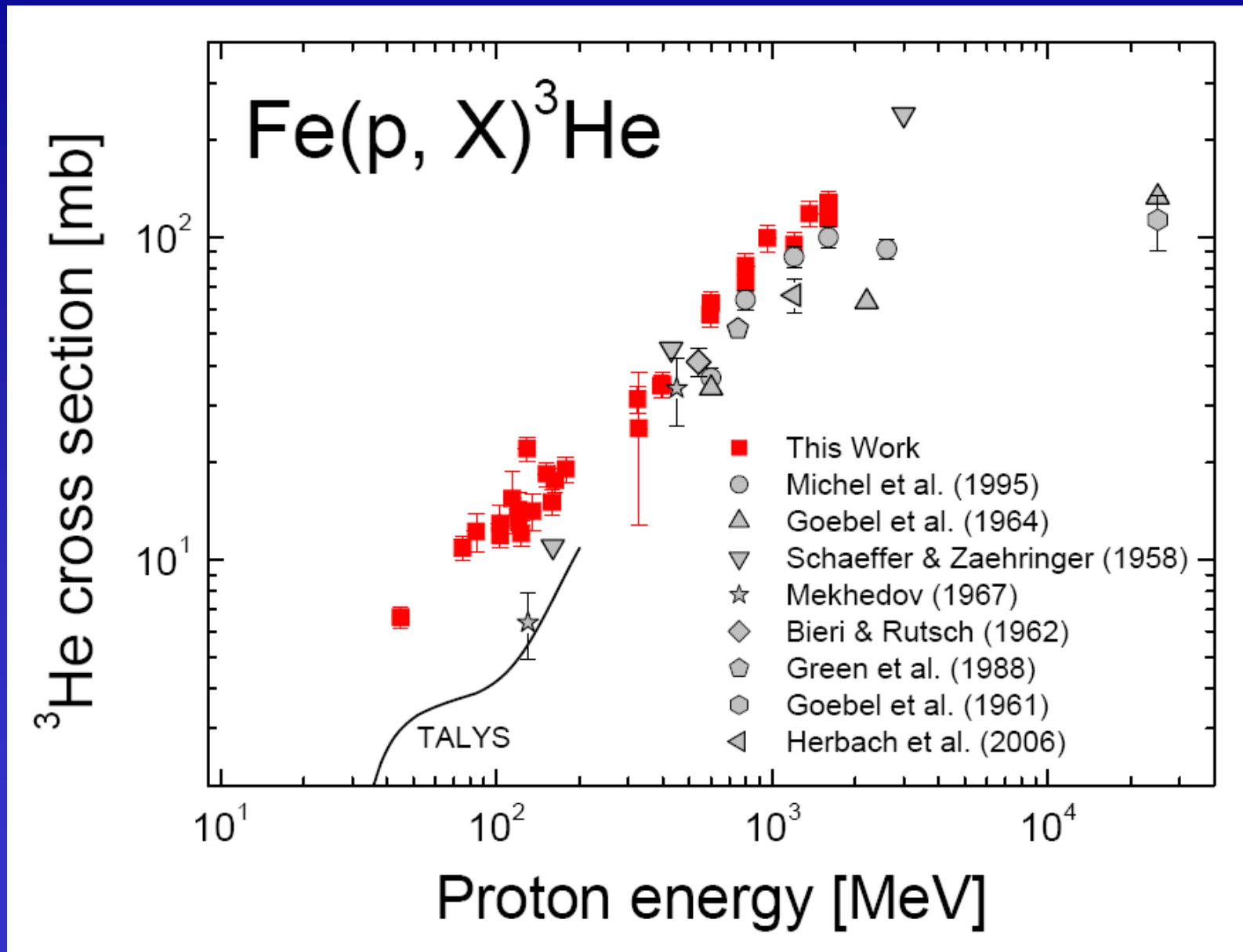
Production of Helium from Lead



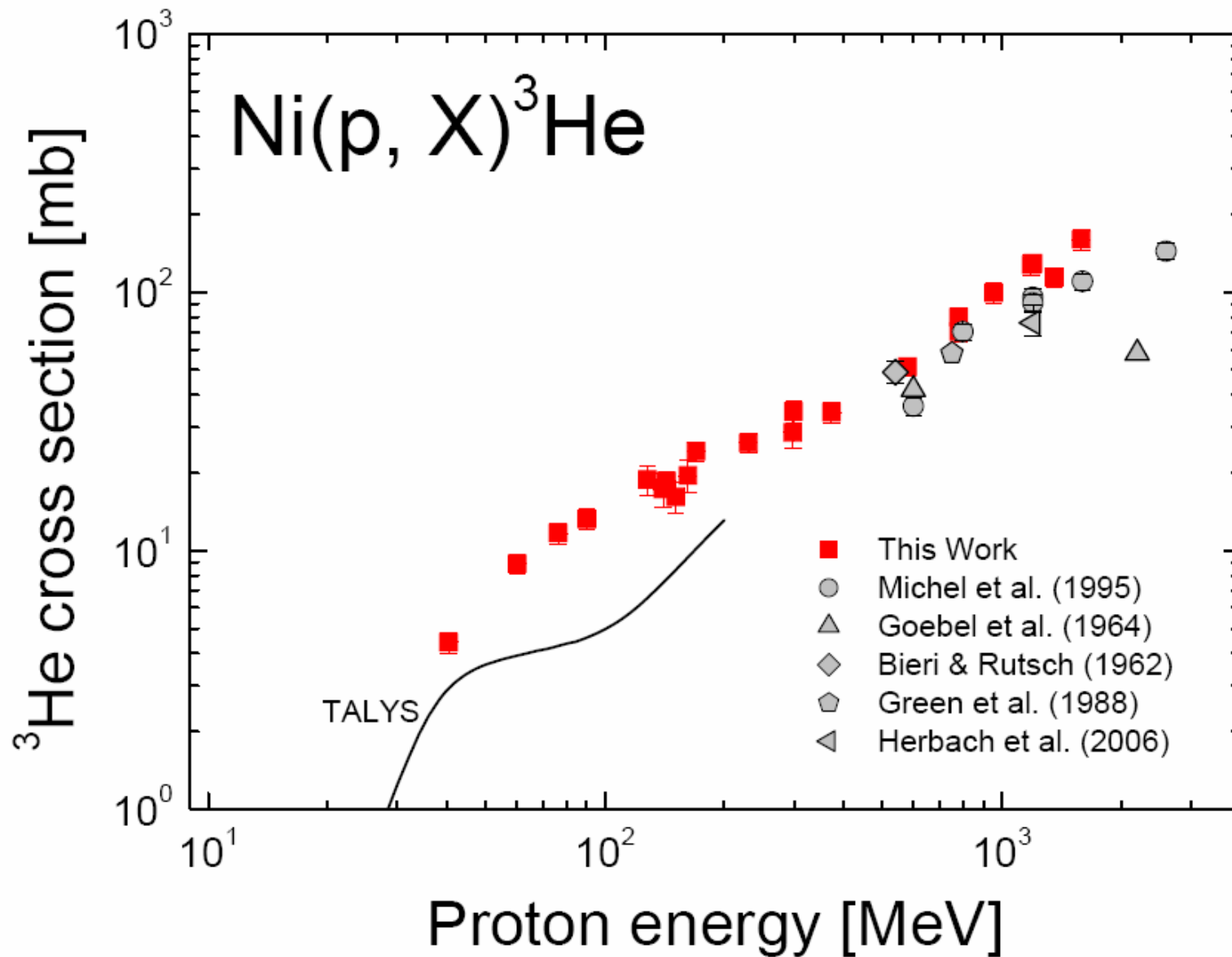




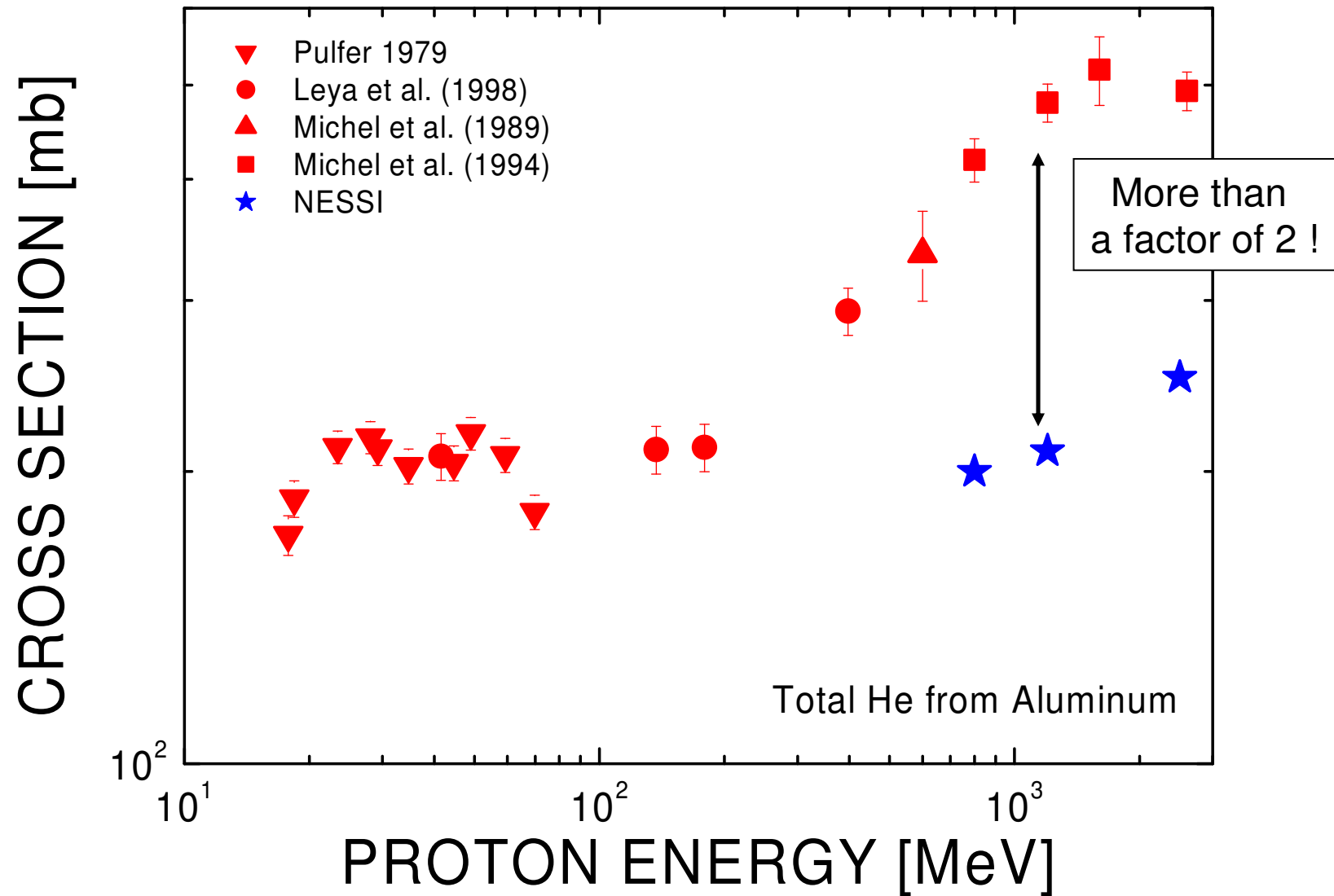
Production of Helium



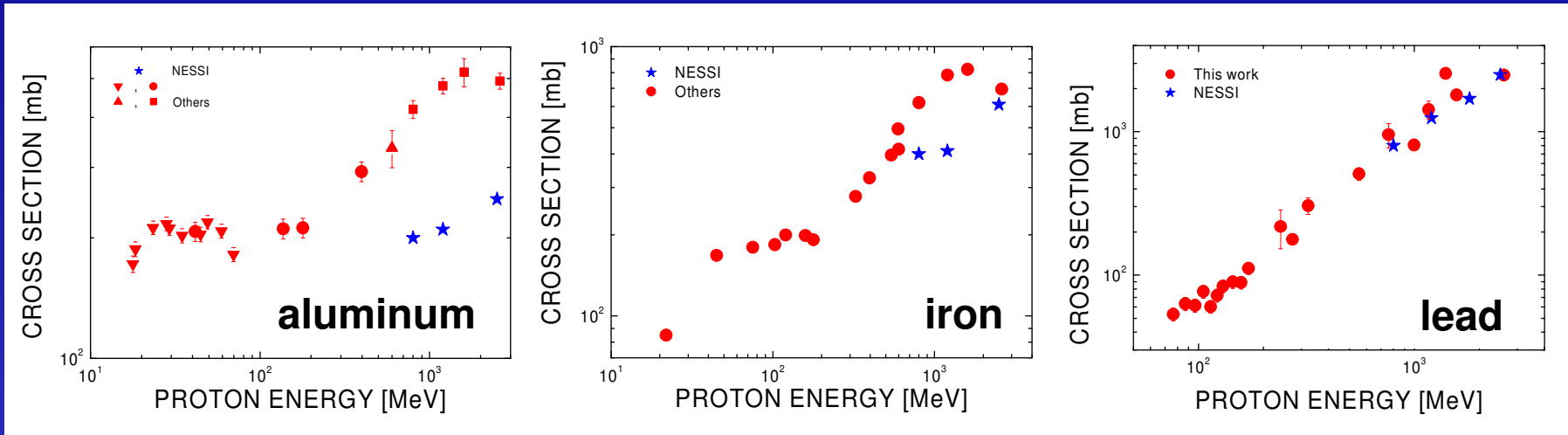
Production of Helium



Production of Helium from Aluminum

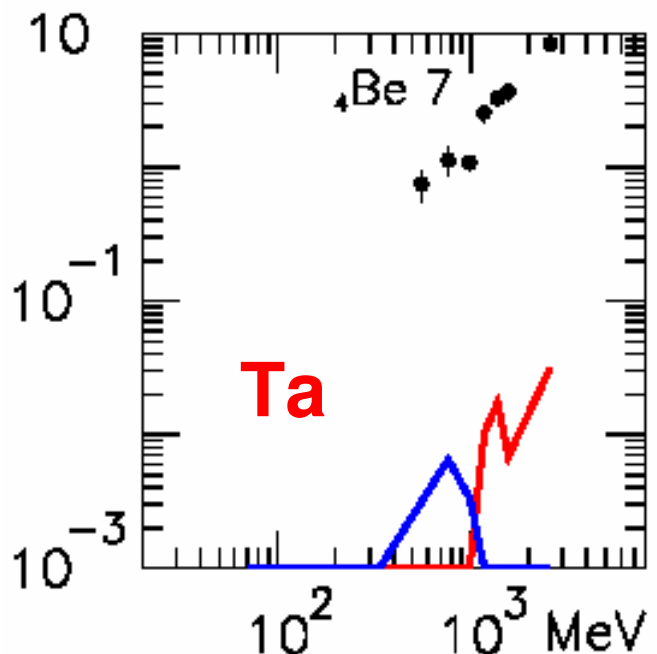
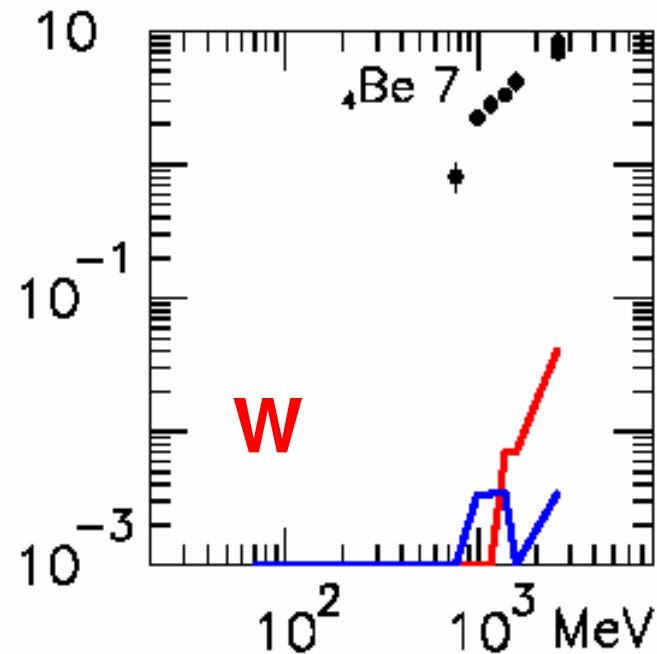
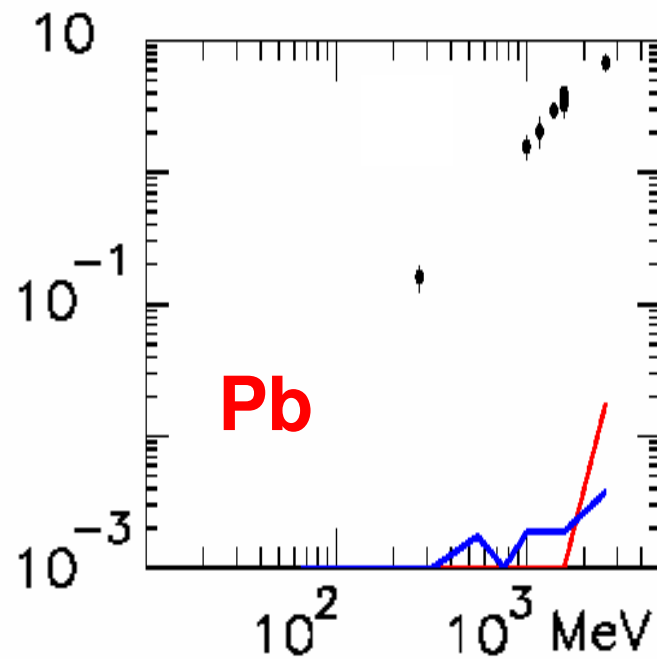
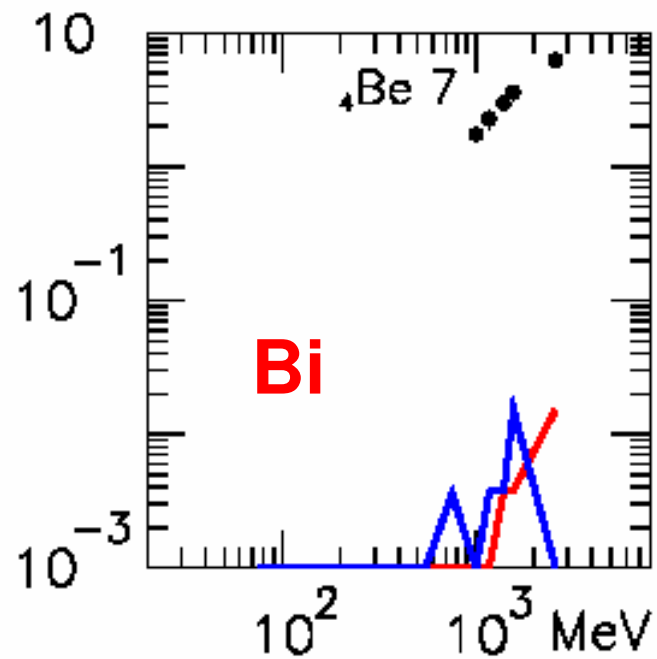


Production of Helium from Al, Fe, and Pb



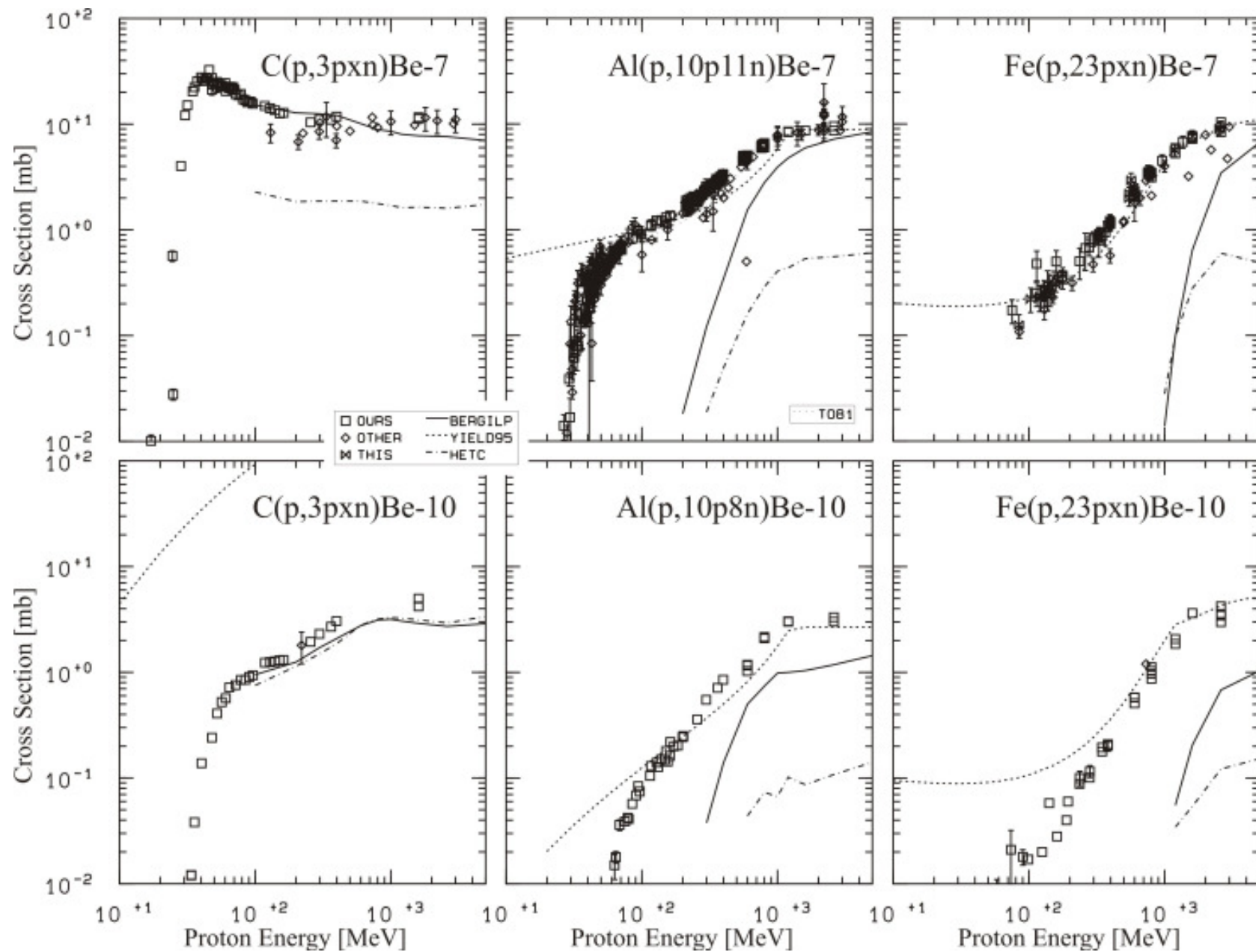
Decreasing evaporation peak

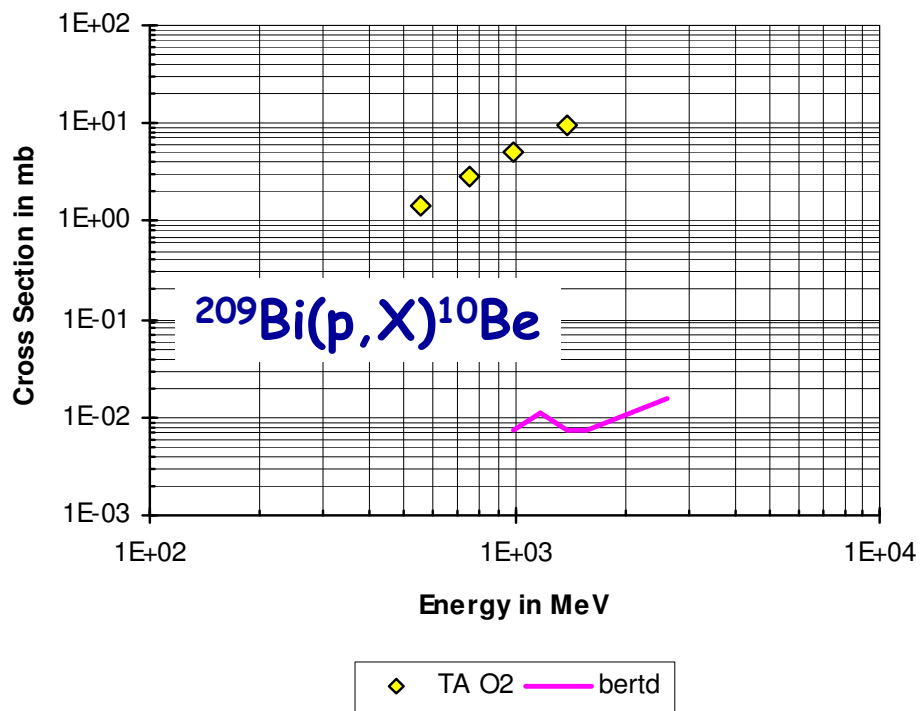
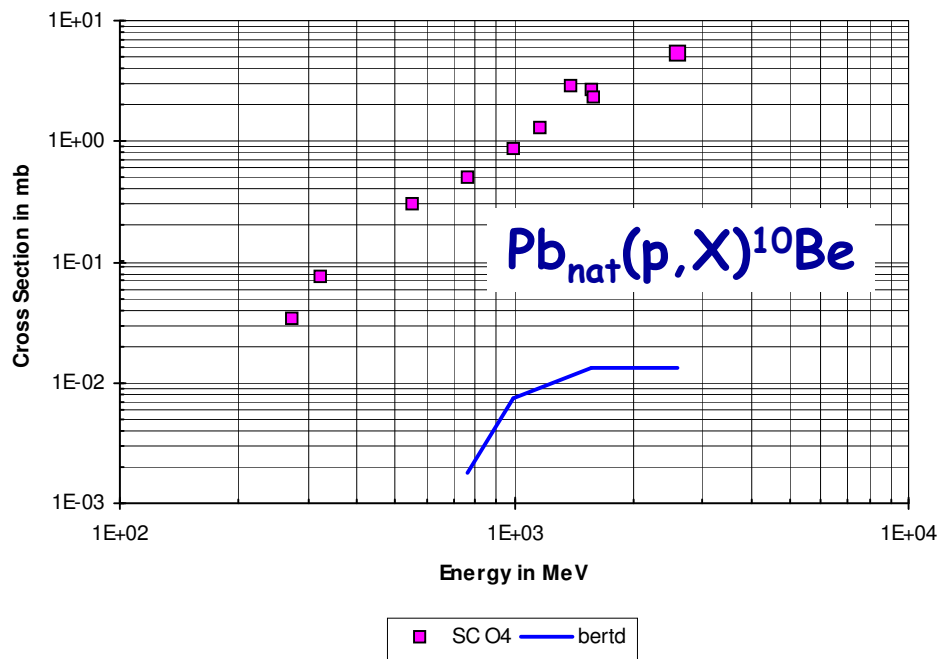
The agreement becomes better with decreasing relevance of the evaporation peak



Production
of ${}^7\text{Be}$
and ...

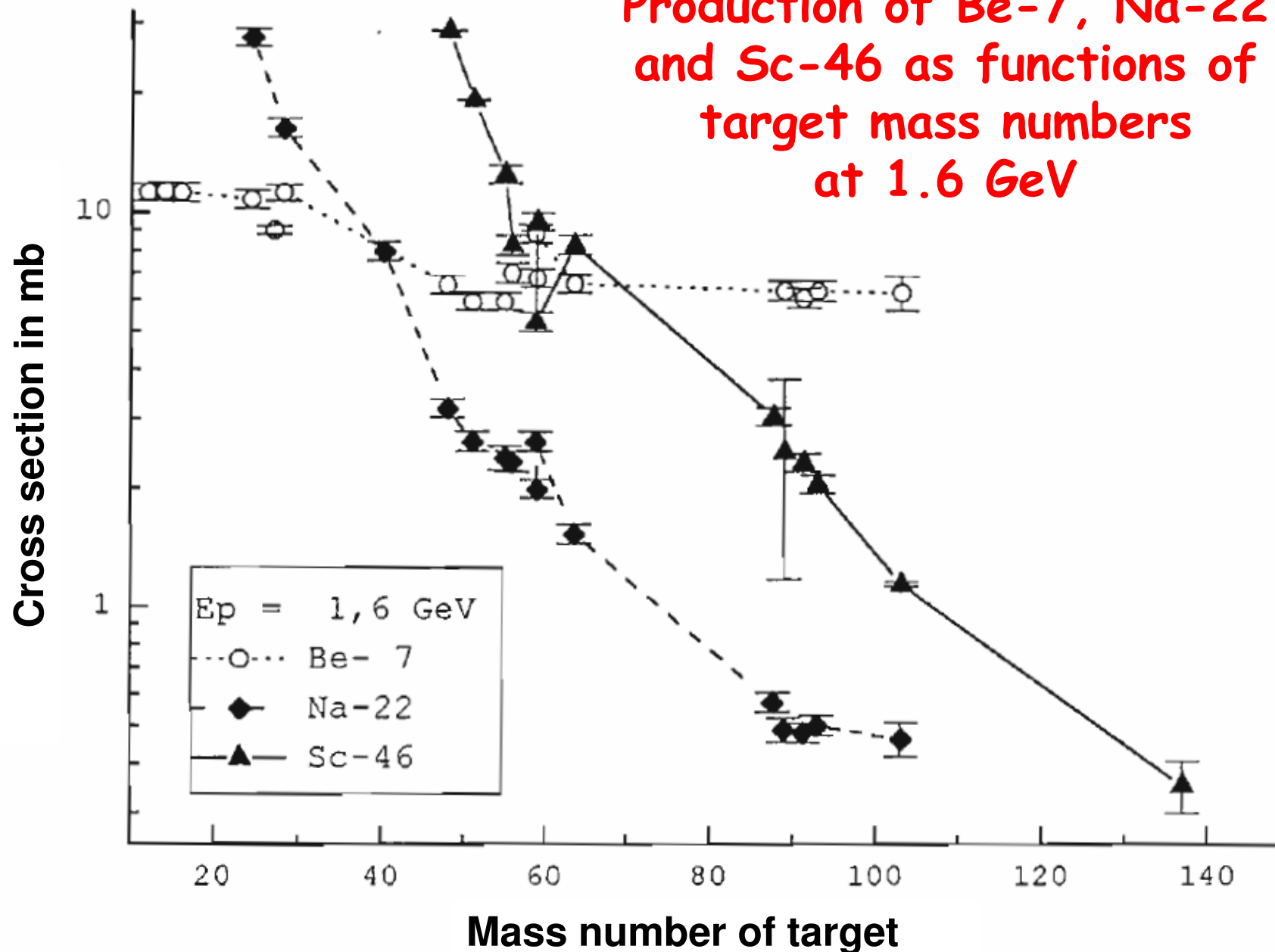
— INCL4 + ABLA
— Bertini-Dresner-PE

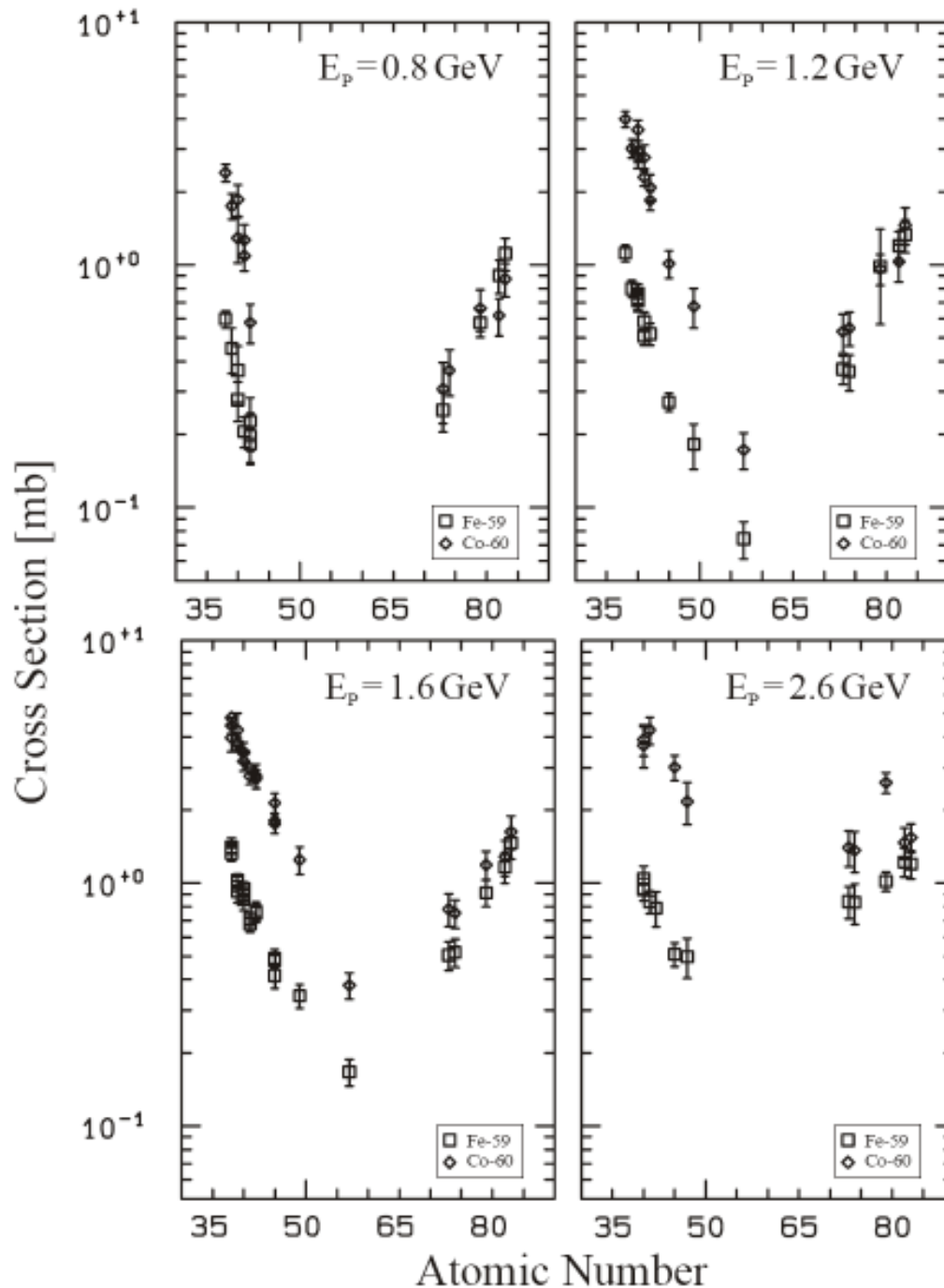




Production of light residual nuclides by fragmentation ?

Production of Be-7, Na-22 and Sc-46 as functions of target mass numbers at 1.6 GeV





**Systematic
of the production
of Fe-59 and Co-60
as functions of the
target element
atomic numbers**

Product	Targets
H-3	Fe, Ni, Pb
He-3	Al, Mg, Si, Fe, Ni, Pb, Bi
He-3cum	Al, Mg, Si, Fe, Ni, Pb, Bi
He-4	Al, Mg, Si, Fe, Ni, Pb, Bi
Li-6	N
Li-7	N, O
Be-7	C, N, O, F, Mg, Al, Si, Ca, Ti, V, Mn, Fe, Co, Ni, Cu, Y, Nb, Mo, Ag, Zr, Ta, W, Au, Pb, Bi
Be-9	N
Be-10	C, N, O, Mg, AL, Si, Ti, Mn, Fe, Co, Ni, Cu, Ta, W, Re, Ir, Pb, B
B-10	N, O
B-11	N,O
C-11	O
C-14	O, Si, Fe, Ni
O-15	O
F-18	Na, Mg, Al, Cu, Ag, Pb
Ne-20	Mg, Al, Si, Fe, Ni, Pb, (Bi)
Ne-21	Mg, Al, Si, Fe, Ni, Pb, (Bi)
Ne-22	Mg, Al, Si, Fe, Ni, Pb, (Bi)
Ne-22cum	Mg, Al, Si, Fe, Ni, Pb, (Bi)
Na-22	Na, Mg, Al, Si, S, K, Ca, Ti, V, Mn, Fe, Co, Ni, Cu, Zn, Sr, Y, Zr, Nb, Mo, Rh, Ag, Ta, W, Au, Pb, Bi
Na-24	Mg, Al, Si, Ca, Sc, Ti, V, Mn, Fe, Co, Ni, Cu, Zn, Zr, Nb, Ag, La, Ta, Au, Pb
Al-26	Mg, Al, Si, Ca, Ti, Mn, Fe, Ni, Pb, (Bi)
Mg-28	Si, Ca, Ti, V, Mn, Fe, Co, Ni, Cu, Pb

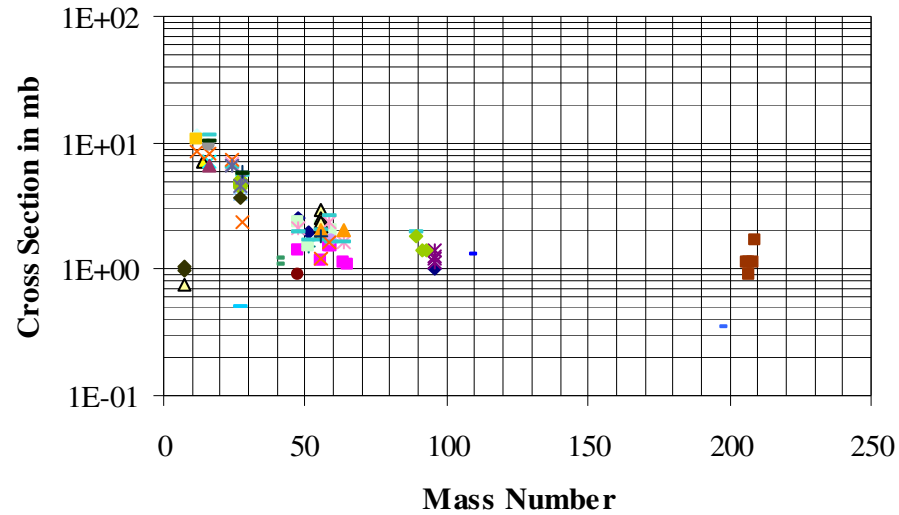
Systematic of the production of low-mass residues

contribution to
NUDATRA WP 5.4
(2007)

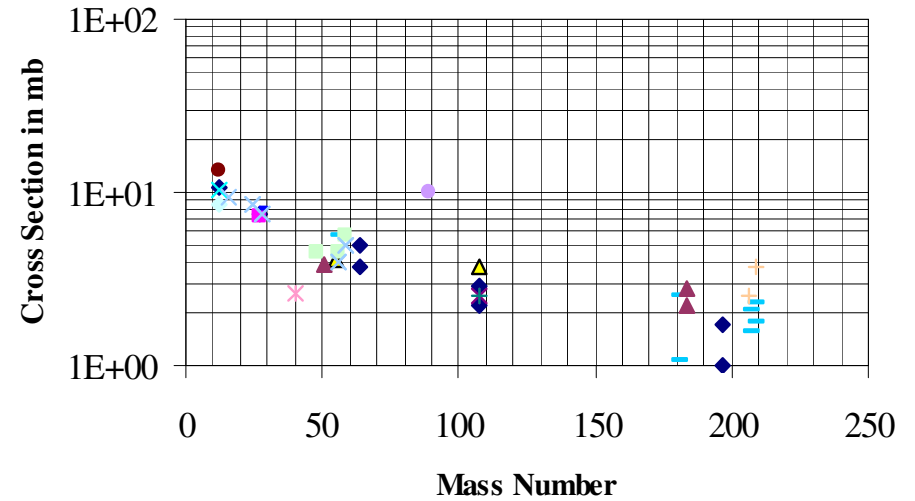
SIXTH FRAMEWORK PROGRAMME EURATOM Management of Radioactive Waste	
	
Project acronym:	EUROTRANS
Project full title:	EUROpean Research Programme for the TRANSmutation of High Level Nuclear Waste in an Accelerator Driven System
Contract no.:	F 6W-CT-2004-516520
Domain:	ERAC NUDATRA
Workpackage N°:	5.4.1
Identification N°:	05.7
Type of document:	Final report
Title:	Infra-structure and status of helium production cross sections in intermediate mass targets at 600MeV
Classification Level:	PU or PP or 19 / or 20
Reference:	17.04.01.001000005_05.7
Status:	Final

Production of Be-7

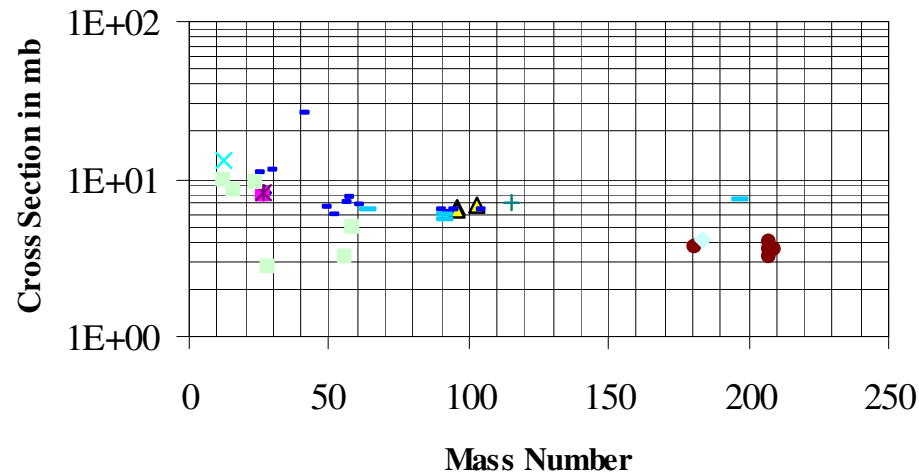
Be-7 for $E_p=500$ MeV



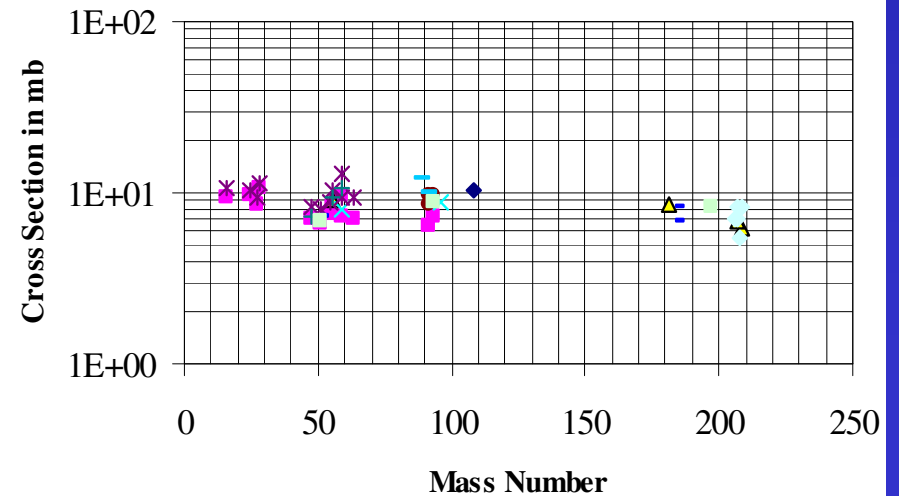
Be-7 for $E_p=1000$ MeV



Be-7 for $E_p=1500$ MeV

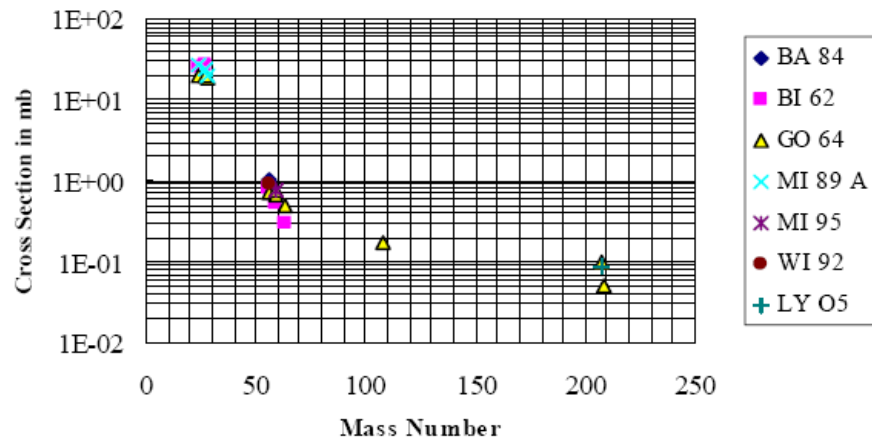


Be-7 for $E_p=2600$ MeV

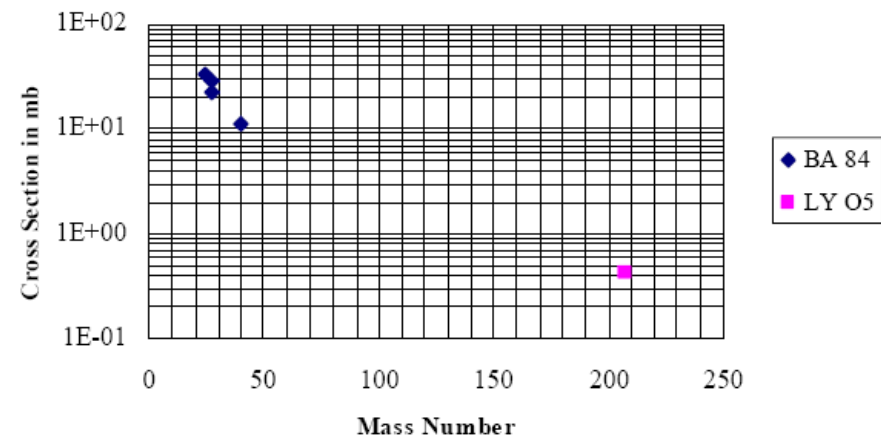


Production of Ne-21

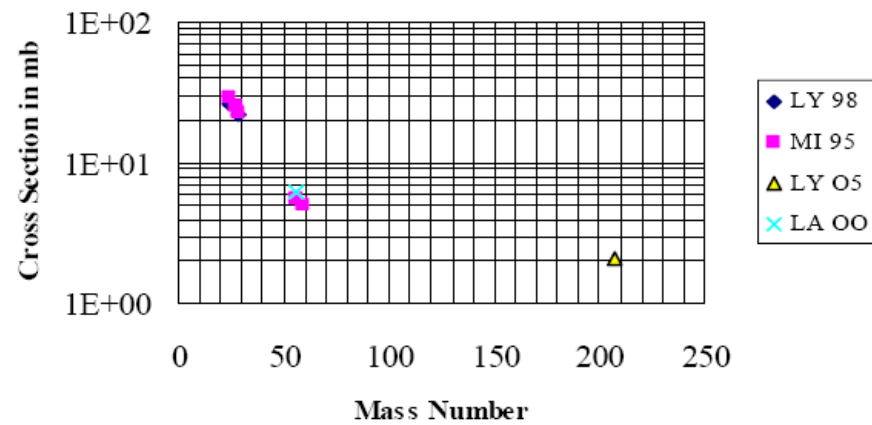
Ne-21 for $E_p=600$ MeV



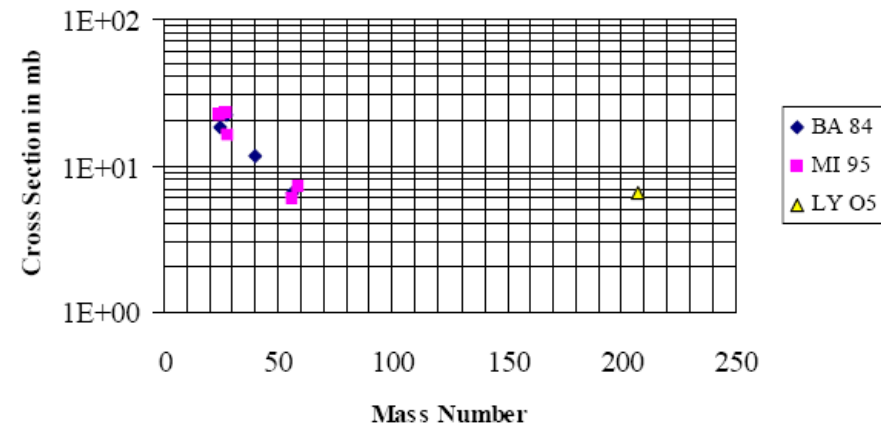
Ne-21 for $E_p=1000$ MeV



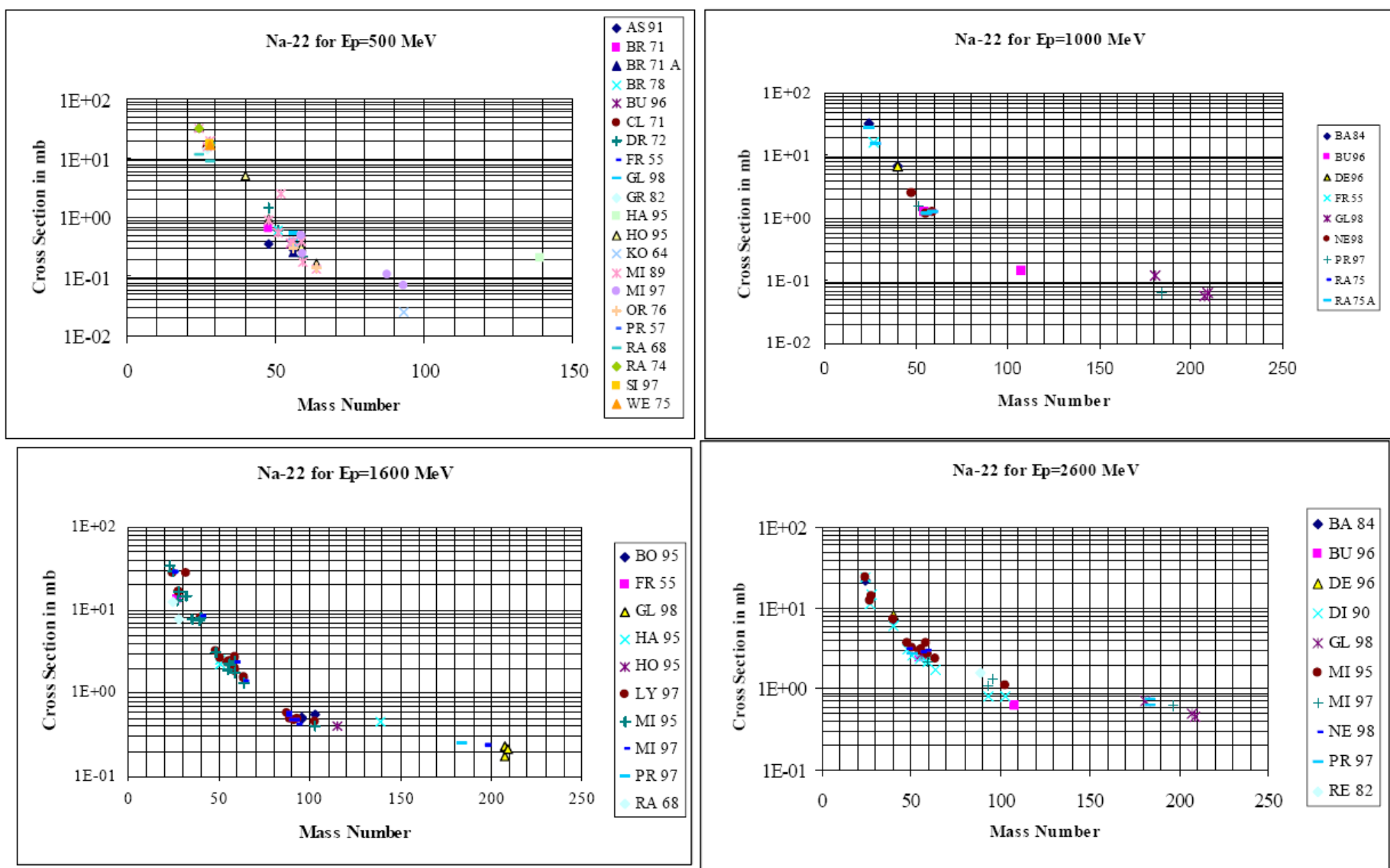
Ne-21 for $E_p=1600$ MeV



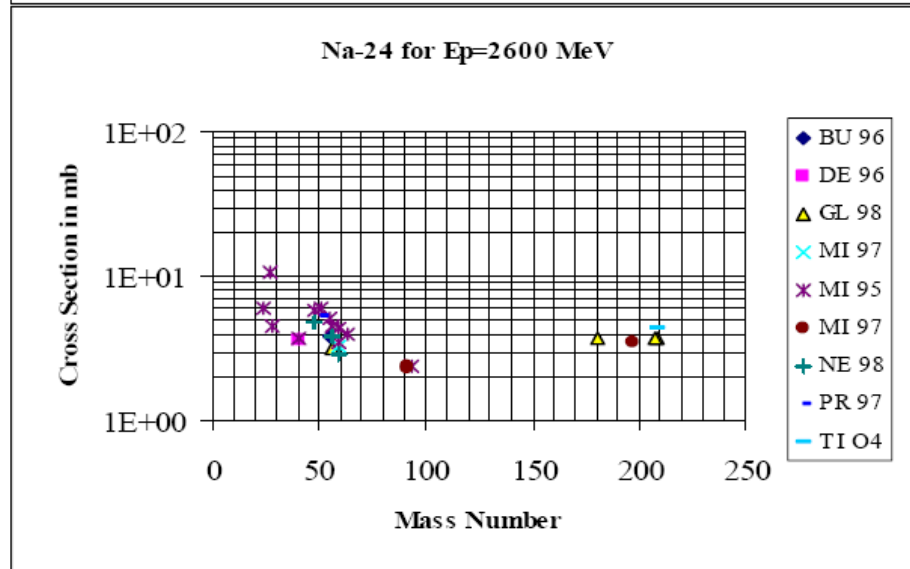
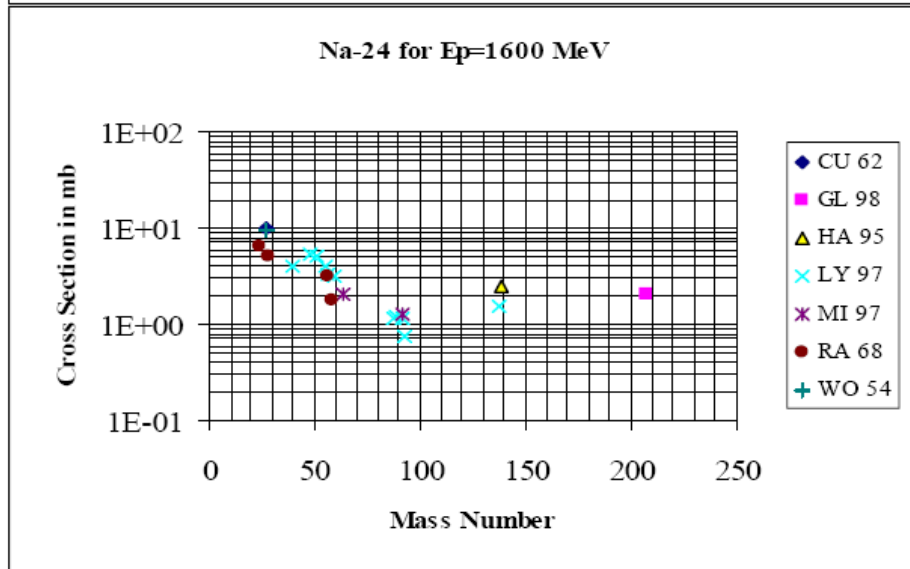
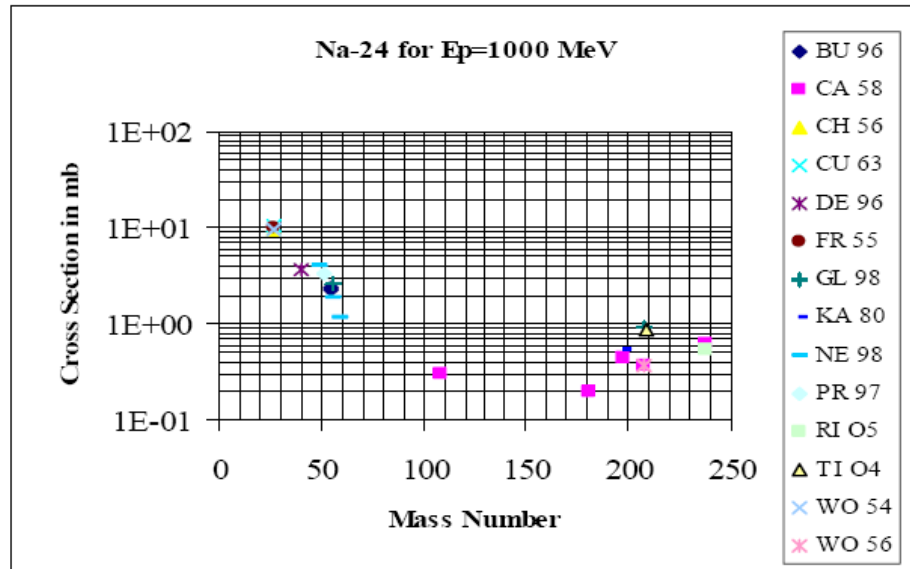
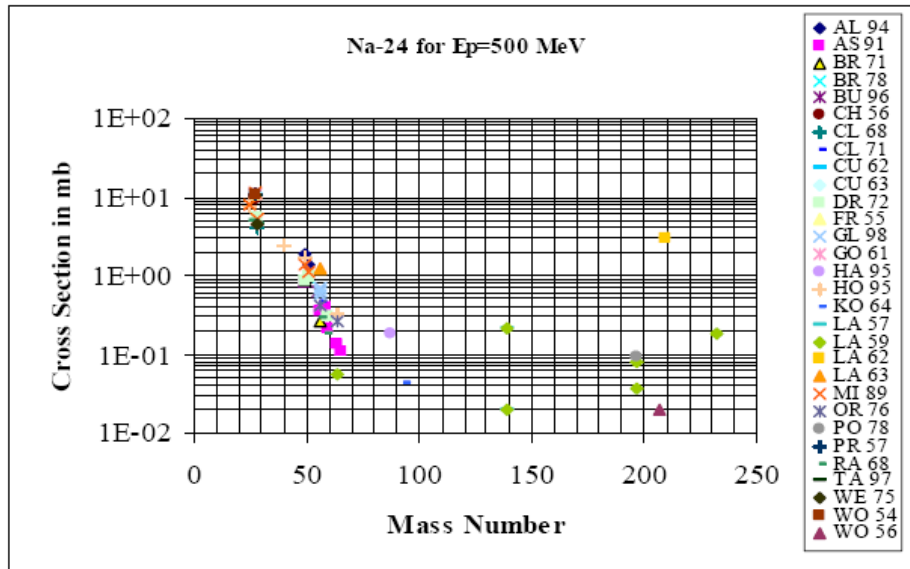
Ne-21 for $E_p=2600$ MeV



Production of Na-22

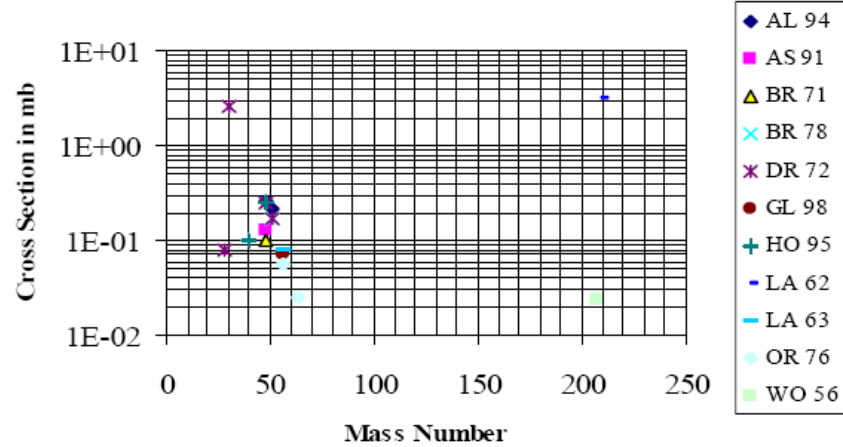


Production of Na-24

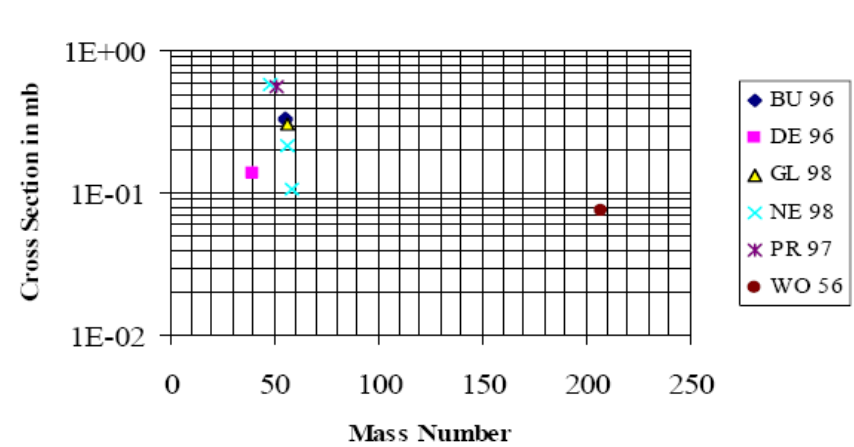


Production of Mg-28

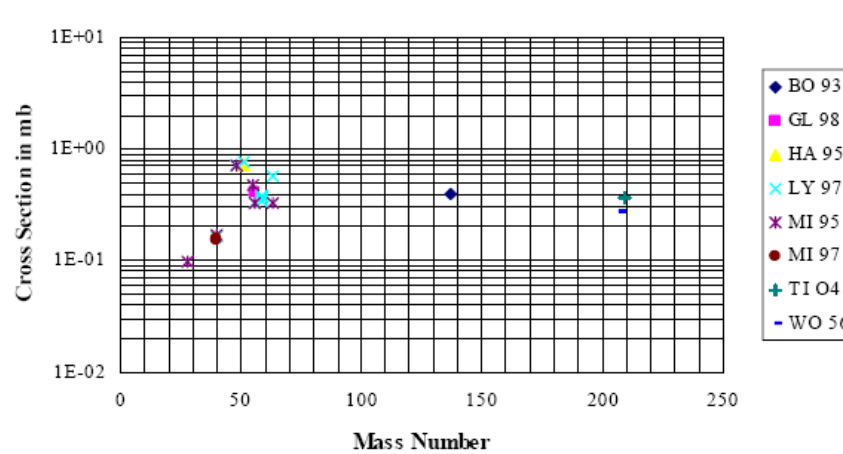
Mg-28 for $E_p=600$ MeV



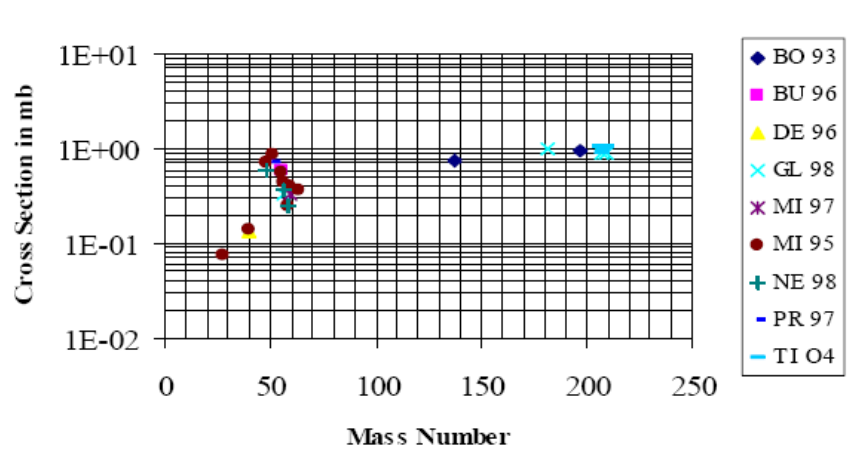
Mg-28 for $E_p=1000$ MeV



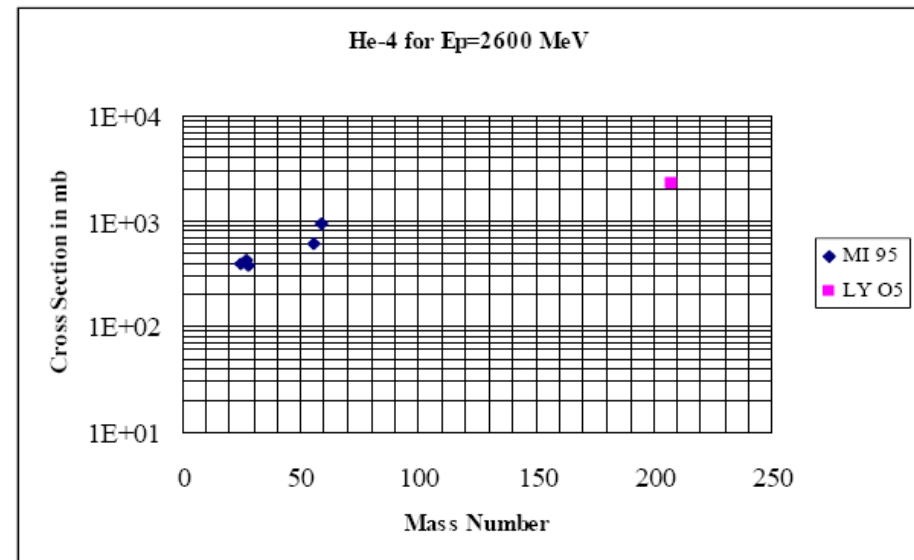
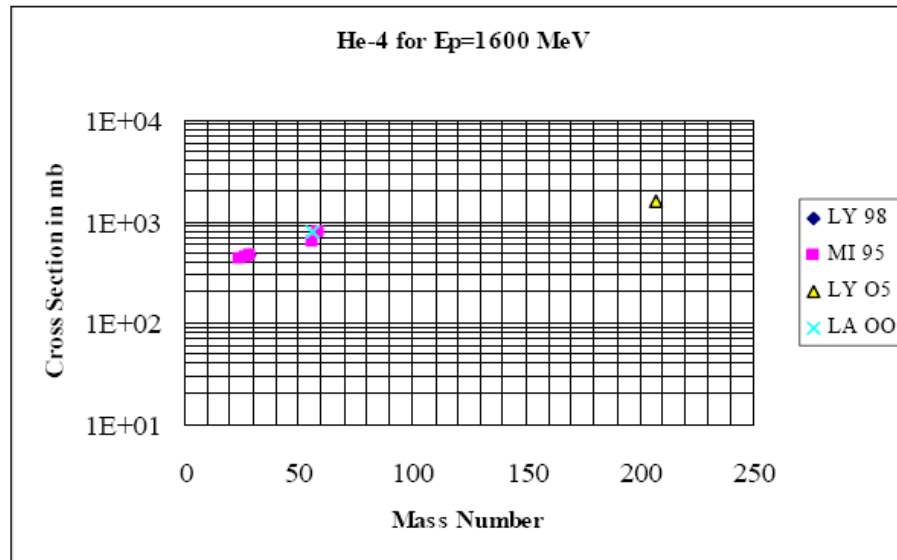
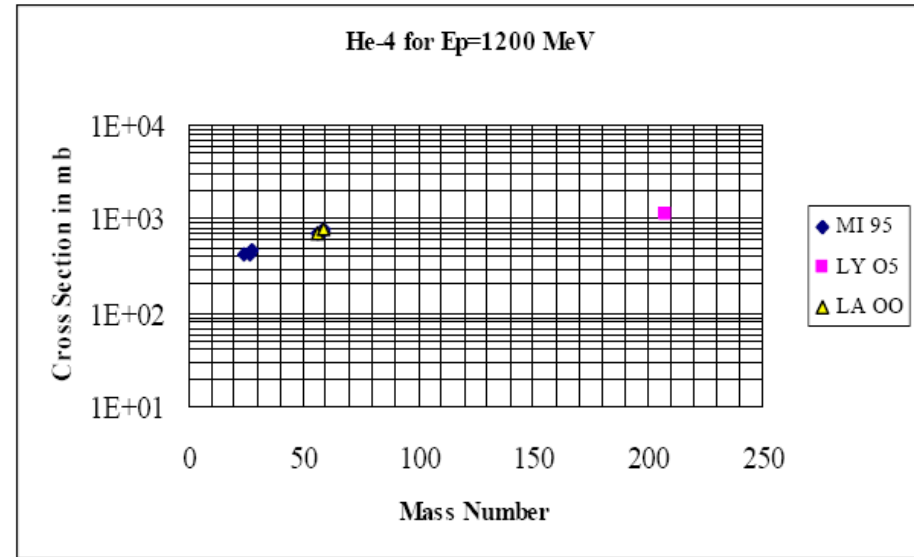
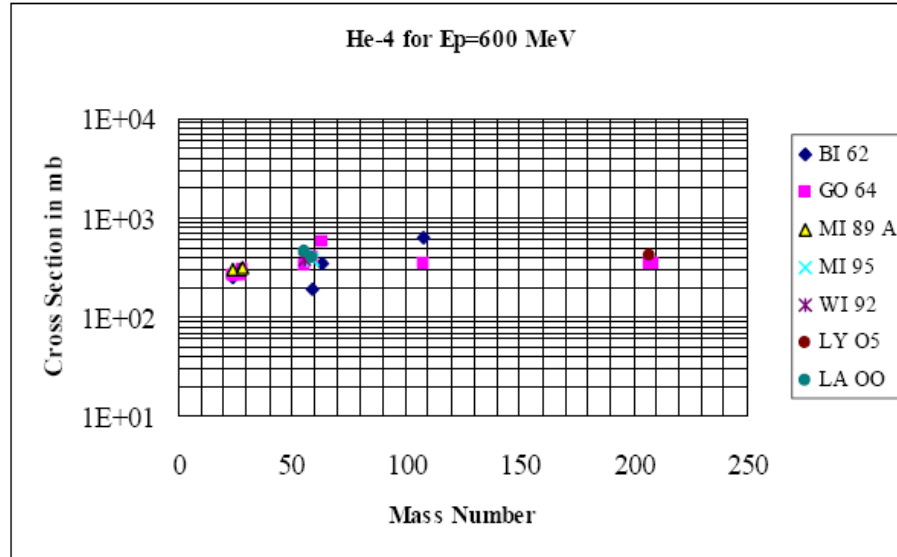
Mg-28 for $E_p=1600$ MeV



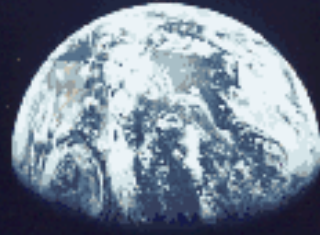
Mg-28 for $E_p=2600$ MeV



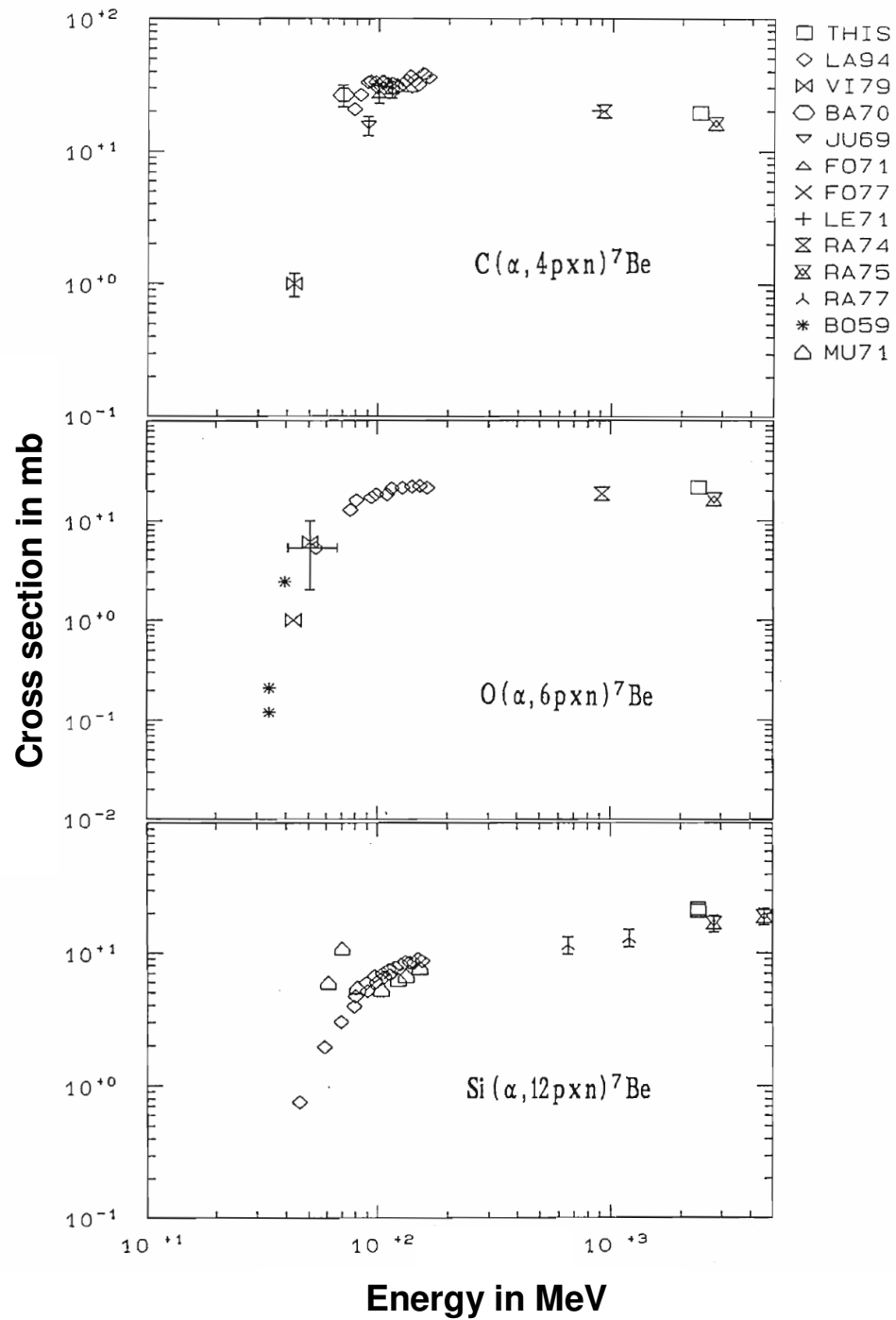
Production of He-4



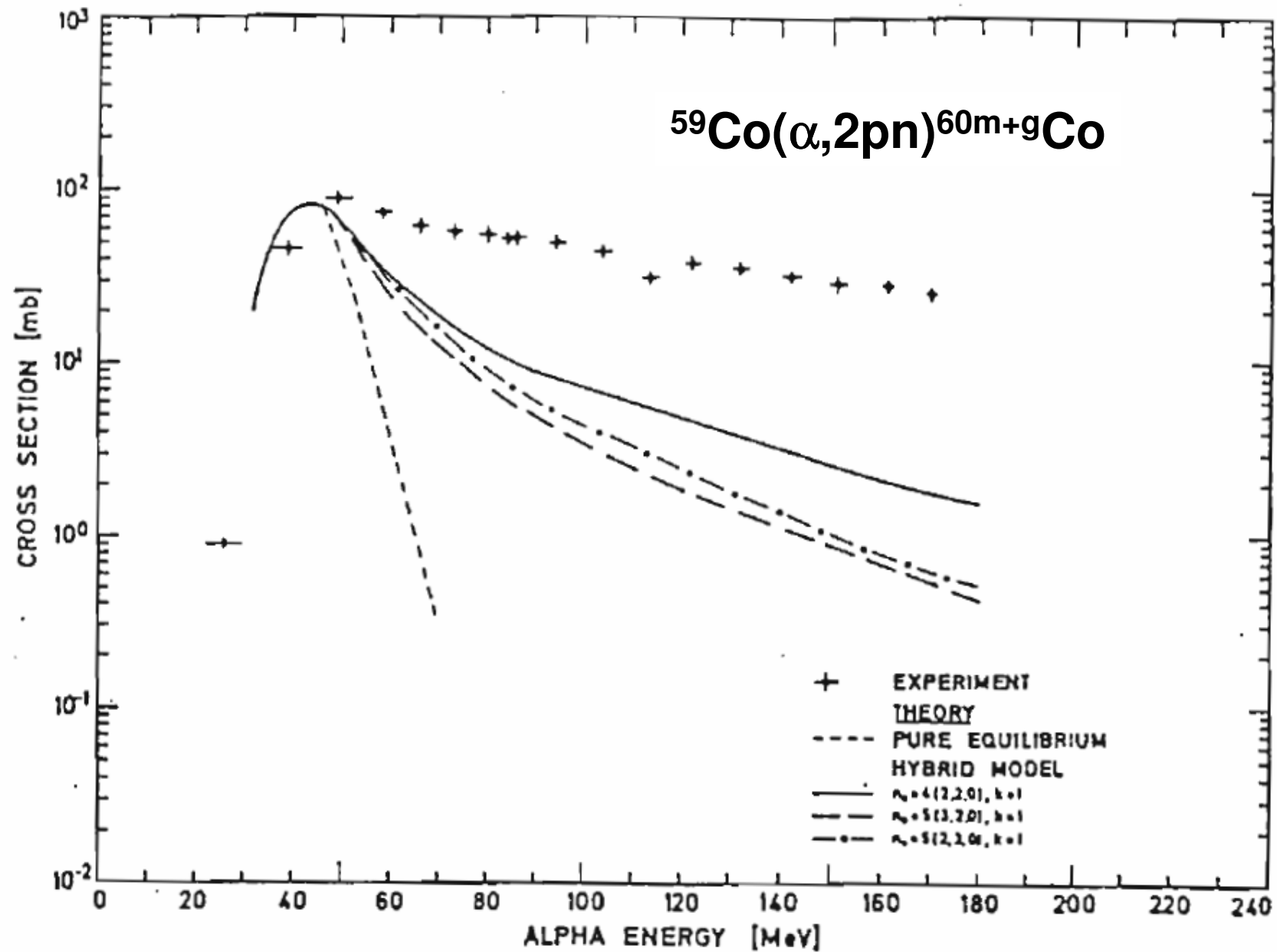
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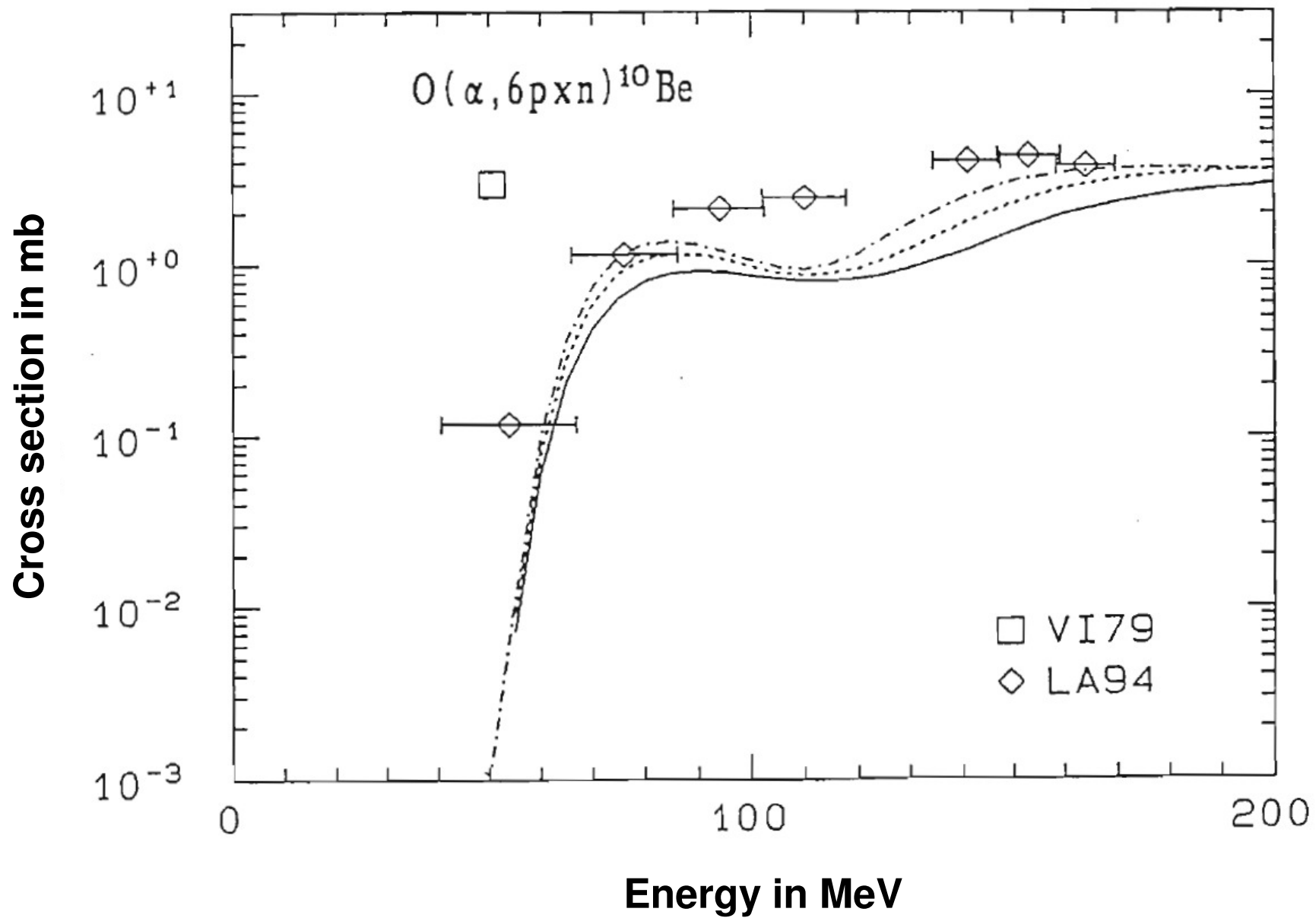


- **Relevance of data for the production of residual nuclides at medium energies**
- **The 1997 NEA intercomparison**
- **Modeling of cosmogenic nuclides in extraterrestrial matter**
- **Candidate data for testing the performance of model codes**
 - ❖ **Thin-target data**
 - **proton-induced reactions**
 - **alpha-induced reactions**
 - **neutron-induced reactions**
 - ❖ **Thick-target data**
- **conclusions**



Production of Be-7 by 4He -induced reactions



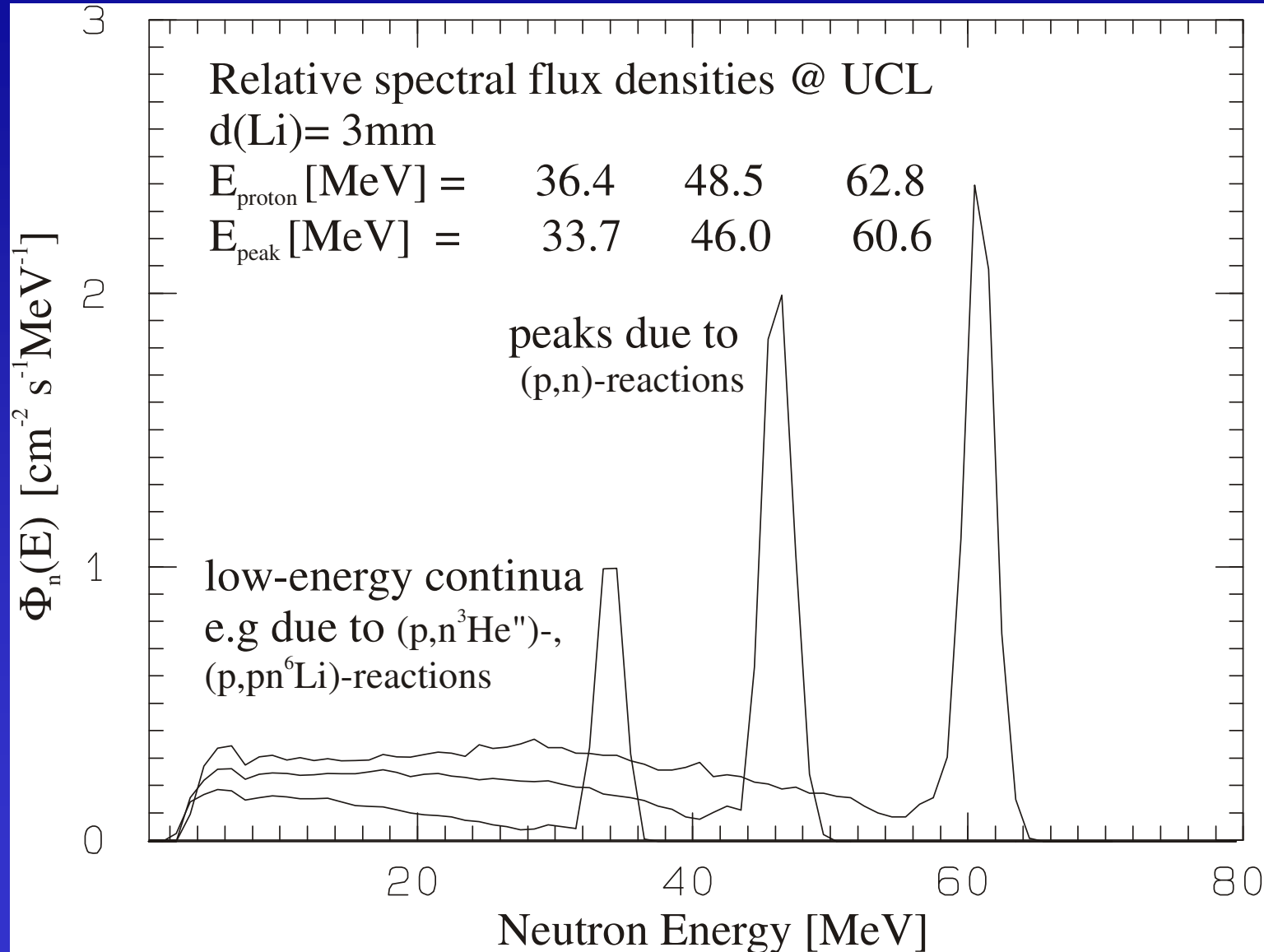


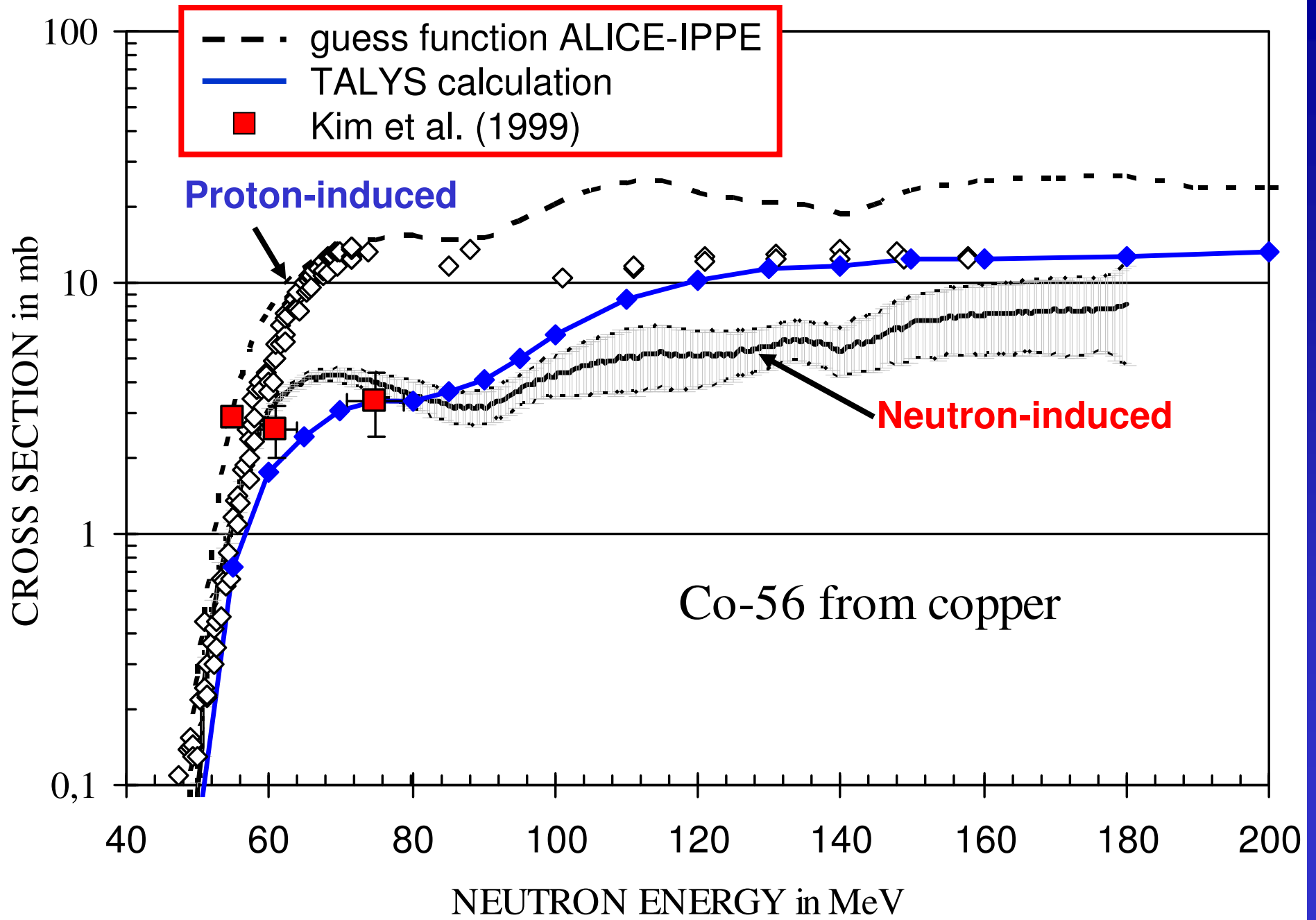
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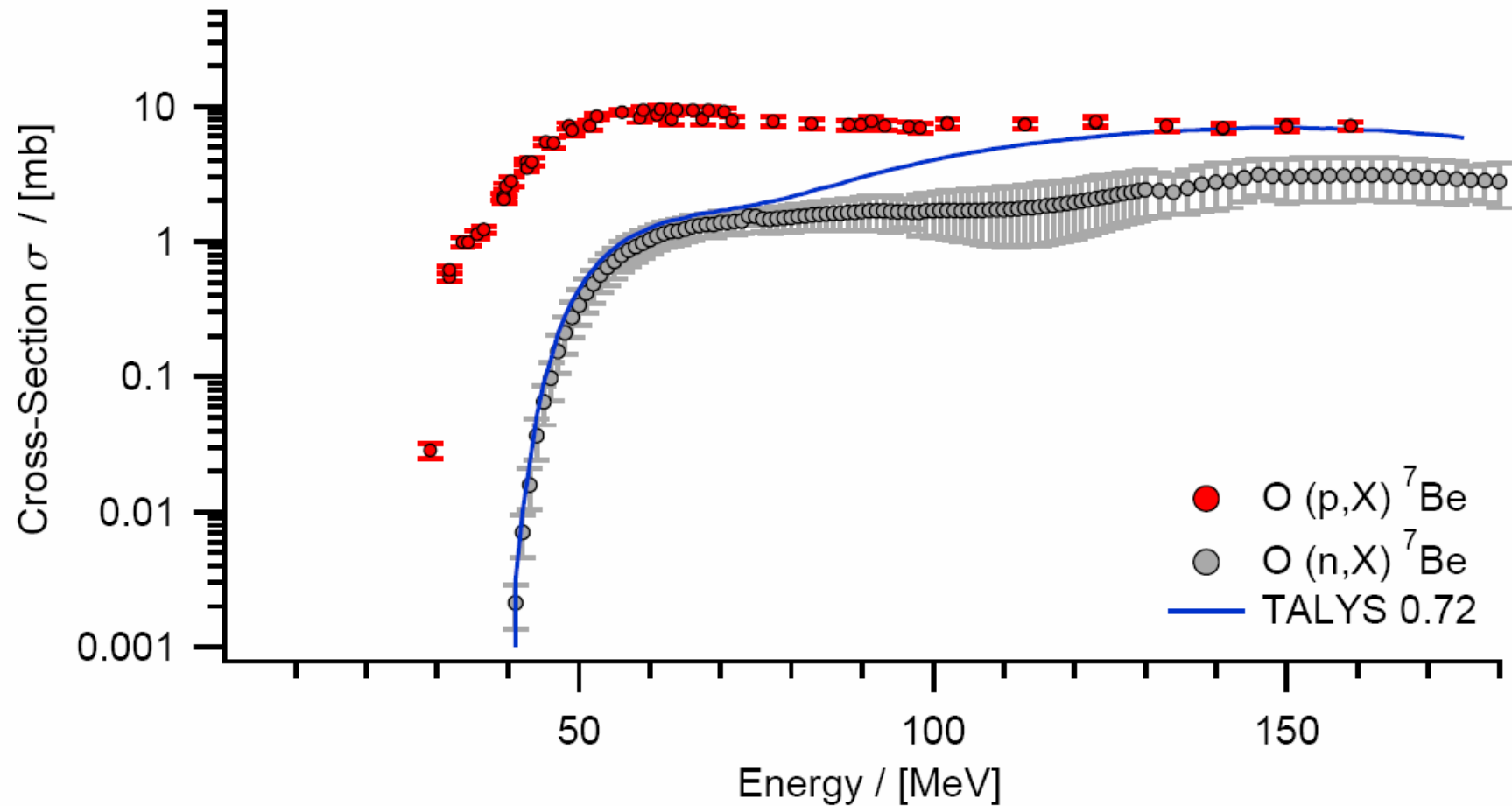
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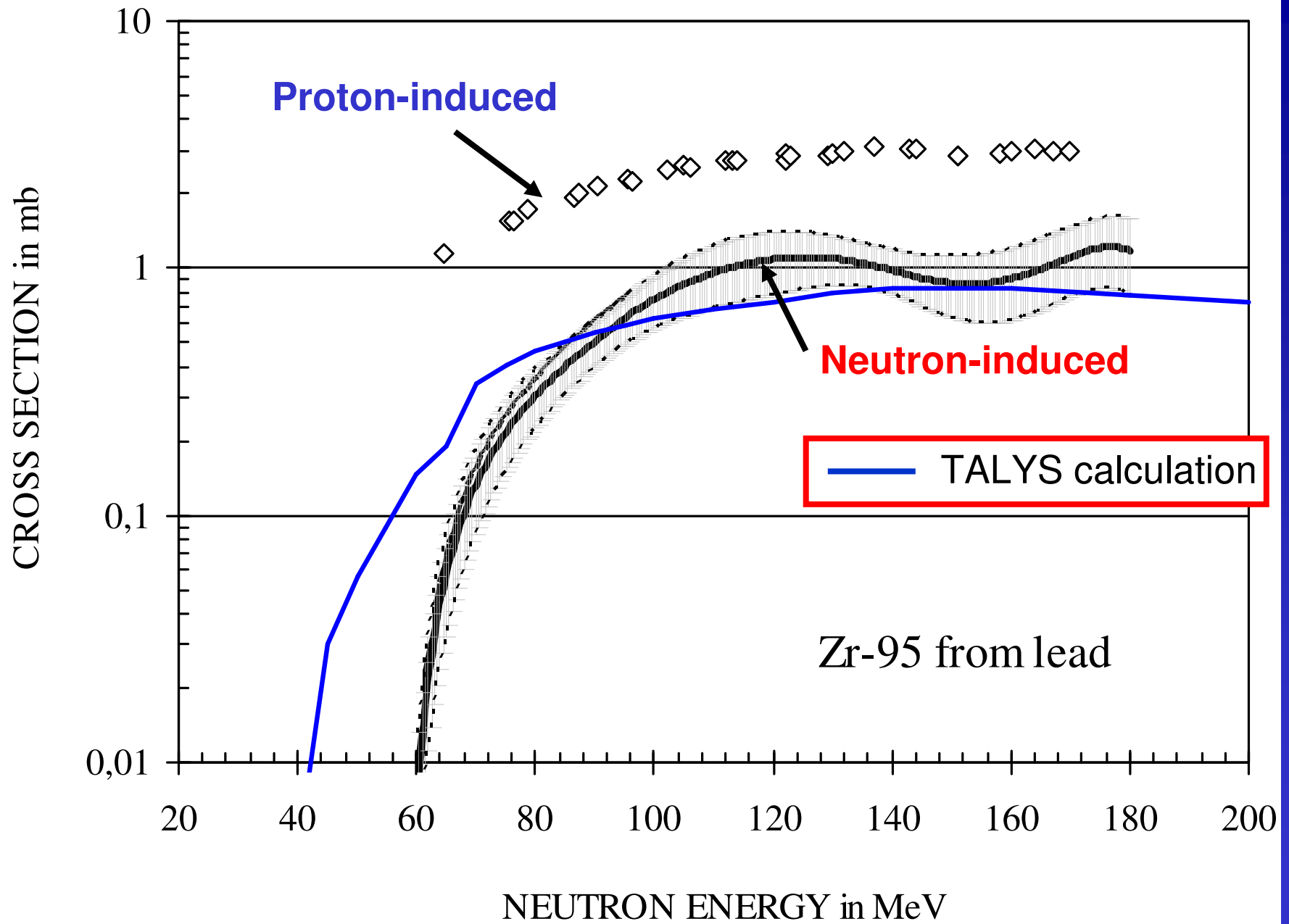
Experiments with Quasi Mono-Energetic Neutrons Produced by the ${}^7\text{Li}(p,n)$ -Reaction

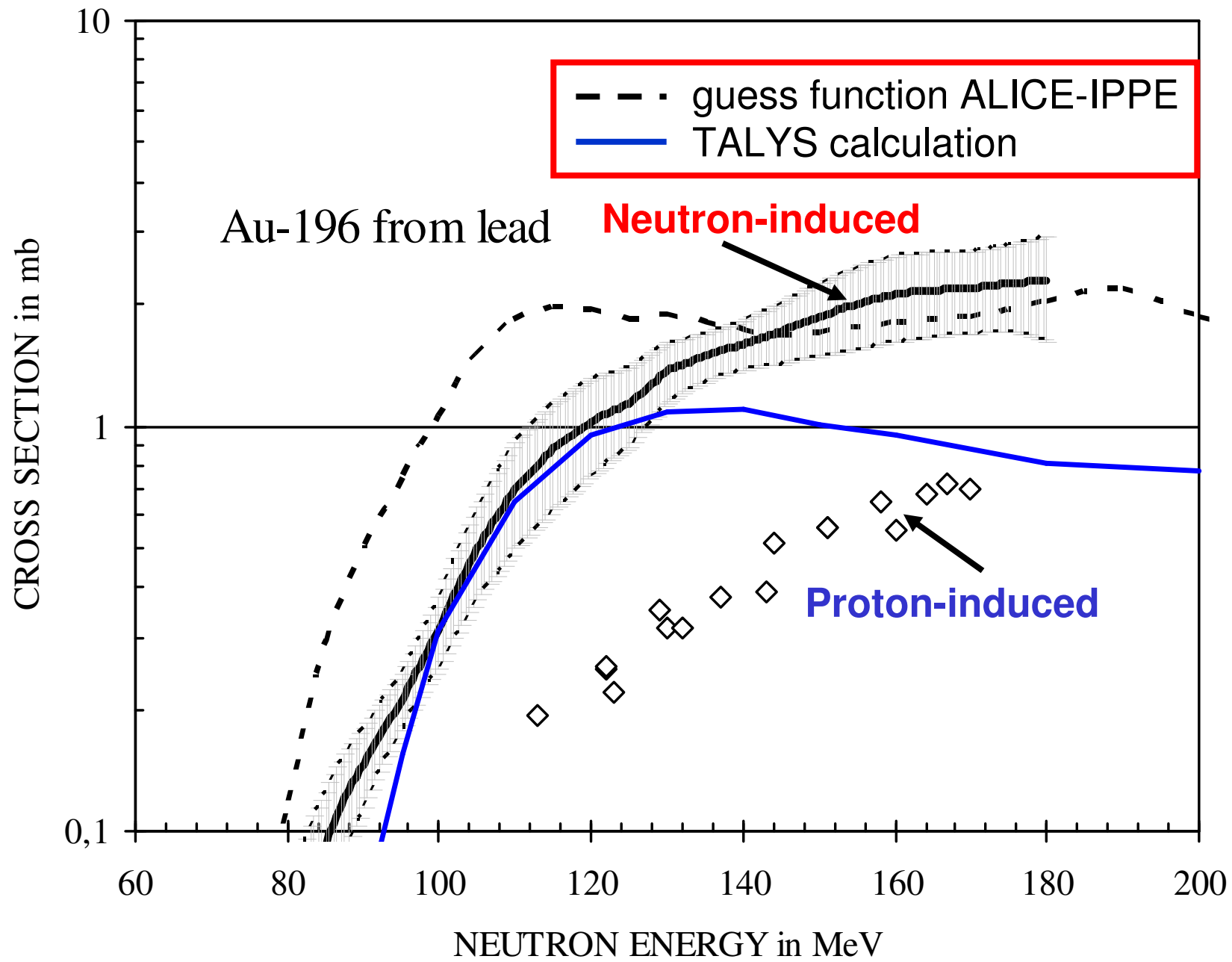




Production of Be-7 from oxygen







Contents

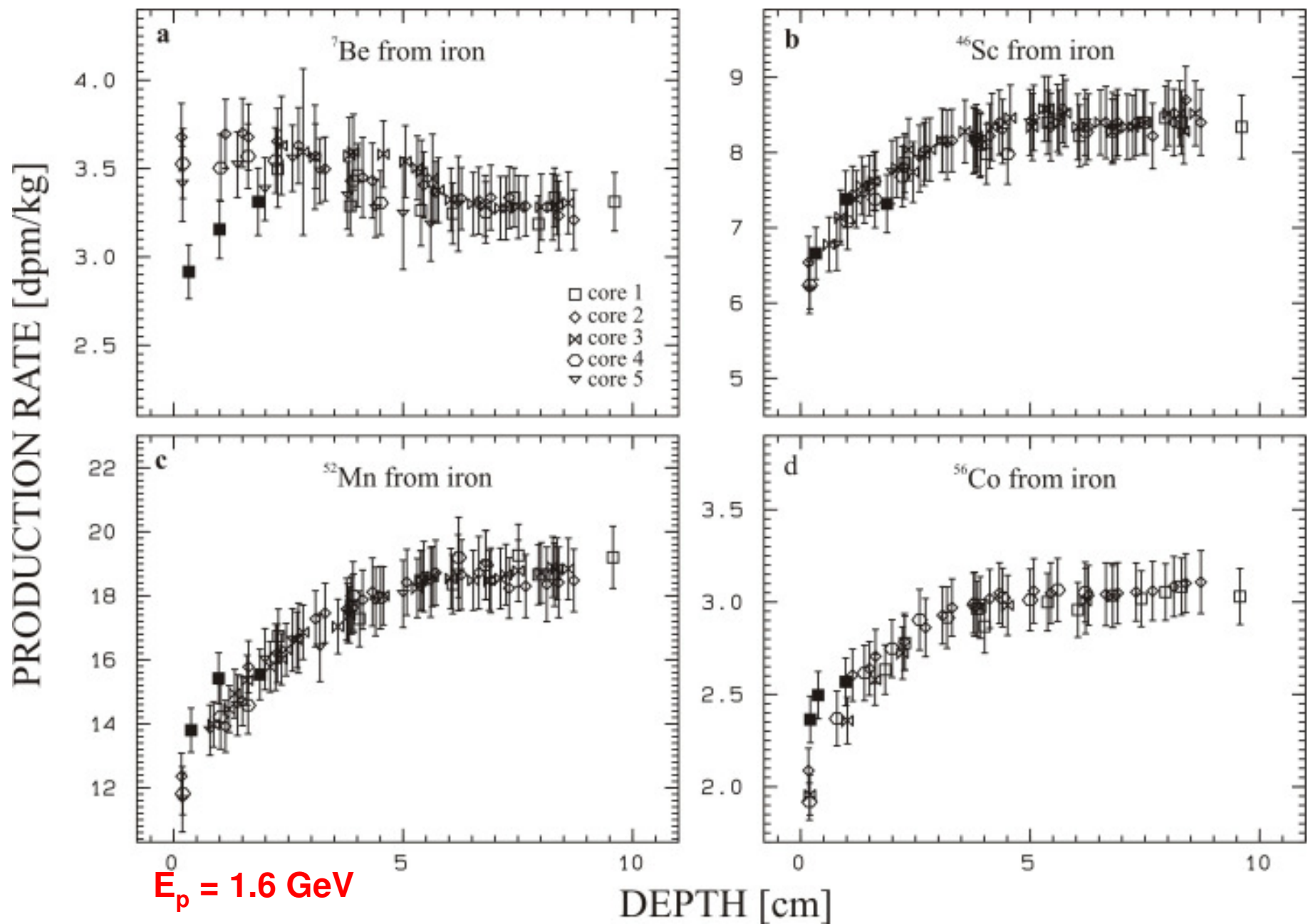


- **Relevance of data for the production of residual nuclides at medium energies**
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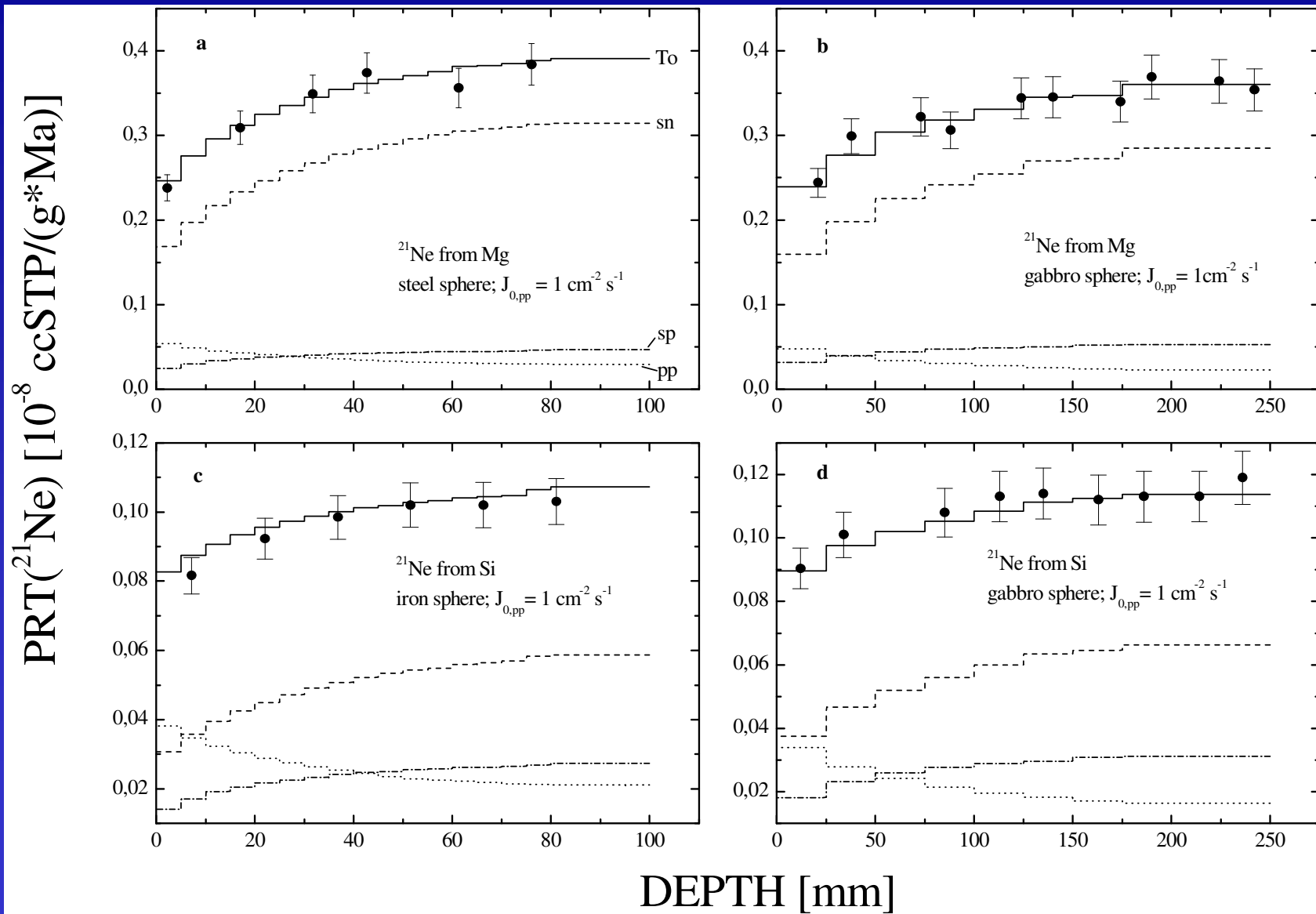


Thick target data as candidates for testing the performance of models and codes?

- Target matrix Fe
- $R = 10$ cm
- $E_p = 1.6$ GeV
- 870 individual targets
- 28 target elements



Production of ^{21}Ne in Simulation Experiments



Conclusions



- **Relevance of data for the production of residual nuclides at medium energies**
- **The 1997 NEA intercomparison**
- **Modeling of cosmogenic nuclides in extraterrestrial matter**
- **Candidate data for testing the performance of model codes on the basis of recent achievements and ongoing work**
 - ❖ **Thin-target data**
 - ❖ **Thick-target data**

Conclusions



Candidate data for testing the performance of model codes

➤ Comprehensive, systematic and representative coverage of

- ❖ particle types: protons, neutrons, He-4
- ❖ target elements: from C,N,O to U
- ❖ product nuclides: $3 \leq A \leq A_T + 1$
- ❖ energies: 10 – 10,000 MeV

to test the systematic of production cross sections.

➤ Special reactions for testing extremes of nuclear reactions:

- ❖ unbalanced nucleons in the exit channel,
- ❖ competition between spallation and fission,
- ❖ competition between evaporation, PE decay, and fission,
- ❖ (multi-)fragmentation

for testing extremes of nuclear reactions.



The End