



**The Abdus Salam  
International Centre for Theoretical Physics**



**1944-20**

**Joint ICTP-IAEA Workshop on Nuclear Reaction Data for Advanced  
Reactor Technologies**

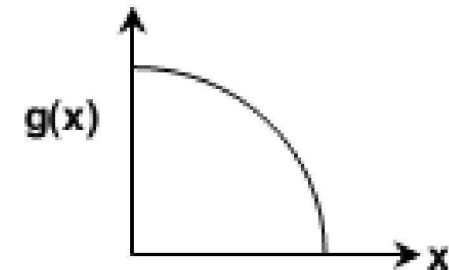
*19 - 30 May 2008*

**Cross Section Measurements and Uncertainties of Cross Section Data.**

BORELLA Alessandro  
*Institute For Reference Materials and Measurements  
EC JRC IRMM  
Retiesweg 111  
B-2440 Geel  
BELGIUM*

## Simple Monte Carlo Example

Evaluate  $G = \int_0^1 g(x) dx$ , with  $g(x) = \sqrt{1-x^2}$

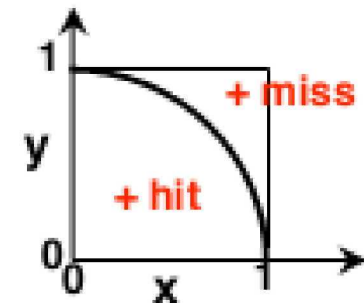


### Simulation approach:

"darts game"

For  $k = 1, \dots, N$ : choose  $\hat{x}_k, \hat{y}_k$  randomly in  $(0,1)$ ,  
 if  $\hat{x}_k^2 + \hat{y}_k^2 \leq 1$ , tally a "hit"

$$G = [\text{area under curve}] = (1 \cdot 1) \cdot \frac{\text{number of hits}}{N}$$



- **Monte Carlo approach to particle transport**
  - Random Numbers are generated to solve particle transport
  - Vs. deterministic codes
  
  - Relies on:
    - Physics and particle transport equations
    - Data (cross sections) for the various process
      - Where data are not available relies on models
  
  - The history are tracked by generating random number (RN):  
so-called “random walks”
  - According to the value of the RN, the XS and the physics, the particle fate will be different
  - The source particle and possible secondary particles are tracked in the medium according to the geometry and composition
  - Quantities of interested (the so-called tallies) are computed
  
  - More info and documentation: <http://mcnp-green.lanl.gov/>

**MCNP is a code developed by LANL to simulate the transport of neutrons, gamma rays and electrons by the Monte Carlo method. It simulates a coupled transport, i.e., it also accounts for transport of secondary particle resulting the interaction of primary particles.**

**MCNPX extends the capacities of MCNP to other particles (e.g. charged particles, heavy ions, pions etc.)**

## Problems that can be solved/Fields of application

- **Reactor calculations**
  - Burnup
  - Criticality
- **Medical application**
  - Radiography
- **Detector**
  - e.g. C6D6
- **High energy physics (MCNPX)**
- **Space applications**

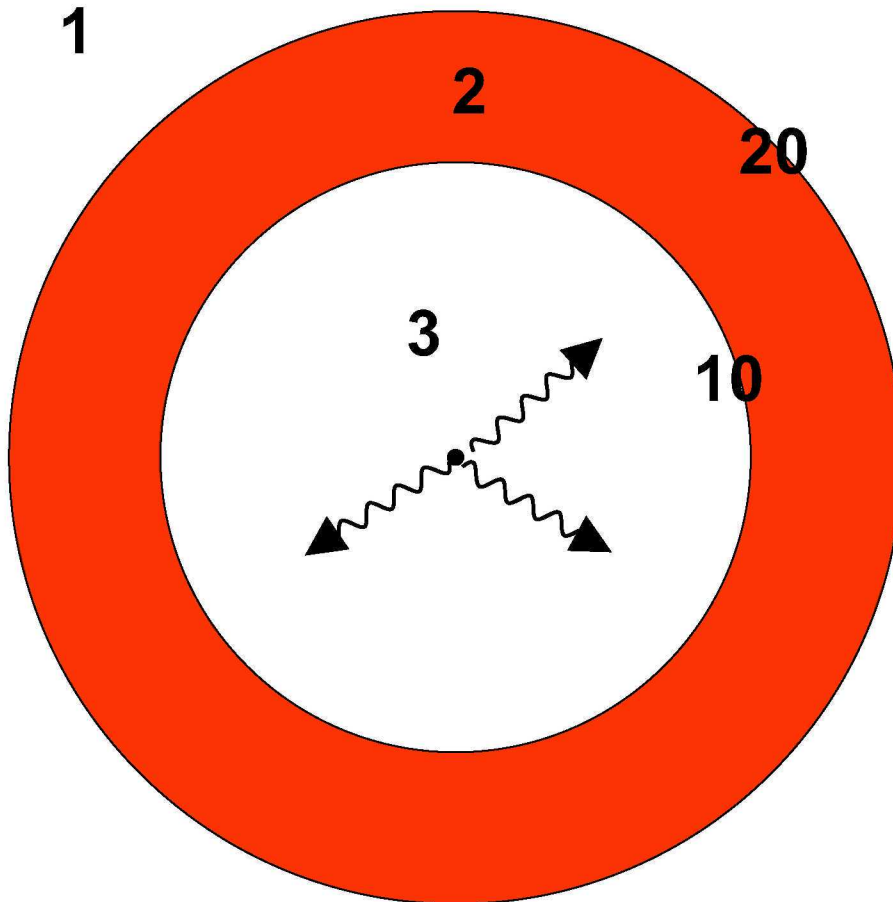


Application	# Groups	Percent
Medical (BNCT, proton therapy, etc.)	50	15
Spacecraft, Cosmic Rays, SEE, propulsion	42	12
Detectors, experiments, Threat Reduction	39	11
ATW, ADS, Energy Amplifiers	37	11
Fuel cycles, beginning to end, including storage	32	9
Accelerator Shielding and Health Physics	28	8
Theoretical Physics	23	7
Neutron Production for Scattering	21	6
Isotope Production	14	4
Radiography	12	4
MCNPX/MCNP code development	11	3
Homeland Security	10	3
Materials studies (IFMIF)	6	2
Radioactive Ion Beams	5	1
Irradiation Facilities	4	1
Neutrino Targets	4	1
Light Sources, electron machines	3	1

- **Examples**
  - Gamma ray transport
  - Neutron transport
- **Structure of the input file**
  - Cell
  - Surfaces
  - Materials (XS)
  - Source definition
  - Tally
  - Additional options
- **Output files**
  - Binary
  - Out file
  - Other files (mctal)

- F1: Surface Current
- F2: Surface Fluence
- F4: Cell Fluence
- F5: Detector Fluence
- F6: Energy Deposition
- F7: Fission Energy Deposition
- F8: Pulse Height



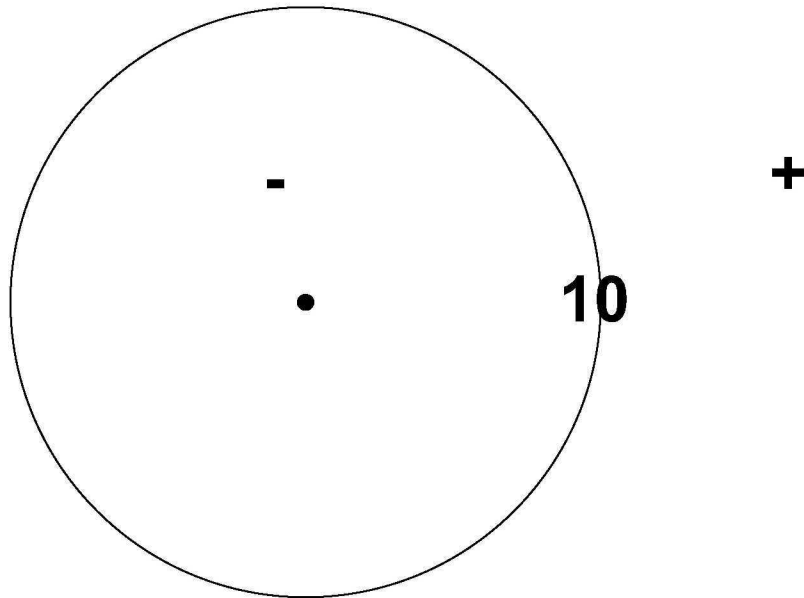


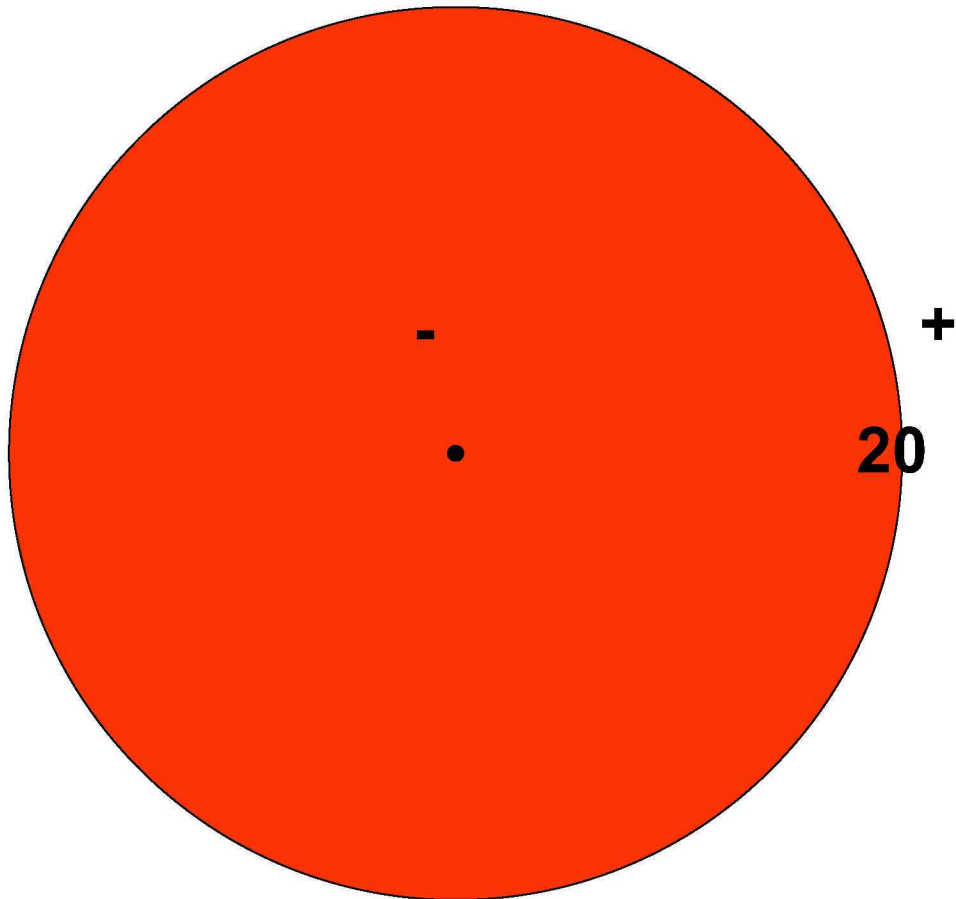
**Point source in (0,0,0)**  
**Isotropic source**

$$E_{\gamma} = 1 \text{ MeV}$$

**Response in a sphere**  
**Deposited energy**  
 **$R_{in}=10 \text{ cm}$   $R_{out}=15 \text{ cm}$**

**C<sub>6</sub>D<sub>6</sub>**  
**BaF<sub>2</sub>**





example 1: g source response for c6d6

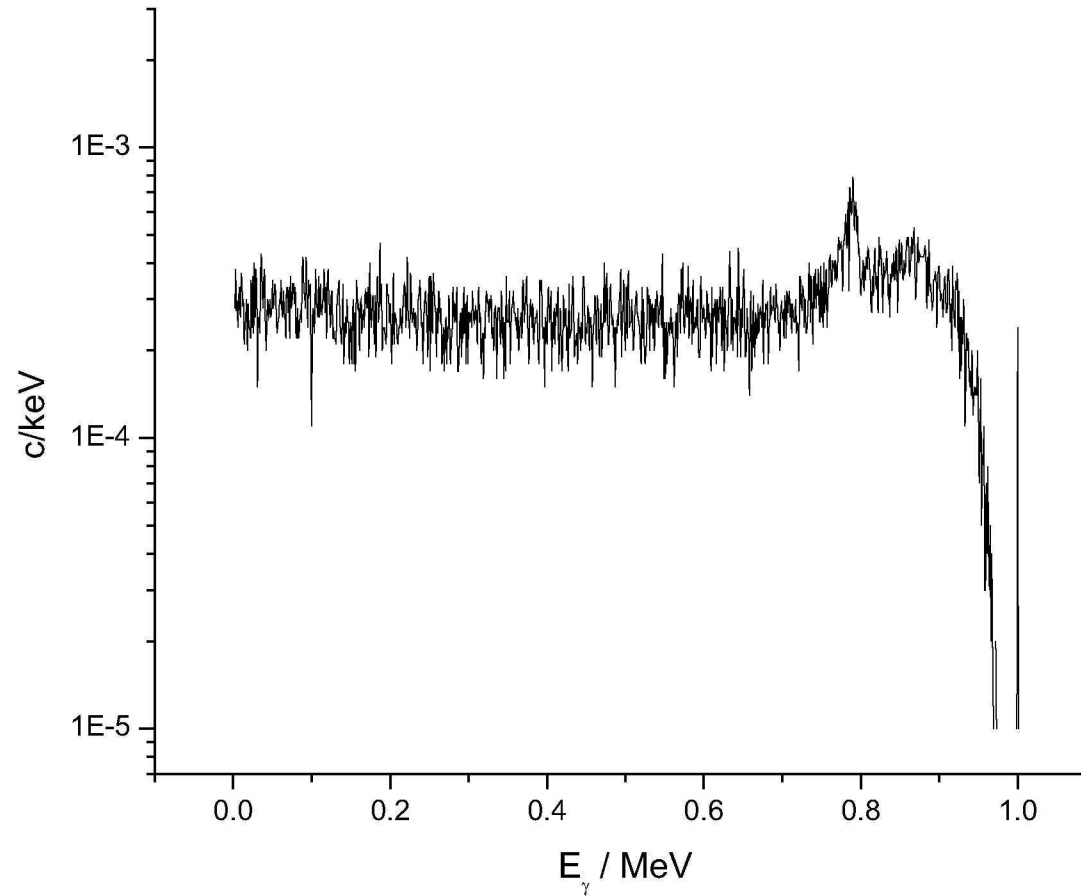
```

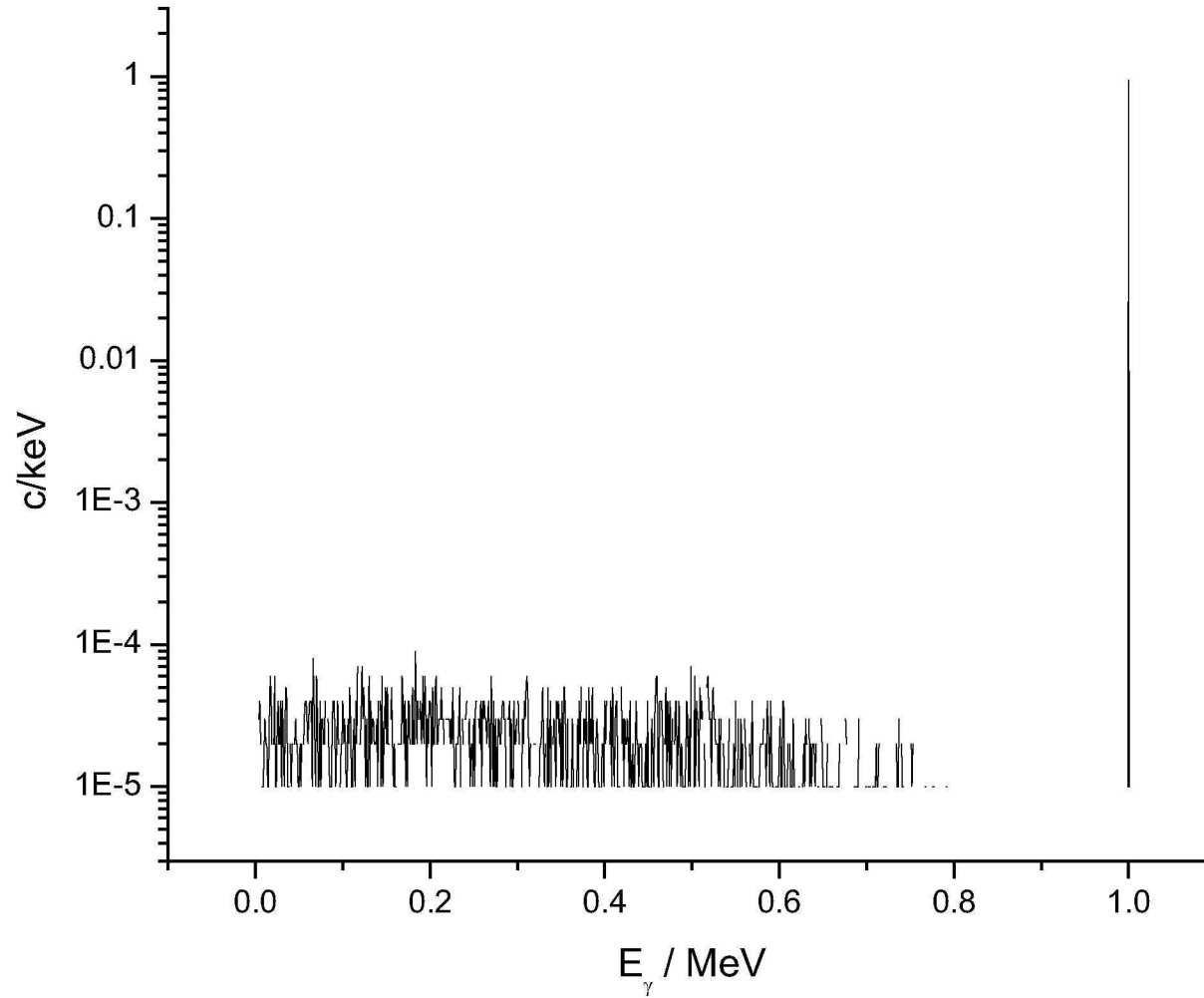
c
c -----
1  0          20          imp:n,p,e=0          $ univers
2  1 -1.00    -20 10     imp:n,p,e=1          $ univers
3  0          -10       imp:n,p,e=1          $ univers
c

c -----
c surface definitions for PM
10  so  10.0          $ global univers
20  so  15.0          $ global univers
c

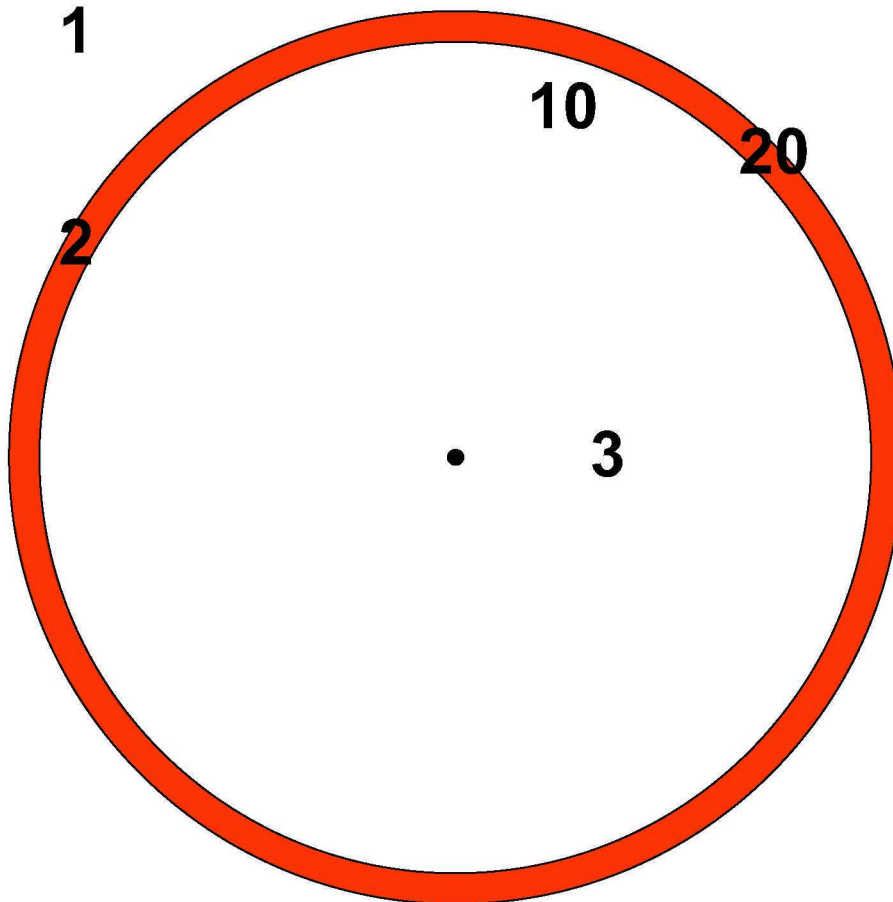
c -----
MODE P E
c
c      ZZAAA  AT1      ZZAAA  AT2
M1      6012  0.5      1002  0.496      1001  0.004          $ c6d6
c
SDEF  ERG=1  PAR=2  POS=0  0  0
c
c -----
F8:P,E  2
E8      0 999I 1.0          $ 1 keV/bin
c
NPS      1e5

```









**Point source in (0,0,0)**  
**Isotropic source**

**$E_n = 0.1-1$  MeV**

**Flux in a sphere vs E**  
**Flux through surfaces**  
 **$R_{in} = 10$  cm  $R_{out} = 10.5$  cm**

**Co+Ag**

**$\phi(E)$  spectrum ?**

example 1: n transport response for co+ag

```

c -----
1  0          30          imp:n,p,e=0          $ univers
2  1 -8.90    -20 10     imp:n,p,e=1          $ univers
3  0          -10        imp:n,p,e=1          $ univers
4  0          20 -30     imp:n,p,e=1          $ univers

```

```

c -----
c surface definitions for PM
10  so  10.0          $ global univers
20  so  10.5          $ global univers
30  so  100.0

```

```

c -----
MODE N
M1  27059  1.0          $ Co

```

```

c -----
SDEF  ERG=D1 PAR=1 POS=0 0 0
SI1  H .1 1
SP1   0 1

```

```

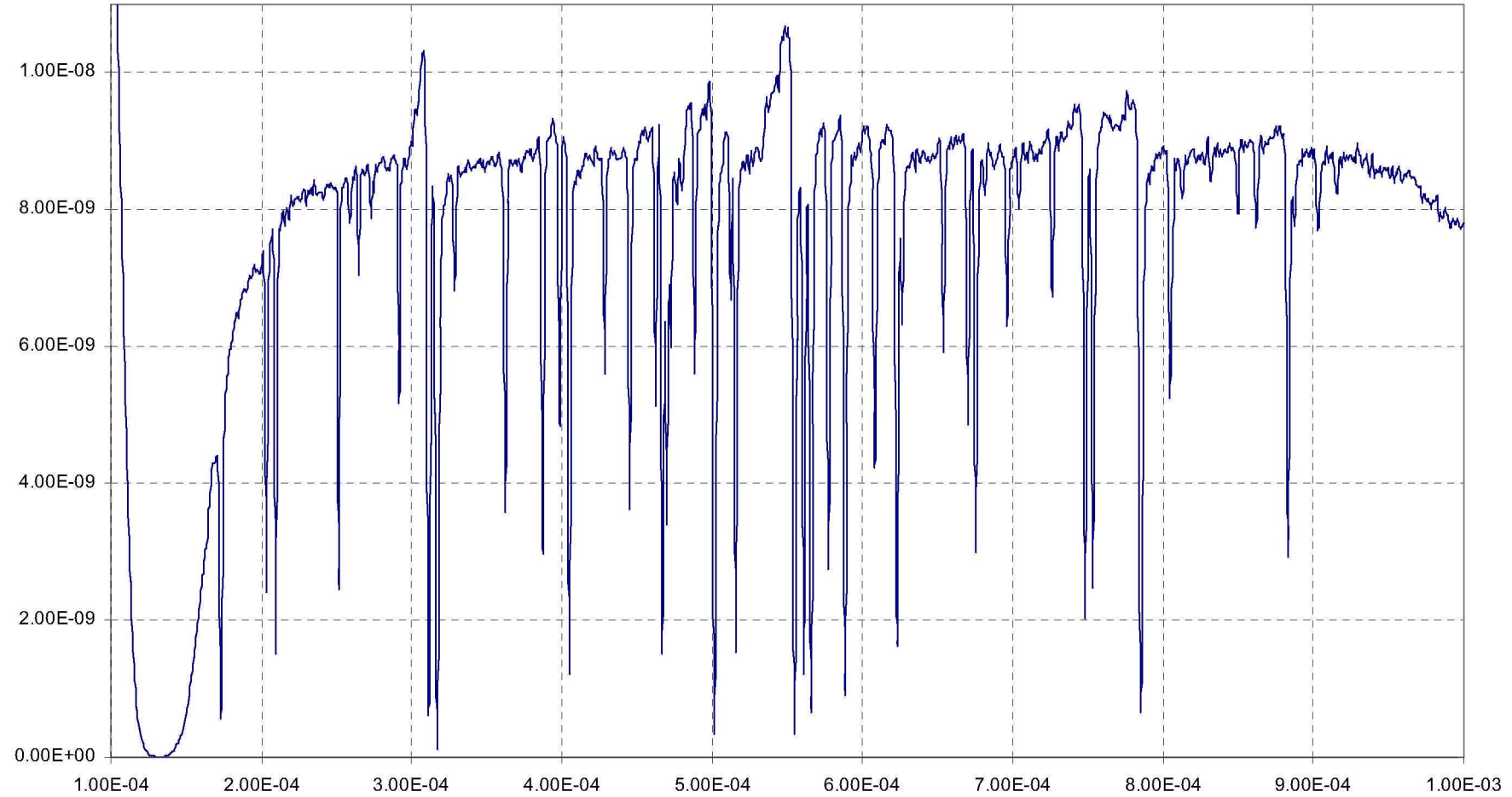
c -----
F12:n  30          $surface flux 2
T12    1.0e00 999I 1.0e03          $ 10 ns bins

```

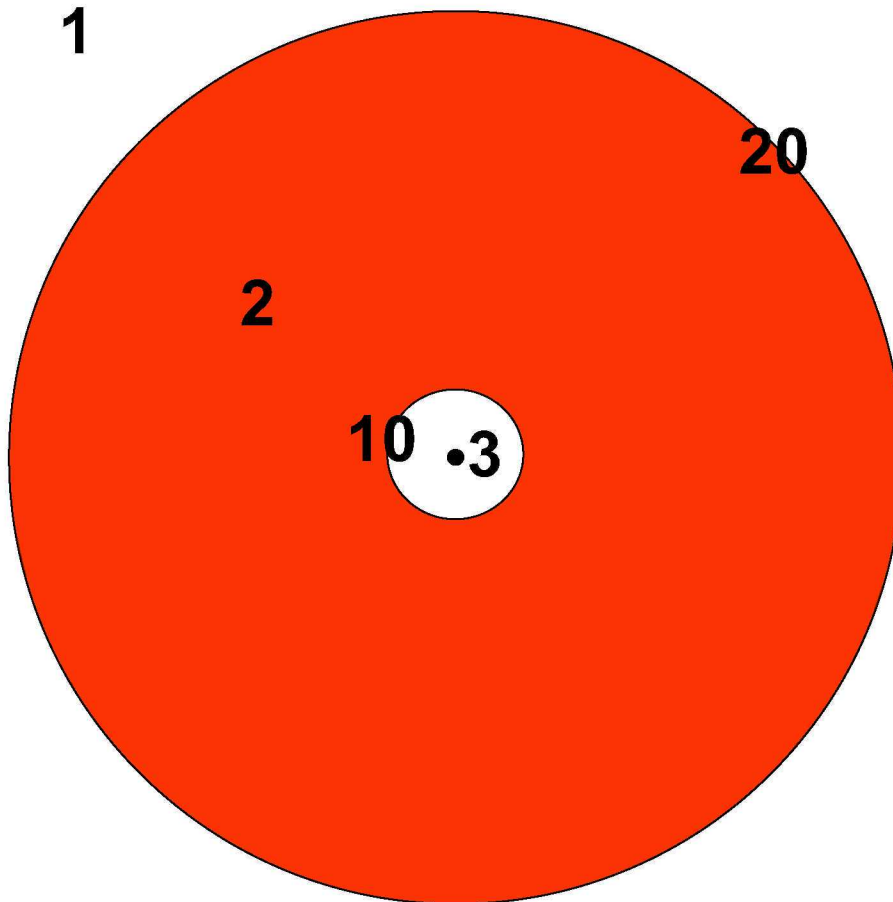
```

c -----
NPS    1e07

```



$E_n / \text{MeV}$

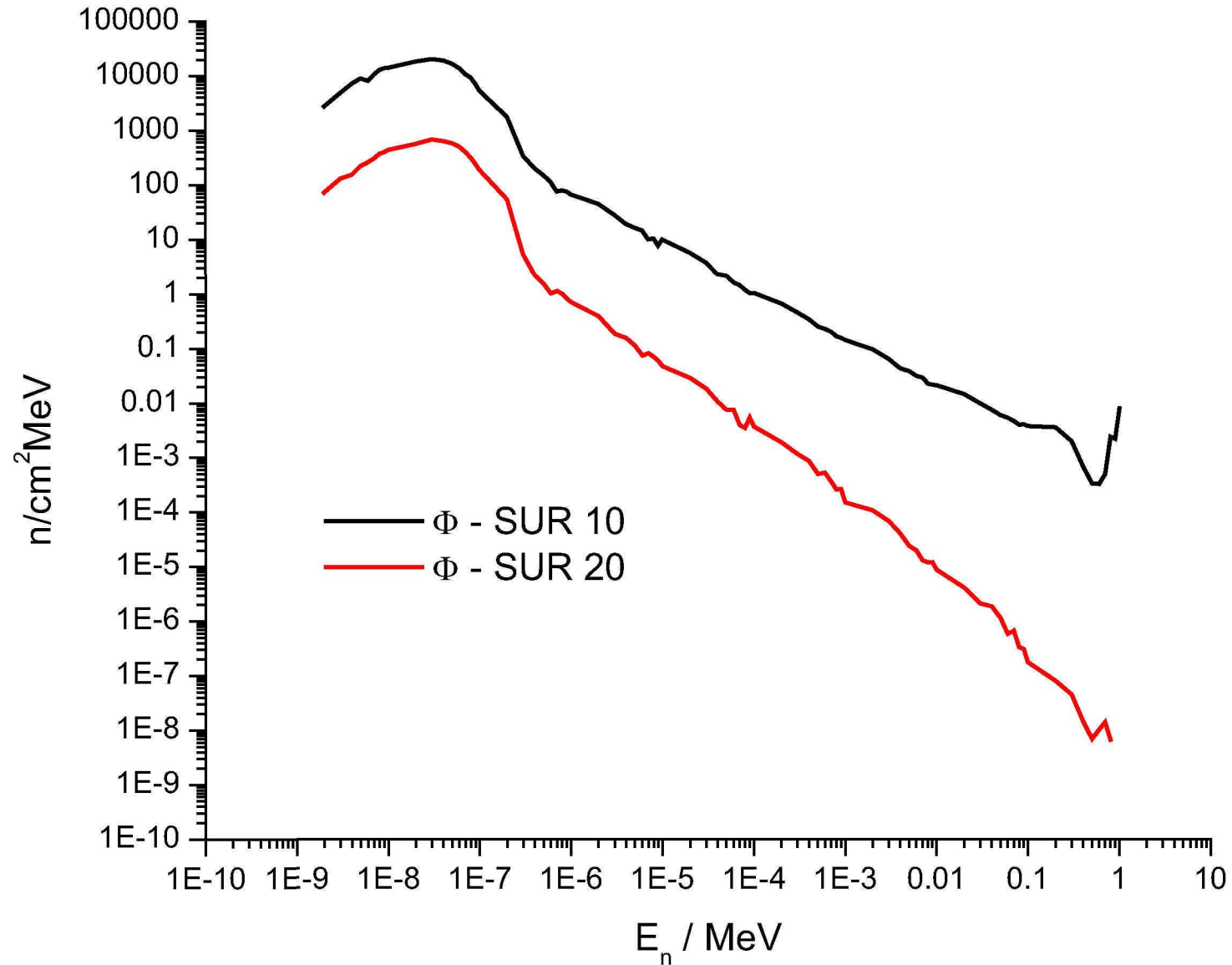


**Point source in (0,0,0)**  
**Isotropic source**

**$E_n = 1$  MeV**

**Flux in a sphere vs E**  
**Flux through surfaces**  
 **$R_{in} = 10$  cm  $R_{out} = 50$  cm**

**H<sub>2</sub>O**



## Example 2 – flux in volumes

