



1953-41

International Workshop on the Frontiers of Modern Plasma Physics

14 - 25 July 2008

Patterns of sound radiation behind point-like charged obstacles in plasma flows.

H. Pecseli University of Oslo Institute for Physics Norway



Patterns of sound radiation behind point-like charged obstacles in plasma flows

P. Guio Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, United Kingdom

W.J. Miloch University of Oslo, Institute of Theoretical Astrophysics, Box 1029 Blindern, N-0315 Oslo, Norway and School of Physics, The University of Sydney, Sydney, NSW 2006, Australia

H.L. Pécseli University of Oslo, Department of Physics, Box 1048 Blindern, N-0316 Oslo, Norway

J. Trulsen University of Oslo, Institute of Theoretical Astrophysics, Box 1029 Blindern, N-0315 Oslo, Norway

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



Basic form of the dielectric function with isothermal electrons and cold ions:

$$\varepsilon(\omega,k) = 1 + \frac{1}{(k\lambda_D)^2} - \left(\frac{\Omega_{pi}}{\omega}\right)^2$$

Dispersion relation, $\varepsilon(k,\omega) = 0$:

$$\omega(k) = \frac{C_s |k|}{\sqrt{1 + (k\lambda_{De})^2}}$$

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



Basic assumtions:

- 1. All the relevant waves can be described by the (normalized) dispersion relation $\omega(k) = |k|/(1+k^2)$, as found before.
- 2. The wavepattern has reached a stationary stage.
- 3. We consider only the "far field", which can be described by geometrical optics.



Definitions of angles and directions:



 $\ensuremath{\textcircled{C}}$ Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



In a moving frame:

$$\Omega(\mathbf{k}) = \omega(\mathbf{k}) - \mathbf{U} \cdot \mathbf{k}$$

The group velocity for the isotropic case $\omega = \omega(|k|)$:

$$\nabla_k \Omega(\mathbf{k}) = \mathbf{u}_g - \mathbf{U} = u_g \frac{\mathbf{k}}{\mathbf{k}} - \mathbf{U}$$

 $\nabla_k \Omega$ - component parallel to $AC = u_g \cos(\psi) - U$ $\nabla_k \Omega$ - component perpendicular to $AC = u_g \sin(\psi)$

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



The direction of the group velocity $\nabla_k \Omega(\mathbf{k})$ must be parallel to *AP*. We then find:

$$\tan(\theta) = \frac{u_g \sin(\psi)}{U - u_g \cos(\psi)}$$

We have not yet used the form of $\omega = \omega(k)$, only its isotropy.

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



Use $u_g = 1/(1+k^2)^{3/2}$. Also, for given *U* we have the relation between *k* and ψ :

$$\Omega(k,\psi) = \omega(k) - Uk\cos(\psi) = 0 \quad gives$$
$$\cos(\psi) = \frac{1}{U\sqrt{1+k^2}}$$

We find: $\tan(\theta) = \frac{\sqrt{U^2(1+k^2)-1}}{U^2(1+k^2)^2-1}$

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



For U < 1 (i.e. Mach number < 1) we have a limiting minimum wavenumber (by $cos(\Psi) = 1/U\sqrt{(1+k^2)} < 0$)

$$k_0 = \frac{\sqrt{1 - U^2}}{U}$$

We have no such wavenumber for the supersonic case, U > 1

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008







The maximum opening angle θ_{max} in the Mach cone, for varying U:



© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008



Phase-fronts determined by:



NB. No distinction here between two or three spatial dimensions! Also, the same results are found for moving single charges or dipoles, for instance.

ICTP Trieste, 21 July 2008





Importance of ion Landau damping, here with $T_i = T_e$, assuming quasi-neutrality in the analysis:





Limiting case of fluid vs. kinetic model? Here $T_e = 25 T_i$.



© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008





$T_e = 80 T_i$. Insulating and conducting materials, here with $U = 2.5 C_s$.





References:

- J. R. Sanmartin and S. H. Lam, Phys. Fluids 14, 62 (1971)
- K. R. Svenes and J. Trøim, Planet. Space Sci. 42, 81 (1994)
- P. Guio, W. J. Miloch, H. L. Pécseli, and J. Trulsen, Phys. Rev. E 78, 016401 (2008)
- P. Guio and H. L. Pécseli, Phys. Plasmas 10, 2667 (2003)
- W. J. Miloch, H. L. Pécseli, and J. Trulsen, Nonlinear Processes Geophys. 14, 575 (2007)
- W. J. Miloch, H. L. Pécseli, and J. Trulsen, Phys. Rev. E 77, 056408 (2008)

© Research Section for Plasma and Space Physics



Thank you for your attention!

© Research Section for Plasma and Space Physics

ICTP Trieste, 21 July 2008