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Quantum Plasmas – Recent Developments

M. Marklund University of Umea Dept. of Physics Sweden



Mattias Marklund Department of Physics Umeå University Sweden

#### Quantum Plasmas – Recent Developments

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#### Collaborators

- > L. Stenflo, G. Brodin, V. Bychkov, B. Eliasson,J. Zamanian (Umeå U.)
- > P. K. Shukla (Ruhr-U. Bochum, also S. Ali)
- > G. Manfredi (U. Strasbourg)
- > C. S. Liu (U. Maryland)



#### Overview

- > Why look at quantum plasma effects?
- > Schrödinger's description.
- > Non-relativistic single electron dynamics.
- > Paramagnetic electrons.
- > From micro to macro physics.
- > MHD regime.
- > Conclusions what the future might bring.





From G. Manfredi, Fields Inst. Comm 46, 263 (2005)





#### UMEA RRS HHL KRS

#### Quantum plasmas

- > Manifold applications (Pines, 1961; Kremp et al. 2005)
  - Condensed matter systems (Gardner, SIAM J. Appl. Math., 1994).
  - Astrophysical environments (Harding & Lai, Rep. Prog. Phys., 2006).
  - Ultracold plasmas (Rydberg states) (Li et al., PRL, 2005).
  - Nanostructured materials (Craighead, Science, 2000).
  - Laser-plasmas (Glenzer et al., PRL, 2007).
  - Spintronics.
- > Interesting fundamental aspects of matter dynamics.
- > Quantum to classical transition?
- > Collective quantum systems.



Electron properties described by complex scalar wavefunction  $\psi$  ( $|\psi|^2$  = probability)

 $i\hbar\frac{\partial\psi}{\partial t} = H\psi$ 

where we have the Hamiltonian operator

$$H = -\frac{\hbar^2}{2m_e}\nabla^2 - e\phi$$

and  $\phi$  is the *external* electrostatic potential and e being the magnitude of the electron charge.



- > Nice approach: allows for easy generalizations, new interactions can be incorporated in Hamiltonian.
- > Microscopic equations of motion

$$\frac{dF}{dt} = \frac{\partial F}{\partial t} + \frac{1}{i\hbar}[F,H]$$

for some operator F, [F, H] Poisson brackets. Example:

$$\boldsymbol{v} \equiv rac{d\boldsymbol{x}}{dt} = [\boldsymbol{x}, H] = rac{\boldsymbol{p}}{m_e}$$

for previous scalar electron description.



> Quantum statistical dynamics

 $W(t, \boldsymbol{x}, \boldsymbol{p}) = \frac{1}{(2\pi\hbar)^3} \int d^3 y \, \exp(i\mathbf{p} \cdot \boldsymbol{y}/\hbar) \langle \psi^*(t, \boldsymbol{x} + \boldsymbol{y}/2) \psi(t, \boldsymbol{x} - \boldsymbol{y}/2) \rangle$ 

$$\left(\frac{\partial}{\partial t} + \frac{\mathbf{p}}{m} \cdot \nabla\right) W + \frac{2e}{\hbar}\phi \sin\left(\frac{\hbar}{2} \overleftarrow{\nabla} \cdot \overrightarrow{\nabla}_p\right) W = 0$$

> Fluid moments from quantum kinetic equation, or from summing up particle contributions in Madelung picture.



> Quantum pressure (gradient of Bohm-de Broglie potential)

$$m\frac{d\boldsymbol{v}}{dt} \sim \frac{\hbar^2}{2m} \nabla \left(\frac{\nabla^2 \sqrt{n}}{\sqrt{n}}\right)$$

> Gives higher order dispersion (spreading of electronic wave function).



#### Plasmonic devices

- > *Surface plasmon polaritons* propagating on conductor surface.
- > Classically, surface is sharp.
- > Broadening of surface layer due to quantum dispersion.
- > Finite width gives damping!







#### Spin contributions

Electron properties described by complex *spinor* wavefunction (spin degrees of freedom)



and the Pauli Hamiltonian operator

$$H = \frac{1}{2m_e} \left( \frac{1}{i\hbar} \boldsymbol{\nabla} + \frac{e}{c} \boldsymbol{A} \right)^2 + \mu_B \boldsymbol{B} \cdot \boldsymbol{\sigma} - e\phi$$

Here *A* is the vector potential,  $\sigma$  is the spin operator,  $\mu_B = e\hbar/2m_e$  is the Bohr magneton.

#### Quantum equations of motion for electrons

$$\boldsymbol{v} = \frac{1}{m_e} \left( \boldsymbol{p} + \frac{e}{c} \boldsymbol{A} \right)$$

$$m_e \frac{d\boldsymbol{v}}{dt} = -e(\boldsymbol{E} + \boldsymbol{v} \times \boldsymbol{B}) - \frac{2\mu_B}{\hbar} \boldsymbol{\nabla}(\boldsymbol{B} \cdot \boldsymbol{S})$$

$$rac{dm{S}}{dt}=rac{2\mu_B}{\hbar}m{B} imesm{S}$$
 where  $m{S}=(\hbar/2)m{\sigma}$  .



#### Multiparticle theory

- > For spin systems: Decompose spinor wave function  $\psi_{\alpha} = \sqrt{n_{\alpha}} \exp(iS_{\alpha}/\hbar)\varphi_{\alpha}$ ,  $\varphi_{\alpha}$  unit spinor .
- > Electron fluid equations (not complete!)

$$m_e \left(\frac{\partial}{\partial t} + \boldsymbol{v}_e \cdot \boldsymbol{\nabla}\right) \boldsymbol{v}_e = -e \left(\boldsymbol{E} + \boldsymbol{v}_e \times \boldsymbol{B}\right) - \frac{\boldsymbol{\nabla} p_e}{n_e} - \frac{2\mu_B}{\hbar} S_a \boldsymbol{\nabla} B^a$$

 $\left(\frac{\partial}{\partial t} + \boldsymbol{v}\cdot\boldsymbol{\nabla}\right)\boldsymbol{S} = \frac{2\mu_B}{\hbar}\boldsymbol{B}\times\boldsymbol{S} + \text{thermal and nonlinear spin terms}$ 



#### Maxwell's equations

Due to the intrinsic magnetization, given by

$$oldsymbol{M} = -rac{2\mu_B n_e}{\hbar}oldsymbol{S},$$

Ampère's law is modified according to $abla imes oldsymbol{B} = \mu_0 (oldsymbol{j} + oldsymbol{
abla} imes oldsymbol{M}) + rac{1}{c^2} rac{\partial oldsymbol{E}}{\partial t}.$ 

Gives dynamic spin contribution to Maxwell's equations.



#### MHD regime

Single fluid dynamics (for lowest order coherent spin)

$$egin{aligned} &
ho rac{dm{v}}{dt} = -m{
abla} \left( rac{B^2}{2\mu_0} - m{M} \cdot m{B} 
ight) + m{B} \cdot m{
abla} \left( rac{1}{\mu_0} m{B} - m{M} 
ight) - m{
abla} p + rac{\hbar^2 
ho}{2m_e m_i} \left( rac{
abla^2 \sqrt{
ho}}{\sqrt{
ho}} 
ight) \ &rac{\partial m{B}}{\partial t} = m{
abla} imes \left\{ m{v} imes m{B} - rac{\left[ m{
abla} imes (m{B} - \mu_0 m{M}) 
ight] imes m{B}}{e\mu_0 m_i} - M_a m{
abla} B^a 
ight\} \end{aligned}$$

For high densities /low temperature: only electrons above Fermi level contribute to magnetization; implies reduced magnetization.

For high densities Model magnetization using Brillouin function for spin-1/2 /low temperature: particles (for long enough time scales)

$$\boldsymbol{M} = \mu_B n_e \tanh x \, \hat{\boldsymbol{B}}$$

where the Zeeman energy  $x = \mu_B B / k_B T_e$  gives the degree of alignment through the Brillouin function. For high temp., magnetization  $\longrightarrow 0$ .

# Instabilities and ferrofluids

<u>Ferrofluids</u> Nanostructured paramagnetic fluids. Formalism similar to the above applicable.

Normal field instability - saturated by gravity and surface tension (Cowley & Rosensweig 1967).





#### Spin kinetics

Spin dependent distribution function: f = f(t, x, v, s)

$$\frac{\partial f}{\partial t} + \frac{d\boldsymbol{x}}{dt} \cdot \nabla_{\boldsymbol{x}} f + \frac{d\boldsymbol{v}}{dt} \cdot \nabla_{\boldsymbol{v}} f + \frac{d\boldsymbol{s}}{dt} \cdot \nabla_{\boldsymbol{s}} f = 0$$

(see also Cowley et al. Phys. Fluids 1986; Kulsrud et al. Nucl. Fusion, 1986)

Quantum equations of motion gives semiclassical dynamics.

See Gert Brodin's talk.



### E Spin kinetics

Formal structure: Wigner matrix from density matrix

 $W_{ab}(t, \boldsymbol{x}, \boldsymbol{p}) = \frac{1}{(2\pi\hbar)^3} \int d^3 y \, \exp(i\boldsymbol{p} \cdot \boldsymbol{y}/\hbar) \rho_{ab}(t, \boldsymbol{x} + \boldsymbol{y}/2, \boldsymbol{x} - \boldsymbol{y}/2)$ 

Distribution function defined by

 $f(t, \boldsymbol{x}, \boldsymbol{p}, \boldsymbol{s}) = \frac{1}{2}(1 + \boldsymbol{s} \cdot \boldsymbol{\sigma})_{ab} W_{ab}$ 

Evolution determined by governing equation for density matrix. Lowest order terms gives semiclassical kinetic equation. Generalized spin Wigner equation (J. Zamanian et al., submitted; Arnold & Steinrück, ZAMP, 1989).

#### Issues in relativistic quantum plasmas

rate

rate

 $\boldsymbol{\chi}$ 

> Pair production: "Schwinger mechanism" for fields with spatial and temporal variation.

> Temporal compression: *increased* production rate.

> Spatial compression: *lower* production rate.

> Strong fields + relativistic plasma particles: can computational models be developed?

Quantum field theoretical models (Melrose, 2008)?

#### Gigagauss laboratory fields

- > Currently, gigagauss laboratory fields generated in solid-laser interactions.
- > Look for quantum plasma effects:
  - > Landau quantization.
  - > Spin effects.

> ...



#### Conclusions

- New important effects appear from collective quantum domain.
- Wide ranging possibilities for applications.
  - Nanomaterials.
  - Astroplasmas.
  - Ultracold plasmas.
- Interesting future possibilities:
  - Theoretical development: Dense, relativistic plasmas using computationally viable models.
  - Solitons and plasmonics.