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Transport and mixing in plasma turbulence.

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Transport and mixing in plasma turbulence

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Outline

Motivation: In magnetically confined plasmas turbulent – anomalous – transport is the dominant mechanism for transport of mass, energy and momentum! (Balescu, *Aspects of Anomalous Transport in Plasmas,* CRC Press 2005)

- Turbulent dispersion and mixing of (passive) particles:
 - in drift wave turbulence
 - in global interchange turbulence in the edge/scrape-off-layer, SOL
- Relation between the passive particle diffusion and bulk plasma transport – the turbulent density flux
- Passive tracer particles/fields are used to model impurity transport
- Inertia effects; pinching and clustering in 2D drift wave turbulence
- Particle mixing and transport in strongly intermittent turbulence in SOL: Curvature pinch in inhomogeneous magnetic fields.

Particle dispersion in plasma turbulence

Vorticity



Drift-wave turbulence

Hasegawa Wakatani Equations PRL **50**, 682 (1983) Particles convected by fluctuating *ExB* velocity



Comm.Nonlinear Sci. Numer. Simul. 8, 477 (2003)³

Dynamics of passive tracers

- Investigation of mixing and diffusion properties of turbulence provides a diffusion coefficient – turbulent flux via Fick's law.
- Passive tracers or passive fields are also widely applied for modelling impurity ion transport in plasmas :: impurities are all materials besides the bulk plasma species; but here is does not include dust particles, only impurity ions are treated!
- Impurities, e.g., originate from sputtering off plasma facing components, PFC
- Passive tracers/fields do not contribute to charge neutrality
 and do not dynamically react back on the turbulence
- Severe condition: impurity density should be much smaller than density of bulk ions and contribution to quasi-neutrality condition much smaller than each of the contributing terms. (Naulin et al Phys. Scripta T122, 129 (2006)).

Passive scalar dynamics is a classical problem in fluid dynamics Falkovich et al. Rev. Mod. Phys. **73**, 913 (2001); Warhaft, Ann. Rev. Fluid Mech. **32**, 203 (2000)

Tracing particles in the turbulence

Up to 100.000 particles are adverted in resistive drift-wave turbulence – Hasegawa-Wakatani (PRL 1983) model 2D.

$$\vec{x}(t) = \vec{x}_0 + \int_0^t \vec{v}(\vec{x}(t'), t') dt$$

Principal component of $\vec{v} = (u, v)$ is the $E \times B$ -velocity, \vec{v}_E :: Ideal inertia-less particles.

Inertial effects : adding the polarization drift, $\vec{v_p}$

$$ec{v}_p = -\zeta \left(rac{\partial}{\partial t} + (ec{v}_E \cdot
abla)
ight)
abla arphi$$

 $\zeta = \frac{eM}{qm_i} \frac{\rho_s}{L_n}$

Important for heavier impurities!

Inertia effects make the advection velocity compressible!

Particle dispersion



The mean square particle displacement radially and poloidally. x - radial direction,

y - poloidal direction

The radial particle displacement. Fit: $At^{\beta} + B$ for t > 400; $\beta = 1$. Asymptotically normal diffusion!

Diffusion coefficient $D_x = \langle X^2(t) \rangle / 2t$

Particle trapping in moving vortical structures:

Diffusion coefficient and flux



Comparison between D_x for tracer particles and D from the flux Γ ; D = - $\Gamma/d_x n_0$

For the present case – HWe, fluctuations around a frozen background profile - passive particle diffusion really mimics bulk plasma transport!

Evolution of impurities as a passive field

The impurities are treated as a passive scalar advected by the turbulent fluctuations, i.e., the the impurities do not act back on the turbulence or the background plasma profile.

$$\partial_t n_{imp} + \nabla \cdot (\vec{v} n_{imp}) = \mu \nabla^2 n_{imp}$$

The influence of inertia enters via the polarization.

$$(\partial_t + \mathbf{v}_E \cdot \nabla) n_{imp} = \zeta \nabla \cdot (n_{imp} (\partial_t + \mathbf{v}_E \cdot \nabla) \nabla \phi) + \mu \nabla^2 n_{imp}.$$

Restriction $n_{imp} \ll n!$

Lagrangian invariant: $(\partial_t + \mathbf{v}_E \cdot \nabla)(\ln n_{imp} - \zeta \omega) \approx 0$ Turbulent mixing will homogenize the Lagrangian invariant : $\ln n_{imp} - \zeta \omega \approx const.$ The initially homogeneous impurity density field will granulate.

Clustering/aggregation of inertial impurities



The impurity equation may be written as:

$$D_t(\ln n_{imp} - \zeta \omega) = o(\zeta^2)$$

 $n_{imp}/n_{imp0}\sim\zeta\omega$

Positive impurities ($\varsigma > 0$) (this case) cluster in positive vortices

Negative impurities ($\varsigma > 0$) will cluster in negative vortices

Priego et al Phys. Plasmas **12**, 062312 (2005)

Impurity density and vorticity



Scatter plot of impurity density and vorticity, $\zeta = 0.05$, $\zeta = 0.01$, and $\zeta = 0.002$.

Linear regression: $\theta/\theta_0 = 1 + K\omega$; $K = 0.82\zeta$ ($\theta \equiv n_{imp}$)

Priego et al Phys. Plasmas **12**, 062312 (2005) 10

Impurity pinch

Finite inertia also introduce a pinch effect: the (positive) impurities are transported up the density gradient – negative pinch velocity



Specific properties of the HWe ?

Turbulence and transport in the edge/SOL

In the edge/scrape-off-layer (SOL) region turbulence and transport is strongly intermittent and characterized by:

 Iarge-amplitude, radially propagating blob-like structures of particles and heat, generated close to the last closed flux surface (LCFS),
 resulting in asymmetric conditional wave forms, and skewed and flattened PDFs with broad tails
 results in localized power loads at PFCs.

Observed under a variety of conditions (linear to tororoidal devices) : see, e.g., Zweben Phys. Fluids **28** 974 (1985); Antar *et al.* PoP **10**, 419 (2003); Boedo *et al* PoP **10**, 1670 (2003); Zweben *et al.* Nucl. Fus. **44**, 134 (2004); Grulke *et al* PoP **13**, 012306 (2006); Garcia *et al.* PPCF **48**, L1 (2006). S.J. Zweben *et al.* PPCF **49**, S1 (2007)

Blob propagation in Alcator, C-Mod

OBSERVATIONS OF DENSITY BLOBS AT THE OUTBOARD MIDPLANE OF A LCATOR C-MOD (D $_{\ensuremath{\mathbb{R}}}$ - LIGHT) O. G RULKE ET AL POP 13, 012306 (2006).

Review: S.J. Zweben et al. PPCF 49, S1 (2007)





Inferred observations in the poloidal/toroidal direction

—Magnetic field line

Radial velocity: 0.05 – 0.1 of the sound speed

Simulations of Edge-SOL

Risø ESEL code: interchange dynamics at the outboard midplane of a toroidally magnetized plasma. B-field gradient and curvature. Global evolution.



Garcia et al, PRL **92**, 165003 (2004); Phys. Plasmas **12**, 062309 (2005); Phys. Scripta **T122**, 89 (2006)

Energetics and energy transfer



Bursting : Kinetic energy contained by the mean, U, and fluctuating, K, motions.

The collective energy transfer terms $F_{_{p}}$ and $F_{_{v}}$

Spatial structure during a burst



Formation and propagation of density blob. Particle density (left) and vorticity (right) during a burst ($\Delta t = 500$), radial blob velocity < $0.02c_s$.

Particle density flux



Re-scaled PDF of particle density flux, $\Gamma = (n - \overline{n})v_x$, measured at the probes, P_i . Exponential tails: flux dominated by strong bursts.

Dynamics of impurity ions

The passive tracer particles model impurity dynamics, in the limit of no back-reaction on the plasma dynamics:

Impurity density is much lower than the plasma particle density. (Naulin PRE '05; Priego *et al* PoP '05; Naulin *et al* Physica Scripta '06)

Particles are advected as:

$$\frac{d\vec{x}}{dt} = \vec{v} = \frac{1}{B(x)}\,\hat{z} \times \nabla\phi$$

Finite inertia effects are neglected; \vec{v} is compressible due to the spatial dependence of B(x)

Garcia et al EPS 2005

Particle dynamics



released inside LCFS

Variogram of the particle motion, $- - - \tau^2; - - - \tau; - - \tau^{1.4}$

Step size PDF



PDF of the radial displacement, Δx , over $\Delta t = 50$; all particles. $\langle \Delta x \rangle = -0.08$, standard deviation, $\sigma = 1.02$, skewness, S = 0.4, and kurtosis, K = 10.7. **Broad exponentially decaying tails.**

Long steps are almost equally probable in both in- and outgoing directions.

Particle dispersion

Evolution of the impurity density

Evolution of the impurity/tracer particle density N₀ averaged over y.

Arrival times

The relative number of particles passing through a radial plane versus time; first passage. Particles released inside LCFS, 39 < x < 41.

Velocity of the front of the particles > $0.02c_s$, typical blob speed.

Evolution of the impurity density

Density profile $N_0(x) \propto B(x)$ independent of release position. The transport is not "Fickian" diffusion. It can be described by an effective pinch:

$$\left(rac{\partial}{\partial t}+rac{1}{B}\,\hat{z} imes
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abla
ight)rac{N}{B}=0,$$

N/B is a Lagrangian invariant: Effective turbulent mixing: N/B uniformly distributed in space.

Uniform distribution within few burst times

Impurities are effectively mixed by the turbulence in the SOL within a few burst periods. Even if originating far out in the SOL they will quickly penetrate across the LCFS into the edge plasma. Corresponding to the so-called inward (curvature) pinch.

Summary

- Dynamics of passive particles in turbulence:
- Diffusion coefficient mimics bulk plasma transport for a "fluctuation" model
- Modelling impurity transport by passive particles:
- Clustering/aggregation of inertial impurities and "inertial pinch"
- Edge/SOL turbulence and transport in a magnetically confined plasmas is bursty/intermittent with broad tailed PDFs and is not diffusive in the Fickian sense.
- No parametrized diffusion type equation: Transport characterisation calls for a universal PDF
- Impurities are effectively mixed in SOL and penetrates the LCFS.
- Impurity pinch: curvature pinch

Classical Particle Dispersion

Particle dispersion in drift wave turbulence

Particle dispersion in 2D drift wave turbulence Hasegawa-Wakatani equations (HWE): the resistive drift wave instability (PRL 50, 682 (1983)):

$$\partial_t n + \partial_y \varphi + \{\varphi, n\} = -C (n - \varphi) + \mu_n \nabla^2 n$$
$$\partial_t \nabla^2 \varphi + \{\varphi, \nabla^2 \varphi\} = -C (n - \varphi) + \mu_\varphi \nabla^4 \varphi$$

$$1/C = 1/k_{\parallel}^2 L_{\parallel}^2 L_{\parallel} = (L_n T_e/m_e c_s v_{ei})^{1/2}$$

$$\{\varphi, \psi\} \equiv \hat{z} \times \nabla \varphi \cdot \nabla \psi = \frac{\partial \varphi}{\partial x} \frac{\partial \psi}{\partial y} - \frac{\partial \psi}{\partial x} \frac{\partial \varphi}{\partial y}$$

$$u = -\frac{\partial \varphi}{\partial y}; v = \frac{\partial \varphi}{\partial x}$$

Normalization: $\rho_s = c_s / \Omega_i$ for lengths; L_n / c_s for the times; $c_s = \sqrt{T_e/m_i}$; $L_n = |(\nabla n_0(x)/n_0(x))^{-1}|$ $(T_e/e) (\rho_s/L_n)$ for potential $n_0 \rho_s/L_n$ for density; $\mu_n = \mu_{\phi} = \mu$.

> Naulin et al. Phys. Plasmas 6,4575 (1999) Basu et al Phys. Plasmas **10**, 2696 (2003); Comm.Nonlinear Sci. Numer. Simul. **8**, 477 (2003)

Energy spectrum

Spectrum is isotropic

Running diffusion coefficient

Running diffusion coefficient: $D_x = \langle X^2(t) \rangle/2t$, for varying adiabaticity parameter C.

Turbulent particle density flux

Particle density flux: $\Gamma = n v_x = n v_{ExB}$

Bursty flux !

$$p_G = \frac{1}{\pi} \frac{\sqrt{1 - \gamma^2}}{\sigma_n \sigma_{v_x}} K_0\left(\frac{|\Gamma|}{\sigma_n \sigma_{v_x}}\right) \exp\left(-\gamma \frac{\Gamma}{\sigma_n \sigma_{v_x}}\right)$$

$$\gamma$$
 is the correlation: $\gamma = -\frac{\langle v_x n \rangle}{\langle v_x^2 \rangle^{1/2} \langle n^2 \rangle^{1/2}} = \cos \alpha.$

Turbulent particle density flux

Particle density flux: flux surface averaged

The probability distribution function for the plasma flux across the magnetic field is strongly *non-Gaussian*, i.e., strong bursts are dominating!

Edge/SOL turbulence transport in JET

B. Gonçalves: 11th European Fusion Physics Workshop, Heraklion, Crete 8-11 Dec. 2003

Radial velocity PDF

Re-scaled PDF of the radial particle velocity coarse grained over time intervals $\Delta t = 50 \cdot 2^{m-1}$; particles released inside LCFS

Re-scaled PDF of the turbulent radial ExB-velocity recorded at the probes $P_1 - P_7$