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9th Workshop on Three-Dimensional Modelling of Seismic Waves Generation, Propagation and their Inversion

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Seismic waves physics: a (partially) guided tour

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The Abdus Salam International Centre for Theoretical Physics



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Seismic waves physics: a (partially) guided tour

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(Seismic) Wave physics

Basic physical concepts

What is a wave? Discrete and continuous models Born of wave equation Dispersion

Basic physical concepts 2

PDE: Poisson, diffusion and wave equation Scattering and diffusion















 Small perturbations of a _______
 Linear ________
 Harmonic Oscillation

 stable equilibrium point
 restoring force
 Oscillation

 Coupling of
 the disturbances can
 Item disturbances can

harmonic oscillators

the disturbances can propagate, superpose and stand

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Normal modes of the system



Monoatomic 1D lattice



Interatomic potential

Now we consider a monatomic 1-D lattice in the x-direction. The lattice atoms are very close to eqilibrium. Let us examine a single i-th atom and find the r_i potential as a function of displacement from equilibrium, $U(r_i)$.

We expand this potential into a Taylor's series:

$$U(r_{i}) = U(r_{0}) + (r_{i} - r_{0}) \left(\frac{dU}{dr_{i}}\right)_{r_{0}} + \frac{1}{2}(r_{i} - r_{0})^{2} \left(\frac{d^{2}U}{dr_{i}^{2}}\right)_{r_{0}} + \frac{1}{6}(r_{i} - r_{0})^{3} \left(\frac{d^{3}U}{dr_{i}^{3}}\right)_{r_{0}} + \dots$$

The first term of this expansion is just the equilibrium binding energy (= const). The second term is the slope of the potential at its minimum (= 0). The fourth and higher terms become increasingly smaller. We are therefore left with the third term as the only significant change in the potential energy for a small displacement $u = r_i - r_o$. This has the form

$$\Delta U = \frac{1}{2}Cu^2 \quad (C = \frac{d^2U}{dr_i^2} \text{ at } r_i = r_0)$$

representing the *harmonic approximation*, since it is the same as the energy stored in a spring, or the potential energy of a harmonic oscillator. Our simple model of the dynamic crystal structure should therefore be a "ball and spring" model, with the lengths of the springs equivalent to the equilibrium separations of the ion cores.

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Monatomic 1D lattice

Let us examine the simplest periodic system within the context of harmonic approximation (F = dU/du = Cu) - a one-dimensional crystal lattice, which is a sequence of masses m connected with springs of force constant *C* and separation *a*.



Note that we neglected hereby the interaction of the *n*-th atom with all but its nearest neighbors. A similar equation should be written for each atom in the lattice, resulting in N coupled differential equations, which should be solved simultaneously (N - total number of atoms in the lattice). In addition the boundary conditions applied to end atoms in the lattice should be taken into account.

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(Seismic) Wave physics

Basic physical concepts

What is a wave? Discrete and continuous models Born of wave equation Continuum mechanics Sound waves





Consider a source causing a perturbation in the gas medium rapid enough to cause a pressure variation and not a simple molecular flux.

The regions where compression (or rarefaction), and thus the density variation of the gas, occurs are big compared to the mean free path (average distance that gas molcules travel without collisions).

 \mathbf{M} The perturbation fronts are planes and the displacement induced in the gas, X, depends only on x & t (and not on y, z).





The conventional unit for pressure is bar=10⁵N/m² and the pressure at the equilibrium is: 1atm=1.0133bar

The pressure perturbations associated to the sound wave passage are typically of the order of 10⁻⁷bar, thus very small if compared to the value of pressure at the equilibrium.

One can thus assume that:

$$P=P_0+\Delta P \rho=\rho_0+\Delta \rho$$

where ΔP and $\Delta \rho$ are the values of the (small) perturbations of the pressure and density from the equilibrium.

Wave equation



Sound wave equation - 1



The gas moves and causes density variations

Let us consider the displacement field, s(x,t) induced by sound



and considering a unitary area perpendicular to x, direction of propagation, one has that the quantity of gas enclosed in the old and new volume is the same

$$\rho_{0}\Delta \mathbf{x} = \rho \Big[\mathbf{x} + \Delta \mathbf{x} + \mathbf{s}(\mathbf{x} + \Delta \mathbf{x}) - \mathbf{x} - \mathbf{s}(\mathbf{x}) \Big]$$

where, since $\Delta \mathbf{x}$ is small, $\mathbf{s}(\mathbf{x} + \Delta \mathbf{x}) \approx \mathbf{s}(\mathbf{x}) + \frac{\partial \mathbf{s}}{\partial \mathbf{x}} \Delta \mathbf{x}$
$$\rho_{0}\Delta \mathbf{x} = (\rho_{0} + \Delta \rho) \Big[\Delta \mathbf{x} + \frac{\partial \mathbf{s}}{\partial \mathbf{x}} \Delta \mathbf{x} \Big] = \rho_{0}\Delta \mathbf{x} + \rho_{0} \frac{\partial \mathbf{s}}{\partial \mathbf{x}} \Delta \mathbf{x} + \Delta \rho \Delta \mathbf{x} + \dots$$

Wave equation





thus, neglecting the second-order term, one has:

$$\Delta \rho = -\rho_0 \, \frac{\partial \mathbf{s}}{\partial \mathbf{x}}$$

relation between the variation of displacement along x with the density variation. The minus sign is due to the fact that, if the variation is positive the volume increases and the density decreases.

If the displacement field is constant the gas is simply translated without perturbation.







Density variations cause pressure variations

The pressure in the medium is related to density with a relationship of the kind $P=f(\rho)$, that at the equilibrium is $P_0=f(\rho_0)$.

$$\mathsf{P} = \mathsf{P}_0 + \Delta \mathsf{P} = \mathsf{f}(\rho) = \mathsf{f}(\rho_0 + \Delta \rho) \approx \mathsf{f}(\rho_0) + \Delta \rho \mathsf{f}'(\rho_0) = \mathsf{P}_0 + \Delta \rho \kappa$$

and neglecting second-order terms:

$$\Delta \mathbf{P} = \mathbf{K} \Delta \boldsymbol{\rho}$$

$$\kappa = f'(\rho_0) = \left(\frac{d\mathbf{P}}{d\rho}\right)_0$$



Sound wave equation - 3



Pressure variations generate gas motion



The gas in the volume is accelerated by the different pressure exerted on the two sides...

$$P(x,t) - P(x + \Delta x,t) \approx -\frac{\partial P}{\partial x} \Delta x = -\frac{\partial (P_0 + \Delta P)}{\partial x} \Delta x = -\frac{\partial \Delta P}{\partial x} \Delta x$$
$$= \rho_0 \Delta x \frac{\partial^2 s}{\partial t^2} \quad \text{for Newton's 2nd law}$$

thus:

$$\rho_0 \frac{\partial^2 \mathbf{s}}{\partial \mathbf{t}^2} = -\frac{\partial \Delta \mathbf{P}}{\partial \mathbf{x}}$$





Using 1, 2 and 3 we have

$$\rho_{0} \frac{\partial^{2} \mathbf{s}}{\partial t^{2}} = -\frac{\partial \Delta P}{\partial \mathbf{x}} = -\frac{\partial \left(\kappa \Delta \rho\right)}{\partial \mathbf{x}} = -\frac{\partial \left[\kappa \left(-\rho_{0} \frac{\partial \mathbf{s}}{\partial \mathbf{x}}\right)\right]}{\partial \mathbf{x}}$$

thus:

1	$\partial^2 \mathbf{S}$ _	$\partial^2 \mathbf{S}$
ĸ	$\frac{1}{\partial t^2}$ –	$\partial \mathbf{x}^2$

i.e. the typical wave equation, describing a perturbation traveling with velocity $v = \sqrt{\kappa}$





From the sound wave equation

$$\prime = \sqrt{\kappa} = \sqrt{\left(\frac{\mathrm{dP}}{\mathrm{d\rho}}\right)_{\mathrm{O}}}$$

Newton computed the derivative of the pressure assuming that the heat is moving from one to another region in a such rapid way that the temperature cannot vary - isotherm PV=constant i.e. P/ρ =constant, thus

$$v = \sqrt{\left(\frac{dP}{d\rho}\right)_{0}} = \sqrt{\left(constant\right)_{0}} = \sqrt{\left(\frac{P}{\rho}\right)_{0}}$$

called isothermal sound velocity





Laplace correctly assumed that the heat flux between a compressed gas region to a rarefied one was negligible, and, thus, that the process of the wave passage was adiabatic $PV_{\gamma}=constant$, $P/\rho_{\gamma}=constant$, with γ ratio of the specific heats: C_p/C_v

$$\mathbf{v} = \sqrt{\left(\frac{dP}{d\rho}\right)_{0}} = \sqrt{\left(\frac{\gamma}{\rho} cons tan t\rho^{\gamma}\right)_{0}} = \sqrt{\gamma\left(\frac{P}{\rho}\right)_{0}}$$

called adiabatic sound velocity



Using the ideal gas law

PV=nRT=NkT

one can write the velocity on many ways:

$$\mathbf{v} = \sqrt{\frac{\gamma P}{\rho}} = \sqrt{\frac{\gamma P V}{\rho V}} = \sqrt{\frac{\gamma n R T}{m}} = \sqrt{\frac{\gamma N k T}{N m_{mol}}} = \sqrt{\frac{\gamma K T}{m_{mol}}} = \sqrt{\frac{\gamma R T}{weight_{mol}}}$$

showing that it depends on temperature only. If the "dry" air is considered (biatomic gas $\gamma=7/5$) one has:

 $v=331.4+0.6T_c m/s$ (temperature measured in Celsius)



BC and 1D systems











Sound velocity depends on the compressibility of the medium.

If the medium has a bulk modulus B and density at the equilibrium is ρ , the sound speed is: $\mathbf{v} = (\mathbf{B}/\rho)^{1/2}$

that can be compared with the velocity of transversal waves on a string:

 $v = (F/\mu)^{1/2}$

Thus, velocity depends on the elastic of the medium (B or F) and on inertial (ρ or μ) properties







Small perturbations of a	Linear	Harmonic
stable equilibrium point	restoring force	Oscillation
Coupling of harmonic oscillators	the disturbances can propagate , superpose and stand	

WAVE: organized propagating imbalance, satisfying differential equations of motion

$$\frac{\partial^2 \gamma}{\partial x^2} = \frac{1}{v^2} \frac{\partial^2 \gamma}{\partial t^2}$$

General form of LWE

(Seismic) wave propagation

Basic physical concepts

What is a wave? Discrete and continuous models Born of wave equation

Dispersion discreteness stiffness geometry boundaries



Dispersion relation



In classical mechanics, the Hamilton's principle the perturbation scheme applied to an averaged Lagrangian for an harmonic wave field gives a characteristic equation: Δ(ω,k_i)=0

Transverse wave in a string $\left(\frac{\partial^2}{\partial x^2} - \frac{\mu}{F}\frac{\partial^2}{\partial t^2}\right)\phi = 0 \Rightarrow \omega = \pm kc$

Acoustic wave

$$\left(\frac{\partial^2}{\partial x^2} - \frac{\rho}{B}\frac{\partial^2}{\partial t^2}\right)\phi = 0 \Rightarrow \omega = \pm kc$$



Longitudinal wave in a rod $\left(\frac{\partial^2}{\partial t^2} - \frac{\rho}{\partial t^2}\right) \phi = 0 \Rightarrow \omega = \pm kc$

$$\left(\frac{\partial}{\partial x^2} - \frac{\rho}{E}\frac{\partial}{\partial t^2}\right)\phi = 0 \Rightarrow \omega = \pm$$





- In physics, the dispersion relation is the relation between the energy of a system and its corresponding momentum. For example, for massive particles in free space, the dispersion relation can easily be calculated from the definition of kinetic energy: $E = \frac{1}{2}mv^2 = \frac{p^2}{2m}$
- For electromagnetic waves, the energy is proportional to the frequency of the wave and the momentum to the wavenumber. In this case, Maxwell's equations tell us that the dispersion relation for vacuum is linear: ω=ck.
- The name "dispersion relation" originally comes from optics. It is possible to make the effective speed of light dependent on wavelength by making light pass through a material which has a non-constant index of refraction, or by using light in a non-uniform medium such as a waveguide. In this case, the waveform will spread over time, such that a narrow pulse will become an extended pulse, i.e. be dispersed.







- In optics, dispersion is a phenomenon that causes the separation of a wave into spectral components with different wavelengths, due to a dependence of the wave's speed on its wavelength. It is most often described in light waves, but it may happen to any kind of wave that interacts with a medium or can be confined to a waveguide, such as sound waves. There are generally two sources of dispersion: material dispersion, which comes from a frequency-dependent response of a material to waves; and waveguide dispersion, which occurs when the speed of a wave in a waveguide depends on its frequency.
- In optics, the phase velocity of a wave v in a given uniform medium is given by: v=c/n, where c is the speed of light in a vacuum and n is the refractive index of the medium. In general, the refractive index is some function of the frequency of the light, thus n = n(f), or alternately, with respect to the wave's wavelength $n = n(\Lambda)$. For visible light, most transparent materials (e.g. glasses) have a refractive index n decreases with increasing wavelength Λ (dn/d Λ <0, i.e. $dv/d\Lambda$ >0). In this case, the medium is said to have normal dispersion and if the index increases with increasing wavelength the medium has anomalous dispersion.



Group velocity



Another consequence of dispersion manifests itself as a temporal effect. The phase velocity is the velocity at which the phase of any one frequency component of the wave will propagate. This is not the same as the group velocity of the wave, which is the rate that changes in amplitude (known as the envelope of the wave) will propagate. The group velocity v_g is related to the phase velocity by, for a homogeneous medium (here λ is the wavelength in vacuum, not in the medium):

$$\mathbf{v}_{g} = c \left(\mathbf{n} - \lambda \frac{d\mathbf{n}}{d\lambda} \right)^{-1} = \mathbf{v} - \lambda \frac{d\mathbf{v}}{d\lambda}$$

and thus in the normal dispersion case v_g is always < v !







The group velocity itself is usually a function of the wave's frequency. This results in group velocity dispersion (GVD), that is often quantified as the group delay dispersion parameter (again, this formula is for a uniform medium only): If D is less than zero, the medium is said to have positive dispersion. If D is greater than zero, the medium has negative dispersion.





Airy Phase -

wave that arises if the phase and the change in group velocity are stationary and gives the highest amplitude in terms of group velocity and are prominent on the seismogram.





M Discrete systems: lattices

Boundary waves: plates and rods

Discontinuity interfaces are intrinsic in their propagation since they allow to store energy (not like body waves)!

Stiff systems: rods and thin plates





Monatomic 1D lattice - continued

 $U_{n} = A e^{i(kx_{n}-\omega t)}$, Now let us attempt a solution of the form:

where x_n is the equilibrium position of the *n*-th atom so that $x_n = na$. This equation represents a traveling wave, in which all atoms oscillate with the same frequency ω and the same amplitude A and have a wavevector k. Now substituting the guess solution into the equation and canceling the common quantities (the amplitude and the time-dependent factor) we obtain

$$M(-\omega^2)\mathbf{e}^{ikna} = -\mathbf{C}[2\mathbf{e}^{ikna} - \mathbf{e}^{ik(n+1)a} - \mathbf{e}^{ik(n-1)a}].$$

This equation can be further simplified by canceling the common factor e^{ikna}, which leads to

$$M\omega^{2} = C\left(2 - e^{ika} - e^{-ika}\right) = 2C(1 - \cos ka) = 4C\sin^{2}\frac{ka}{2}$$

We find thus the dispersion relation for the frequency:

$$\varpi = \sqrt{\frac{4C}{M}} \left| \sin \frac{ka}{2} \right|$$

which is the relationship between the frequency of vibrations and the wavevector k. The dispersion relation has a number of important properties.









Long wavelength limit. The long wavelength limit implies that $\lambda >> a$. In this limit ka << 1. We can then expand the sine in ' ω ' and obtain for the positive frequencies: $\omega = \sqrt{\frac{C}{M}}ka$. We see that the frequency of vibration is proportional to the wavevector. This is equivalent to the statement that velocity is independent of frequency. In this case:

$$v_p = \frac{\omega}{k} = \sqrt{\frac{C}{M}}a$$
. This is the velocity of sound for the one dimensional lattice which is consistent with the expression we obtained earlier for elastic waves.





Monatomic 1D lattice – continued

Finite chain – Born – von Karman periodic boundary condition.

Unlike a continuum, there is only a finite number of distinguishable vibrational modes. But how many?

Let us impose on the chain ends the Born – von Karman periodic boundary conditions specified as following: we simply join the two remote ends by one more spring in a ring or device in the figure below forcing atom *N* to interact with ion 1 via a spring with a spring constant *C*. If the atoms occupy sites *a*, 2*a*, …, *Na* The boundary condition is $u_{N+1} = u_1$ or $u_N = u_0$.

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With the displacement solution of the form $u_n = A \exp[i(kna-wt)]$, the periodic boundary condition requires that $\exp(\pm ikNa) = 1$, which in turn requires 'k' to have the form:

$$k = \frac{2\pi}{a} \frac{n}{N}$$
 (n – an integer), and $-\frac{N}{2} \le n \le \frac{N}{2}$, or

$$k = \pm \frac{2\pi}{Na}, \ \pm \frac{4\pi}{Na}, \ \pm \frac{6\pi}{Na}, \ \dots, \pm \frac{\pi}{a}$$
 (N values of k).

$$n = 1$$











We can treat the motion of this lattice in a similar fashion as for the monatomic lattice. However, in this case, because we have two different kinds of atoms, we should write two equations of motion: u^{2}

$$M_{1} \frac{d^{2}u_{n}}{dt^{2}} = -C(2u_{n} - u_{n+1} - u_{n-1})$$
$$M_{2} \frac{d^{2}u_{n+1}}{dt^{2}} = -C(2u_{n+1} - u_{n+2} - u_{n})$$

In analogy with the monatomic lattice we are looking for the solution in the form of traveling mode for the two atoms:

$$\begin{bmatrix} u_n \\ u_{n+1} \end{bmatrix} = \begin{bmatrix} A_1 e^{ikna} \\ A_2 e^{ik(n+1)a} \end{bmatrix} e^{-i\omega t} \quad \text{in matrix form}$$

Substituting this solution into the equations of the previous slide we obtain:

$$\begin{bmatrix} 2C - M_1 \omega^2 & -2C \cos ka \\ -2C \cos ka & 2C - M_2 \omega^2 \end{bmatrix} \begin{bmatrix} A_1 \\ A_2 \end{bmatrix} = 0.$$

This is a system of linear homogeneous equations for the unknowns A_1 and A_2 . A nontrivial solution exists only if the determinant of the matrix is zero. This leads to the secular equation

$$(2\mathbf{C}-\mathbf{M}_{1}\omega^{2})(2\mathbf{C}-\mathbf{M}_{2}\omega^{2})-4\mathbf{C}\cos^{2}\mathbf{k}\mathbf{a}=0.$$



Dispersion



This is a quadratic equation, which can be readily solved:

$$\omega^{2} = C\left(\frac{1}{M_{1}} + \frac{1}{M_{2}}\right) \pm C_{1} \sqrt{\left(\frac{1}{M_{1}} + \frac{1}{M_{2}}\right)^{2} - \frac{4\sin^{2}ka}{M_{1}M_{2}}}$$

Depending on sign in this formula there are two different solutions corresponding to two different dispersion curves, as is shown in the figure:

The lower curve is called the *acoustic* branch, while the upper curve is called the *optical* branch.

The acoustic branch begins at k = 0 and $\omega = 0$, and as $k \Rightarrow 0$:

With increasing k the frequency increases in a linear fashion. This is why this branch is called *acoustic*: it corresponds to elastic waves, or sound. Eventually, this curve saturates at the edge of the Brillouin zone. On the other hand, the optical branch Has a nonzero frequency at zero k,



$$\omega_{o} = \sqrt{2C\left(\frac{1}{M_{1}} + \frac{1}{M_{2}}\right)}$$

and it does not change much with k.





Another feature of the dispersion curves is the existence of a forbidden gap between $\omega_a = (2C/M_1)^{1/2}$ and $\omega_o = (2C/M_2)^{1/2}$ at the zone boundaries ($k = \pm \pi/2a$). The forbidden region corresponds to frequencies in which lattice waves cannot propagate through the linear chain without attenuation. It is interesting to note that a similar situation also exists in the energy band scheme of a solid to be discussed later.

The distinction between the acoustic and optical branches of lattice vibrations can be seen most clearly by comparing them at k = 0 (infinite wavelength). As follows from the equations of motion, for the acoustic branch $\omega = 0$ and $A_1 = A_2$. So, in this limit the two atoms in the cell have the same amplitude and phase. Therefore, the molecule oscillates as a rigid body, as shown in the left figure for the acoustic mode.



On the other hand, for the optical vibrations, by substituting ω_0 we obtain for k = 0:

$$M_1A_1 + M_2A_2 = 0$$
 $(M_1/M_2 = -A_2/A_1).$

This implies that the optical oscillation takes place in such a way that the center of mass of a molecule remains fixed. The two atoms move in out of phase as shown. The frequency of these vibrations lies in the infrared region $(10^{12} \text{ to } 10^{14} \text{ Hz})$ which is the reason for referring to this branch as *optical*. If the two atoms carry opposite charges, we may excite a standing wave motion with the electric field of a light wave.

Dispersion







Dispersion

35




Oiscrete systems: lattices



Discontinuity interfaces are intrinsic in their propagation since they allow to store energy (not like body waves)!







$$k_{x} = k \sin(i) = \omega \frac{\sin(i)}{\alpha} = \frac{\omega}{c}$$

$$k_{z} = k \cos(i) = \sqrt{k^{2} - k_{x}^{2}} = \omega \sqrt{\left(\frac{1}{\alpha}\right)^{2} - \left(\frac{1}{c}\right)^{2}} = \frac{\omega}{c} \sqrt{\left(\frac{c}{\alpha}\right)^{2} - 1} = k_{x}r_{\alpha}$$

In current terminology, k_x is k!

SH Waves in plates: Geometry

In an elastic half-space no SH type surface waves exist. Why? Because there is total reflection and no interaction between an evanescent P wave and a phase shifted SV wave as in the case of Rayleigh waves. What happens if we have a layer delimited by two free boundaries, i.e. a homogeneous plate?



Repeated reflection in the layer allow interference between incident and reflected SH waves: SH reverberations can be totally trapped.





$$\mathbf{u}_{y} = \mathbf{A} \exp[i(\omega \mathbf{t} + \omega \eta_{\beta} \mathbf{z} - \mathbf{k} \mathbf{x})] + \mathbf{B} \exp[i(\omega \mathbf{t} - \omega \eta_{\beta} \mathbf{z} - \mathbf{k} \mathbf{x})]$$



$$\mathbf{u}_{y} = \mathbf{A} \exp[i(\omega t + kr_{\beta}z - kx)] + \mathbf{B} \exp[i(\omega t - kr_{\beta}z - kx)]$$

The formal derivation is very similar to the derivation of the Rayleigh waves. The conditions to be fulfilled are: free surface conditions

$$\sigma_{zy}(0) = \mu \frac{\partial u_{y}}{\partial z} \bigg|_{0} = i k r_{\beta} \mu \Big\{ A \exp[i(\omega t - kx)] - B \exp[i(\omega t - kx)] \Big\} = 0$$

$$\sigma_{zy}(2h) = \mu \frac{\partial u_{y}}{\partial z} \bigg|_{2h} = i k r_{\beta} \mu \Big\{ A \exp[i(\omega t + k r_{\beta} 2h - kx)] - B \exp[i(\omega t - k r_{\beta} 2h - kx)] \Big\} = 0$$











http://www.ee.iastate.edu/~hsiu/descriptions/paral.html

Dispersion

Seismic Wave physics







Oiscrete systems: lattices

Boundary waves: plates and rods

Discontinuity interfaces are intrinsic in their propagation since they allow to store energy (not like body waves)!

Stiff systems: rods and thin plates







We we with the second s









Stiffness in a vibrating string introduces a restoring force proportional to the bending angle of the string and the usual stiffness term added to the wave equation for the ideal string. Stiff-string models are commonly used in piano synthesis and they have to be included in tuning of piano strings due to inharmonic effects.



6

Waves in plates



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=

In low frequency plate waves, there are two distinct type of harmonic motion. These are called symmetric or **extensional** waves and antisymmetric or **flexural** waves.







satisfying the boundary conditions $\tau_{yy} = \tau_{xy} = 0$ on $y = \pm h$ gives the Rayleigh-Lamb equations:

$$\frac{\tanh(\beta h)}{\tanh(\alpha h)} = \left[\frac{4\omega^2 \alpha \beta}{c^2 \left(\omega^2 / c^2 + \beta^2\right)^2}\right]^{\pm 1}$$

- + ... extensional waves
- ... flexural waves

$$\alpha = \left| \frac{\omega}{c} \right| \sqrt{1 - \frac{c^2}{c_p^2}} \quad , \beta = \left| \frac{\omega}{c} \right| \sqrt{1 - \frac{c^2}{c_s^2}}$$





consider the extensional waves

$$\frac{\tanh\left[2\pi fh\sqrt{1/c^2-1/c_s^2}\right]}{\tanh\left[2\pi fh\sqrt{1/c^2-1/c_p^2}\right]} = \frac{4\sqrt{1-c^2/c_s^2}\sqrt{1-c^2/c_p^2}}{\left(2-c^2/c_s^2\right)^2}$$

If we let
$$kh = \frac{2\pi fh}{c} >> 1$$
 (high frequency)

then both tanh functions are $\cong 1$

and we find
$$(2-c^2/c_s^2)^2 = 4\sqrt{1-c^2/c_s^2}\sqrt{1-c^2/c_p^2}$$

so we just have Rayleigh waves on both stress-free surfaces:







In contrast for $kh \ll 1$ (low frequency) tanh $(\alpha h) \cong \alpha h$ we find tanh $(\beta h) \cong \beta h$

and the Rayleigh-Lamb equation reduces to

$$\left(2 - c^2 / c_s^2\right)^2 = 4\left(1 - c^2 / c_p^2\right)$$

which can be solved for c to give

$$c = c_{plate} = \sqrt{\frac{E}{\rho(1 - v^2)}}$$



Waves in plates



In low frequency plate waves, there are two distinct type of harmonic motion. These are called symmetric or **extensional** waves and antisymmetric or **flexural** waves.



Dispersion



Lamb waves



Lamb waves are waves of plane strain that occur in a free plate, and the traction force must vanish on the upper and lower surface of the plate. In a free plate, a line source along y axis and all wave vectors must lie in the x-z plane. This requirement implies that response of the plate will be independent of the in-plane coordinate normal to the propagation direction.







Three types of elastic waves can propagate in rods: (1) **longitudinal waves**, (2) **flexural waves**, and (3) **torsional waves**. Longitudinal waves are similar to the symmetric Lamb waves, flexural waves are similar to antisymmetric Lamb waves, and torsional waves are similar to horizontal shear (SH) waves in plates.



Dispersion

Seismic Wave physics

Torsional modes dispersion













(Seismic) wave propagation

Basic physical concepts

What is a wave? Discrete and continuous models Born of wave equation Dispersion

Basic physical concepts 2

PDE: Poisson, diffusion and wave equation Navier-Stokes equation Scattering and diffusion Mathematic reference: Linear PDE

Classification of Partial Differential Equations (PDE)

Second-order PDEs of two variables are of the form:

$$a\frac{\partial^{2} f(x, y)}{\partial x^{2}} + b\frac{\partial^{2} f(x, y)}{\partial x \partial y} + c\frac{\partial^{2} f(x, y)}{\partial y^{2}} + d\frac{\partial f(x, y)}{\partial x} + e\frac{\partial f(x, y)}{\partial y} = F(x, y)$$

$$b^{2} - 4ac < 0 \qquad \text{elliptic} \qquad \text{LAPLACE equation}$$

$$b^{2} - 4ac = 0 \qquad \text{parabolic} \qquad \text{DIFFUSION equation}$$

$$b^{2} - 4ac > 0 \qquad \text{hyperbolic} \qquad \text{WAVE equation}$$

Elliptic equations produce stationary and energy-minimizing solutions

Parabolic equations a **smooth-spreading flow** of an initial disturbance

Hyperbolic equations a propagating disturbance

<u>Initial conditions</u>: starting point for propagation problems

Boundary conditions: specified on domain boundaries to provide the interior solution in computational domain



(i) Dirichlet condition : u = f on ∂R (ii) Neumann condition : $\frac{\partial u}{\partial n} = f$ or $\frac{\partial u}{\partial s} = g$ on ∂R

(iii) Robin (mixed) condition :
$$\frac{\partial u}{\partial n} + ku = f$$
 on ∂R

Boundary and Initial conditions

Seismic Wave physics





Hyperbolic Equation $b^2 - 4ac = 0 - 4(1)(-c^2) > 0$: Hyperbolic $\frac{\partial^2 u}{\partial t^2} = v^2 \frac{\partial^2 u}{\partial x^2}, \quad 0 \le x \le a, \quad 0 \le t$

I.C.s
$$\begin{cases} u(x,0) = f_1(x) \\ u_t(x,0) = f_2(x) \end{cases} & 0 \le x \le a \\ u_t(x,0) = f_2(x) \end{cases}$$

B.C.s
$$\begin{cases} u(0,t) = g_1(t) \\ u(a,t) = g_2(t) \end{cases} & t > 0 \end{cases}$$

Seismic Wave physics





Steady-state two-dimensional heat conduction equation is prototypical elliptic PDE

Laplace equation - no heat generation

$$\frac{\partial^2 T}{\partial x^2} + \frac{\partial^2 T}{\partial y^2} = \theta$$

Poisson equation - with heat source

$$\frac{\partial^2 T}{\partial x^2} + \frac{\partial^2 T}{\partial y^2} = f(x, y)$$



Heat transfer in a one-dimensional rod

I.C.s
$$u(x,0) = f(x)$$
 $0 \le x \le a$
B.C.s
$$\begin{cases} u(0,t) = g_1(t) \\ u(a,t) = g_2(t) \end{cases}$$
 $0 \le t \le T$

Seismic Wave physics







Navier-Stokes Equations



$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$$

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + v \left(\frac{\partial^2 u}{\partial x^2} + \frac{\partial^2 u}{\partial y^2} \right)$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} = -\frac{1}{\rho} \frac{\partial p}{\partial y} + v \left(\frac{\partial^2 v}{\partial x^2} + \frac{\partial^2 v}{\partial y^2} \right)$$

(Seismic) wave propagation

Basic physical concepts

What is a wave? Discrete and continuous models Born of wave equation Dispersion

Basic physical concepts 2

PDE: Poisson, diffusion and wave equation Navier-Stokes equation Scattering and diffusion







The space occupied by the material will be called the **domain**.

Solids are materials that have a more or less intrinsic configuration or shape and do not conform to their domain under nominal conditions.

Fluids do not have an intrinsic shape; gases are fluids that will completely fill their domain (or container) and liquids are fluids that form a free surface in the presence of gravity.

Fluid mechanics assumptions



- Classical fluid mechanics, like classical thermodynamics, is concerned with **macroscopic phenomena** (bulk properties) rather than microscopic (molecular-scale) phenomena.
- The molecular makeup of a fluid will be ignored in all that follows, and the crucially important physical properties of a fluid, e.g., its mass density, ρ, and specific heat, Cp, among others, must be provided from outside of this theory. It is assumed that these physical properties, along with flow properties, e.g., the pressure, P, velocity, v, temperature, T, etc., are in principle definable at every point in space, as if the fluid was a smoothly varying continuum, rather than a swarm of very fine, discrete particles (molecules).



Continuity equation



Conservation of matter

The total mass of fluid flowing, in unit of time, through a surface S, has to be equal to the decrease, in unit time, in the mass of fluid in the volume V:

$$\oint \rho \mathbf{v} dS = -\frac{\partial}{\partial t} \int \rho dV$$

$$\frac{\partial \rho}{\partial t} + \operatorname{div}(\rho \mathbf{v}) = \frac{\partial \rho}{\partial t} + \rho \operatorname{div}(\mathbf{v}) + \mathbf{v} \cdot \operatorname{grad}(\rho) = 0$$

that can be compared with what we obtained considering 1D sound waves:

$$\Delta \rho = -\rho_0 \, \frac{\partial \mathbf{S}}{\partial \mathbf{X}} \qquad \text{The gas moves and causes density variations}$$









The fluid in the volume is accelerated by the total force acting on it:

$$\rho \frac{d\mathbf{v}}{dt} = -grad(\mathbf{P}) - \rho grad(\phi) + \mathbf{f}_{visc}$$

Fluid moves from high-pressure areas to low-pressure areas. Moving implies that fluid moves in direction of largest change in pressure

External forces that act at a distance; we can suppose that they are conservative (like gravity and electricity)

Internal force due to the fact that in a flowing fluid there can also be a shearing stress, and it is called the **viscous** force







Newton's law

The fluid in the volume is accelerated by the total force acting on it:

$$\rho \frac{d\mathbf{v}}{dt} = -grad(\mathbf{P}) - \rho grad(\phi) + \mathbf{f}_{visc}$$

that can be compared with what we obtained considering for 1D sound waves:

$$\rho_0 \frac{\partial^2 s}{\partial t^2} = -\frac{\partial \Delta P}{\partial x}$$
 Pressure variations generate gas motion





In mathematics and continuum mechanics, including fluid dynamics, the substantive derivative (sometimes the Lagrangian derivative, material derivative or advective derivative), written D/Dt, is the rate of change of some property of a small parcel of fluid.

Note that if the fluid is moving, the substantive derivative is the rate of change of fluid within the small parcel, hence the other names advective derivative and fluid following derivative. Advection is transport of a some conserved scalar quantity in a vector field.

$$\mathbf{v}_{x} \frac{\partial \mathbf{f}}{\partial \mathbf{x}} + \mathbf{v}_{y} \frac{\partial \mathbf{f}}{\partial \mathbf{y}} + \mathbf{v}_{z} \frac{\partial \mathbf{f}}{\partial \mathbf{z}} = \mathbf{v} \cdot \operatorname{grad}(\mathbf{f})$$







The functional relationship between **density**, **pressure and temperature**:

 $P=P(\rho,T)$ or equivalently, $\rho=\rho(P,T)$

with T the absolute temperature in Kelvin.

The archetype of an equation of state is that of an ideal gas, $P=\rho RT/M$

where R=8.31 (Joule moles⁻¹ K⁻¹) is the universal gas constant and M is the molecular weight (kg/mole).

If the composition of the material changes, then the appropriate equation of state will involve more than three variables, for example the concentration of salt if sea water, or water vapor if air.

An important class of phenomenon may be described by a reduced equation of state having state variables density and pressure alone,

$$P=P(\rho)$$
 or equivalently, $\rho = \rho(P)$

and the fluid is said to be **barotropic**.

Navier-Stokes



The temperature of the fluid will change as pressure work is done on or by the fluid, and yet temperature need not appear as a separate, independent state variable provided conditions approximate one of two limiting cases:

1) If the fluid is a fixed mass of ideal gas, say, that can readily exchange heat with a heat reservoir having a constant temperature, then the gas may remain **isothermal** under pressure changes;

2) the other limit, which is more likely to be relevant, is that heat exchange with the surroundings is negligible because the time scale for significant conduction is very long compared to the time scale (lifetime or period) of the phenomenon. In that event the system is said to be **adiabatic** and in the case of an ideal gas the density and pressure are related by the well-known adiabatic law.

$$\Delta \rho = \frac{\partial \rho}{\partial \mathsf{P}} \Delta \mathsf{P}$$

that can be compared with what we obtained considering sound waves:

$$\Delta P = \kappa \Delta \rho = c^2 \Delta \rho$$
 Density variations cause pressure variations

Navier-Stokes


Incompressible fluids



In many cases of the flow of fluids their density may be supposed invariable, i.e. constant throughout the volume and its $\rho = constant$ motion and we speak of incompressible flow

Conservation of matterdiv (v)=0Euler equation $\frac{\partial v}{\partial t} + \Omega \times v = -grad \left(\frac{1}{2}v^2 + \frac{P}{\rho} + \phi \right)$

The conditions under which the fluid can be considered incompressible are:



i.e. the time taken by a sound signal to traverse distances must be small compared with that during which flow changes appreciably





Having considered gravity waves whose length is small compared with the depth of the liquid, let us now discuss the opposite limiting case of waves whose length is large compared with the depth. These are called **long waves**.

Let us examine the propagation of long waves in a channel that is supposed to be along the x-axis, and of infinite length. The cross-section of the channel may have any shape, and may vary along its length. We denote the cross-sectional area of the liquid in the channel by S = S(x,t). The depth and width of the channel are supposed small in comparison with the wavelength.

We shall here consider longitudinal waves, in which the liquid moves along the channel. In such waves the velocity component v_x along the channel is large compared with the components v_y , v_z . We denote v_x by v simply, and omit small terms.







Viscosity is a measure of the resistance of a fluid to deform under shear stress. It is commonly perceived as "thickness", or resistance to pouring. Viscosity describes a fluid's internal resistance to flow and may be thought of as a measure of fluid friction. Thus, water is "thin", having a lower viscosity, while vegetable oil is "thick" having a higher viscosity. All real fluids (except superfluids) have some resistance to shear stress, but a fluid which has no resistance to shear stress is known as an **ideal fluid**



Strain as a measure of



To understand deformation due to shear, picture two flat plates with a fixed spacing, h, between them:



Fluids are qualitatively different from solids in their response to a shear stress. Ordinary fluids such as air and water have no intrinsic configuration, and hence fluids **do not develop a restoring force** that can provide a static balance to a shear stress. When the shear stress is held steady, and assuming that the geometry does not interfere, the **shear deformation rate**, may also be steady or have a meaningful timeaverage.









Viscosity of Newtonian fluids depends only on temperature and pressure, e.g.:

$$\eta(\mathbf{T},\mathbf{P}) = \eta_{0} e^{\left[\frac{\Delta E}{R} \left(\frac{T_{0}-T}{T_{0}T}\right)\right]} e^{\beta\left(\mathbf{P}-\mathbf{P}_{0}\right)}$$

Where: η_o :viscosity at T_o and P_o (reference temperature and pressure)

- ΔE : activation energy for flow
- R: gas constant
- β : material property [m²/N]





The structure of some polymers, especially filled polymers or concentrated suspensions can be sufficiently rigid that it permits the material to withstand a certain level of deforming stress without flowing. The maximum stress that can be sustained without flow is called the "**yield stress**" and this type of behavior is called "**plasticity**"









Viscosity is a measure of the resistance of a fluid to deform under shear stress. It is commonly perceived as "thickness", or resistance to pouring. Viscosity describes a fluid's internal resistance to flow and may be thought of as a measure of fluid friction. Thus, water is "thin", having a lower viscosity, while vegetable oil is "thick" having a higher viscosity. All real fluids (except superfluids) have some resistance to shear stress, but a fluid which has no resistance to shear stress is known as an **ideal fluid**



In the general case there is another term that depends on other derivatives of the velocity and the general expression is:

$$\eta \left(\frac{\partial \mathbf{v}_{\mathbf{y}}}{\partial \mathbf{x}} + \frac{\partial \mathbf{v}_{\mathbf{x}}}{\partial \mathbf{y}} \right) + \eta' \delta_{ij} \left(\mathsf{div}(\mathbf{v}) \right)$$

so two constants are required: the "first coefficient of viscosity" or "shear viscosity coefficient" and "second coefficient of viscosity".

The component of the viscous force per unit volume in the direction of the rectangular coordinate x_j is:

$$(\mathbf{f}_{visc})_{i} = \sum_{j=1,3} \frac{\partial \tau_{ij}}{\partial \mathbf{x}_{j}} = \sum_{j=1,3} \frac{\partial}{\partial \mathbf{x}_{j}} \left[\eta \left(\frac{\partial \mathbf{v}_{i}}{\partial \mathbf{x}_{j}} + \frac{\partial \mathbf{v}_{j}}{\partial \mathbf{x}_{i}} \right) \right] + \frac{\partial}{\partial \mathbf{x}_{i}} \left(\eta' \operatorname{div}(\mathbf{v}) \right)$$
$$\mathbf{f}_{visc} = \eta \Delta \mathbf{v} + (\eta + \eta') \operatorname{grad}(\operatorname{div}(\mathbf{v}))$$

Navier-Stokes





Newton's law + Conservation of matter + Viscosity

$$\rho \frac{\partial \mathbf{v}}{\partial t} + \rho (\mathbf{v} \cdot \mathbf{grad}) \mathbf{v} = -\mathbf{grad}(\mathbf{P}) - \rho \mathbf{grad}(\mathbf{\phi}) + \eta \Delta \mathbf{v} + (\eta + \eta') \mathbf{grad}(\operatorname{div}(\mathbf{v}))$$

and in the incompressible case...

$$\frac{\partial \mathbf{\Omega}}{\partial t} + \operatorname{rot}(\mathbf{\Omega} \times \mathbf{v}) = \frac{\eta}{\rho} \Delta \mathbf{\Omega}$$







The study and determination of the response of continuous, perfectly elastic solids subjected to applications of forces



(Seismic) wave propagation

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Basic concepts of EM wavefield (5)

Extinction and **emission** are two main types of the interactions between an electromagnetic radiation field and a medium (e.g., the atmosphere).

Extinction is due to absorption and scattering.

- Absorption is a process that removes the radiant energy from an electromagnetic field and transfers it to other forms of energy.
 Scattering is a process that does not remove energy from the radiation field, but redirect it. Scattering can be thought of as absorption of radiant energy followed by re-emission back to the electromagnetic field with negligible conversion of energy, i.e.can be a "source" of radiant energy for the light beams traveling in other directions.
 - Scattering occurs at all wavelengths (spectrally not selective) in the electromagnetic spectrum, for any material whose refractive index is different from that of the surrounding medium (optically inhomogeneous).



The amount of scattered energy depends strongly on the ratio of: particle size (a) to wavelength (λ) of the incident wave

When (a < $\lambda/10$), the scattered intensity on both forward and backward directions are equal. This type of scattering is called Rayleigh scattering.

For (a > λ), the angular distribution of scattered intensity becomes more complex with more energy scattered in the forward direction. This type of scattering is called Mie scattering





Scattering of EM wavefield (2)







Single Scattering





FIGURE 3.18. Scattering regimes. [Adapted from Wallace and Hobbs (1977). Reprinted by permission of Academic Press.]



For (a >> λ), the Scattering characteristics are determined from explicit Reflection, Refraction and Diffraction: Geometric "Ray" Optics

Seismic Wave physics



Composition of the scatterer (n) is important!

The interaction (and its redirection) of electromagnetic radiation with matter May or may not occur with **transfer of energy**, i.e., the scattered radiation has a slightly different or the same wavelength.







Scattering and Absorption



When the photon is absorbed and re-emitted at a different wavelength, this is absorption.



Transmissivity of the Earth's atmosphere





In single scattering, the properties of the scatterer are important, but multiple scattering erases these effects - eventually **all** wavelengths are scattered in **all** directions.



Works for turbid media: clouds, beer foam, milk, etc...

Example: when a solid has a very low temperature, phonons behave like waves (long mean free paths) and heat propagate following ballistic term. At higher temperatures, the phonons are in a diffusive regime and heat propagate following Maxwell law.

(Seismic) wave propagation

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The governing parameters for the seismic scattering are:

wavelength of the wavefield (or wavenumber k) λ (10⁰-10⁶ m) correlation length, or dimension, of the heterogeneity a (10[?]-10⁵ m) distance travelled in the heterogeneity L (10⁰-10⁶ m)

With special cases:

- a = L homogeneous region
- a $\gg \lambda$ ray theory is valid
- $a \approx \lambda$ strong scattering effects



Seismic Scattering (1)





Wave propagation problems can be classified using the parameters just introduced.

This classification is crucial for the choice of technique to calculate synthetic seismograms

(Adapted from Aki and Richards, 1980)





Let us consider a **perturbed** model: reference+perturbation (in elastic parameters)

$$\rho = \rho_0 + \epsilon \delta \rho \quad \lambda = \lambda_0 + \epsilon \delta \lambda \quad \mu = \mu_0 + \epsilon \delta \mu$$

resulting in a velocity perturbation

 $c = c_0 + \varepsilon \delta c$

solution: **Primary** field + **Scattered** field $\mathbf{u} = \mathbf{u}_0 + \mathbf{u}_1(\delta\rho, \delta\lambda, \delta\mu)$

satisfying equations of motion: $\rho_{0}\ddot{\mathbf{u}}_{i}^{0} - (\lambda_{0} + \mu_{0}) (\nabla \cdot \mathbf{u}^{0})_{,i} - \mu_{0} \nabla^{2} \mathbf{u}_{i}^{0} = 0$ $\rho_{0}\ddot{\mathbf{u}}_{i} - (\lambda \nabla \cdot \mathbf{u})_{,i} - [\mu(\mathbf{u}_{i,j} + \mathbf{u}_{j,i})]_{,j} = 0$ $\rho_{0}\ddot{\mathbf{u}}_{i}^{1} - (\lambda_{0} + \mu_{0}) (\nabla \cdot \mathbf{u}^{1})_{,i} - \mu_{0} \nabla^{2} \mathbf{u}_{i}^{1} = Q_{i}$





How does a point-like perturbation of the elastic parameters affect the wavefield?

Perturbation of the different elastic parameters produce characteristic radiation patterns. These effects are used in diffraction tomography to recover the perturbations from the recorded wavefield.

(Figure from Aki and Richards, 1980)





Correlation distance



When velocity varies in all directions with a finite scale length, it is more convenient to consider spatial fluctuations

Autocorrelation function (a is the **correlation distance**):

$$N(\mathbf{r}_{1}) = \frac{\left\langle \frac{\delta c(\mathbf{r})}{c_{0}(\mathbf{r})} \frac{\delta c(\mathbf{r} + \mathbf{r}_{1})}{c_{0}(\mathbf{r} + \mathbf{r}_{1})} \right\rangle}{\left\langle \left(\frac{\delta c(\mathbf{r})}{c_{0}(\mathbf{r})} \right)^{2} \right\rangle} = \begin{cases} e^{-|\mathbf{r}_{1}|/a} \\ e^{-(|\mathbf{r}_{1}|/a)^{2}} \end{cases}$$

Power Spectra of scattered waves

$$\left\langle \left| \mathbf{u}_{1} \right|^{2} \right\rangle \propto \begin{cases} k^{4} \left(1 + 4k^{2}a^{2}\sin^{2}\frac{\theta}{2} \right)^{2} \\ k^{4} \exp \left(-k^{2}a^{2}\sin^{2}\frac{\theta}{2} \right) \end{cases}$$

∝ k⁴ if ka <<1 (Rayleigh scattering) if ka is large (forward scattering)

Seismic wavefield



$$\frac{\Delta I}{I} \propto \begin{cases} k^4 a^3 L \left(1 + 4k^2 a^2\right)^{-1} \\ k^2 a L \left(1 - e^{k^2 a^2}\right)^{-1} \end{cases}$$

but violates the energy conservation law and it is valid if (<0.1)

the **perturbations** (P &A) are function of the **wave parameter**:

$$D = \frac{4L}{ka^2}$$

 $D = \begin{cases} 0 & \text{phase perturbation} \\ \infty & \text{phase = amplitude} \end{cases}$

when D<1, geometric ray theory is valid

Seismic wavefield



Seismic Scattering (2)





Wave propagation problems can be classified using the parameters just introduced.

This classification is crucial for the choice of technique to calculate synthetic seismograms

(Adapted from Aki and Richards, 1980)







Multiple scattering process leads to **attenuation** (spatial loss non a true dissipative one) and **energy mean free path**



 $\sigma(\theta)$ is the **differential scattering cross-section** and after a wave has travelled a distance x, the energy is reduced by an amount of

$$e^{-\Sigma x} \quad \Sigma = \int_{-1}^{+1} \sigma(\cos\theta) d\cos\theta$$

and the average path length between scattering events is

$$l = \int_0^\infty e^{-\Sigma x} dx = \frac{1}{\Sigma}$$

Seismic wavefield







forward scattering tendency

$$\Sigma' = \int_{-1}^{+1} (\cos \theta) \sigma(\cos \theta) d\cos \theta \begin{cases} > 0 \text{ forward} \\ \approx 0 \text{ isotropic} \\ < 0 \text{ backward} \end{cases}$$

Multiple scattering randomizes the phases of the waves adding a diffuse (incoherent) component to the average wavefield. Statistical approaches can be used to derive **elastic radiative transfer equations**

> **Diffusion constants** use the definition of a diffusion (transport) mean free path

$$d = \frac{cl^{*}}{3} \quad l^{*} = \frac{l}{\Sigma - \Sigma'} \text{ (acoustic)}$$
$$d = \frac{l}{1 + 2K^{3}} \left(\frac{c_{p}l_{p}^{*}}{3} + 2K^{2}\frac{c_{s}l_{s}^{*}}{3} \right) \text{ (elastic)}$$

for non-preferential scattering l* coincides with energy mean free path, l for enhanced forward scattering l*>l

Experiments for ultrasound in materials can be applied to seismological problems...

Seismic wavefield







How is a propagating wavefield affected by random heterogeneities?

Synthetic seismograms

Correlation length: 10km

Synthetic seismograms for a global model with random velocity perturbations.



When the wavelength is long compared to the correlation length, scattering effects are difficult to distinguish from intrinsic attenuation.



Correlation length: 20km





Seismic Scattering Classification



Wave propagation problems can be classified using the parameters just introduced.

This classification is crucial for the choice of technique to calculate synthetic seismograms

(Adapted from Aki and Richards, 1980)











Tsunami physics









Fabio Romanelli

Seismic Wave physics 106

Dispersion & Non linearity



The dynamics of water waves in shallow water is described mathematically by the Korteveg – de Vries (KdV) equation

u=u(x,t) measures the elevation at time t and position x, i.e. the height of the water above the equilibrium level

Dispersive term

$$u_{t} + u_{xxx} = 0$$

Nonlinearity

$$u_{t} + u u_{x} = 0$$

KdV

$$\mathbf{u}_{\mathsf{t}} + \mathbf{u}_{\mathsf{x}\mathsf{x}\mathsf{x}} + \mathbf{u} \ \mathbf{u}_{\mathsf{x}} = \mathbf{0}$$












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