



**The Abdus Salam
International Centre for Theoretical Physics**



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Winter College on Optics in Environmental Science

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Adaptive Optics: Systems and Applications

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Adaptive Optics: Systems & Applications

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William Herschel Telescope with GLAS Rayleigh
Laser Guide Star

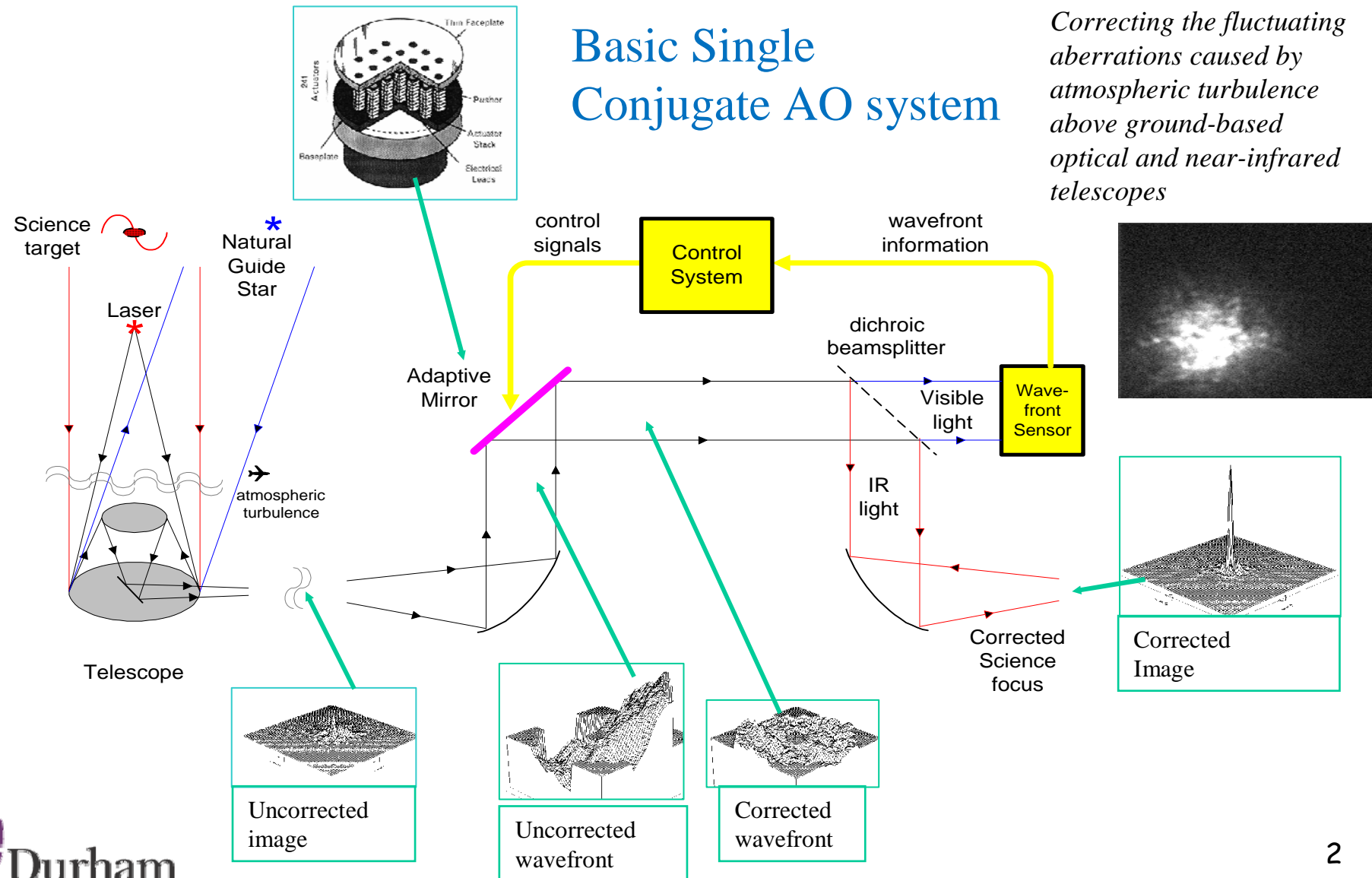
Photo: Tibor Agocs, Isaac Newton Group of Telescopes

Wavefront Correctors

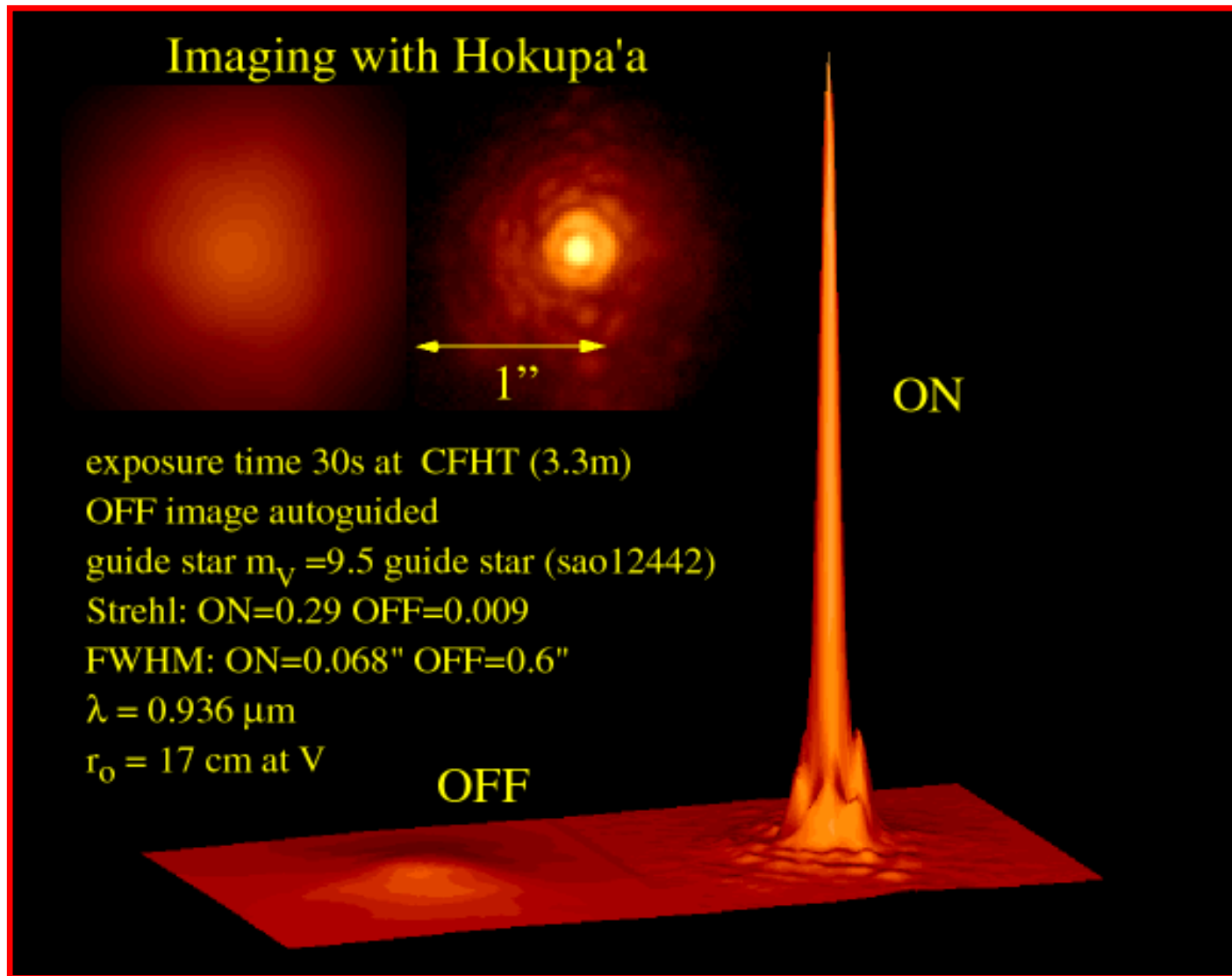
Generic Astronomical Adaptive Optics

Basic Single Conjugate AO system

Correcting the fluctuating aberrations caused by atmospheric turbulence above ground-based optical and near-infrared telescopes

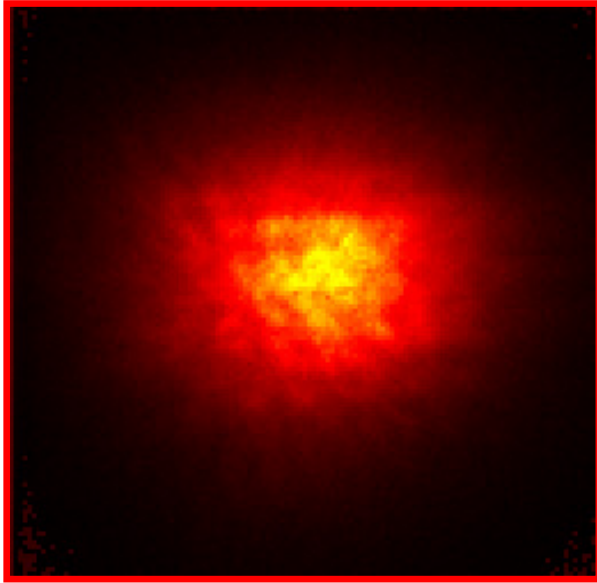


University of Hawaii AO System



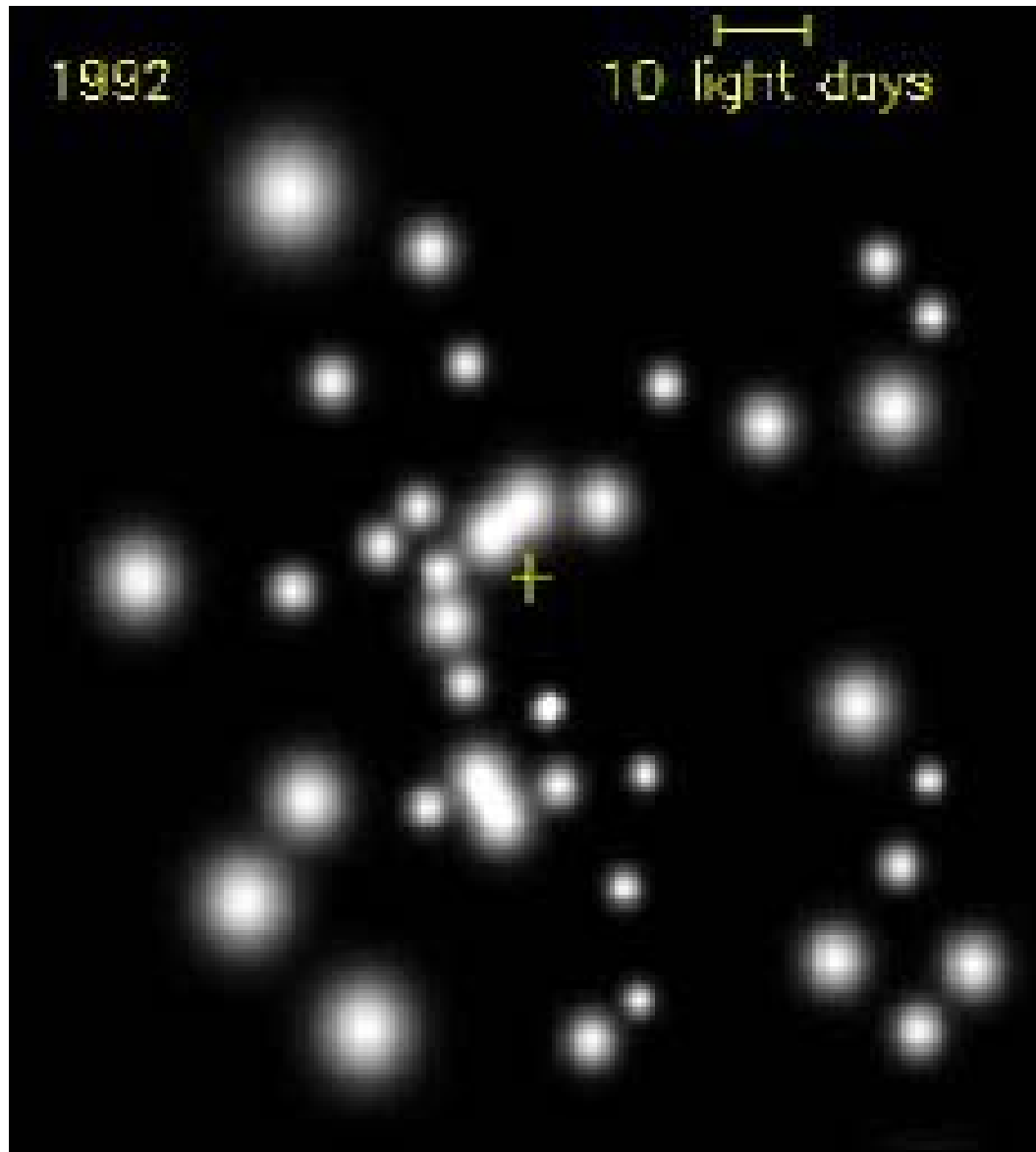
USAF - Starfire Optical Range

Albuquerque, NM, USA



Asteroid - Vesta





Star orbiting
a black hole at
the centre of
the Milky Way
(courtesy of ESO)

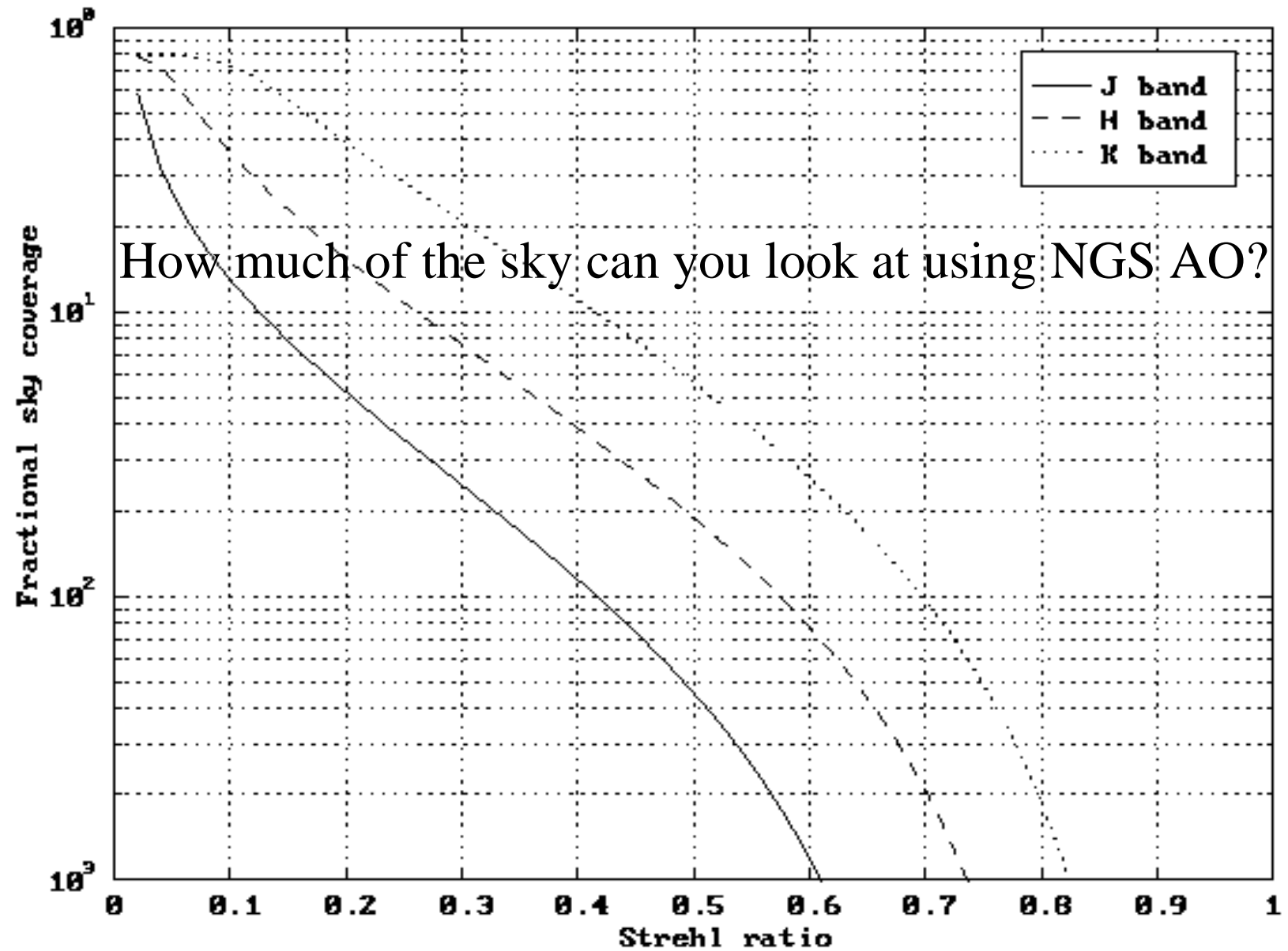
Astronomical AO

- State of the Art

- All major astronomical observatories have adaptive optics systems.
- The technology is maturing
- The number of science papers exploiting AO is increasing

... so what's the problem?

- We can only look at a tiny fraction of the sky
 - Requires Laser Guide Stars
- Field of view is very narrow. Requires...
 - MCAO (multiconjugate AO)
 - GLAO (ground layer AO)
 - MOAO (multi-object AO)
- Systems all work in the infrared
 - Visible AO requires XAO (Extreme Adaptive Optics)

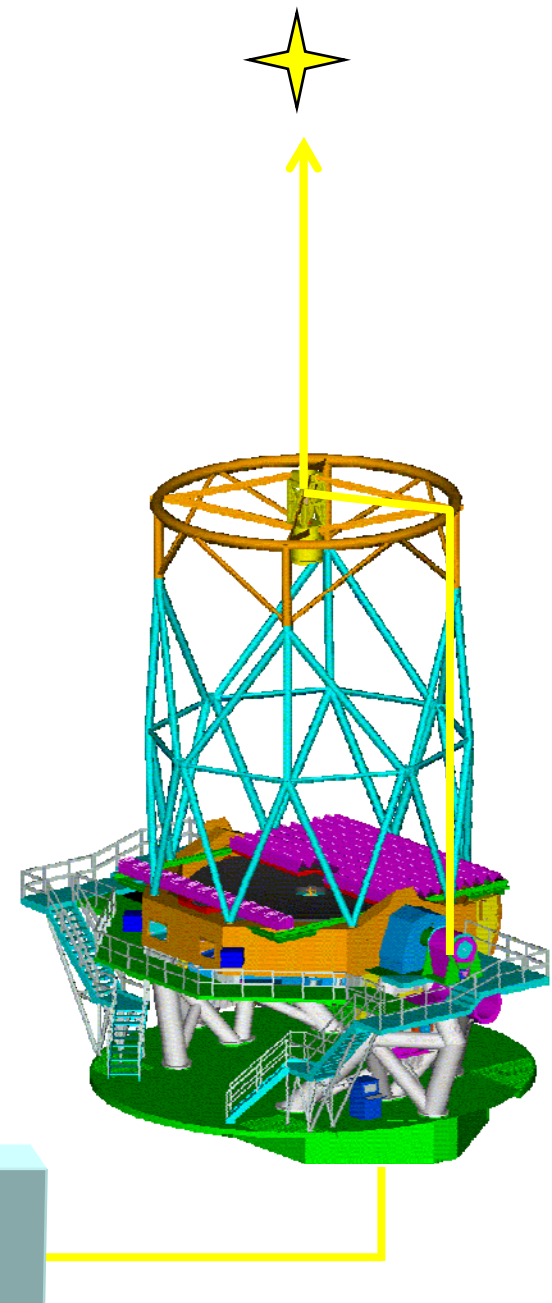


(Courtesy: Brent Ellerbroek and David Tyler, Phillips Laboratory)

A: Laser Beacons or Laser Guide Stars

The solution to the lack of suitable guide stars is to create an artificial one by shining a high-power laser into the atmosphere

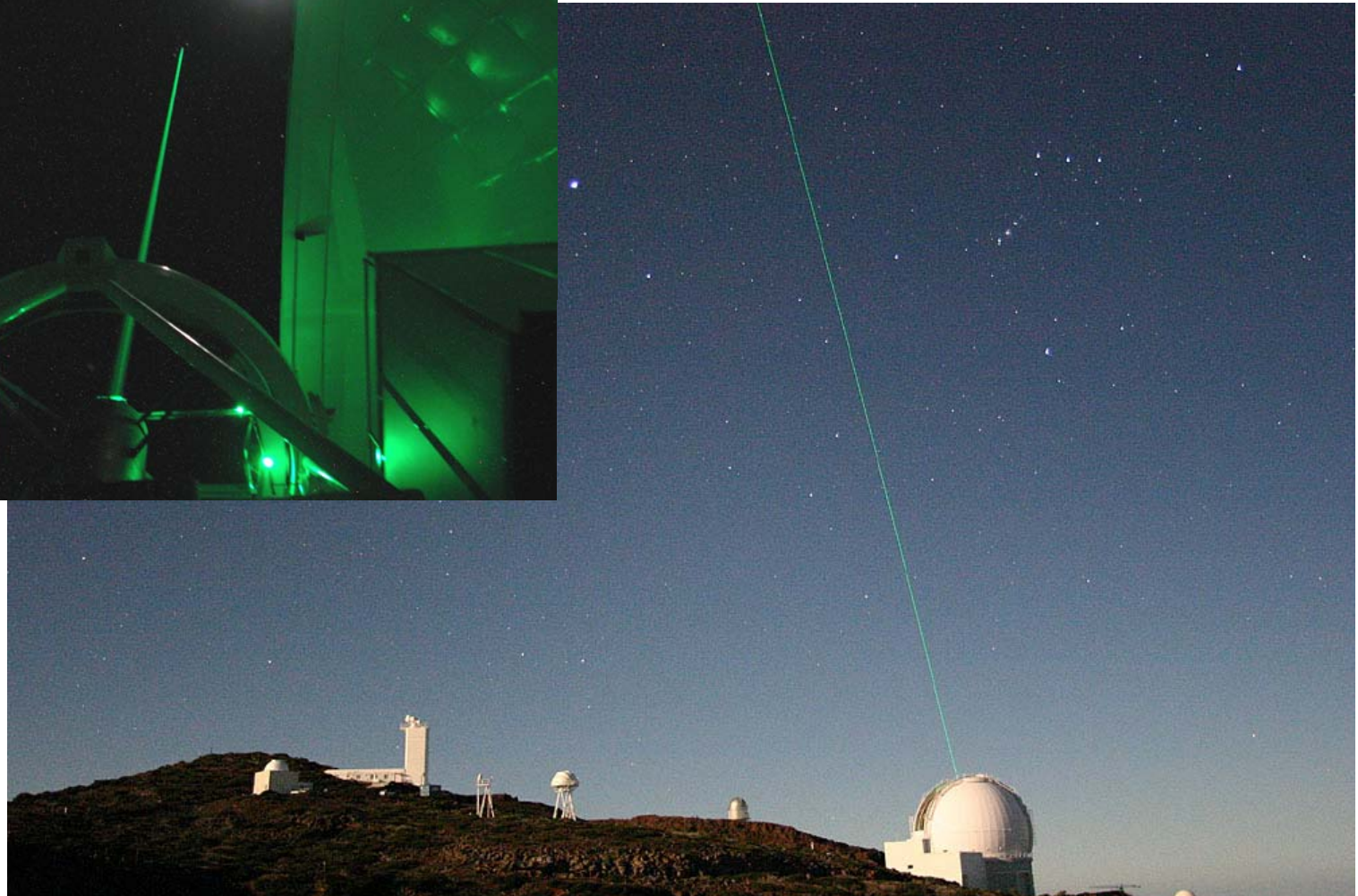
(Courtesy of Gemini)



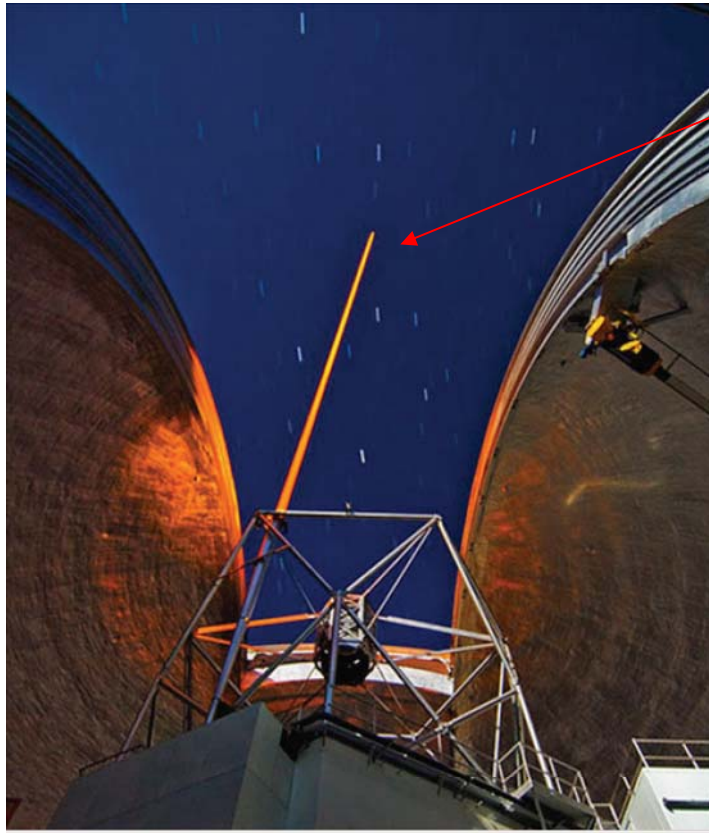
Sodium Layer (~90km).
This picture shows the aurora and sodium layer
photographed from space



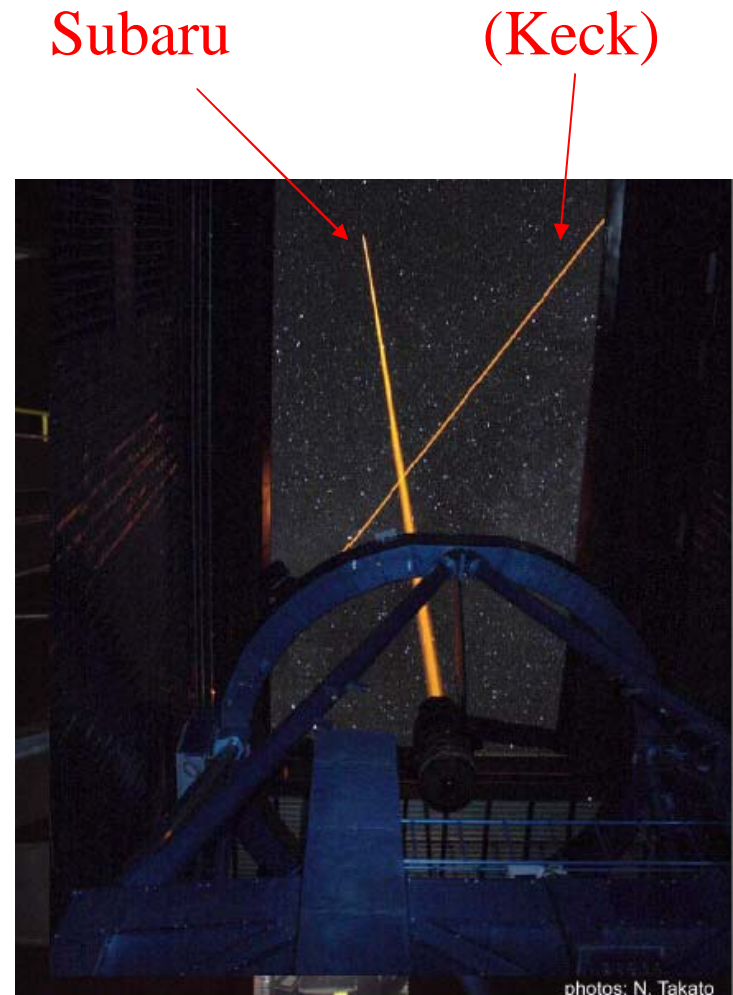
Durham's Rayleigh Laser Guide Star



Other LGS Systems



Keck

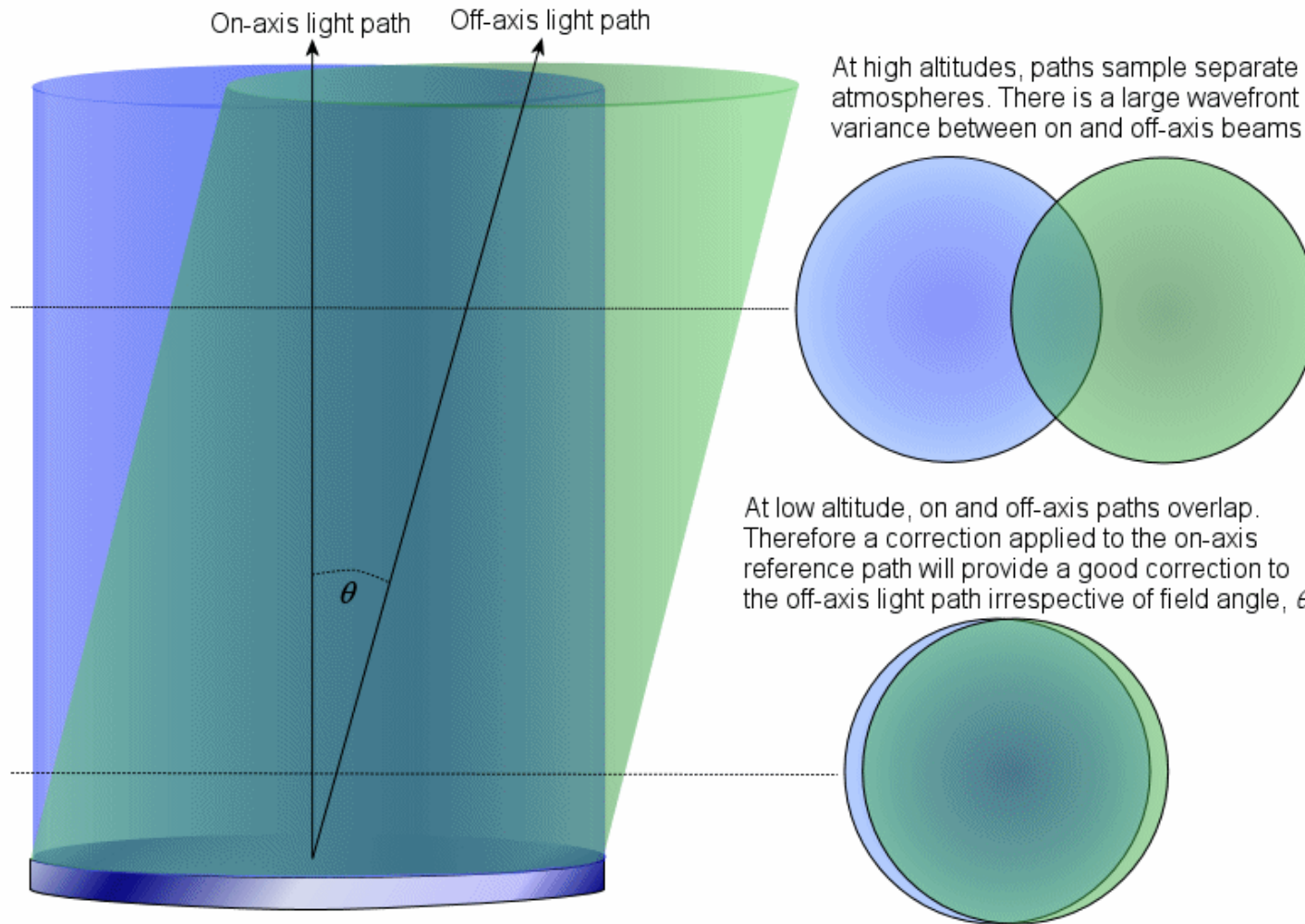


B. Wide-Field Adaptive Optics

There are 3 solutions under development

- MCAO - Multi-conjugate AO
- GLAO - Ground Layer
- MOAO - Multi-object AO

Angular Anisoplanatism



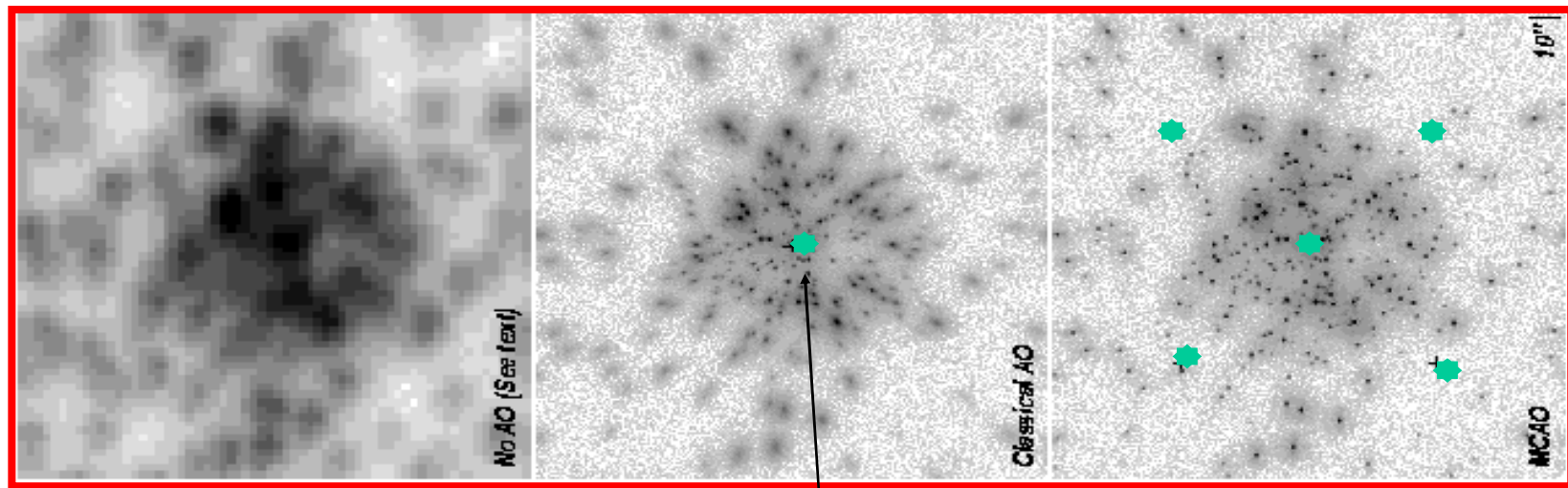
Multi-Conjugate AO

Simulation showing how correction degrades off axis, and how MCAO can correct this.

No AO

Classical AO

MCAO



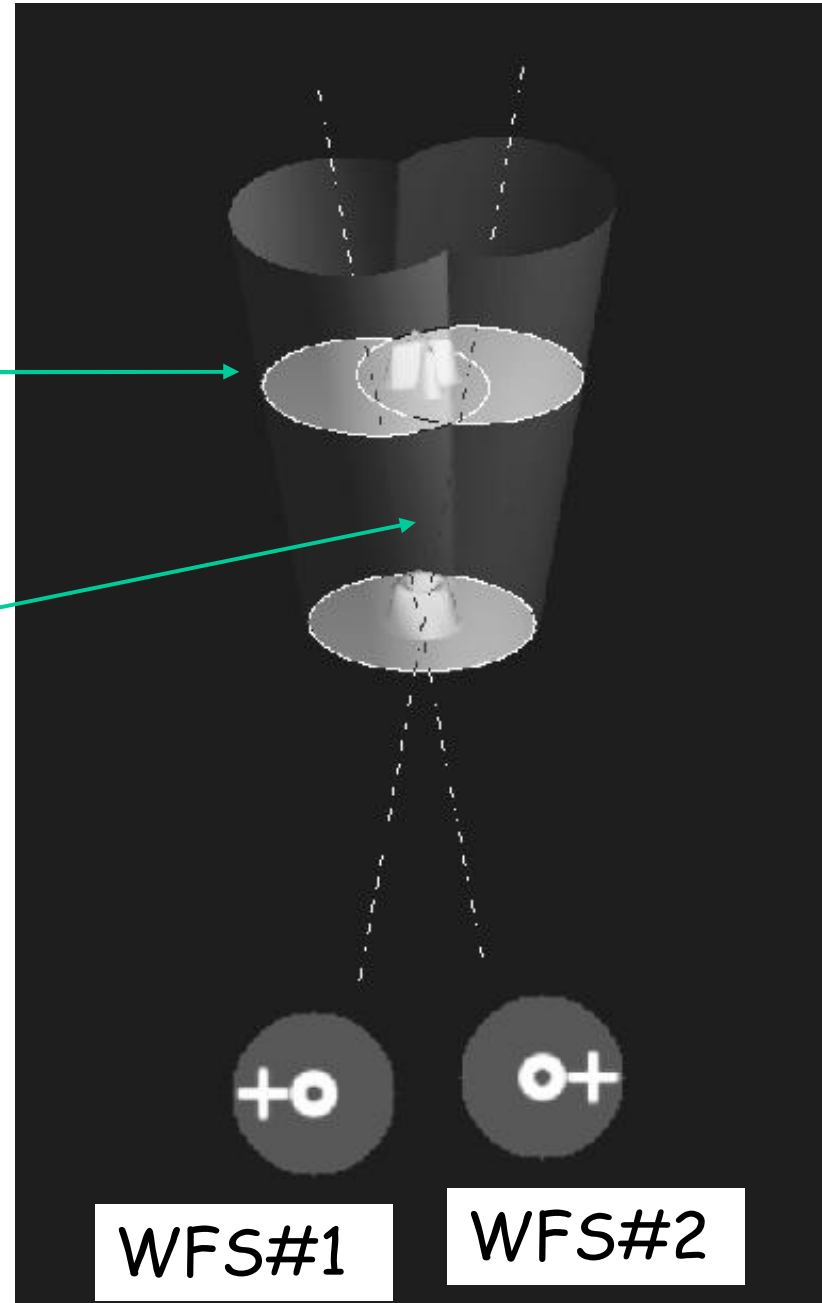
Star sizes are magnified,
and f.o.v $\sim 2'$

LGS

Tomography

Altitude Layer
(phase aberration = $+$)

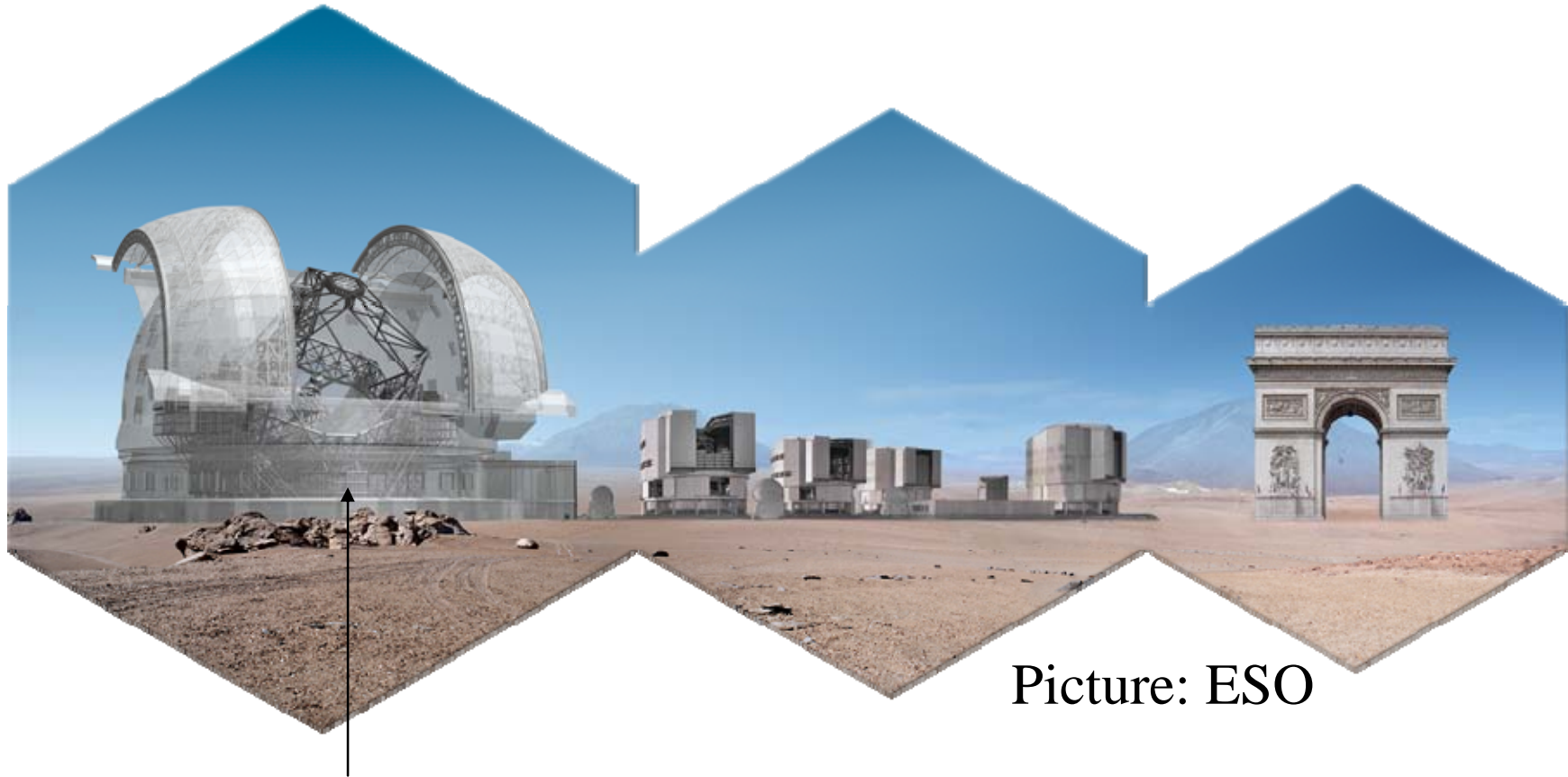
Ground Layer = Pupil
(phase aberration = 0)



Courtesy of F. Rigaut, Gemini

C. Extreme AO (XAO)

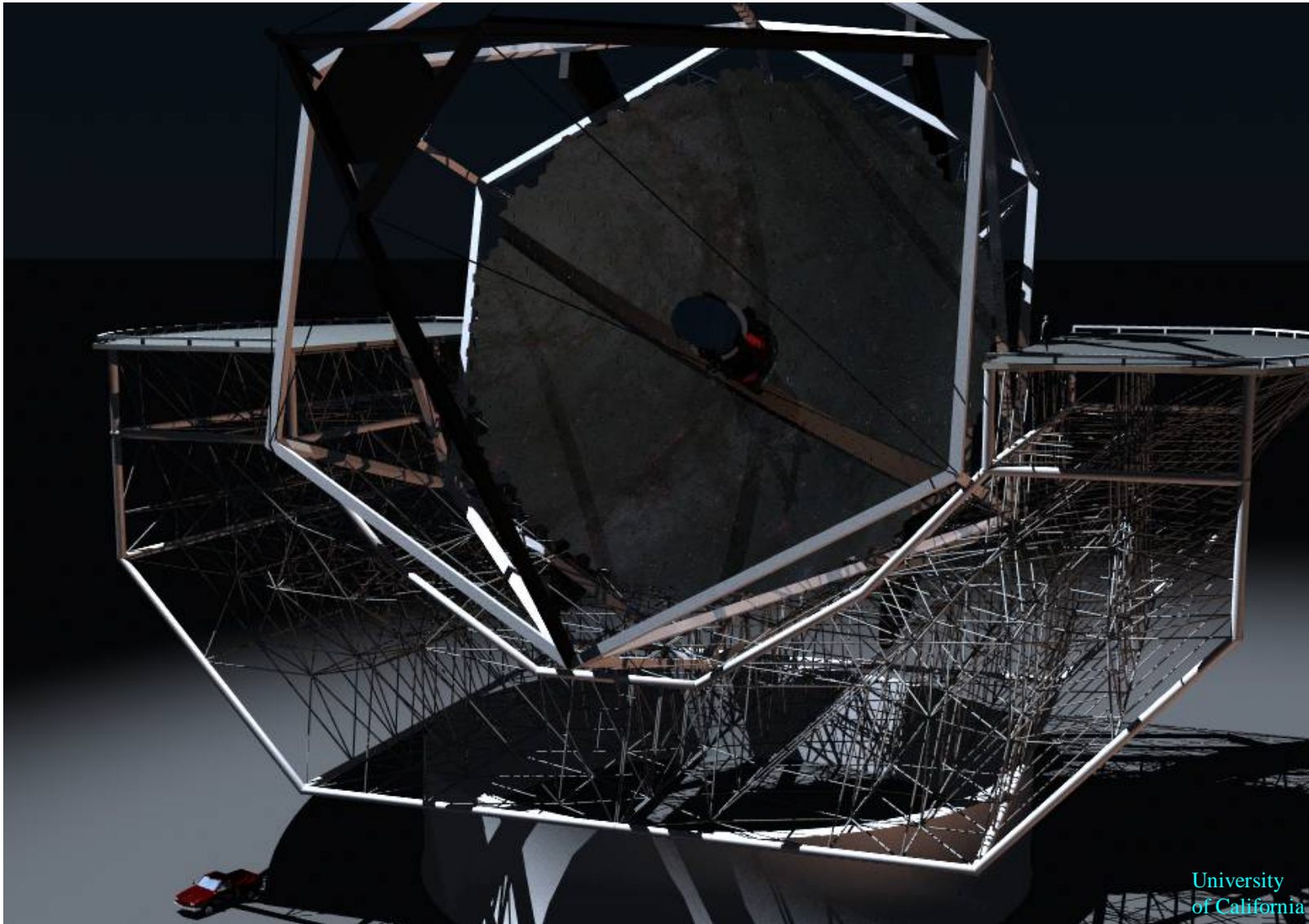
- In theory - this is the same as conventional AO - but with "more of everything"
 - More actuators in the DM
 - More wavefront sensing elements
- Current generation of AO systems have $\sim 10^2$ actuators.
- Visible AO requires $10^3 - 10^4$
- E.g. A 42m ELT (extremely large telescope) requires $\sim (D/r_0)^2 \sim (42/0.15)^2 \sim 10^5$ actuators.



Picture: ESO

The European Extremely Large Telescope (E-ELT) is being designed with a 42-metre primary mirror -and will be the first telescope with AO built-in: telescope mirror 4 will be a deformable mirror with over 5000 actuators.

TMT - 30m (USA)



University
of California

Non-Astronomical (and Non-Military) Applications of Adaptive Optics

Other Applications

OPHTHALMIC OPTICS

- Retinal Imaging
- Optometry

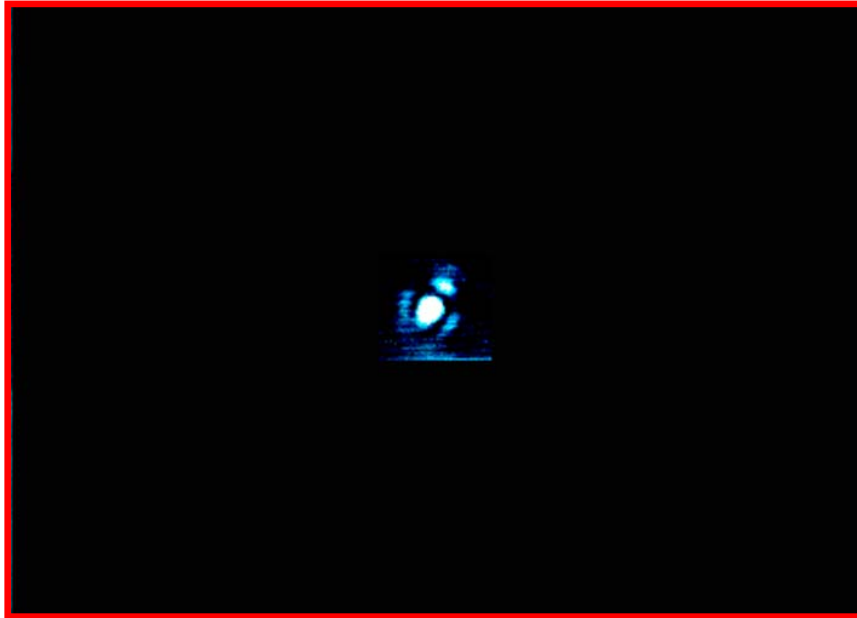
LASERS

- High power laser spot control
- Optical storage
- Optical Communications

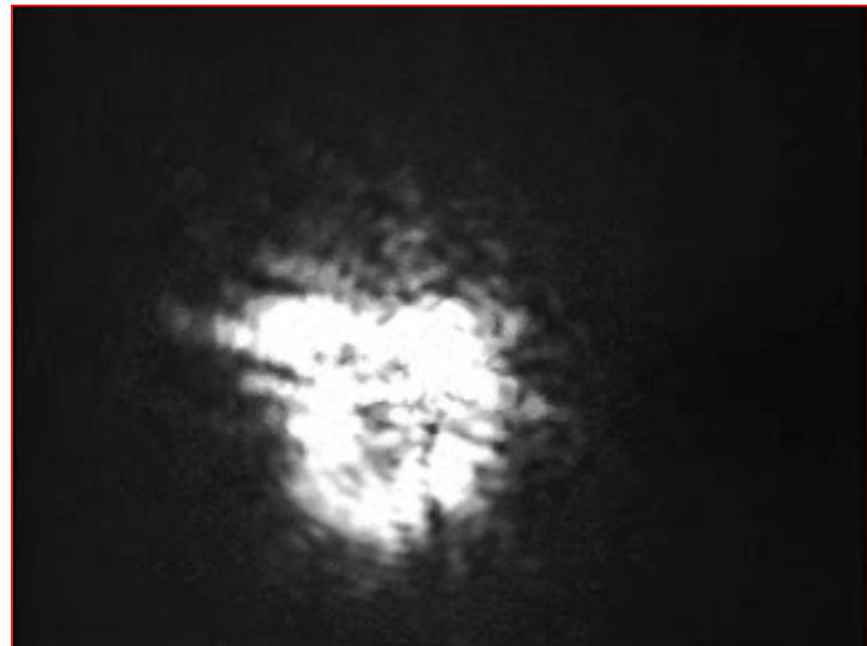
IMAGING

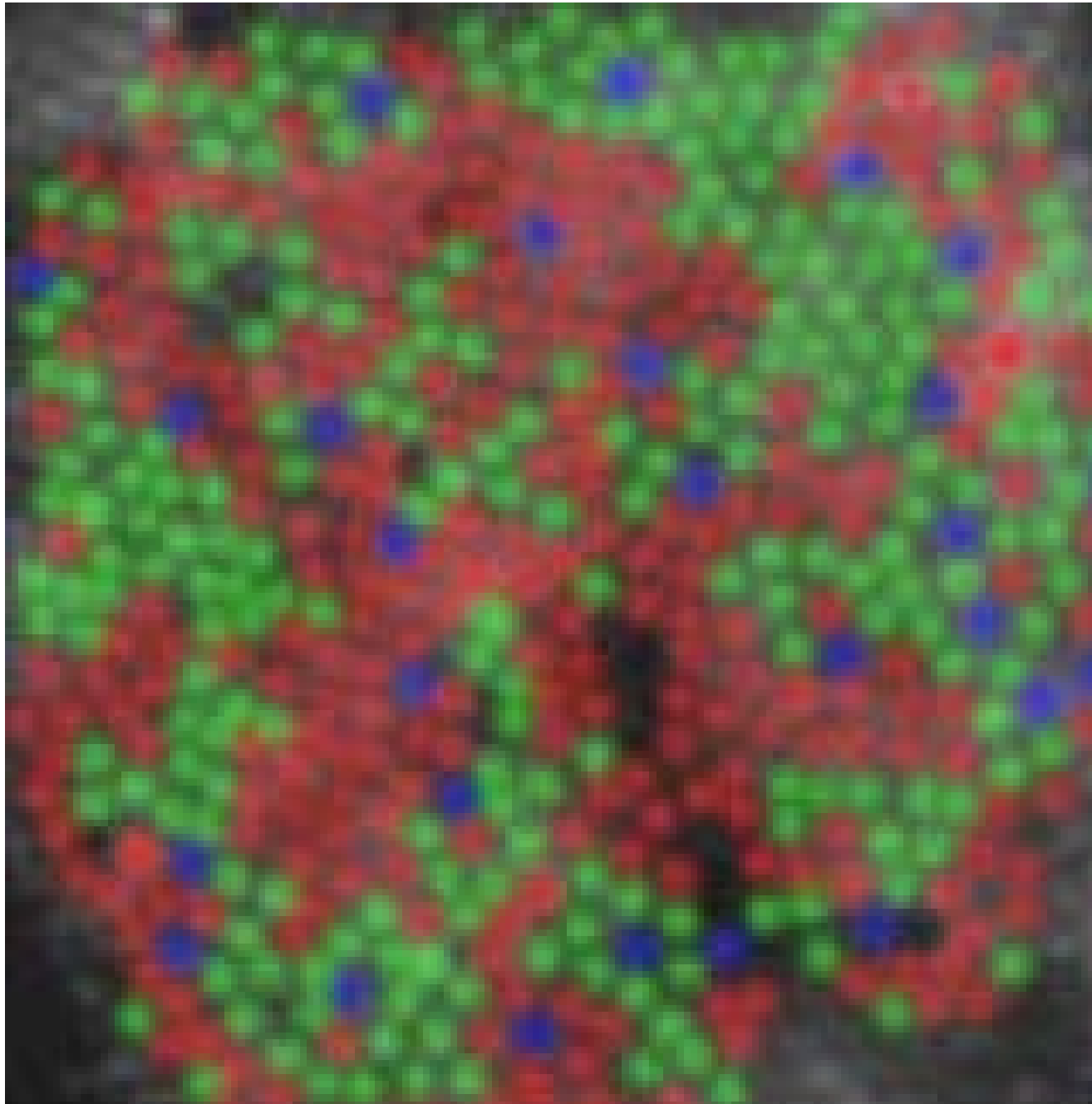
- Consumer optics

E.g. Optical Communications



E.g. – 100m propagation path length at Durham







The quick brown fox jumps over the lazy dog

Fig. 1

The quick brown fox jumps over the lazy dogs

Fig. 2

The quick brown fox jumps over the lazy dogs

Fig. 3

Wavefront Control

Adaptive
Optics

Adaptive Spectacles



Conclusions

- Astronomical AO is maturing and most observatories now have common-user systems
- We are working towards...
 - Full sky AO
 - Visible AO
 - Wide field AO
 - AO for Extremely Large Telescopes
- AO is being exploited in other areas of optics (e.g. ophthalmology) but exploitation also involves the exploitation of individual parts of a system - rather than the whole