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Workshop Towards the Neutrino Technologies

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Sub-keV Neutrino and Dark Matter Physics with Ultra-Low-Energy Germanium Detector

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Sub-keV Neutrino and Dark Matter Physics with Ultra-Low-Energy Germanium Detector

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NUTECH-09



- TEXONO Overview
 - Neutrino Magnetic Moment Search
 - $\sin^2 \theta_W$, g_A/g_V Measurement on $\bar{\nu_e}e^-$ Scattering
- vN coherent scattering
 - Basic : cross-section, spectrum.
 - Non Standard Interaction & Unparticle.
- Detector R & D.
- 4×5 g Detector
 - Data, threshold, Efficiencies.
 - Quenching factor.
 - Dark Matter Search.
- 500 g PCGe detector.
 - Data, threshold, Efficiencies.
- Plans
 - Channeling effect, Quenching factor Measurement.
 - Jin-Ping Underground Lab.
- Summary



TEXONO collaboration & Physics Programs

- Taiwan(AS, INER, KSNPS)
- China(IHEP, CIAE, THU, NKU, SCU, LNU)
- Turkey(METU), India(BHU)



- [1] Magnetic Moment search at ~ 10 keV range \rightarrow [PRD <u>75</u> 2007].
- [2] $sin^2\theta_W$ measurement at ~ MeV range \rightarrow Results will be published.
- [3] WIMP Search at sub keV range \rightarrow [PRD <u>79</u> 2009].
- [3] $\bar{\nu_e} N$ Coherent Scattering \rightarrow Goal.

Kuo-Sheng Reactor Plant





- $\phi_{\nu} = 6.4 \times 10^{12} \text{cm}^{-2} \text{s}^{-1}$
- · Lab : 28 m from core,
- 30 mwe concrete.



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Kuo-Sheng Neutrino Lab





Neutrino Magnetic Moment



- $(\frac{d\sigma}{dT})_{\text{SM}} \propto \text{const.}$
- $(\frac{d\sigma}{dT})_{MM} = \frac{\pi \alpha^2 \mu_{\nu}^2}{m_e^2} (\frac{1}{T} \frac{1}{E_{\nu}})$
- * $(d\sigma/dT)_{MM} >> (d\sigma/dT)_{SM}$ at $T << E_{\nu}$
- 1 kg HPGe detector at reactor 571/128 days ON/OFF data.
- Background level ~ 1 cpd
- $\mu_{\nu} < 7.4 \times 10^{-11} \mu_B$ 90% C. L.

ref: [PRD <u>75</u> 2007]





$\bar{\nu_e}e^-$ Scattering at MeV



- $\sigma_{measured}/\sigma_{\bar{\nu_e}e^-}$ = 1.18 \pm 0.29(stat) \pm 0.08(sys)
- $\sin^2 \theta_W$ = 0.264 \pm 0.04(stat) \pm 0.01(sys)



vN Coherent Scattering

• Probe $\bar{\nu_e}N$ Coherent Scattering below ~ 500 eV.

$$\begin{split} \nu + \mathsf{N} &\to \nu + \mathsf{N} \\ \bullet \ (\frac{\mathrm{d}\sigma}{\mathrm{d}t})_{\mathsf{S}\mathsf{M}} = \frac{\mathsf{G}_{\mathsf{F}}^2 \mathfrak{m}_{\mathsf{N}}}{4\pi} [Z(1 - 4 \mathsf{sin}^2 \theta_W) - \mathsf{N}]^2 [1 - \frac{\mathsf{M}_{\mathsf{N}} \mathsf{T}_{\mathsf{N}}}{2\mathsf{E}_{\nu}^2}] \\ &\to \mathsf{N}^2 \text{ enhancement} \end{split}$$

- Low recoil energy : ~ 1.9 keV for $E_{\nu} = 8$ MeV, Ge $\rightarrow \sim 500$ eV after Quenching Factor.
- A fundamental neutrino interaction never been experimentally observed.
- An important interaction/energy loss channel in astrophysics media.
- new detection channel for neutrinos → reactor monitoring.
 → compact and portable(?).
- WIMP search in sub-keV range and improve magnetic moment limit on same detector.



vN Coherent Scattering



- Take Q. F. = 0.2, extrapolate 1 cpkkd to eV level
 ⇒ signal/noise > 1 at 300 eV
- At threshold = 100 eV \Rightarrow 3.1 count day⁻¹ kg⁻¹ At threshold = 200 eV \Rightarrow 0.23 count day⁻¹ kg⁻¹ At threshold = 300 eV \Rightarrow 0.0084 count day⁻¹ kg⁻¹



Non Standard Interaction & Unparticle

- Non Standard Interaction:
 - Parametrize by ε^{fP}_{αβ},
 - f = e, u or d.
 α, β = ν-flavors.
 P = L, R.





$$\varepsilon^{\mathsf{fP}}_{\alpha\beta} 2\sqrt{2} \mathsf{G}_{\mathsf{F}}(\bar{\nu}_{\alpha}\gamma_{\mu}L\nu_{\beta})(\bar{\mathsf{f}}\nu^{\mu}\mathsf{P}\mathsf{f})$$

 $\varepsilon_{\mu e}^{\,qV} < 7.7 \times 10^{-4}$ from limit of $\mu^- \text{Ti} \to e^- \text{Ti}$ conversion

[Present Limits	a Eactory	
		Viacioly	Je I _{th} =400eV
			$(^{76}\text{Ge}_{T_{th}=100eV})$
ϵ_{ee}^{dV}	$-0.5 < \epsilon_{ee}^{dV} < 1.2$	$ \epsilon_{ee}^{dV} < 0.002$	$ \epsilon_{ee}^{dV} < 0.003$
			$(\epsilon_{ee}^{dV} < 0.001)$
$\epsilon_{\tau e}^{dV}$	$ \epsilon_{\tau e}^{\mathrm{dV}} < 0.78$	$ \epsilon_{\tau e}^{\mathrm{dV}} < 0.06$	$ \epsilon_{\tau e}^{\mathrm{dV}} < 0.032$
			$(\epsilon_{ au e}^{ dV} < 0.020)$
€ ^{uV} _{ee}	$-1.0 < \varepsilon_{ee}^{uV} < 0.61$	$ \epsilon_{ee}^{uV} < 0.002$	$ \epsilon_{ee}^{uV} < 0.003$
			$(\epsilon_{ee}^{uV} < 0.001)$
ε ^{uV} _{τe}	$ \epsilon_{\tau e}^{\mathrm{uV}} < 0.78$	$ \epsilon_{\tau e}^{\mathrm{uV}} < 0.06$	$ \epsilon_{\tau e}^{\mathrm{uV}} < 0.036$
			$(\epsilon_{\tau e}^{uV} < 0.023)$



[J. Barranco, et. al., hep-ph/0508299]

Non Standard Interaction & Unparticle



[J. Barranco, et. al., hep-ph/0901.2099]

- Parametrize by d: dimension, $\lambda_{if}^{\alpha\beta}$: coupling.
 - $\rightarrow \alpha, \beta$: neutrino flavors,
 - $\rightarrow f = e, u, d$
- Bounds on λ_{ie} by νe^- experiments.
- νN Coherent Scattering set limits on $\lambda_{iu,d}$.



Ge Detector

- ULE-HPGe → developed for soft X-rays detection
 → easy & inexpensive & robust operation
- threshold ~100 eV achievable.
- $\bar{\nu_e}N$ coherent scattering, CDM search, μ_{ν} search.
- possible reactor monitoring and other new detection channel.
- · Prototype detectors:
 - 5 g(at KS & Y2L), 4×5 g(at KS & Y2L), 10 g(at AS & CIAE), 180 g segmented(at AS).
- 500 g single element PCGe at KS.







5 g Detector

4×5 g Detector

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4×5 g Ge Detector







- Event Selection :
 - Anti-Compton, cosmic veto, PSD.
- Random trigger events :
 - · Deadtime.
 - Anti-Compton & cosmic veto cut efficiencies.



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Sub-keV Neutrino and Dark Matter Physics with Ultra-Low-Energy

4×5 g Detector : Threshold & Trigger Efficiency



- Fraction of pulser-amplitude > threshold.
- Noise edge $\sim 300~eV.$
- Trigger threshold \sim 100 eV.





4×5 g Detector : PSD Efficiency



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4×5 g Detector : PSD Efficiency

$$f = \frac{(\varepsilon_{PSD} \times P + f_N \times N)}{p + N}$$

- ϵ_{PSD} : real PSD efficiency.
- P: Physical events count, N: Noise count.
- f_{N} fraction of noise survived, measured: $f_{\text{N}} \sim 10^{-4}.$
- f: Estimate PSD efficiency by this method.
- $\epsilon_{PSD} > f$, conservative choice.



Anti-Compton Tag events to evaluate PSD efficiency.

4×5 g Detector : PSD Efficiency



- ACV tag method & spectrum-edge fitting method consistent.
- Trigger(50% eff.) \sim 100 eV.
- PSD(50% eff.) \sim 220 eV.



Quenching Factor



CIAE neutron beam facility



- Adopt TRIM approach.
- Quenching Factor ~ 0.2 at keV.



Background





Dark Matter Search

WIMP Spin-Independent Cross Section



Standard conservative analysis

 \rightarrow WIMP rate < total events rate (No background subtraction)

- Adopt standard astrophysics parameters.
- Optimal Interval Method [Yellin, PRD 2002]

Dark Matter Search

WIMP Spin-Dependent Cross Section



- Formalisms [Tovey, et. al., PLB 2000]
- · Ge Matrix elements [Dimitrov, et. al., PRD 1995]

Published : S. T. Lin, et. al., PRD 79, 061101(R) 2009



Point Contact Ge Detector





- Proposed by Luke (80's), successful demonstration by CoGeNT (2007).
- Position sensitive from drift-profile pulse shape.
- Dual-electrode readout for PSD.
- Utra-Low-Background specification.
- 500 g, Data taking in Nuclear Power Plant since November 2008.
- 900 g under construction.

Trigger & PSD Effeciency



- Trigger Eff. from background data.
- Trigger Eff. from pulser.
- Trigger threshold (Eff. = 50%) \sim 280 eV
- + PSD threshold (Eff. = 50%) \sim 320 eV



PSD Effeciency





200

Nal Timing(50ns)

Jin-Ping Underground Lab





- Sichuang, China.
- Tallest Peak 4193 m.
- Max. Rock Overburden: 2375 m.
- Road Tunnel Distance: 17.5 km.
- Fraction of tunnel with >1500 m Rock Overburden: >70%.



Jin-Ping Underground Lab





Channeling Effect





Summary

- Point Contact Ge detector :
 - · 500 g PC-Ge detector had been installed.
 - Preliminary result : Threshold ~ 300 eV.
- Physics :
 - Magnetic Moment to $10^{-11}\mu_B$ level.
 - vN coherent scattering, threshold < 500 eV.
 - WIMP search at low mass region.
- Plans :
 - Quenching Factor mesurement of Ge.
 - · Channeling Effect measurement of Csl.
 - · Further PSD analysis and background suppression.
 - Further background understanding & WIMP search
 → move in Sichuan underground Lab. (>2 km rock)



Thanks for invitation.



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