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From Core to Crust: Towards an Integrated Vision of Earth's Interior

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Molten Earth: Magma in the Deep Mantle

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MOLTEN EARTH: MAGMA IN THE DEEP MANTLE

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Molten Earth

- Very little of it is molten today
- Melt produced in shallow mantle escapes rapidly to the surface
- Responsible for formation of crust
- Long-term evolution of atmosphere, ocean
- Deep Earth?
- Early Earth?



Early Earth

May have been completely molten

- Kinetic energy of accretion
- Core formation
- Archean Lavas
 - Hotter
 - Melting at great depth



Canup (2004) Icarus

Deep Earth

Ultra Low Velocity Zones

- Located at core-mantle boundary (2890 km depth)
- ~10 km thick
- Collosal velocity anomaly (>20 % in V_S)
- Possibly partial melts
- Lateral heterogeneity in the deep mantle



Williams & Garnero (1996) Science

Magma Dynamics Today

- Mostly in upper 100 km
 Large volume contrast
 ~10 % between liquid and coexisting solids
 - Liquid and crystalline structures differ
- Consequences:
 - Buoyancy
 - Steep melting slope
 - (Clapeyron Equation)
- Higher Pressure?



Structure of Silicates

- Tetrahedrally coordinated
 - SiO₄ tetrahedra
 - Most abundant minerals in Earth's crust
- Octahedrally coordinated
 - SiO₆ octahedra
 - Stabilized by elevated pressure (P>10 GPa)
 - Much denser than tetrahedrally coordinated silicates
 - Most abundant "minerals" in mantle (MgSiO₃ perovskite)
- Liquids?
 - Can silicate liquids exhibit octahedral coordination?



Chao et al. (1962) JGR



Silicate Liquid Structure?

- Low pressure experiment reveals tetrahedral coordination
- No experimental data above a few GPa
- Previous studies on glasses
- Suggest gradual transition from dominantly tetrahedral to dominantly octahedral over several tens of GPa
- Liquids?





Equation of State of Mantle Solids

- Thermal pressure
 - Decreases on compression
 - Thermal effects "squeezed out"
 - $\square P_{TH} \sim \gamma 3 N k_B T / V$
- Grüneisen parameter γ
 - Decreases on compression
 - Controls magnitude of thermal pressure
 - Controls adiabatic gradient
 - $\Box \gamma = (dlnT/dlnV)_{S}$
- Liquids? Virtually unknown at lower mantle conditions





N. de Koker, Bayreuth

Density Functional Theory

Charge density in Mg₂SiO₄ wadsleyite

Density Functional Theory

- □ Schrödinger Equation
 - Exact but insoluble
 - Central quantity: many electron wave function
- □ Kohn-Sham Equations
 - Exact and an excellent starting point for approximation
 - Central quantity: charge density
- Exchange-correlation hole
 - e⁻ digs a hole of reduced charge density about itself
 - Approximate form of this interaction
 - Local Density Approximation (LDA)



R. Q. Hood, Cambridge

Density Functional Theory

Predictive power

- No free parameters
- No a priori assumptions regarding shape of charge density or nature of bonding
- Scope
 - Entire pressuretemperature range of planets (and stars)
 - Entire periodic table
- Accuracy
 - Tested via comparison with experimental data

Circles: Karki et al., 1997, Am. Min. Squares: Murakami et al., 2006, EPSL



Theory of Planetary Materials

- Earth occupies an intermediate regime pressure-temperature regime in which material behavior is remarkably rich
- Pressure is large ~1 Mbar
 - Bond deformation pressure: 1 eV/Å³~1.6 Mbar
 - Bulk Modulus ~ 1.5 Mbar
- But not so large that simple theories apply, e.g. for stars:
 - Thomas-Fermi







First Principles Molecular Dynamics

Example: MgSiO₃; Two-fold compression: $V/V_x=0.5$; 6000 K Initial condition: pyroxene structure, Maxwellian velocities

First Principles Molecular Dynamics

- Dynamics
 - Newtonian trajectory
 - Forces on nuclei from density functional theory
 - Born-Oppenheimer limit
 - Efficient solution of electronic problem
- Statistical mechanics
 - Thermodynamic properties via ergodic hypothesis
 - Nosé Thermostat (NVT)
- Computation
 - Plane-wave pseudopotential method
 - VASP: Kresse, Hafner, Furthmüller
- Systems:

 $\begin{array}{l} \mathsf{MgSiO}_3, \, \mathsf{Mg}_2\mathsf{SiO}_4, \, \mathsf{MgSiO}_3\text{-}\mathsf{H}_2\mathsf{O} \\ \mathsf{MgO}\text{-}\mathsf{SiO}_2, \, \mathsf{Fe}_2\mathsf{SiO}_4 \end{array}$



MgSiO₃; P~0 GPa; 3000 K

Liquid Structure

- Octahedral coordination at base of mantle
- Gradual change in mean coordination number
- Nearly linear in volume
- Five-fold most abundant at intermediate compression





Stixrude & Karki (2005) Science

Volume V/V_X

Equation of State

- Differs fundamentally from solid equation of state
- Isotherms diverge on compression
- Smooth function
- Excellent agreement
 with available
 experimental data



Volume V/V_X



Stixrude & Karki (2005) Science

Grünseisen Parameter

- Increases on compression
- Differs from all mantle crystalline phases
- Caused by change in liquid structure on compression
- Excellent agreement
 with available
 experimental data





Stixrude & Karki (2005) Science

Melting Curve

- Integrate Clapeyron equation using FPMD results
- MgSiO₃ melts at 5400±600
 K at mantle's base
- Agrees with Hugoniot datum
- Disagrees with previous studies based on the laserheated diamond anvil cell
 - Detection of melt
- Disagrees with Lindemann Law
 - Liquid structure changes on compression

 $\partial \ln T_M$





Stixrude & Karki (2005) Science

- Ultra Low Velocity Zones
 - Located at core-mantle boundary (136 GPa)
 - ~10 km thick
 - Colossal shear velocity anomaly ($\Delta V_{s} > 20$ %)
 - Large longitudinal velocity anomaly ($\Delta V_P \sim 10$ %)
- Partial Melt
 - Thermodynamically stable
 - Denser than surroundings
 - Satisfy seismic observations



Williams & Garnero (1996) Science

- Base of mantle may be molten
- Assume freezing point depression of 1300 K
- Based on ionic MD (Zhou & Miller, 1997). Water not considered
- Solidus at 4100 K
- Identical to modern estimates of temperature at core-mantle boundary
- Agrees with previous experimentally based extrapolations



Stixrude et al. (2009) EPSL

- Liquid likely denser than coexisting solids at base of mantle
- Small volume contrast
- Sensible iron partitioning
- DFT level of theory essential
- Excellent agreement
 with available shock
 wave data



- Partial melt
- Wave velocity of silicate liquids ~20 % lower than that at mantle's base
- Seismological observation finds ULVZ ~10 % slower
- Precise amount of partial melt required depends sensitively on texture



Stixrude & Karki (2005) Science Stixrude et al. (2009) EPSL

Previous Studies

- Concept of magma "ocean" with a crystalline silicate floor comes from studies of moon
- Magma ocean adiabat roughly parallel to that of present mantle
- Shallower than solidus
- Very high potential temperatures required for complete melting
- Crystallization begins at base





Wood et al. (2006) Nature Walter & Tronnes (2004) EPSL

- Deep turbulent convection
- Isentropic (adiabatic)
 temperature gradient
- Much higher than thought because γ is much larger than previously assumed
- Adiabat steeper than melting curve at depth
- Complete melting much easier than previously thought



James Garry, Fastlight Used with permission

 $\partial \ln T$

 $\partial \ln T$

- Compare temperature at depth T_{ad}(P) to melting temperature T_M(P)
- Adiabatic gradient is a material property
- Particular adiabat T_{ad}(P) is specified by temperature at the surface
- \Box Potential temperature T_0
- Increase in temperature originates in gravitational self-compression





Pressure

Temperature

- Potential temperature is also the temperature that any parcel of the interior would cool to on adiabatic ascent
- □ The temperature increment due to gravitational self
 -compression: ΔT(P)=T_{ad}(P)-T₀



Pressure

Temperature

- \Box Temperature $T_{cd}(P)$ is the sum of two energy sources
- Gravitational self -compression sets $\Delta T(P)$
- Other energy sources set T_0
 - Accretional Energy
 - Core Formation
 - Radioactive decay



- Complete melting much easier than previously thought
- \Box $\Delta T(P)$ much greater than previously thought
 - Steep liquid-state adiabats (large γ)
- □ T₀~2450 K sufficient to melt entire mantle
- T₀~2000 K (Archean plumes?) produces melt in lower mantle
- Crystallization of magma ocean begins at midmantle depths



Molten Earth

Lower Magma Layer ULVZ a remnant? Source of chondritic

complement?

Source of lower mantle chemical heterogeneity?

Volatile reservoir?

Reaction with core?



Fossils of Molten Earth?

- Geochemically
 observable Earth, i.e.
 lavas, not chondritic
 after all
- Hidden non-chondritic reservoir?
- Possible remnant of lower magma layer



Boyet & Carlson. (2005) Science

Fossils of Molten Earth

- Lateral variations in shear and bulk sound velocity may be anti-correlated in the lowermost mantle
- Some anomalies have sharp edges
- This cannot be caused by lateral variations in temperature alone
- Lateral variations in composition seem to be required
- Melting processes a natural way to produce compositional heterogeneity
- Lower magma layer may be the source



Ishii & Tromp (1999) Science

Molten Earth



Labrosse et al. (2007) Nature



- Large at low pressure
- H₂O in melt is denser and less compressible than pure H₂O
- □ Volume of solution ≤0 at all ~100 GPa pressures
- Solution behaves nearly ideally at high pressure

Mookherjee et al. (2008) Nature

 $MgSiO_3-H_2O$ (10 wt. %)

Influence of Water on Melts

~0 GPa

2

 \cap

Н

Mg





 $MgSiO_3-H_2O$

Complete miscibility of silicate liquids and water over nearly entire mantle

 ∂G_{sol}

dP

 $=V_{sol}\leq 0$





b





Composition Mookherjee et al. (2008) Nature

Shen & Keppler (1997) Nature P ~ 1.5 GPa

Initial Water Content of Earth

- Unlimited solubility of water in deep magma ocean has implications for
 - Deep mantle solidus
 - Origin of atmosphere
 - Present water content of interior
- Why is the magmatically probed mantle much drier than Cl chondrites?
 - Most water never accreted
 - Accreted and then removed
 - Currently stored at depth?



Fluid inclusions in a chondritic meteorite Zolensky et al. (1999) Science

First Principles Petrology of the Deep Earth

- Predict solution properties with first principles molecular dynamics
- \square MgO-SiO₂ join
- Large, negative volume of mixing approaches zero on compression
- Enthalpy of solution is large and positive at low pressure and silica-rich end
- Everywhere negative at high pressure
- No immiscibility at depth





De Koker et al. (to be published)

Liquid Fundamental

Thermodynamic Relation

- New analytical formulation F(V,T)
- Accounts self-consistently for all thermodynamic properties
- Based on our FPMD results
- Birch-Murnaghan finite strain theory
- Generalized Rosenfeld-Tarazona thermal contributions
- Electronic contributions are important



de Koker & Stixrude (2009) GJI



Liquid Fundamental Thermodynamic Relation

- New formulation correctly captures liquidgas transition
- Important for early
 Earth heat transfer
- Prediction for Mg₂SiO₄ at 6000 K
 - Vapor pressure = 0.09
 GPa (900 bars)
 - Heat of vaporization = 1130 kJ/mol





Conclusions

- First principles molecular dynamics simulations a powerful means of studying silicate liquids
- Predict gradual four-fold to six-fold Si-O coordination change over mantle pressure range
- Predict large Grüneisen parameter of silicate liquids
- ULVZ may be partially molten
- Early Earth may have been completely molten
- Began crystallizing from the mid-mantle
- Unlimited solubility of water in magma ocean
- Deep mantle may be a reservoir of primordial water