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The Optical Model

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# The Optical Model 

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#### Abstract

The basic features of elastic and inelastic scattering within the framework of the spherical and deformed nuclear optical models are discussed. The calculation of cross sections, angular distributions and other scattering quantities is described and the distorted-wave Born approximation calculations is also reviewed.


## 1 Introduction

A nuclear reaction is initiated when a nucleon or nucleus collides with another nucleon or nucleus. Reactions are characterized in first place by the incoming nuclei and the outgoing reaction products. Examples of the usual notation for this are ${ }^{14} \mathrm{C}(\mathrm{n}, \mathrm{n}){ }^{14} \mathrm{C}$, for the elastic scattering of neutrons on ${ }^{12} \mathrm{C},{ }^{56} \mathrm{Fe}(\mathrm{p}, \mathrm{t})^{54} \mathrm{Fe}$, for the pickup by a proton of two neutrons from ${ }^{56} \mathrm{Fe}$, and ${ }^{235} \mathrm{U}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)$ for inelastic neutron scattering from ${ }^{235} \mathrm{U}$.

A complete description of a nuclear reaction involves other observable quantities beside the incoming nuclei and the outgoing reaction products. Among these are the relative energy of the incoming and outgoing nuclei and the scattering angle of the outgoing products. When the nuclei/nucleons involved have spin and/or excited states, their polarizations and/or excitation energies can also be observed.

The characteristics of the reactions induced by a given pair of incident nucleons/nuclei can be summarized in distributions of the occurrence of the reaction products, called cross sections. Quantitatively, the cross section $\sigma_{p}$ for the production of a product $p$ is defined as

$$
\begin{equation*}
\sigma_{p}=\frac{\text { number of particles } \mathrm{p} \text { produced per unit time }}{\text { number of incident particles per unit time per unit area }} . \tag{1}
\end{equation*}
$$

Cross sections have the dimension of area. The information obtained from cross sections often depends quite strongly on the internal structure of the initial and final nuclei. In fact, the comparison of experimental scattering observables with those obtained from various nuclear models can teach us a great deal about the structure of individual nuclei. After having used such a comparison to determine the model parameters appropriate for a given system, one hopes to use the same parameters to predict cross sections in other energy ranges or in neighboring systems.

At low energies and for all but the lightest nuclear systems, nuclear reactions occur on two very distinct time scales. Direct reactions occur promptly, on a time scale of the same magnitude as the time it takes the projectile nucleus to pass by the target nucleus. Compound nuclear reactions, which involve the formation of a quasi-bound intermediate complex, occur on a time scale that is at least several orders of magnitude larger. A naive application of the uncertainty relation, $\Delta E \Delta t \geq \hbar$, would lead one to expect their energy scales to be inversely related. This is, in fact, the case. The contributions of direct reactions to the cross sections vary smoothly with energy. Compound nuclear reactions make contributions to the cross sections that fluctuate rapidly with energy.


Figure 1: ${ }^{58} \mathrm{Ni}$ at low incident energy, taken from the data of Ref. 1

The difference in the energy dependence of the direct and compound nucleus contributions to the cross section is clearly seen in Fig. 1, which displays the total neutron cross section on ${ }^{58} \mathrm{Ni}$ at extremely low incident neutron energy. One observes a direct reaction cross section - the result of elastic scattering of the neutron, in this case - that varies slowly with energy, except where it is punctuated by a faster variation due to the presence of a compound nuclear state of ${ }^{59} \mathrm{Ni}$ of about the same energy. At such low energies, separation of the direct and compound nucleus cross sections is a fairly straightforward (although often grueling) task.

At higher energies, the density of compound nucleus states becomes so large that the individual contributions can no longer be resolved. It then becomes impossible to distinguish the slow energy dependence of the direct contribution from the rapid variations of the compound nucleus one. An example of this is given in Fig. 2, where the total cross section for neutrons incident on ${ }^{58} \mathrm{Ni}$ is again shown, but now at higher energies. The fluctuations in the cross section, called Ericson fluctuations, [3] do not permit the determination of the contribution to the cross section of each individual compound nuclear state. Instead, only the average properties of the compound nucleus contribution to the cross section can be determined. It is in this context that the optical potential plays a crucial role in the separation of the two contributions.

The principal objective of the optical model is to describe just the prompt, direct reactions in a nuclear collision. To separate the direct reactions from the compound-nucleus ones (the-


Figure 2: ${ }^{58} \mathrm{Ni}$ in a small incident energy range close to 5 MeV , taken from the data of Ref. 2
oretically), one assumes that the compound-nucleus reactions do not contribute to the average wave function and scattering amplitudes, due to their rapid fluctuations in energy. Note that the compound-nucleus reactions still DO contribute to the average cross sections, which are, for the most part, proportional to the squares of the amplitudes. The energy-averaged amplitudes, however, are associated with the scattering amplitudes for the prompt component of the scattering. The optical model potential is defined as the potential which furnishes the energy-averaged scattering amplitudes.

In a wider context, the optical potential can be considered an effective potential that takes into account all of the physical processes one does not want to take into account explicitly. The most important of these are the rapidly fluctuating compound-nucleus contributions to the scattering. But direct processes are also included at times. One example of this is the use of an effective spherical optical model potential to take into account the coupling to excited states of the target. Another example is the deuteron optical potential, which usually contains the contribution of direct deuteron breakup.

As well as being fundamental for the calculation of direct reaction observables, optical model calculations are also used to produce the transmission coefficients essential for the analysis of compound nucleus cross sections within the Hauser-Feshbach statistical theory. They are thus one of the first and most important steps in the evaluation of nuclear cross sections.

## 2 Formal development of the optical model

To derive the optical model from 'first principles', one begins by partitioning the Hilbert space of states into a component $\mathcal{P}$ containing the prompt states and an orthogonal component $\mathcal{Q}$ that contains the closed channels of the intermediate compound complex.[4] As a concrete example, one may consider $\mathcal{P}$ to be the subspace consisting of a nucleon scattering on ${ }^{58} \mathrm{Ni}$, while $\mathcal{Q}$ consists of the ground and excited states of the nucleus ${ }^{59} \mathrm{Ni}$ (and other processes, such as $\gamma$ emission, that have been neglected). The projection operators, $P$ and $Q$, onto the subspaces $\mathcal{P}$ and $\mathcal{Q}$,
respectively, which satisfy the properties

$$
\begin{gather*}
P=P^{\dagger} \\
P^{2}=P \quad Q=Q^{\dagger}  \tag{2}\\
P+Q=1,
\end{gather*}
$$

are then used to decompose the state vector of the system, $\Psi$, and the Schrödinger equation it satisfies,

$$
\begin{equation*}
(E-H) \Psi=0 . \tag{3}
\end{equation*}
$$

The prompt component of the state vector is $P \Psi$, while the slower component is $Q \Psi$, with

$$
\begin{equation*}
\Psi=P \Psi+Q \Psi . \tag{4}
\end{equation*}
$$

We can multiply the Schrödinger equation on the left by $P$ or by $Q$ and use the decomposition of the wave vector to write the equation as two coupled equations,

$$
\begin{equation*}
\left(E-H_{P P}\right) P \Psi=V_{P Q} Q \Psi \tag{5}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(E-H_{Q Q}\right) Q \Psi=V_{Q P} P \Psi \tag{6}
\end{equation*}
$$

where

$$
\begin{equation*}
H_{P P} \equiv H_{0 P}+V_{P P} \equiv P H_{0} P+P V P, \quad V_{P Q} \equiv P H Q, \quad \text { etc. }, \tag{7}
\end{equation*}
$$

and we have assumed that the contributions to the Hamiltonian of the internal degrees of freedom and the kinetic energy, both contained in $H_{0}$, do not couple the $\mathcal{P}$ and $\mathcal{Q}$ subspaces. We may formally solve the first of these, Eq. (5), as

$$
\begin{equation*}
P \Psi_{i}=\phi_{i}^{(+)}+\frac{1}{E^{(+)}-H_{P P}} V_{P Q} Q \Psi_{i} \tag{8}
\end{equation*}
$$

in which the $(+)$ denotes an incoming wave boundary condition, the vector $\phi_{i}^{(+)}$satisfies the Schrödinger equation in the $\mathcal{P}$ subspace,

$$
\begin{equation*}
\left(E-H_{P P}\right) \phi_{i}^{(+)}=0 \tag{9}
\end{equation*}
$$

with an incoming wave in channel i alone (and none in the $\mathcal{Q}$ subspace) and $P \Psi_{i}$ and $Q \Psi_{i}$ are the components of the full wave vector that evolve from this incoming wave. The solution $P \Psi_{i}$, when substituted into the second coupled equation, Eq. (6), yields

$$
\begin{equation*}
\left(E-H_{Q Q}-W_{Q Q}\right) Q \Psi_{i}=V_{Q P} \phi_{i}^{(+)}, \tag{10}
\end{equation*}
$$

where

$$
\begin{equation*}
W_{Q Q} \equiv V_{Q P} \frac{1}{E^{(+)}-H_{P P}} V_{P Q} . \tag{11}
\end{equation*}
$$

We can decompose the $\mathcal{P}$-subspace Greens function into its real and imaginary parts as

$$
\begin{equation*}
\frac{1}{E^{(+)}-H_{P P}}=\frac{P . P .}{E-H_{P P}}-i \pi \delta\left(E-H_{P P}\right), \tag{12}
\end{equation*}
$$

where $P . P$. represents the principal part. The open channels in the $\mathcal{P}$ subspace thus make a negative imaginary contribution to $W_{Q Q}$, which results in singularities in the wave vector in the lower half of the complex E plane.

Eq. (10) can be solved to obtain the $\mathcal{Q}$-subspace component of the wave vector as

$$
\begin{equation*}
Q \Psi_{i}=\frac{1}{E-H_{Q Q}-W_{Q Q}} V_{Q P} \phi_{i}^{(+)} \tag{13}
\end{equation*}
$$

which then permits the expression of the $\mathcal{P}$-subspace component of the wave vector as

$$
\begin{equation*}
P \Psi_{i}=\phi_{i}^{(+)}+\frac{1}{E^{(+)}-H_{P P}} V_{P Q} \frac{1}{E-H_{Q Q}-W_{Q Q}} V_{Q P} \phi_{i}^{(+)} \tag{14}
\end{equation*}
$$

A careful analysis of the last expression leads one to the scattering matrix $\mathcal{I}_{f i}$ giving the transition amplitude in the $\mathcal{P}$ subspace,

$$
\begin{equation*}
\mathcal{T}_{f i}=\mathcal{T}_{f i}^{(P)}+\left\langle\phi_{f}^{(-)}\right| V_{P Q} \frac{1}{E-H_{Q Q}-W_{Q Q}} V_{Q P}\left|\phi_{i}^{(+)}\right\rangle \tag{15}
\end{equation*}
$$

The first term in this expression is the direct scattering amplitude associated with scattering in the $\mathcal{P}$ subspace alone. The second term describes the slower processes that result from coupling through the states of the $\mathcal{Q}$ subspace. The first term varies slowly as a function of energy while the second term varies rapidly.

The energy average of the $\mathcal{P}$-subspace wave vector can now be written as

$$
\begin{equation*}
\left\langle P \Psi_{i}\right\rangle=\phi_{i}^{(+)}+\frac{1}{E^{(+)}-H_{P P}} V_{P Q}\left\langle\frac{1}{e_{Q Q}}\right\rangle V_{Q P} \phi_{i}^{(+)} \tag{16}
\end{equation*}
$$

where

$$
\begin{equation*}
e_{Q Q}=E-H_{Q Q}-W_{Q Q} \tag{17}
\end{equation*}
$$

is the only rapidly varying function of the energy in the expression. The average wave vector can be written in a Schrödinger-equation-like form by multiplying both sides of the expression, Eq. (16), by $E^{(+)}-H_{P P}$,

$$
\begin{equation*}
\left(E-H_{P P}\right)\left\langle P \Psi_{i}\right\rangle=V_{P Q}\left\langle\frac{1}{e_{Q Q}}\right\rangle V_{Q P} \phi_{i}^{(+)} \tag{18}
\end{equation*}
$$

Using Eq. (16) again to rewrite the wave vector $\phi_{i}^{(+)}$as

$$
\begin{equation*}
\phi_{i}^{(+)}=\frac{1}{1+\left(E^{(+)}-H_{P P}\right)^{-1} V_{P Q}\left\langle 1 / e_{Q Q}\right\rangle V_{Q P}}\left\langle P \Psi_{i}\right\rangle \tag{19}
\end{equation*}
$$

substituting this in Eq. (18) and performing a bit of algebra, one finally obtains the optical model equation,

$$
\begin{equation*}
\left[E-H_{P P}-V_{P Q} \frac{1}{\left\langle 1 / e_{Q Q}\right\rangle^{-1}+W_{Q Q}} V_{Q P}\right]\left\langle P \Psi_{i}\right\rangle=0 . \tag{20}
\end{equation*}
$$

The optical potential can thus be written as

$$
\begin{equation*}
V_{o p t}=V_{P P}+V_{P Q} \frac{1}{\left\langle 1 / e_{Q Q}\right\rangle^{-1}+W_{Q Q}} V_{Q P} \tag{21}
\end{equation*}
$$

To conclude the formal development of the optical model, one must evaluate the average value $\left\langle 1 / e_{Q Q}\right\rangle$. The simplest way of doing this is to average the quantity $1 / e_{Q Q}$ over a normalized Lorentzian density,

$$
\begin{equation*}
\left\langle\frac{1}{e_{Q Q}}\right\rangle=\int d E_{0} \frac{\rho\left(E, E_{0}\right)}{E_{0}-H_{Q Q}-W_{Q Q}} \tag{22}
\end{equation*}
$$

where

$$
\begin{equation*}
\rho\left(E, E_{0}\right)=\frac{\Delta}{2 \pi} \frac{1}{\left(E-E_{0}\right)^{2}+(\Delta / 2)^{2}} \tag{23}
\end{equation*}
$$

Assuming the quantity $1 / e_{Q Q}$ to have no poles in the upper half of the complex E plane (due to causality, it should have them only in the lower half-plane), we can perform the integral by closing the contour and calculating residues in the upper half plane to obtain

$$
\begin{equation*}
\left\langle\frac{1}{e_{Q Q}}\right\rangle=\frac{1}{E+i \Delta / 2-H_{Q Q}-W_{Q Q}} \tag{24}
\end{equation*}
$$

and hence

$$
\begin{equation*}
V_{o p t}=V_{P P}+V_{P Q} \frac{1}{E-H_{Q Q}+i \Delta / 2} V_{Q P} \tag{25}
\end{equation*}
$$

The optical potential is obviously energy-dependent, non-local and complex due to the energyaveraged propagator $\left(E-H_{Q Q}+i \Delta / 2\right)^{-1}$ in the second term. Its imaginary part is negative, resulting in a potential that is absorptive. The flux of particles leaving the scattering region is, in this case, smaller than the incident flux, with the remaining fraction of the flux being absorbed by the potential. It is through its imaginary part that the optical potential takes into account the flux that is lost from the states of the $\mathcal{P}$ subspace to the states of the $\mathcal{Q}$ subspace.

The optical scattering matrix can easily be derived in the same manner. One obtains

$$
\begin{equation*}
\langle\mathcal{T}\rangle_{f i}=\mathcal{T}_{f i}^{(P)}+\left\langle\phi_{f}^{(-)}\right| V_{P Q}\left\langle\frac{1}{e_{Q Q}}\right\rangle V_{Q P}\left|\phi_{i}^{(+)}\right\rangle \tag{26}
\end{equation*}
$$

with $\left\langle 1 / e_{Q Q}\right\rangle$ given by Eq. (24). Observe that the second, rapidly fluctuating term does not vanish completely. Indeed it should not vanish in general, for its average contribution describes the loss of flux from the prompt channels to the long-lived compound-nucleus states.

## 3 Low-energy neutron scattering

At low relative energies, a collision between charged nuclei or a nucleus and a charged nucleon is dominated by the Coulomb force, which keeps the two beyond the range of nuclear interaction. Only neutrons can enter sufficiently close to a nucleus at such energies to feel the effects of the nuclear force.

Several factors also simplify the description of low-energy neutron scattering. The centripetal barrier keeps all but the $l=0$ s-wave contribution effectively out of the reach of the nuclear interaction for energies smaller than about 50 keV . In addition, with few exceptions, nuclei have no excited states at energies lower than about 20 keV . The prompt component of neutron scattering then reduces to s-wave elastic scattering in this energy range.

The optical model equation for the s-state wave function $\psi_{0}$ is

$$
\begin{equation*}
\left(E_{c m}-T-U_{o p t}\right) \frac{\psi_{0}}{r}=0 \tag{27}
\end{equation*}
$$

which can be reduced to

$$
\begin{equation*}
\frac{d^{2} \psi_{0}}{d r^{2}}+\left[k^{2}-\frac{2 \mu}{\hbar^{2}} U_{o p t}\right] \psi_{0}=0 \tag{28}
\end{equation*}
$$

where the wavenumber is $k=\sqrt{2 \mu E_{c m} / \hbar^{2}}, \mu$ is the reduced mass and $E_{c m}$ the center-of-mass energy.

To solve this equation numerically, one develops the solution, $\psi_{0, i n t}(r)$, starting from $r=0$, using the condition that the wave function vanishes at the origin, $\psi_{0, i n t}(r=0)=0$ and one of many possible numerical methods (Cowell, Numerov, modified Numerov, Runge-Kutta, etc.). The equation is solved numerically out to a radius $r_{m}$, beyond which the optical potential can be neglected. For values of the radius equal to or larger than this matching radius, the solution to the differential equation that satisfies the incoming wave boundary condition takes the form

$$
\begin{equation*}
\psi_{0, e x t}(r)=\frac{i}{2}\left(e^{-i k r}-S_{0} e^{i k r}\right) \quad r \geq r_{m} \tag{29}
\end{equation*}
$$

One requires, at the matching radius $r_{m}$, that this external wave function and its derivative be the continuous extensions of the numerical wave function obtained in the the internal region and of its derivative. This results in two equations,

$$
\begin{equation*}
a_{0} \psi_{0, \text { int }}\left(r_{m}\right)=\frac{i}{2}\left(e^{-i k r_{m}}-S_{0} e^{i k r_{m}}\right) \tag{30}
\end{equation*}
$$

and

$$
\begin{equation*}
a_{0} \frac{d}{d r} \psi_{0, i n t}\left(r_{m}\right)=\frac{k}{2}\left(e^{-i k r_{m}}+S_{0} e^{i k r_{m}}\right) \tag{31}
\end{equation*}
$$

whose solution yields the amplitude of the internal wave function, $a_{0}$, and the S-matrix element, $S_{0}$.

Once the S-matrix is known, the cross sections can be calculated. For the case of s-wave scattering, these are

$$
\begin{align*}
\sigma_{t o t} & =\frac{2 \pi}{k^{2}}\left(1-\operatorname{Re} S_{0}\right) \\
\sigma_{e l} & =\frac{\pi}{k^{2}}\left|S_{0}-1\right|^{2}  \tag{32}\\
\text { and } \quad \sigma_{r} & =\frac{\pi}{k^{2}}\left(1-\left|S_{0}\right|^{2}\right)=\frac{\pi}{k^{2}} T_{0},
\end{align*}
$$

where $T_{0}$ is the s-wave transmission coefficient. The reaction cross section and the transmission coefficient $T_{0}$ are non-zero when the S-matrix element $S_{0}$ is smaller than one in magnitude. This occurs when flux is absorbed by the long-lived compound-nucleus states. Care must be taken, however, when comparing the optical model reaction cross section to the experimental one. A part of the flux absorbed by the compound nucleus can later be re-emitted in the elastic channel, in which case it should rightly be considered part of the elastic cross section.

Of the three cross sections, only the total one can be compared directly with experimental data, as it is the only one that is linear in the scattering amplitude (here the S-matrix element
$\left.S_{0}\right)$. The S-matrix element can be written in general as the sum of an average and a fluctuating part, $S=S_{\text {ave }}+S_{f l}$. The average elastic cross section then has the form

$$
\begin{equation*}
\left.\left\langle\sigma_{e l}\right\rangle=\frac{\pi}{k^{2}}\left|S_{a v e}-1\right|^{2}+\left.\frac{\pi}{k^{2}}\langle | S_{f l}\right|^{2}\right\rangle . \tag{33}
\end{equation*}
$$

The first term alone gives the elastic cross section of the optical model. The second term contributes to the optical model reaction cross section.

Other scattering quantities of physical interest can also be calculated for low-energy neutron scattering. At extremely low energies - below the resonance region - the elastic cross section is observed to approach a constant value, $\sigma_{e l}^{0}$. This value is used to calculate the scattering radius, $R^{\prime}=\sqrt{\sigma_{e l}^{0} / 4 \pi}$.

In the resonance region, s-wave and p-wave strength functions can be defined. The s-wave strength function, $s_{0}$, relates the average neutron partial width $\left\langle\Gamma_{0}\right\rangle$ and spacing $D_{0}$ of the resonances to the optical model absorption. One has, approximately,

$$
\begin{equation*}
s_{0}=\frac{\left\langle\Gamma_{0}\right\rangle}{D_{0}}\left(\frac{E_{0}}{E_{c m}}\right)^{1 / 2} \approx \frac{1-\left|S_{0}\right|^{2}}{2 \pi \sqrt{E_{c m}}} \tag{34}
\end{equation*}
$$

where $E_{0}$ is usually taken to be 1 eV . The factor $\sqrt{E_{c m}}$, proportional to the s-wave penetrability, $k R^{\prime}$, cancels the energy dependence of the neutron partial width, so that the strength function varies slowly with the incident neutron energy. The p-wave strength function, $s_{1}$, relating the average partial width and spacing of the $l=1$ reonances is defined analogously in terms of the p -wave S -matrix elements and penetrability.

Adjustment of the optical model parameters at low energy to reproduce the s-wave and pwave strength functions, the scattering radius and the total cross section is known as the SPRT method.[5] A good fit to these observables is important in determining the low energy behavior of the optical cross sections and the transmission coefficients, which is important, in turn, in determining the behavior of compound nucleus cross section calculations near threshold.

## 4 The phenomenological optical potential

The formal derivation of the optical potential presented in Section 2 might suggest that it could be calculated directly. Although a good deal of work has indeed been done in this direction, the resulting potentials are often difficult to calculate and still not sufficiently precise. They also have the drawback of being non-local, which can greatly complicate solution of the corresponding Schrödinger equation.[6, 7, 8]

Instead, phenomenological optical model potentials are normally used to compare and fit to experimental data. With few exceptions, these potentials are taken to be local. However, the qualitative characteristics of the geometry and the general trend of the energy dependence of the phenomenological potentials are quite similar to those found in microscopic potentials. Both types of potentials are, after all, trying to describe the same physical processes.

In the empirical potentials, the functional form is usually determined by a limited set of parameters that are adjusted to obtain a best fit with the experimental data. Over the years, a
standard form of the phenomenological optical model potential has evolved, which permits the parametrization of the scattering of a light particle (neutron, proton, deuteron, triton, ${ }^{3} \mathrm{He}$ or alpha) from a given nucleus. This is

$$
\begin{array}{rlrl}
U_{o p t}(r)= & & \\
& +V_{C}(r) & & \text { a Coulomb term, } \\
& -V f_{V}(r) & & \text { a real volume term, } \\
& +V_{s} g_{V}(r) & & \text { a real surface term, }  \tag{35}\\
& -i W_{s} g_{W}(r) & & \text { an imaginary surface term, } \\
& -i W_{v} f_{W}(r) & & \text { an imaginary volume term } \\
& -d_{s o} \vec{l} \cdot \vec{s} V_{s o} h_{V_{s o}}(r) & & \text { a real spin-orbit term, } \\
& +i d_{s o} \vec{l} \cdot \vec{s} W_{s o} h_{W_{s o}}(r) & & \text { and an imaginary spin-orbit term, }
\end{array}
$$

where the spin-orbit constant is $d_{s o}=\left(\hbar / m_{\pi} c\right)^{2} \approx 2 \mathrm{fm}^{2}, m_{\pi}$ being the pion mass.
The Coulomb term is usually taken to be the interaction of a point charge with a uniformly charged sphere of radius $R_{c}$,

$$
V_{C}(r)= \begin{cases}\left(\frac{3}{2}-\frac{r^{2}}{2 R_{c}^{2}}\right) Z_{p} Z_{t} e^{2} / R_{c} & r \leq R_{c}  \tag{36}\\ Z_{p} Z_{t} e^{2} / r & r>R_{c}\end{cases}
$$

where $Z_{p}$ and $Z_{t}$ are the projectile and target charge, respectively. Although this potential neglects the surface diffusivity of the nuclear charge distribution, it is a reasonable approximation in the case of the scattering of light particles from nuclei.

The real and imaginary volume terms are normally taken to be of Wood-Saxon form,

$$
\begin{equation*}
f_{i}=\frac{1}{1+\exp \left[\left(r-R_{i}\right) / a_{i}\right]} \quad i=V, W \tag{37}
\end{equation*}
$$

where $R_{i}$ and $a_{i}$ are the radii and the diffusivities, respectively, of the two terms. The WoodSaxon form factor, shown in Fig. 3, can be thought of as a smoothed step function, falling from one for values of the radius $r$ smaller than the radius $R_{i}$ to zero for values of $r$ greater than $R_{i}$, in a few multiples of the diffusivity $a_{i}$.


Figure 3:

The real volume potential reflects the average interaction of the projectile with the nucleons of the target nucleus. The Wood-Saxon form factor it uses is quite similar in form to the nucleon
density of a saturated nucleus $(A \geq 30)$. (For lighter nuclei, a Gaussian geometry is sometimes used.) The strength of the real volume potential is roughly proportional to the mass of the projectile and decreases with the incident energy, in qualitative agreement with the results of calculations of the nuclear mean field.[9]

The imaginary volume potential takes into account the loss of projectile particles due to collisions with the nucleons of the target. It is zero at low energies, for which the projectile does not have sufficient energy to knock out a target nucleon. At higher energies, it increases slowly with the incident energy, as the phase space available for nucleon knockout increases. At even higher energies, both the real and imaginary volume potentials for nucleon scattering are fairly well described by the impulse approximation, in which the the target density is simply folded with the nucleon-nucleon cross section.[10, 11]

The real and imaginary surface terms of the optical potential are taken to be either the derivative of a Wood-Saxon,

$$
\begin{equation*}
g_{i}(r)=-4 a_{i} \frac{d}{d r} f_{i}(r)=4 \frac{\exp \left[\left(r-R_{i}\right) / a_{i}\right]}{\left(1+\exp \left[\left(r-R_{i}\right) / a_{i}\right]\right)^{2}} \quad i=V, W \tag{38}
\end{equation*}
$$

or a Gaussian,

$$
\begin{equation*}
g_{i}(r)=\exp \left[\frac{\left(r-R_{i}\right)^{2}}{a_{i}^{2}}\right] \quad i=V, W \tag{39}
\end{equation*}
$$

In either case, the potential peaks at a radius $R_{i}$ and falls to zero within a few multiples of the diffusivity $a_{i}$. A derivative Wood-Saxon form factor with diffusivity $a_{W S}$ is almost indistinguishable from a Gaussian form factor with diffusivity $a_{G}=2.21 a_{W S}$, as shown in Fig. 4.


Figure 4: $a_{W S}=0.5 \mathrm{fm}$. and $a_{G}=1.105 \mathrm{fm}$, respectively, are shown.

The imaginary surface term of the optical potential takes into account the absorption due to the coupling to the quasi-bound compound nucleus states through the excitation of low-energy collective modes, which have their couplings concentrated in the nuclear surface. Similar manybody effects can also be invoked to justify the presence of a real surface term. However, given the imaginary surface term, the existence of the real term can be shown to follow directly, by using a dispersion relation based on the causality of the optical potential (no singularities in the energy upper halfplane).[12] The dispersion relation shows that an energy-dependent imaginary
potential $W(r, E)$ necessarily leads to a contribution $\Delta V(r, E)$ to the real potential given by

$$
\Delta V(r, E)=\frac{\text { P.P. }}{\pi} \int_{-\infty}^{\infty} \frac{W\left(r, E^{\prime}\right)}{E^{\prime}-E} d E^{\prime}
$$

Obviously, if the imaginary term is a surface one, the real term resulting from the dispersion relation will be a surface one as well.

Both the real and imaginary spin-orbit terms of the optical potential are taken to have a Thomas form factor,

$$
\begin{equation*}
h_{i}(r)=-\frac{1}{r} \frac{d}{d r} f_{i}(r)=\frac{1}{r a_{i}} \frac{\exp \left[\left(r-R_{i}\right) / a_{i}\right]}{\left(1+\exp \left[\left(r-R_{i}\right) / a_{i}\right]\right)^{2}} \quad i=V_{s o}, W_{s o} . \tag{40}
\end{equation*}
$$

Like the surface imaginary term, the Thomas form factor,shown in Fig. 5, yields potentials which peak at a radius near $R_{i}$ and fall to zero in a few multiples of the diffusivity $a_{i}$.


Figure 5:
The Thomas form factor, as well as the spin-orbit potential itself, can be derived (for spin $1 / 2$ particles) by performing a reduction of a Dirac equation with Wood-Saxon potentials to an equivalent Schrödinger equation.[9] The spin-orbit interaction and the Thomas form factor can then be interpreted as but another manifestation of the volume interaction of the incident particle with the nucleons of the target nucleus.

The phenomenological optical potential is thus parametrized in terms of a set of potential strengths and corresponding geometrical parameters. These parameters have been adjusted for many systems and values of the relative energy. Several attempts have been made to adjust a single set of parameters to a wide range of systems by introducing a dependence on the target charge and mass as well as that on the relative energy. The potentials obtained using such sets of parameters are called global optical potentials. Many individual and global optical parameter sets can be found in an old compilation by Perey and Perey.[13] An extensive study of neutron and proton optical potentials has been performed recently by Koning and Delaroche.[14] However, the best modern reference for optical potential parameters is the Reference Input Parameter Library (RIPL-2), available both online from the International Atomic Energy Agency.[15]

For nucleons, typical values of the potential strengths are

$$
\begin{align*}
& V \approx(45-55) \mathrm{MeV}-(0.2-0.3) E, \\
& W_{s} \approx(2-7) \mathrm{MeV}-(0.1-0.3) E  \tag{41}\\
& V_{\text {so }} \approx(4-10) \mathrm{MeV} .
\end{align*} \quad E<8-10 \mathrm{MeV},
$$

Above $8-10 \mathrm{MeV}, W_{s}$ is usually constant or slightly decreasing. $V_{s}$ and $W_{s o}$ can normally be taken to be zero as can $W$ below about 10 MeV . Above about $10 \mathrm{MeV}, W$ is constant or slightly increasing. As mentioned above, for heavier particles, the real volume potential $V$ scales approximately linearly with the mass.

The radii $R_{i}$ characteristically take on values close to that of the radius of the target matter distribution. They are often parameterized in terms of reduced radii $r_{i}$ and the target mass as $R_{i}=r_{i} A_{t}^{1 / 3}$, with the reduced radii in the range $r_{i} \approx 1.2-1.3 \mathrm{fm}$. The diffusivities normally take on values in the range $a_{i} \approx 0.4-0.7 \mathrm{fm}$, except in the case of a Gaussian surface form factor, for which the typical values are slightly larger.

Not all of the optical model parameters are uniquely determined by the experimental data. It has been observed, for example, that fairly wide ranges of the parameters $V, R_{v}, W_{s}$, and $a_{s}$ result in equally good fits to the experimental data if the values of $V R_{v}^{2}$ and $W_{s} a_{s}$ remain constant. These are known as potential ambiguities.

## 5 Partial wave expansion in the single-channel optical model

When angular momenta greater than the s-wave contribute to the scattering, the wave function and the scattering matrix are determined most conveniently when decomposed in angular momentum partial waves.

The partial wave expansion of the scattering wave function of a particle of spin $s[16]$ can be written as

$$
\begin{equation*}
\Psi=\frac{4 \pi}{k r} \sum_{l j n} i^{l} e^{i \sigma_{l}} \psi_{l}^{j}(r) \mathcal{Y}_{l s}^{j n}(\hat{r}) \mathcal{Y}_{l s}^{j n \dagger}(\hat{k}), \tag{42}
\end{equation*}
$$

in terms of the spin-angular functions

$$
\begin{equation*}
\mathcal{Y}_{l s}^{j n}(\hat{r})=i^{l} \sum_{m \nu}\langle s \nu l m \mid j n\rangle Y_{l m}(\hat{r})|s \nu\rangle, \tag{43}
\end{equation*}
$$

where $l$ and $j$ are the orbital and total angular momenta and $|s \nu\rangle$ is an eigenvector of the particle spin. In the expansion of the wave function, $\sigma_{l}$ is the Coulomb phase, $\hat{r}$ denotes the angular variables and $\hat{k}$ the direction of the incident momentum. (The S-matrix element in partial wave $l$ for pure Coulomb scattering of the projectile from the target would be $e^{2 i \sigma_{l}}$.) The factor $i^{l} e^{i \sigma_{l}} \psi_{l}^{j}(r) / k r$ could have been written as simply $\psi_{l}^{j}(r)$ in the partial wave expansion. The form used above simplifies later manipulations.

When the partial-wave expansion of the wave function is substituted in the optical Schrödinger equation, one can extract an independent equation for the wave function $\psi_{l}^{j}$ in each partial wave. One finds

$$
\begin{equation*}
\left\{\frac{d^{2}}{d r^{2}}-\frac{l(l+1)}{r^{2}}+k^{2}-\frac{2 \mu}{\hbar^{2}}\left(U_{c e n}(r)+d_{l}^{j} U_{s o}(r)\right)\right\} \psi_{l}^{j}(r)=0 \tag{44}
\end{equation*}
$$

where the spin-orbit constant is $d_{l}^{j}=d_{s o}(j(j+1)-l(l+1)-s(s+1)) / 2$ and $U_{c e n}$ and $U_{s o}$ are the central and spin-orbit terms of the phenomenological optical model potential discussed in the previous section.

The incoming-wave boundary condition requires that asymptotically the wave function take the form of an incoming plane wave and an outgoing scattering wave,

$$
\begin{align*}
\Psi \rightarrow \exp (i \vec{k} \cdot \vec{r} & +i \eta \log (k r-\vec{k} \cdot \vec{r})) \sum_{\nu}|s \nu\rangle\langle s \nu|  \tag{45}\\
& +\frac{1}{r} \exp (i k r-i \eta \log (2 k r)) \sum_{\nu \nu^{\prime}}\left|s \nu^{\prime}\right\rangle f_{\nu^{\prime} \nu}(\theta)\langle s \nu|,
\end{align*}
$$

where the $f_{\nu^{\prime} \nu}(\theta)$ are the spin-projected matrix elements of the elastic scattering amplitude and $\eta$ is the Coulomb parameter, $\eta=\mu Z_{p} Z_{t} e^{2} / \hbar^{2} k$. To be consistent with this expression and satisfy the differential equation, the wave function $\psi_{l}^{j}$ must have the asymptotic form,

$$
\begin{equation*}
\psi_{l}^{j}(r) \rightarrow F_{l}(r)+\left(G_{l}(r)+i F_{l}(r)\right) C_{l}^{j}=\frac{i}{2}\left(H_{l}^{-}(r)-H_{l}^{+}(r) e^{2 i \sigma_{l}} S_{l}^{j}\right) e^{-i \sigma_{l}} \tag{46}
\end{equation*}
$$

where $C_{l}^{j}=\left(S_{l}^{j}-1\right) / 2 i, F_{l}$ and $G_{l}$ are the regular and irregular Coulomb wave functions, respectively, and $H_{l}^{ \pm}=e^{\mp i \sigma_{l}}\left(G_{l} \pm i F_{l}\right)$ are the linear combinations of these that asymptotically contain only incoming ( $H_{l}^{-}$) or outgoing ( $H_{l}^{+}$) waves. $S_{l}^{j}$ is the nuclear part of the S-matrix element and $e^{2 i \sigma_{l}}$ the Coulomb part.

The S-matrix elements, $S_{l}^{j}$, are obtained in the same manner as $S_{0}$ is obtained in the case of low-energy neutron scattering. In the internal region, the differential equation for each partial wave, Eq. (44), is solved numerically out to the radius, $r_{m}$. The numerical solution and its derivative are matched there to the wave function in the external region, given by Eq. (46), and to its derivative, to obtain the ampitude in the internal region, $a_{l}^{j}$, and the $S$-matrix element, $S_{l}^{j}$.

The only novelty to the solution here is deciding with which partial wave to stop the calculation, for $l$ and $j$ extend to infinity. The calculation is normally stopped when the nuclear S-matrix elements are sufficiently close to one. This occurs when the centripetal barrier no longer permits the projectile to enter the range of nuclear interaction with the target. For partial waves of larger $l$, the scattering reduces to pure Coulomb scattering (or for neutrons, no scattering at all), as is evident from Eq. (46).

When the asymptotic form of the partial wave function, $\psi_{l}^{j}$, of Eq. (46), is substituted in the partial wave expansion of the total wave function, Eq. (42), and the result is compared to the expected form of the asymptotic wave function, Eq. (45), one can extract the partial wave expansion of the scattering amplitude,

$$
\begin{equation*}
f(\theta)=\frac{4 \pi}{2 i k} \sum_{l j n}\left(e^{2 i \sigma_{l}} S_{l}^{j}-1\right) \mathcal{Y}_{l s}^{j n}(\hat{r}) \mathcal{Y}_{l s}^{j n \dagger}(\hat{k}) \tag{47}
\end{equation*}
$$

or,in terms of its spin-projected matrix elements,

$$
\begin{equation*}
f_{\nu^{\prime} \nu}(\theta)=\frac{4 \pi}{2 i k} \sum_{\substack{l j n \\ m m^{\prime}}}\left(e^{2 i \sigma_{l}} S_{l}^{j}-1\right) Y_{l m^{\prime}}(\hat{r}) Y_{l m}^{*}(\hat{k})\left\langle l m^{\prime} s \nu^{\prime} \mid j n\right\rangle\langle j n \mid l m s \nu\rangle . \tag{48}
\end{equation*}
$$

Due to the slow convergence of the Coulomb term, it is convenient to write these amplitudes in a form in which the Coulomb contribution has been summed exactly,

$$
\begin{array}{r}
f_{\nu^{\prime} \nu}(\theta)=\frac{4 \pi}{2 i k} \sum_{\substack{l j n \\
m m^{\prime}}}\left[\left(e^{2 i \sigma_{l}}-1\right)+e^{2 i \sigma_{l}}\left(S_{l}^{j}-1\right)\right] Y_{l m^{\prime}}(\hat{r}) Y_{l m}^{*}(\hat{k}) \\
=\delta_{\nu^{\prime} \nu} f_{C}(\theta)+\frac{4 \pi}{2 i k} \sum_{\substack{l j n \\
m m^{\prime}}} e^{2 i \sigma_{l}}\left(S_{l}^{j}-1\right) Y_{l m^{\prime}}(\hat{r}) Y_{l m}^{*}(\hat{k}) \\
 \tag{49}\\
\times\left\langle l m^{\prime} s \nu^{\prime} \mid j n\right\rangle\langle j n \mid l m s \nu\rangle \\
\\
\end{array}
$$

where

$$
\begin{equation*}
f_{C}(\theta)=-\frac{\eta}{2 k \sin ^{2} \theta / 2} \exp \left[-i \eta \log \left(\sin ^{2} \theta / 2\right)+2 i \sigma_{0}\right] \tag{50}
\end{equation*}
$$

is the Coulomb amplitude.
For spin- 0 particles, there is only one amplitude. This is

$$
\begin{equation*}
f(\theta)=f_{00}(\theta)=f_{C}(\theta)+\frac{1}{2 i k} \sum_{l}(2 l+1) e^{2 i \sigma_{l}}\left(S_{l}^{l}-1\right) P_{l}(\cos \theta) \tag{51}
\end{equation*}
$$

For spin- $1 / 2$ particles there are two distinct amplitudes. They are

$$
\begin{align*}
A(\theta) & =f_{\frac{1}{2} \frac{1}{2}}=f_{-\frac{1}{2}-\frac{1}{2}}  \tag{52}\\
& =f_{C}(\theta)+\frac{1}{2 i k} \sum_{l} e^{2 i \sigma_{l}}\left[(l+1)\left(S_{l}^{l+\frac{1}{2}}-1\right)+l\left(S_{l}^{l-\frac{1}{2}}-1\right)\right] P_{l}(\cos \theta)
\end{align*}
$$

and

$$
\begin{align*}
B(\theta) & =f_{\frac{1}{2}-\frac{1}{2}}=f_{-\frac{1}{2} \frac{1}{2}}  \tag{53}\\
& =\frac{1}{2 i k} \sum_{l} e^{2 i \sigma_{l}}\left[S_{l}^{l+\frac{1}{2}}-S_{l}^{l-\frac{1}{2}}\right] P_{l}^{1}(\cos \theta) .
\end{align*}
$$

The differential elastic cross section for an unpolarized incident beam is obtained by averaging the squared magnitude of the scattering amplitudes over the initial values of the projectile spin and summing over the final ones. The general expression that results is

$$
\begin{equation*}
\frac{d \sigma}{d \Omega}=\frac{1}{2 s+1} \sum_{\nu \nu^{\prime}}\left|f_{\nu^{\prime} \nu}(\theta)\right|^{2} . \tag{54}
\end{equation*}
$$

For spin-0 particles, this is

$$
\begin{equation*}
\frac{d \sigma}{d \Omega}=|f(\theta)|^{2} \quad s=0 \tag{55}
\end{equation*}
$$

For spin- $1 / 2$ particles, it is

$$
\begin{equation*}
\frac{d \sigma}{d \Omega}=|A(\theta)|^{2}+|B(\theta)|^{2} \quad s=1 / 2 . \tag{56}
\end{equation*}
$$

For particles of spin-1/2 and greater, one can define vector and tensor spin observables in terms of other combinations of the amplitudes. In particular, for particles of spin- $1 / 2$, the vector polarization, $P(\theta)$, is

$$
\begin{equation*}
P(\theta)=\frac{2 \operatorname{Im} A^{*}(\theta) B(\theta)}{d \sigma / d \Omega} \tag{57}
\end{equation*}
$$

The fraction of flux absorbed from each partial wave is given by the transmission coefficient, $T_{l}^{j}$, defined as

$$
\begin{equation*}
T_{l}^{j}=1-\left|S_{l}^{j}\right|^{2} \tag{58}
\end{equation*}
$$

When the S-matrix element is unitary, no flux is absorbed and the transmission coefficient is zero. When absorption is complete, the transmission coefficient is one. These quantities are essential for calculating statistical reaction cross sections. Quite often, optical model calculations are a mere preliminary to statistical model calculations and are performed only to obtain the transmission coefficients.

The total flux lost in the scattering is related to the reaction cross section through the equation

$$
\begin{equation*}
\sigma_{r}=-\frac{1}{v} \oint \vec{\jmath} \cdot d \vec{a} \tag{59}
\end{equation*}
$$

where it is understood that the probability current,

$$
\begin{equation*}
\vec{\jmath}=\frac{\hbar}{2 i \mu}\left(\Psi^{\dagger} \nabla \Psi-\left(\nabla \Psi^{\dagger}\right) \Psi\right) \tag{60}
\end{equation*}
$$

is integrated over a surface which tends to infinity. The reaction cross section can be expressed in terms of the transmission coefficients as

$$
\begin{equation*}
\sigma_{r}=\frac{1}{2 s+1} \frac{\pi}{k^{2}} \sum_{l j}(2 j+1) T_{l}^{j} \tag{61}
\end{equation*}
$$

For charged particles, integration of the differential elastic cross section of Eq. (54) leads to an infinite result, due to the infinite range of the Coulomb interaction. For neutrons, it yields the elastic cross section,

$$
\begin{equation*}
\sigma_{e l}=\int d \Omega \frac{d \sigma}{d \Omega}=\frac{\pi}{2 k^{2}} \sum_{l j}(2 j+1)\left|1-S_{l}^{j}\right|^{2} \tag{62}
\end{equation*}
$$

This is often called the shape elastic cross section to distinguish it form the compound elastic one.

For neutrons, a total cross section can also be defined as the sum of the elastic and reaction cross sections,

$$
\begin{equation*}
\sigma_{t o t}=\sigma_{e l}+\sigma_{r}=\frac{\pi}{k^{2}} \sum_{l j}(2 j+1)\left(1-\operatorname{Re} S_{l}^{j}\right) \tag{63}
\end{equation*}
$$

The total cross section takes into account all flux lost from the incident plane wave, either by scattering or by absorption. Comparing the expression for the total cross section with that of the


Figure 6: $n+{ }^{58}$ Ni total cross section, identified by their EXFOR access numbers, are shown together with two optical model calculations.
scattering amplitude, $A(\theta)$, one sees that the optical theorem is explicitly verified by the partial wave expansion,

$$
\begin{equation*}
\sigma_{t o t}=\frac{4 \pi}{k} \operatorname{Im} A\left(\theta=0^{\circ}\right) \tag{64}
\end{equation*}
$$

As observed earlier, when it exists, the total optical cross section is the average of an amplitude and can thus be compared directly with the energy-averaged experimental data. This is done in Fig. 6, where a selection of the experimental measurements of the $n+{ }^{58} N i$ total cross section is shown together with optical model calculations using the parameters of A. Prince[17] and those used in the exercises. Although there is a great deal of dispersion in the low energy data, the calculations succeed in following its trend.


Figure 7: $n+{ }^{58} N i$ elastic cross section, identified by their EXFOR access numbers, are shown together with two optical model calculations.

The optical elastic and reaction cross sections involve the average of a squared amplitude and cannot be compared directly with the energy-averaged experimental data. The compound elastic cross section is part of the optical reaction cross section rather than the elastic cross section.

The experimental elastic cross section can thus greatly exceed the optical component of the cross section. This is illustrated in Fig. 7, in which a selection of the experimental measurements of $n+{ }^{58} N i$ are compared to optical model calculations using the parameters of Prince[17] and of the exercises. At energies sufficiently high for the elastic compound nucleus cross section to have dropped to zero (which usually occurs at an energy of the order of a few MeV ), the differential and integral (when it exists) optical elastic cross sections can be compared with the energy-averaged experimental data. Note that the elastic cross section for neutron-induced scattering can also be compared to the experimental data at extremely low incident energies, where it is customarily expressed as a scattering radius $R^{\prime}$.

At high energies, the reaction cross section can also be compared to experimental data. However, the reaction cross section cannot be measured directly, making the data for such a comparison scarce.

## 6 The generalized optical potential

The single-channel or spherical optical model treates the target nucleus as if it were spherical. But nuclei are often deformed. An all nuclei, whether spherical or deformed, are susceptible to shape oscillations. Most deformed and many spherical nuclei possess low-lying collective states that are easily excited in a collision. As these excitations are prompt reaction modes, one would expect their description to lie within the scope of a generalized optical model. The standard extension of the optical model takes into account the expected deviation from spherical symmetry by modifying the radii $R_{i}$ of the terms in the optical model potential.

A vibrational nucleus possesses a spherically symmetric ground state. Its excited states undergo shape oscillations about the spherical equilibrium mode.[20] To take these into account, the radii of the terms in the potential are expressed as

$$
\begin{align*}
R_{i} & =R_{0 i}\left(1+\sum_{\lambda \mu} a_{\lambda \mu} Y_{\lambda \mu}(\hat{r})\right)  \tag{65}\\
& =R_{0 i}\left(1+\sum_{\lambda} \frac{\beta_{\lambda}}{\sqrt{2 \lambda+1}} \sum_{\mu}\left(b_{\lambda \mu}^{\dagger}+(-)^{\mu} b_{\lambda-\mu}\right) Y_{\lambda \mu}(\hat{r})\right)
\end{align*}
$$

where the $b_{\lambda \mu}^{\dagger}$ and $b_{\lambda \mu}$ are the creation and annihilation operators of nuclear phonons and the $\beta_{\lambda}$ are the amplitudes of their respective shape oscillations. One usually expands the optical potential to first or second order in the creation and annihilation operators,

$$
\begin{align*}
U_{o p t}(r, \hat{r})=U_{o p t}(r) & +\sum_{i} \frac{\partial U_{o p t}}{\partial R_{i}} R_{0 i} \sum_{\lambda \mu} a_{\lambda \mu} Y_{\lambda \mu}(\hat{r})  \tag{66}\\
& +\sum_{i} \frac{\partial^{2} U_{o p t}}{\partial R_{i}^{2}} R_{0 i}^{2}\left(\sum_{\lambda \mu} a_{\lambda \mu} Y_{\lambda \mu}(\hat{r})\right)^{2}
\end{align*}
$$

thereby taking into account the direct excitation of one- and two-phonon states. The vibrational model including one-phonon states is called the first-order vibrational model, while that
containing the two-photon states as well is known as the second-order model.
The nucleus ${ }^{58} \mathrm{Ni}$ serves as an example of a typical vibrational nucleus. Two neutrons from a doubly magic configuration, it has a spherically symmetric $J=0^{+}$ground state and a $J=2^{+}$ excited state at $E_{x}=1.454 \mathrm{MeV}$ that can be considered a one-quadrupole-phonon vibrational state. At about twice the energy of the one-phonon state, in particular, at $E_{x}=2.459,2.776$, and 2.943 MeV , one finds a trio of states with $J=4^{+}, 2^{+}$, and $0^{+}$, respectively, which can be interpreted as the two-phonon states. The fact that the first two these (but not the third) decay almost exclusively to the $J=2^{+}$excited state corroborates such an interpretation, but also shows its limitations.

A statically deformed nucleus possesses rotational excited states.[20] When the nucleus possesses axial symmetry, the radii are replaced by

$$
\begin{equation*}
R_{i}\left(\theta^{\prime}\right)=R_{0 i}\left(1+\sum_{\lambda} \beta_{\lambda} Y_{\lambda 0}\left(\theta^{\prime}\right)\right) \tag{67}
\end{equation*}
$$

where $\beta_{\lambda}$ is the static deformation of mutipolarity $\lambda$ and the angle $\theta^{\prime}$ is in the body-fixed frame. This substitution could be extended to the general case of triaxial nuclei without too much difficulty (at least at this point). The model is then called the assymetric rotational model in contrast to the (axially) symmetric one. In either case, the potential obviously depends on the orientation of the principal axes of the target.

When the deformation of the nucleus is large, expansion of the potential in a Taylor series is not a good approximation. It is better to expand it directly in multipoles as

$$
\begin{equation*}
U_{o p t}\left(r, \hat{r}^{\prime}\right)=\sum_{\lambda} U_{\lambda}(r) Y_{\lambda 0}\left(\hat{r}^{\prime}\right) \tag{68}
\end{equation*}
$$

where the multipole potentials are obtained as

$$
\begin{equation*}
U_{\lambda}(r)=\int d \Omega^{\prime} U_{o p t}\left(r, \theta^{\prime}\right) Y_{\lambda 0}\left(\theta^{\prime}\right) \tag{69}
\end{equation*}
$$

In the body-fixed frame, the moments $U_{\lambda \mu}(r)$ with $\mu \neq 0$ vanish. The body-fixed angles $\hat{r}^{\prime}$ are related to the space-fixed ones $\hat{r}$ by a rotation through the angles that define the orientation of the nucleus, which are the collective angular coordinates of the nucleus, $\hat{r}_{i n t}$. For the spherical harmonics, this implies that

$$
\begin{equation*}
Y_{\lambda 0}\left(\hat{r}^{\prime}\right)=\sum_{\mu} Y_{\lambda \mu}(\hat{r}) D_{\mu 0}^{\lambda}\left(\hat{r}_{i n t}\right)=\sum_{\mu} Y_{\lambda \mu}(\hat{r}) Y_{\lambda \mu}^{*}\left(\hat{r}_{i n t}\right) \tag{70}
\end{equation*}
$$

where the $D_{\mu \mu^{\prime}}^{\lambda}$ are rotation matrices with the special value for $\mu^{\prime}=0$ used in the last equality. The optical potential in the rotational model can thus be decomposed as

$$
\begin{equation*}
U_{o p t}\left(\vec{r}, \hat{r}_{i n t}\right)=\sum_{\lambda \mu} U_{\lambda}(r) Y_{\lambda \mu}(\hat{r}) Y_{\lambda \mu}^{*}\left(\hat{r}_{i n t}\right) \tag{71}
\end{equation*}
$$

The nucleus ${ }^{238} \mathrm{U}$ provides an excellent example of a statically deformed nucleus with rotational excitations. Its $0^{+}$ground state possesses static quadrupolar e hexadecapolar deformations
with $\beta_{2}=0.198$ and $\beta_{4}=0.057$. Its first four excited states, at $E_{x}-J^{\pi}=0.044 \mathrm{MeV}-2^{+}, 0.148$ $\mathrm{MeV}-4^{+}, 0.307 \mathrm{MeV}-6^{+}$, and $0.518 \mathrm{MeV}-8^{+}$, iniciate a rotational band that can be traced to at least the $28^{+}$state at $E_{x}=4.516 \mathrm{MeV}$. Each of these states decays exclusively to the next lower state in the rotational band.

Morel elaborate couplings between projectile and target can also be described through appropriate generalizations of the optical potential. One of these was mentioned above - the asymmetric rotational model - obtained by taking into account triaxial deformations of the nucleus. Another is the vibrational-rotational model which couples static deformation to dynamic shape oscillations. More details about these can be found in Refs. [20, 18, 19].

In the generalizations of the optical potential discussed in this section, the introduction of target degrees of freedom leads to a potential that depends on the relative orientation of the target with respect to the projectile. The system is no longer invariant under independent rotations of the target or the projectile and their individual angular momenta are not conserved. However, in all cases, the system continues invariant under a simultaneous rotation of the projectile and target. The total angular momentum thus continues to be a conserved quantity.

## 7 Partial wave expansion in the coupled-channels optical model

The partial wave expansion proceeds in the coupled-channels optical model much as it did in the single-channel one. There are several new features however. The first of these is that the excited states and their angular momentum must now be taken into account.[16, 18, 19]

To include the angular momentum of the target, the spin-angular functions are coupled to the target states to form target-spin-angular functions,

$$
\begin{equation*}
\mathcal{X}_{l s j c}^{J M}(\hat{r})=\sum_{n N_{c}}\left\langle j n I_{c} N_{c} \mid J M\right\rangle \mathcal{Y}_{l s}^{j n}(\hat{r})\left|I_{c} N_{c}\right\rangle, \tag{72}
\end{equation*}
$$

where the $\left|I_{c} N_{c}\right\rangle$ represent the target states with total angular momentum and angular momentum projection $I_{c}$ and $N_{c}$, respectively. In the vibrational model, these are the one- and two-phonon states,

$$
\begin{equation*}
\left|I_{c} N_{c}\right\rangle=b_{I_{c} N_{c}}^{\dagger}|0\rangle \quad \text { and } \quad\left|I_{c} N_{c}\right\rangle=\frac{1}{\sqrt{1+\delta_{I_{1} I_{2}}}}\left[b_{I_{1}}^{\dagger} b_{I_{2}}^{\dagger}\right]_{I_{c} N_{c}}|0\rangle . \tag{73}
\end{equation*}
$$

In the axial symmetric rotational model, they are the rotational states, which can be written in terms of the rotation matrices $D_{N K}^{I}$ as

$$
\begin{align*}
\left\langle\vec{r}_{i n t} \mid I_{c} N_{c}\right\rangle & =\frac{1}{\sqrt{1+\delta_{K 0}}} \sqrt{\frac{2 I_{c}+1}{16 \pi^{2}}}  \tag{74}\\
& \times\left[\chi_{K}\left(\vec{r}_{i n t}\right) D_{N_{c} K}^{I_{c} *}\left(\hat{r}_{i n t}\right)+(-)^{I_{c}-I_{\chi}} \chi_{-K}\left(\vec{r}_{i n t}\right) D_{N_{c},-K}^{I_{c} *}\left(\hat{r}_{i n t}\right)\right]
\end{align*}
$$

where $\chi_{ \pm K}$ is the internal wave function of the rotational band, $I_{\chi}$ its internal angular momentum and $K$ the projection of the total internal angular momentum on the symmetry axis.

The scattering wave function is expanded in a sum over both the excited states and the angular momenta. The expansion can be written as

$$
\begin{equation*}
\Psi=4 \pi \sum_{\substack{l j c c M \\ l^{\prime} j^{\prime} c^{\prime}}} \mathcal{y}_{l^{\prime} s j^{\prime} c^{\prime}}^{J M}(\hat{r}) i^{l^{\prime}} \psi_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r) \frac{e^{i \sigma_{l c}}}{k_{c} r} \mathcal{Y}_{l s j c}^{J M \dagger}(\hat{k}) \tag{75}
\end{equation*}
$$

Factors of $i^{\prime}$ and $e^{i \sigma_{c}} / k_{c} r$, where $k_{c}$ is the wave number in channel $c$, have again been extracted from the wave functions to simplify later manipulations. The Coulomb phase now depends on the channel energy, through $k_{c}$, as well as on the angular momentum.

In the deformed optical model, the wave functions, $\psi_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r)$, couple different values of the angular momenta $l$ and $j$ and different channels $c$ for each value of the total angular momentum $J$. This contrasts with the single-channel optical model, in which the wave function for total angular momentum $j$ also possesses a well defined orbital angular momentum $l$. This is due to the fact that the interaction couples the different partial waves in the deformed model, while it does not do so in the spherical one. (If it did, a partial wave expansion in terms of the wave functions $\psi_{l^{\prime} l}^{j}$ in Eq. (42) would, in general, have been necessary.)

When the partial-wave expansion of the wave function is substituted in the optical Schrödinger equation, the latter can be reduced to a set of coupled equations for each value of the total angular momentum $J$. These are

$$
\begin{align*}
\frac{\hbar^{2}}{2 \mu}\left\{\frac{d^{2}}{d r^{2}}-\frac{l^{\prime}\left(l^{\prime}+1\right)}{r^{2}}\right. & \left.+k_{c^{\prime}}^{2}\right\} \psi_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r)  \tag{76}\\
& -\sum_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}} \mathcal{U}_{l^{\prime} j^{\prime} c^{\prime} c^{\prime}, l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}}^{J}(r) \psi_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}, l j c}^{J}(r)=0
\end{align*}
$$

where the potential matrix elements are calculated with respect to the orthonormal target-spinangular functions,

$$
\begin{equation*}
\mathcal{U}_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r)=\int d^{3} r_{i n t} d \Omega \mathcal{Y}_{l^{\prime} s j^{\prime} c^{\prime}}^{J M \dagger}(\hat{r}) U_{o p t}\left(\vec{r}, \vec{r}_{i n t}\right) \mathcal{l}_{l s j c}^{J M}(\hat{r}) . \tag{77}
\end{equation*}
$$

Although the target-spin-angular functions used to calculate these matrix elements have a welldefined projection $M$ of the the total angular momentum $J$, the matrix elements that result are independent of this value if the system is rotationally invariant. When the system is invariant under time-reversal, the potential matrix is also symmetric under interchange of the primed and unprimed indices.

By writing the matrix elements as matrices,

$$
\begin{align*}
l^{\prime} \delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c} \rightarrow L_{J}, & k_{c^{\prime}} \delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c} \rightarrow K_{J},  \tag{78}\\
\psi_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r) \rightarrow \Psi_{J}(r), & \mathcal{U}_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r) \rightarrow U_{J}(r),
\end{align*}
$$

the Schrödinger equation can be recast in a more familiar form as a matrix equation,

$$
\begin{equation*}
\left\{\frac{d^{2}}{d r^{2}}-\frac{L_{J}\left(L_{J}+1\right)}{r^{2}}+K_{J}^{2}-\frac{2 \mu}{\hbar^{2}} U_{J}(r)\right\} \Psi_{J}(r)=0 . \tag{79}
\end{equation*}
$$

The incoming-wave boundary condition again requires that asymptotically the wave function take the form of an incoming plane wave and an outgoing scattering wave. Here this takes the form

$$
\begin{align*}
\Psi \rightarrow \sum_{c} \exp \left(i \vec{k}_{c} \cdot \vec{r}+i \eta_{c} \log \left(k_{c} r\right.\right. & \left.\left.-\vec{k}_{c} \cdot \vec{r}\right)\right) \sum_{\nu}\left|s \nu I_{c} N_{c}\right\rangle\left\langle s \nu I_{c} N_{c}\right| \\
+\frac{1}{r} \sum_{c c^{\prime}} \exp \left(i k_{c^{\prime}} r\right. & \left.-i \eta_{c^{\prime}} \log \left(2 k_{c^{\prime}} r\right)\right)  \tag{80}\\
& \times \sum_{\nu \nu^{\prime}}\left|s \nu^{\prime} I_{c^{\prime}} N_{c^{\prime}}\right\rangle \bar{f}_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta)\left\langle s \nu I_{c} N_{c}\right|,
\end{align*}
$$

where the $\bar{f}_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta)$ are the target and projectile spin-projected matrix elements of the elastic scattering amplitude.

The asymptotic form expected of the wave function in the partial wave of total angular momentum $J$ can be most easily expressed using an obvious extension of the matrix notation above. To be consistent with Eq. (80) and satisfy the differential equation, the matrix wave function $\Psi_{J}$ must have for its asymptotic form,

$$
\begin{align*}
\Psi_{J}(r) \rightarrow F_{J}(r) & +\left(G_{J}(r)+i F_{J}(r)\right) \bar{C}_{J}  \tag{81}\\
& =\frac{i}{2}\left(H_{J}^{-}(r)-H_{J}^{+}(r) e^{i \sigma_{J}} \bar{S}_{J} e^{i \sigma_{J}}\right) e^{-i \sigma_{J}}
\end{align*}
$$

where $\bar{C}_{J}=\left(\bar{S}_{J}-1_{J}\right) / 2 i, F_{J}$ and $G_{J}$ are the regular and irregular Coulomb wave functions in (diagonal) matrix form, respectively, and $H_{J}^{ \pm}=e^{\mp i \sigma_{J}}\left(G_{J} \pm i F_{J}\right)$ are the linear combinations of these that asymptotically contain only incoming $\left(H_{J}^{-}\right)$or outgoing $\left(H_{J}^{+}\right)$waves. $\bar{S}_{J}$ is the nuclear part of the S-matrix element. One can loosely interpret the half of the Coulomb phase shift $e^{i \sigma_{J}}$ that precedes the nuclear S-matrix in Eq. (81) as the Coulomb deflection accumulated on the incoming half of the 'trajectory,' with the half of the Coulomb phase shift following the nuclear S-matrix then being the Coulomb deflection of the outgoing half of the 'trajectory'.

The S-matrix elements, $S_{J}$, can be obtained by an obvious extension of the method used for the spherical optical model. In the internal region, the differential equation for each partial wave, Eq. (79) is solved numerically out to the radius, $r_{m}$. The numerical solution and its derivative are matched there to the wave function in the external region, given by Eq. (81), and to its derivative, to obtain the amplitudes in the internal region, $a_{J}$, and the S -matrix elements, $S_{J}$. This is repeated for increasing values of $J$ until the value of $S_{J}$ which results is sufficiently close to one.

The important difference between the deformed optical model and the spherical model is that the wave function in partial wave $J$ is not a scalar, as it is in the spherical model, but a matrix. The differential equation that must be solved is also a matrix one. Although the only solution that is normally of interest is the one in which the target is in its ground state in the incoming wave, the complete matrix solution is needed to invert the matching equations and obtain the S-matrix. The calculation is thus much more time consuming than in the spherical case.

To obtain the partial wave expansion of the scattering amplitude, one repeats the procedure used earlier: substitute the asymptotic form of the partial wave function, $\Psi_{J}$ of Eq. (81), in
the partial wave expansion of the total wave function, Eq. (75), and compare the result to the expected form of the asymptotic wave function, Eq. (80). One then finds

$$
\begin{equation*}
\bar{f}(\theta)=\frac{4 \pi}{2 i} \sum_{\substack{l j c J M \\ l^{\prime} j^{\prime} c^{\prime}}}\left(e^{i \sigma_{l^{\prime} c^{\prime}}} \bar{S}_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J} e^{i \sigma_{l c}}-\delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c}\right) \frac{1}{k_{c}} \mathcal{Y}_{l^{\prime} s j^{\prime} c^{\prime}}^{J N}(\hat{r}) \mathcal{Y}_{l s j c}^{J N \dagger}(\hat{k}) \tag{82}
\end{equation*}
$$

Its target and projectile spin-projected matrix elements are

$$
\begin{align*}
\bar{f}_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta)=\frac{4 \pi}{2 i} \sum_{\begin{array}{c}
l j c l^{\prime} j^{\prime} c^{\prime} \\
J M m m^{\prime} n n^{\prime}
\end{array}} & \left(e^{i \sigma_{l^{\prime} c^{\prime}}} \bar{S}_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J} e^{i \sigma_{l c}}-\delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c}\right) \frac{1}{k_{c}} \\
& \times Y_{l^{\prime} m^{\prime}}(\hat{r}) Y_{l m}^{*}(\hat{k})\left\langle l^{\prime} m^{\prime} s \nu^{\prime} \mid j^{\prime} n^{\prime}\right\rangle  \tag{83}\\
& \times\left\langle j^{\prime} n^{\prime} I_{c^{\prime}} N_{c^{\prime}} \mid J M\right\rangle\left\langle J M \mid j n I_{c} N_{c}\right\rangle\langle j n \mid l m s \nu\rangle .
\end{align*}
$$

In the scattering problem considered here, the flux, not the density, is conserved. When a state is excited, the energy that goes to excitation must be taken from the relative motion. The relative velocity thus decreases, as does the flux. To take this into account, the S-matrix and scattering amplitude must be renormalized as

$$
\begin{equation*}
S_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}=\sqrt{\frac{k_{c^{\prime}}}{k_{c}}} \bar{S}_{l^{\prime} j^{\prime} c^{\prime} l j c}^{J} \quad \text { or } \quad S_{J}=K_{J}^{1 / 2} \bar{S}_{J} K_{J}^{-1 / 2} \tag{84}
\end{equation*}
$$

and

$$
\begin{equation*}
f_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta)=\sqrt{\frac{k_{c^{\prime}}}{k_{c}}} \bar{f}_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta) . \tag{85}
\end{equation*}
$$

When the system is time-reversal invariant, the matrix $S_{J}$ is symmetric.
The scattering amplitude with the Coulomb term extracted thus has the form

$$
\begin{align*}
f_{\nu^{\prime} N_{c^{\prime}} \nu N_{c}}(\theta)= & \delta_{\nu^{\prime} \nu} \delta_{N_{c^{\prime}}} N_{c} f_{C c}(\theta) \\
+\frac{4 \pi}{2 i} \sum_{\substack{l j c^{\prime} j^{\prime} c^{\prime} \\
J M m m^{\prime} n n^{\prime}}} & e^{i \sigma_{l^{\prime} c^{\prime}}}\left(S_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}-\delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c}\right) e^{i \sigma_{l c}} \frac{1}{k_{c}} \\
& \times Y_{l^{\prime} m^{\prime}}(\hat{r}) Y_{l m}^{*}(\hat{k})\left\langle l^{\prime} m^{\prime} s \nu^{\prime} \mid j^{\prime} n^{\prime}\right\rangle  \tag{86}\\
& \times\left\langle j^{\prime} n^{\prime} I_{c^{\prime}} N_{c^{\prime}} \mid J M\right\rangle\left\langle J M \mid j n I_{c} N_{c}\right\rangle\langle j n \mid l m s \nu\rangle .
\end{align*}
$$

where

$$
\begin{equation*}
f_{C c}(\theta)=-\frac{\eta_{c}}{2 k_{c} \sin ^{2} \theta / 2} \exp \left[-i \eta_{c} \log \left(\sin ^{2} \theta / 2\right)+2 i \sigma_{0 c}\right] \tag{87}
\end{equation*}
$$

is the Coulomb amplitude in channel $c$.
Once the scattering amplitude is known, calculating cross sections is a simple matter. The differential cross sections for an unpolarized incident beam and target are obtained by averaging the squared magnitude of the scattering amplitudes over the initial values of the projectile and target spin and summing over the final values. The differential elastic cross section for a collision in which the target is initially in its ground state is then given by

$$
\begin{equation*}
\frac{d \sigma_{e l}}{d \Omega}=\frac{1}{(2 s+1)\left(2 I_{0}+1\right)} \sum_{\substack{\nu^{\prime} N_{0}^{\prime} \\ \nu N_{0}}}\left|f_{\nu^{\prime} N_{0}^{\prime} \nu N_{0}}(\theta)\right|^{2} \tag{88}
\end{equation*}
$$

The differential inelastic cross section for inelastic scattering to an excited state $c$ can be written similarly as

$$
\begin{equation*}
\frac{d \sigma_{c}}{d \Omega}=\frac{1}{(2 s+1)\left(2 I_{0}+1\right)} \sum_{\substack{\nu^{\prime} N_{c}^{\prime} \\ \nu N_{0}}}\left|f_{\nu^{\prime} N_{\mathrm{c}}^{\prime} \nu N_{0}}(\theta)\right|^{2}, \tag{89}
\end{equation*}
$$

where it should be emphasized that the sum over $N_{c}^{\prime}$ refers to a sum over the spin projections of the final state $c$ only.

Due to the infinite range of the Coulomb force, the integrated elastic cross section is finite only when at least one of the two colliding particles is neutral. In the particular case of neutrons incident on a nucleus, integration of the differential cross section of Eq. (88) yields.

$$
\begin{equation*}
\sigma_{e l}=\frac{1}{2\left(2 I_{0}+1\right)} \frac{\pi}{k^{2}} \sum_{l^{\prime} j_{J}^{\prime} l j}(2 J+1)\left|S_{l^{\prime} j^{\prime} c_{0}, l j c_{0}}^{J}-\delta_{l^{\prime} l} \delta_{j^{\prime} j}\right|^{2} \tag{90}
\end{equation*}
$$

The integrated inelastic cross sections exist for for both neutral and charged particles. They take the form

$$
\begin{equation*}
\sigma_{c}=\frac{1}{(2 s+1)\left(2 I_{0}+1\right)} \frac{\pi}{k^{2}} \sum_{l^{\prime} j_{J}^{\prime} l_{j}}(2 J+1)\left|S_{l^{\prime} j^{\prime} c, l j c_{0}}^{J}\right|^{2} \quad c \neq c_{0} \tag{91}
\end{equation*}
$$

Just as in the single-channel problem, the total flux lost from the elastic channel can be related to the reaction cross section through the equation

$$
\begin{equation*}
\sigma_{r}=-\frac{1}{v} \oint \overrightarrow{\jmath_{0}} \cdot d \vec{a}, \tag{92}
\end{equation*}
$$

where the probability current,

$$
\begin{equation*}
\overrightarrow{\jmath_{0}}=\frac{\hbar}{2 i \mu}\left(\Psi_{0}^{\dagger} \nabla \Psi_{0}-\left(\nabla \Psi_{0}\right)^{\dagger} \Psi_{0}\right) \tag{93}
\end{equation*}
$$

is integrated over a surface which tends to infinity, with $\Psi_{0}$ being the ground-state component of the wave function. However, in the coupled-channel problem, it is also possible to define an absorption cross section, which can be related to the total flux lost from all channels, elastic and inelastic, as

$$
\begin{equation*}
\sigma_{a b s}=-\frac{1}{v} \oint \sum_{c} \overrightarrow{\jmath_{c}} \cdot d \vec{a} \tag{94}
\end{equation*}
$$

where the probability current in channel $c, \vec{\jmath}_{c}$, is

$$
\begin{equation*}
\vec{\jmath}_{c}=\frac{\hbar}{2 i \mu}\left(\Psi_{c}^{\dagger} \nabla \Psi_{c}-\left(\nabla \Psi_{c}\right)^{\dagger} \Psi_{c}\right) \tag{95}
\end{equation*}
$$

with $\Psi_{c}$ the component of the wave function that asymptotically occupies state $c$.
Using the asymptotic form of the partial waves, Eq. (81), the expression for the reaction cross section can be reduced to

$$
\begin{equation*}
\sigma_{r}=\frac{1}{(2 s+1)\left(2 I_{0}+1\right)} \frac{\pi}{k^{2}} \sum_{\substack{l^{\prime} j^{\prime} l j \\ c J}}(2 J+1)\left(\delta_{l^{\prime} l} \delta_{j^{\prime} j}-\left|S_{l^{\prime} j^{\prime} c_{0}, l j c_{0}}^{J}\right|^{2}\right) \tag{96}
\end{equation*}
$$

The contribution of each partial wave to the reaction cross section is determined by the fraction of the flux lost from the elastic channel. The absorption cross section can, of course, be reduced to a similiar form, which can be written as

$$
\begin{equation*}
\sigma_{a b s}=\frac{1}{(2 s+1)\left(2 I_{0}+1\right)} \frac{\pi}{k^{2}} \sum_{l j J}(2 J+1) T_{l j c_{0}, l j c_{0}}^{J} \tag{97}
\end{equation*}
$$

where the coupled-channels transmission coefficients have been introduced. These are defined as

$$
\begin{equation*}
T_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}=\delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c}-\sum_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}} S_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}, l^{\prime} j^{\prime} c^{\prime} s^{\prime}}^{S_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}, l j c}^{J}} \tag{98}
\end{equation*}
$$

The similarity of the transmission coefficients to the single-channel ones becomes clear when they are written in matrix form. The transmission matrix in partial wave $J$ can be written in terms of the corresponding S-matrix as

$$
\begin{equation*}
T_{J}=1_{J}-S_{J}^{\dagger} S_{J} \tag{99}
\end{equation*}
$$

Comparison of the form of the reaction and absorption cross sections reveals a simple relation between the two,

$$
\begin{equation*}
\sigma_{r}=\sigma_{a b s}+\sum_{c \neq c_{0}} \sigma_{c} \tag{100}
\end{equation*}
$$

In other words, the elastic channel loses flux to both the prompt inelastic channels and the long-lived compound states. The reaction cross section takes both of these into account.

The absorption cross section and the corresponding transmission coefficients characterize the transition of flux from the prompt channels to the compound states. These are the quantities of principal interest for compound-nucleus calculations. When using coupled-channels transmission coefficients in compound-nucleus calculations, it is quite common to use just the diagonal elements of the transmission matrix and neglect the off-diagonal ones. A careful analysis by Engelbrecht and Weidenmuiller [22] showed that a more correct procedure is to perform the compound-nucleus calculation in a basis in which the transmission coefficients are diagonal and transform the resulting cross sections back to the non-diagonal basis.

For neutral particles, the neutron in particular, the elastic cross section is finite. A total cross section can then be defined as the sum of the elastic and reaction cross sections,

$$
\begin{equation*}
\sigma_{t o t}=\sigma_{e l}+\sigma_{r}=\frac{1}{2 I_{0}+1} \frac{\pi}{k^{2}} \sum_{l j J}(2 J+1)\left(1-\operatorname{Re} S_{l j c_{0}, l j c_{0}}^{J}\right) . \tag{101}
\end{equation*}
$$

The total cross section takes into account the occurence of scattering of any type. It is a measure of the flux lost from the incident plane wave state.

Just as in the case of the elastic cross section, care must be taken when comparing inelastic optical model cross sections with experimental data. At low energies, these cross sections are dominated by their compound nucleus contribution, as shown in Figs. 8 and 9, for neutroninduced excitation of the first excited state in ${ }^{58} \mathrm{Ni}$ and ${ }^{238} \mathrm{U}$, respectively. One observes that the direct process plays a very minor role in the excitation of these states in the first few MeV above threshhold. In Fig. 8, the ${ }^{58} \mathrm{Ni}$ data are compared to optical model calculations using


Figure 8: $n+{ }^{58}$ Ni $E_{x}=1.454 \mathrm{MeV} 2^{+}$inelastic cross section, identified by their EXFOR access numbers, are shown together with two optical model calculations.
the parameters of A. Prince[17] and those of the exercises, both with a phonon amplitude of $\beta_{2}=0.2$. Note the strong influence of the optical model parameters on the direct component of the inelastic ${ }^{58} \mathrm{Ni}$ excitation. The Prince parameters yield an inelastic cross section that is almost twice that of the parameters of the exercises, although both use the same phonon amplitude. The ${ }^{238} \mathrm{U}$ data of Fig. 9 is compared to an optical model calculation using the parameters of Young and Arthur[21], which fits the higher energy data quite well. One notes that the direct excitation cross section of the ${ }^{238} \mathrm{U} 2^{+}$state reaches a value of almost 500 mb . In general, the inelastic excitation of a rotational band can be quite large, demanding a coupled channels method for its precise calculation.


Figure 9: $n+{ }^{238} U E_{x}=0.044 \mathrm{MeV} 2^{+}$inelastic cross section, identified by their EXFOR access numbers, are shown together with an optical model calculation.

## 8 The distorted-wave Born approximation

The distorted-wave Born approximations (DWBA) can be understood as a simple iterative expansion of the Lippmann-Schwinger equation in powers of the potential. It is thus a good approximation when the coupling to the excited states is weak. The Lippmann-Schwinger equation the integral representation of the wave equation, which can be written as

$$
\begin{equation*}
\Psi=\Psi_{0}^{+}+G_{0}^{+} U^{\prime} \Psi \tag{102}
\end{equation*}
$$

where $G_{0}^{+}$is the Green's function of the Schrödinger equation containing the part of the optical potential that does not couple states, having as its wave function $\Psi_{0}^{+}$, and the + superscript means that the wave function $\Psi_{0}^{+}$satisfies incoming-wave bounday conditions while the propagator $G_{0}^{+}$asymptotically contains only outgoing waves.

In the single-channel optical model problem, one can define incoming- and outgoing-wave solutions, $h_{l c}^{j \pm}(r)$, of the wave equation in each channel,

$$
\begin{equation*}
\left\{\frac{d^{2}}{d r^{2}}-\frac{l(l+1)}{r^{2}}+k_{c}^{2}-\frac{2 \mu}{\hbar^{2}}\left(U_{c e n, c}(r)+d_{l}^{j} U_{s o, c}(r)\right)\right\} h_{l c}^{j \pm}(r)=0, \tag{103}
\end{equation*}
$$

where the spin-orbit constant $d_{l}^{j}=d_{s o}(j(j+1)-l(l+1)-s(s+1)) / 2$ is as before. Asymptotically, these solutions have the same behavior as the incoming and outgoing Coulomb waves,

$$
\begin{equation*}
h_{l c}^{j \pm}(r) \rightarrow H_{l c}^{ \pm}(r)=e^{\mp i \sigma_{l c}}\left(G_{l c}(r) \pm i F_{l c}(r)\right) . \tag{104}
\end{equation*}
$$

They are, however, solutions to the optical Schrödinger equation at all values of $r$. Numerically, they can be obtained by solving the differential equation inward from the matching point, using the conditions for matching to the asymptotic Coulomb functions as the initial conditions.

The incoming- and outgoing- wave solutions to the optical Schrödinger equation are not regular at the origin. Through a comparison with the asymptotic form given in Eq. (46), it is easy to convince oneself that a linear combination of the two that is regular is given in terms of the S -matrix as

$$
\begin{align*}
\psi_{l c}^{j+}(r) & =\frac{i}{2}\left(h_{l c}^{j-}(r)-h_{l c}^{j+}(r) e^{2 i \sigma_{l c}} S_{0 l c}^{j}\right) \\
& =\psi_{l c}^{j}(r) e^{i \sigma_{l c}}, \tag{105}
\end{align*}
$$

where the last equality simply makes note of the relationship between the wave function $\psi_{l c}^{j+}$ and the single-channel wave function $\psi_{l c}^{j}$ of the partial wave expansion in Eq. (42). The S-matrix has been relabelled $S_{0}$ to emphasize its relation to the single-channel problem as well.

The single-channel Green's function in channel $c$ can be decomposed in partial waves as

$$
\begin{equation*}
G_{0 c}^{+}\left(\vec{r}, \vec{r}^{\prime}\right)=\frac{1}{r r^{\prime}} \sum_{l j n} \mathcal{Y}_{l s}^{j n}(\hat{r}) g_{l c}^{j+}\left(r, r^{\prime}\right) \mathcal{Y}_{l s}^{j n \dagger}\left(\hat{r}^{\prime}\right), \tag{106}
\end{equation*}
$$

where the partial-wave Green's functions are defined in terms of the regular and outgoing partial wave solutions as

$$
\begin{equation*}
g_{l c}^{j+}\left(r, r^{\prime}\right)=-\frac{2 \mu}{\hbar^{2} k_{c}} \psi_{l c}^{j+}\left(r_{<}\right) h_{l c}^{j+}\left(r_{>}\right) . \tag{107}
\end{equation*}
$$

The channel Green's functions can then be combined into the complete single-channel Green's function appropriate to the coupled-channels problem,

$$
\begin{align*}
G_{0}^{+}\left(\vec{r}, \vec{r}^{\prime}\right) & =\sum_{c N_{c}} G_{0 c}^{+}\left(\vec{r}, \vec{r}^{\prime}\right)\left|I_{c} N_{c}\right\rangle\left\langle I_{c} N_{c}\right|  \tag{108}\\
& =\frac{1}{r r^{\prime}} \sum_{\substack{l j c \\
J M}} \mathcal{Y}_{l s j}^{J M}(\hat{r}) g_{l c}^{j+}\left(r, r^{\prime}\right) \mathcal{Y}_{l s j}^{J M \dagger}\left(\hat{r}^{\prime}\right)
\end{align*}
$$

When the partial wave expansions, Eqs. (75) and (108), are substituted in the LippmannSchwinger equation, Eq. (102), it can be reduced to a set of coupled equations for each partial wave,

$$
\begin{align*}
\psi_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r)= & \psi_{l c}^{j}(r) \delta_{l^{\prime} l} \delta_{j^{\prime} j} \delta_{c^{\prime} c}  \tag{109}\\
& +\int_{0}^{\infty} d r^{\prime} g_{l^{\prime} c^{\prime}}^{j^{\prime}+}\left(r, r^{\prime}\right) \sum_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}} \mathcal{U}_{l^{\prime} j^{\prime} c^{\prime}, l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}}^{J}\left(r^{\prime}\right) \psi_{l^{\prime \prime} j^{\prime \prime} c^{\prime \prime}, l j c}^{J}\left(r^{\prime}\right)
\end{align*}
$$

In matrix notation, this takes a much simpler form,

$$
\begin{equation*}
\Psi_{J}(r)=\Psi_{0 J}(r)+\int_{0}^{\infty} d r^{\prime} G_{0 J}^{+}\left(r, r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{J}\left(r^{\prime}\right) \tag{110}
\end{equation*}
$$

where $\Psi_{0 J}=\Psi_{0 J}^{+} e^{i \sigma_{J}}$ is the coupled-channels (diagonal) matrix form of the single-channel wave function of Eq. (42). Asymptotically, this partial wave equation tends to

$$
\begin{align*}
\frac{i}{2}\left(H_{J}^{-}(r)-\right. & \left.H_{J}^{+}(r) e^{i \sigma_{J}} \bar{S}_{J} e^{i \sigma_{J}}\right) e^{-i \sigma_{J}} \\
= & \frac{i}{2}\left(H_{J}^{-}(r)-H_{J}^{+}(r) e^{i \sigma_{J}} S_{0 J} e^{i \sigma_{J}}\right) e^{-i \sigma_{J}}  \tag{111}\\
& -H_{J}^{+}(r) \frac{2 \mu}{\hbar^{2}} K_{J}^{-1} e^{i \sigma_{J}} \int_{0}^{\infty} d r^{\prime} \Psi_{0 J}\left(r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{J}\left(r^{\prime}\right)
\end{align*}
$$

Extracting the coefficient of the outgoing Coulomb wave, $H_{J}^{+}$, one obtains an expression for the coupled-channels S-matrix,

$$
\begin{equation*}
S_{J}=S_{0 J}+2 i \frac{2 \mu}{\hbar^{2}} K_{J}^{-1 / 2} \int_{0}^{\infty} d r^{\prime} \Psi_{0 J}\left(r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{J}\left(r^{\prime}\right) K_{J}^{-1 / 2} \tag{112}
\end{equation*}
$$

(Recall that $S_{J}=K_{J}^{1 / 2} \bar{S}_{J} K_{J}^{-1 / 2}$.)
The drawback to this form of obtaining the S-matrix is that it first requires knowledge of the full coupled-channels wave function. However, by using this general expression and Eq, (110) above, one has for the zeroth-order approximations (starting values) for the wave function and S-matrix of the partial wave $J$,

$$
\begin{equation*}
\Psi_{J}^{(0)}(r)=\Psi_{0 J}(r) \quad \text { and } \quad S_{J}^{(1)}=S_{0 J} \tag{113}
\end{equation*}
$$

The first-order distorted-wave Born approximation, or just DWBA, to the wave function is then

$$
\begin{equation*}
\Psi_{J}^{(1)}(r)=\Psi_{0 J}(r)+\int_{0}^{\infty} d r^{\prime} G_{0 J}^{+}\left(r, r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{0 J}\left(r^{\prime}\right), \tag{114}
\end{equation*}
$$

while the DWBA to the S-matrix is

$$
\begin{equation*}
S_{J}^{(1)}=S_{0 J}+2 i \frac{2 \mu}{\hbar^{2}} K_{J}^{-1 / 2} \int_{0}^{\infty} d r^{\prime} \Psi_{0 J}\left(r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{0 J}\left(r^{\prime}\right) K_{J}^{-1 / 2} \tag{115}
\end{equation*}
$$

Note that the S-matrix is clearly symmetric when the optical potential coupling matrix $U_{J}^{\prime}$ is symmetric.

The DWBA is at times extended to second order in the coupling. In this case, one obtains for the wave function,

$$
\begin{align*}
\Psi_{J}^{(2)}(r)= & \Psi_{0 J}(r)+\int_{0}^{\infty} d r^{\prime} G_{0 J}^{+}\left(r, r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{0 J}\left(r^{\prime}\right)  \tag{116}\\
& +\int_{0}^{\infty} d r^{\prime} G_{0 J}^{+}\left(r, r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \int_{0}^{\infty} d r^{\prime \prime} G_{0 J}^{+}\left(r^{\prime}, r^{\prime \prime}\right) U_{J}^{\prime}\left(r^{\prime \prime}\right) \Psi_{0 J}\left(r^{\prime \prime}\right)
\end{align*}
$$

and for the S-matrix

$$
\begin{align*}
S_{J}^{(2)}=S_{0 J}+ & 2 i \frac{2 \mu}{\hbar^{2}} K_{J}^{-1 / 2} \int_{0}^{\infty} d r^{\prime} \Psi_{0 J}\left(r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right) \Psi_{0 J}\left(r^{\prime}\right) K_{J}^{-1 / 2} \\
+ & 2 i \frac{2 \mu}{\hbar^{2}} K_{J}^{-1 / 2} \int_{0}^{\infty} d r^{\prime} \Psi_{0 J}\left(r^{\prime}\right) U_{J}^{\prime}\left(r^{\prime}\right)  \tag{117}\\
& \times \int_{0}^{\infty} d r^{\prime \prime} G_{0 J}^{+}\left(r^{\prime}, r^{\prime \prime}\right) U_{J}^{\prime}\left(r^{\prime \prime}\right) \Psi_{0 J}\left(r^{\prime \prime}\right) K_{J}^{-1 / 2}
\end{align*}
$$

The DWBA is usually not extended beyond second order. If higher order terms in the interaction are necessary, it is usually better to resort to other methods, such as a conventional coupled-channels calculations.

The DWBA was developed to approximate the effects of the coupling between channels when that coupling is weak. It assumes that the contribution of the next-order term will always be relatively small compared to the last term included, due to the weakness of the coupling. It usually overestimates the cross section since it neglects the flux lost due to the excitation that it itself describes.

## 9 Reduced matrix elements and form factors

In Eq. (77), the target-spin-angular functions were used to calculate the matrix elements of the optical potential,

$$
\begin{equation*}
\mathcal{U}_{l^{\prime} j^{\prime} c^{\prime}, l j c}^{J}(r)=\int d^{3} r_{i n t} d \Omega \mathcal{Y}_{l^{\prime} s j^{\prime} c^{\prime}}^{J M \dagger}(\hat{r}) U_{o p t}\left(\vec{r}, \vec{r}_{i n t}\right) \mathcal{Y}_{l s j c}^{J M}(\hat{r}) \tag{118}
\end{equation*}
$$

It was noted that although these elements are calculated for a particular value $M$ of the projection of the total angular momentum $J$, the matrix elements that result are independent of this value if the system is rotationally invariant. The representation of these matrix elements will be discussed here.

The most general form of a rotationally-invariant interaction between a projectile and a target couples tensor operators acting on the orbital angular momentum, $i^{\lambda} Y_{\lambda \mu}(\hat{r})$, the spin of
the projectile, $Q_{\sigma \nu}^{p}$, and the angular momentum of the target, $Q_{\kappa \xi}^{t}\left(\vec{r}_{i n t}\right)$, to a scalar,

$$
U_{o p t}\left(\vec{r}, \vec{r}_{i n t}\right)=\sum_{\lambda \sigma \kappa} V_{\lambda \sigma \kappa}^{c^{\prime} c}(r) i^{\lambda} Y_{\lambda \mu}(\hat{r}) Q_{\sigma \nu}^{p} Q_{\kappa \xi}^{t}\left(\vec{r}_{i n t}\right)\left(\begin{array}{ccc}
\lambda & \sigma & \kappa  \tag{119}\\
\mu & \nu & \xi
\end{array}\right) .
$$

In principle, the radial dependence of each term, $V_{\lambda \sigma \kappa}^{c^{\prime} c}(r)$, can depend on the angular momenta of the tensor operators, $\lambda, \sigma$, and $\kappa$, as well as on the initial and final channels, $c$ and $c^{\prime}$.

The interactions in the vibrational and rotational models (with a spherical spin-orbit potential) are simpler than the general one above, as they couple only the orbital and target angular momenta. In these cases, the projectile-spin tensor operator is itself a scalar. This is not true in general.

The matrix elements of the general interaction of Eq. (119), calculated with respect to the target-spin-angular functions, can always be written in the form

$$
\begin{equation*}
\mathcal{U}_{l^{\prime} s^{\prime} j^{\prime} c^{\prime}, l s j c}^{J}(r)=\frac{1}{\sqrt{4 \pi}} \sum_{\lambda \sigma \kappa} G_{l^{\prime} s^{\prime} j^{\prime} c^{\prime}, l s j c}^{J \lambda c^{\prime}}\left\langle s^{\prime}\right| Q_{\sigma}^{p}|s\rangle\left\langle I_{c^{\prime}}\right| Q_{\kappa}^{t}\left|I_{c}\right\rangle V_{\lambda \sigma \kappa}^{c^{\prime} c}(r) . \tag{120}
\end{equation*}
$$

The factor $G_{l^{\prime} s^{\prime} j^{\prime} c^{\prime}, l s j c}^{J \lambda \sigma k}$ is a geometrical/statistical coefficient, which gives the appropriate weight to the angular momenta involved,

$$
\begin{align*}
G_{l^{\prime} s^{\prime} j^{\prime} c^{\prime}, l s j c}^{J \lambda \sigma}= & (-)^{J+I_{c}^{\prime}+j} i^{l^{\prime}+l+\lambda}  \tag{121}\\
& \times \sqrt{\left(2 l^{\prime}+1\right)(2 l+1)(2 \lambda+1)\left(2 j^{\prime}+1\right)(2 j+1)(2 \mu+1)} \\
& \times\left(\begin{array}{ccc}
l^{\prime} & l & \lambda \\
0 & 0 & 0
\end{array}\right)\left\{\begin{array}{ccc}
j^{\prime} & j & \kappa \\
I_{c} & I_{c^{\prime}} & J
\end{array}\right\}\left\{\begin{array}{lll}
l^{\prime} & l & \lambda \\
s^{\prime} & s & \sigma \\
j^{\prime} & j & \kappa
\end{array}\right\} .
\end{align*}
$$

The reduced matrix elements of the projectile and target angular momentum tensor operators, $\left\langle s^{\prime}\right| Q_{\sigma}^{p}|s\rangle$ and $\left\langle I_{c^{\prime}}\right| Q_{\kappa}^{t}\left|I_{c}\right\rangle$ can, in principle, contain information about the nuclear part of the matrix elements. In the macroscopic models discussed, the reduced matrix element for the projectile is just a number, the number 1 , to be precise. The reduced matrix element for the target depend on the oscillation amplitude in the vibrational model and on the deformation parameters in the rotational one.

Models with quite general couplings can be constructed in terms of their reduced matrix elements and form factors. In this way, it is possible to construct microscopic as well as alternative macroscopic models of the nuclear coupling.

## 10 Summary remarks

The objective of the optical model is to describe the fast, direct contribution to nuclear scattering. It makes use of an optical potential having both real and negative imaginary parts. The absorption of flux from the optical wave function, due to the imaginary part of the potential, accounts for the flux lost to the slower, compound nucleus component of the scattering. The single-channel optical model describes the scattering in the elastic channel alone. It is often called the spherical optical model because, in it, the target may be considered to be spherically symmetric, since its structure is never introduced.

Direct reactions that transfer energy as well as momentum are often quite important. Such inelastic scatterings, in the case of the inert projectiles that we are considering ( $\mathrm{n}, \mathrm{p}, \alpha, \mathrm{d}$, etc.), leave the target in an excited state and diminish the asymptotic kinetic energy of the projectile. To describe inelastic scattering, one must introduce at least the basic characteristics of the ground and excites states of the target.

The optical model and optical potential continue to be subjects of intense research. One can find out more about the directions this research is taking in the proceedings of a more or less recent conference[23].

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