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International Centre for Theoretical Physics*



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**Gribov-80 Memorial Workshop on Quantum Chromodynamics and
Beyond'**

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**Discovery of the Perfect Fluidity of Quark-Gluon Plasma at RHIC and
hydrodynamical analysis of recent RHIC data**

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Exploring the Quark Gluon Plasma at the Relativistic Heavy Ion Collider



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- Milestones at RHIC
- Hydrodynamics
- Direct photon results

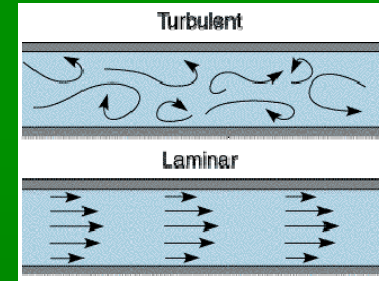


Milestones of RHIC results

- Jet suppression in A+A: new phenomenon
 - Phys. Rev. Lett. 88, 022301 (2002)
- No jet suppression in d+A: new form of matter
 - Phys. Rev. Lett. 91, 072303 (2003)
- Summary of the results: matter is a liquid
 - Nucl. Phys. A 757, 184-283 (2005)
- Elliptic flow scaling: quark degrees of freedom
 - Phys. Rev. Lett. 98, 162301 (2007)
- Heavy quark flow: viscosity near lower limit
 - Phys. Rev. Lett. 98, 172301 (2007)
- Initial temperature, state of matter: no final answer yet

Hydrodynamic scaling predictions

- Hydro predicts scaling (also viscous)
- What does a scaling mean?
 - For example Reynolds number $\rho v r / \eta$
 - Only a combination of parameters matters

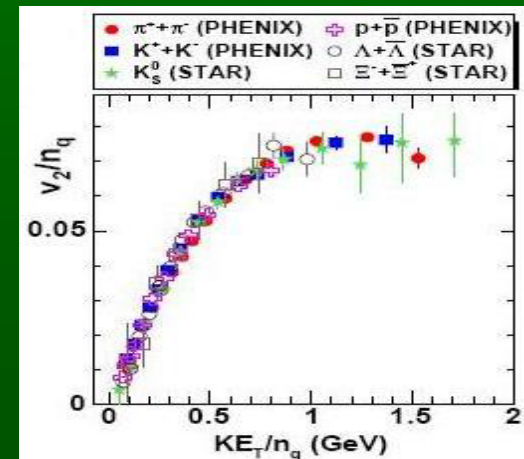
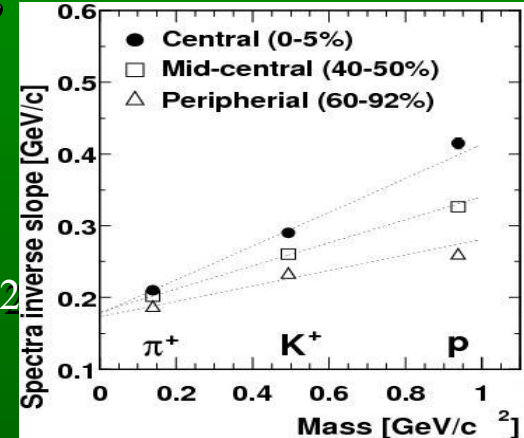
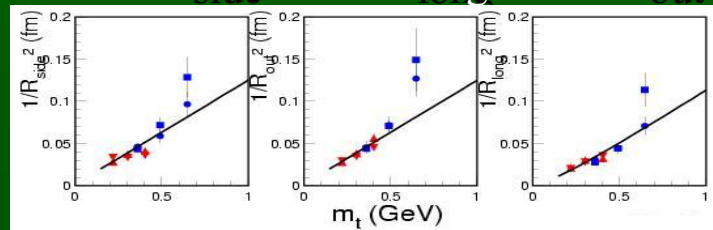


- Collective, thermal behavior →
Loss of information

- Spectra slopes: $N_1 \sim e^{\frac{m_t}{T_{\text{eff}}}}$; $T_{\text{eff}} = T_0 + m u_t^2$

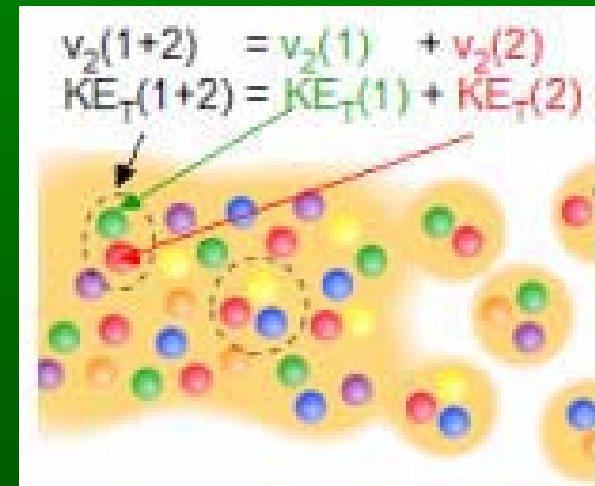
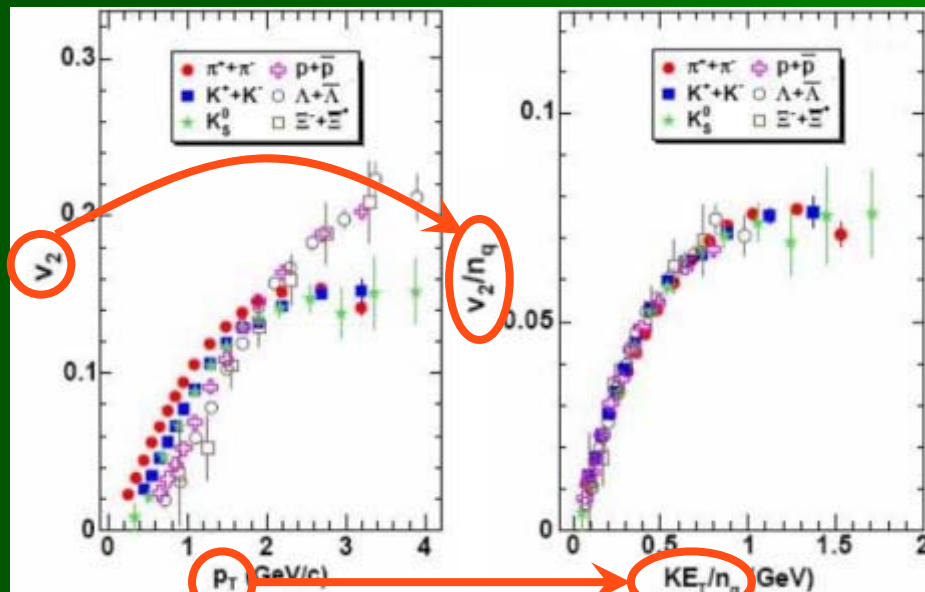
- Elliptic flow: $v_2 = \frac{I_1(w)}{I_0(w)}$; $w \sim KE_T$

- HBT radii: $R_{\text{side}}^2 \approx R_{\text{long}}^2 \approx R_{\text{out}}^2 \sim \frac{1}{m_t}$



Quark degrees of freedom

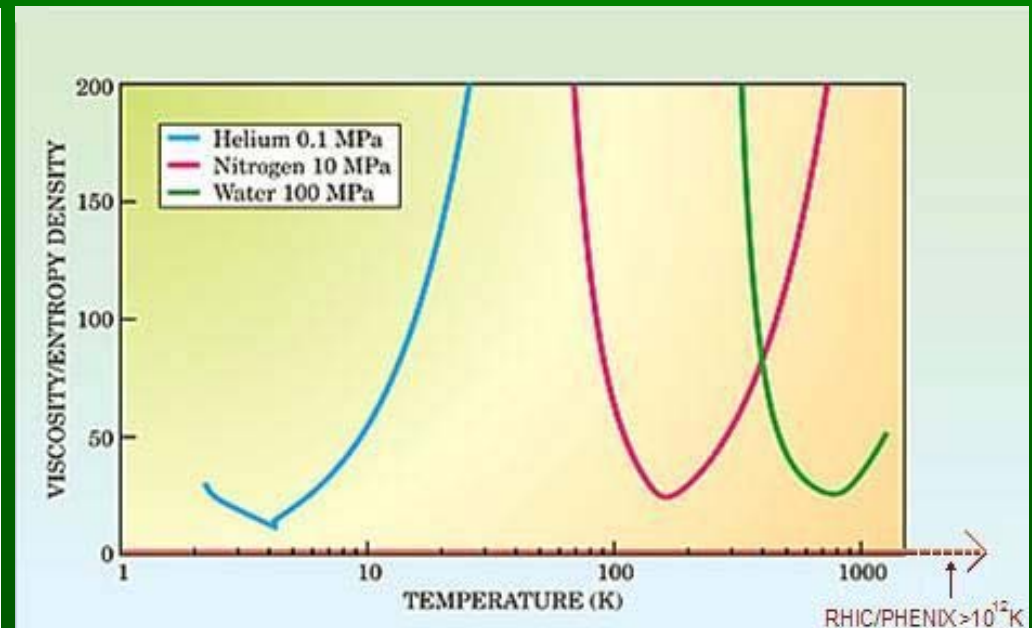
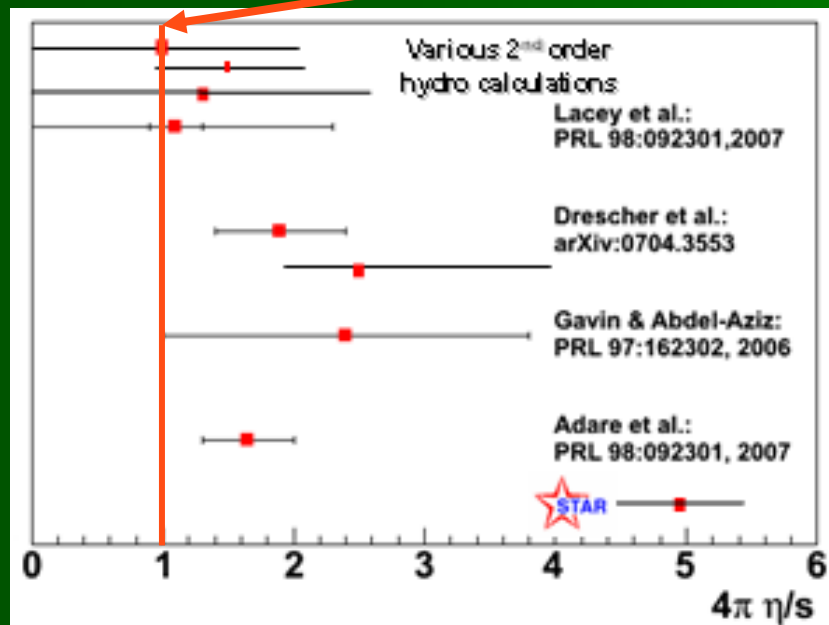
- Elliptic flow (v_2) is a central observable
 - measure of momentum space asymmetry
 - sign of collective behavior
- Measured for several types of particles
- Scaling variable suggested by hydro: KE_t
- Both v_2 and KE_t scale by the number of quarks



- Explanation by quark flow
- Indirect sign of quark degrees of freedom

Superfluidity

- Kinematic viscosity η/s lower than that of superfluid helium
- Conjectured limit: $1/4\pi$ (via AdS/CFT)



Perfect hydro solutions

Solution	Basic prop's	EoS	Observables
Csörgő, Nagy, Csanád Phys.Lett.B 663:306-311, 2008 Phys.Rev.C77:024908,2008	Ellipsoidal, 1D accelerating	$\varepsilon - B = \kappa(p + B)$	$dn/dy, \varepsilon$
Landau Izv. Acad. Nauk SSSR 81 (1953) 51	Cylindr., 1D, accelerating	$\varepsilon = \kappa n T$	None
Hwa-Björken R.C. Hwa, PRD10, 2260,1974 J.D. Bjorken, PRD27, 40(1983)	Cylindr., 1D, non-accelerating	$\varepsilon = \kappa n T$	$dn/dy, \varepsilon$
Bialas et al. A. Bialas, R. A. Janik, and R. B. Peschanski, Phys. Rev. C76, 054901 (2007).	1D, between Landau and Hwa-Björken	$\varepsilon = \kappa n T$	dn/dy
Csörgő, Csernai, Hama, Kodama Heavy Ion Phys. A 21, 73 (2004)	Ellipsoidal, 3D, non-accelerating	$\varepsilon = \kappa n T$	This work does the calculation
???	Ell, 3D, accel.	realistic	???

Where we are

- Revival of interest, new solutions
 - Sinyukov, Karpenko, nucl-th/0505041
 - Pratt, nucl-th/0612010
 - Bialas et al.: Phys.Rev.C76:054901,2007
 - Borsch, Zhdanov: SIGMA 3:116,2007
 - Nagy et al.: J.Phys.G35:104128,2008 and arXiv/0909.4285
 - Liao et al.: arXiv/09092284 and Phys.Rev.C80:034904,2009
 - Mizoguchi et al.: Eur.Phys.J.A40:99-108,2009
 - Beuf et al.: Phys.Rev.C78:064909,2008 (dS/dy as well!)
- Need for solutions that are:
 - accelerating + relativistic + 3 dimensional
 - explicit + simple + compatible with the data

A 3D relativistic solution

- The hydro fields are these:

- $v(s)$ arbitrary, but realistic to choose Gaussian $v(s) = e^{-bs/2}$
 $b < 0$ is realistic

$$n = n_0 \left(\frac{\tau_0}{\tau} \right)^3 v(s)$$

$$T = T_0 \left(\frac{\tau_0}{\tau} \right)^{(3/\kappa)} \frac{1}{v(s)}$$

- Ellipsoidal symmetry

$$s = \frac{x^2}{X(t)^2} + \frac{y^2}{Y(t)^2} + \frac{z^2}{Z(t)^2}$$

$$p = p_0 \left(\frac{\tau_0}{\tau} \right)^{\left(3 + \frac{3}{\kappa}\right)}$$

$$u^\mu = \gamma \left(1, \frac{\dot{X}(t)}{X(t)} x, \frac{\dot{Y}(t)}{Y(t)} y, \frac{\dot{Z}(t)}{Z(t)} z \right)$$

- (thermodynamic quantities const. on the $s = \text{const.}$ ellipsoid)

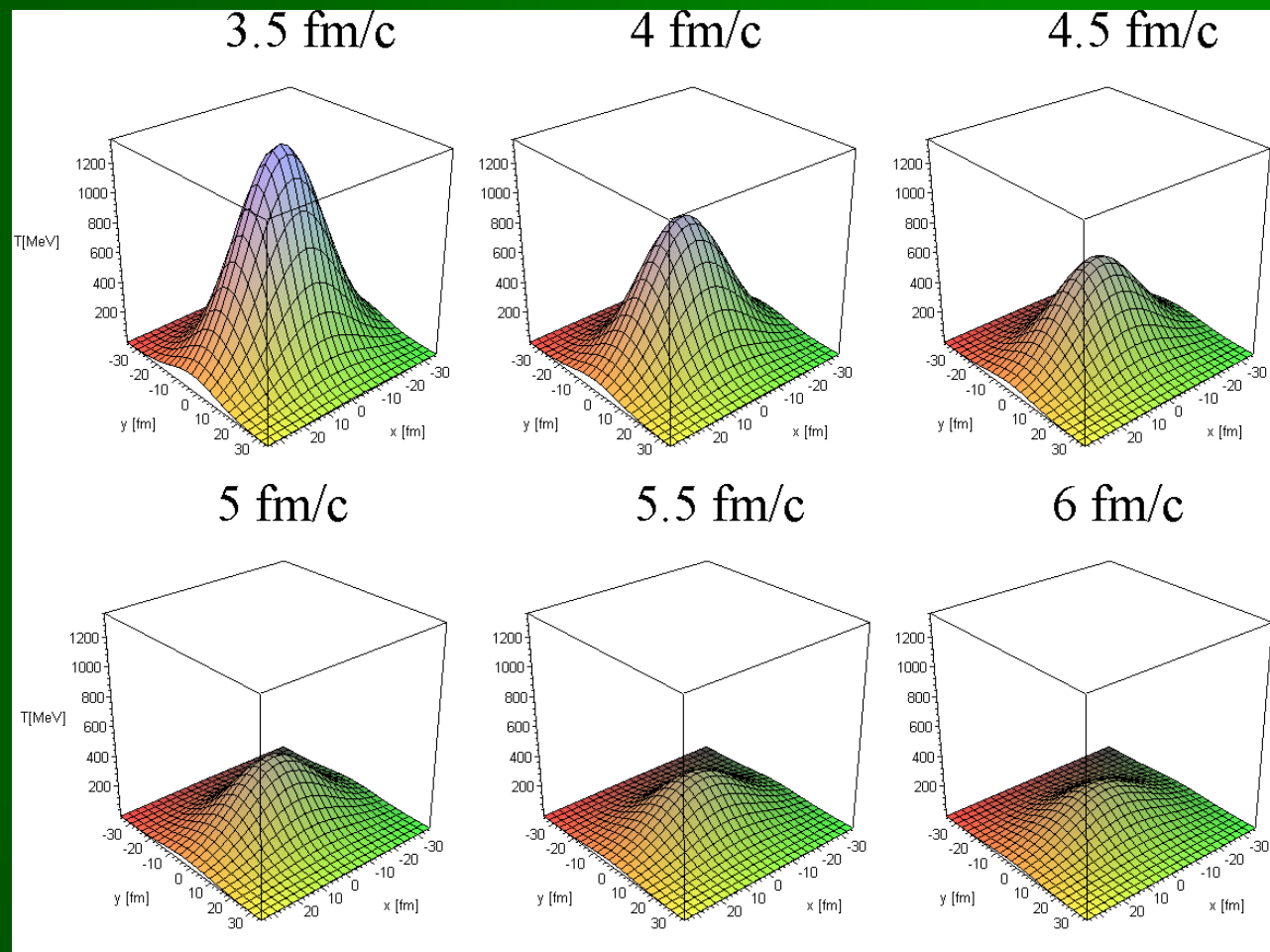
- Directional Hubble-flow (no acceleration)

- $v = Hr$ or $H = v/r$, the Hubble-constants: $\frac{\dot{X}(t)}{X(t)}, \frac{\dot{Y}(t)}{Y(t)}, \frac{\dot{Z}(t)}{Z(t)}$
- $\dot{X}(t), \dot{Y}(t), \dot{Z}(t) : \text{const.}; (\dot{X}, \dot{Y}) \Leftrightarrow (u_t, \varepsilon)$

(T. Csörgő, L. P. Csernai, Y. Hama és T. Kodama, Heavy Ion Phys. A 21, 73 (2004))

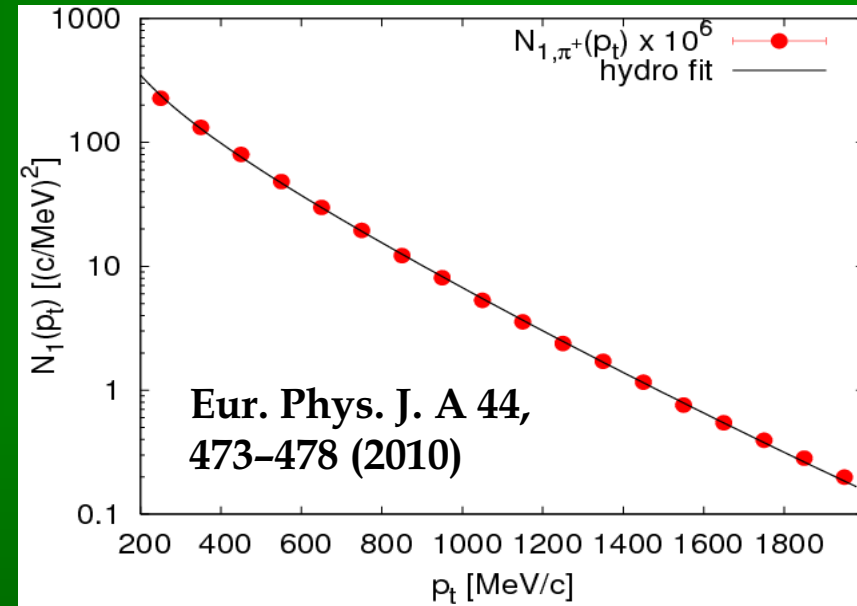
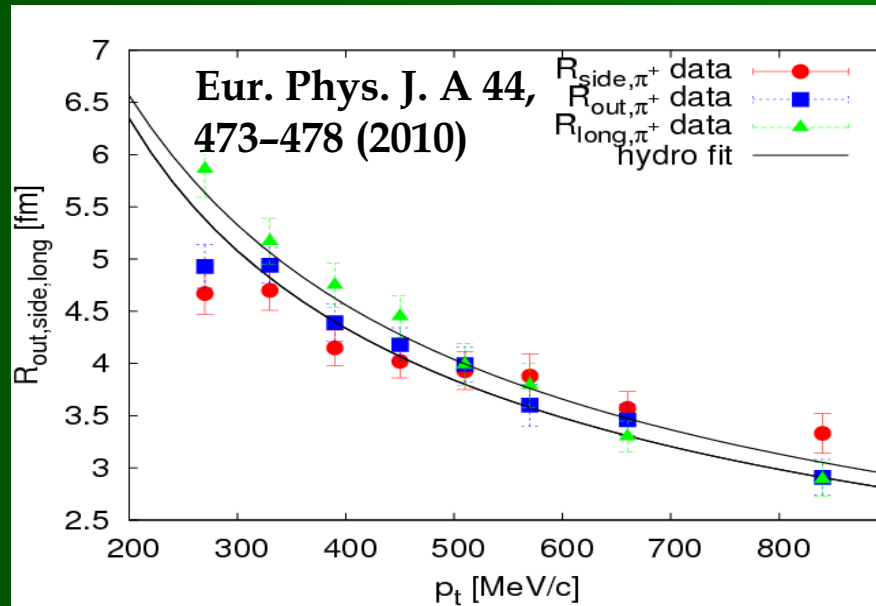
Temperature profile

Transverse temperature profile as a function of time
with an example parameter set:



Single pion spectrum with HBT radii

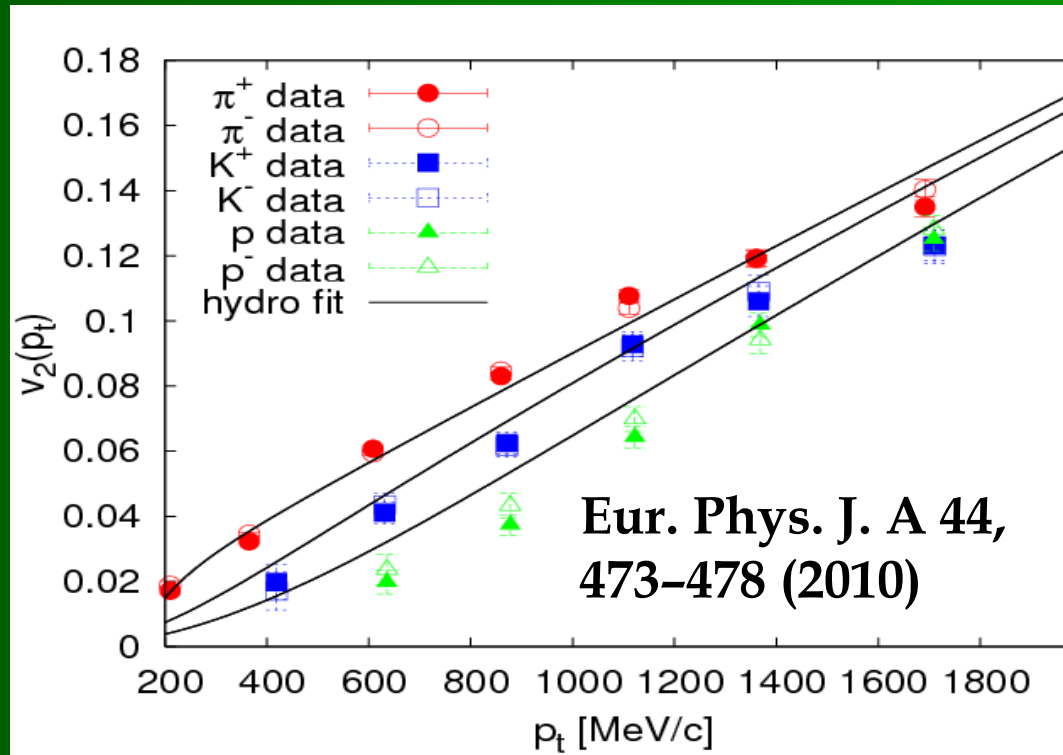
- 0-30% centrality, Au+Au, PHENIX



- T_0 199 ± 3 MeV central freeze-out temp.
- ε 0.80 ± 0.02 momentum space ecc.
- u_t^2/b -0.84 ± 0.1 ($b < 0$) transv. flow/temp. grad
- τ_0 7.7 ± 0.1 freeze-out proper time
- χ^2 171 (24 with theory error)

Elliptic flow

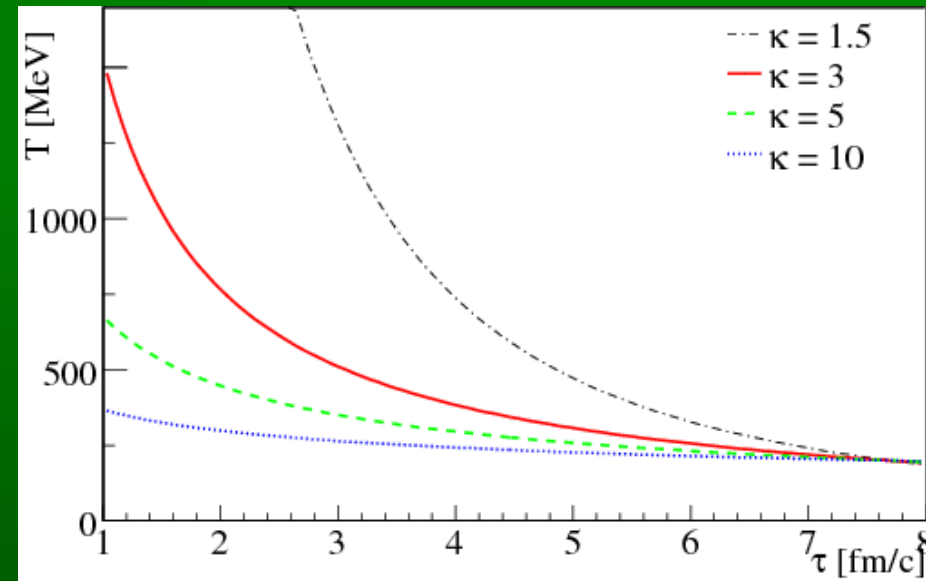
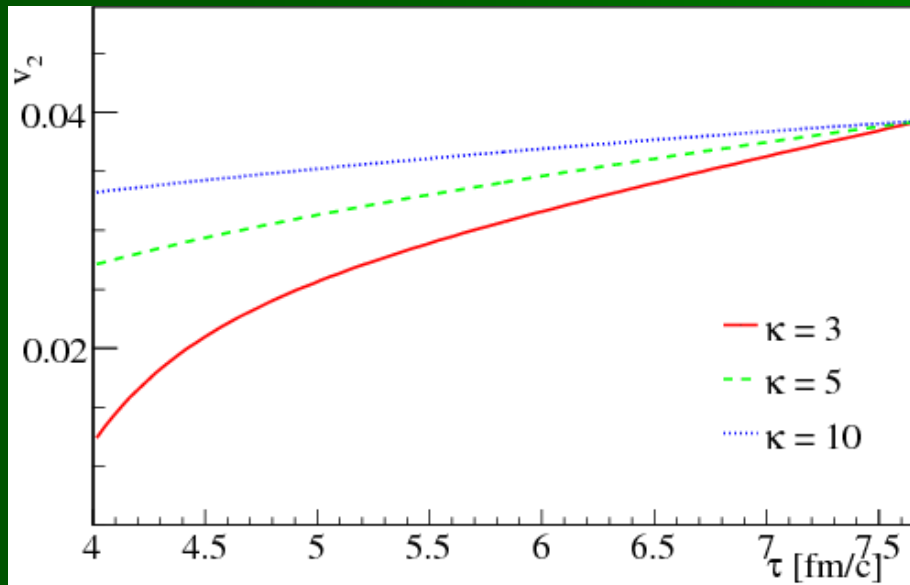
- 0-92% centrality, Au+Au, PHENIX (R.P. method technique)



- T_0 204 ± 7 MeV f.o. temperature
- ε 0.34 ± 0.01 eccentricity
- u_t^2/b -0.34 ± 0.07 ($b < 0$) transv. flow/temp. grad
- χ^2 256 (66 with theory error)

Insensitivity on the EoS

- Hadronic data corresponds to the initial state
- Not directly sensitive to
 - Equation of state
 - Initial condition



- Look for other data...
- Photon-emission rather continuous

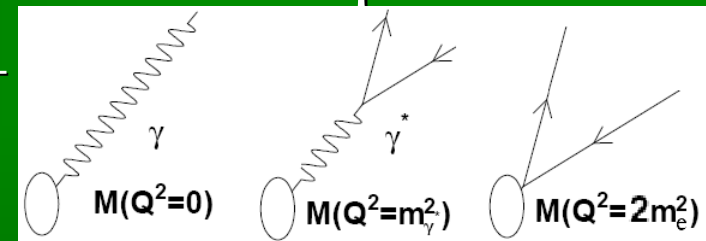
Summary of the fit results

- Freeze-out central temperature ~ 200 MeV
- Hadronic fits don't depend on:
 - EoS (κ or sound velocity)
 - initial temperature
- Other data needed to determine IC and EoS
- After thermalization:
 - thermal/direct emission will play a significant role
 - direct photons 'feel' the evolution of the fireball
- Can we measure direct photons?
- Yes we can!

How to measure direct photons?

- PHENIX measurement done: PRL 104, 132301 (2010)
- Problem: huge background from $h \rightarrow \gamma\gamma$
- Idea: thermal + virtual photon production parallel

– $X \rightarrow e^+e^-$, $X \rightarrow \gamma$ and $X \rightarrow \gamma^* \rightarrow e^+e^-$
from the same process



- Dielectron and real photon production related as:

$$\frac{d^2 n_{ee}}{dm_{ee} dp_t} = \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} L(m_{ee}) S(m_{ee}) \frac{dn_\gamma}{dp_t}, \quad L(m_{ee}) = \sqrt{1 - \frac{4m_e^2}{m_{ee}^2}} \left(1 + \frac{2m_e^2}{m_{ee}^2} \right)$$

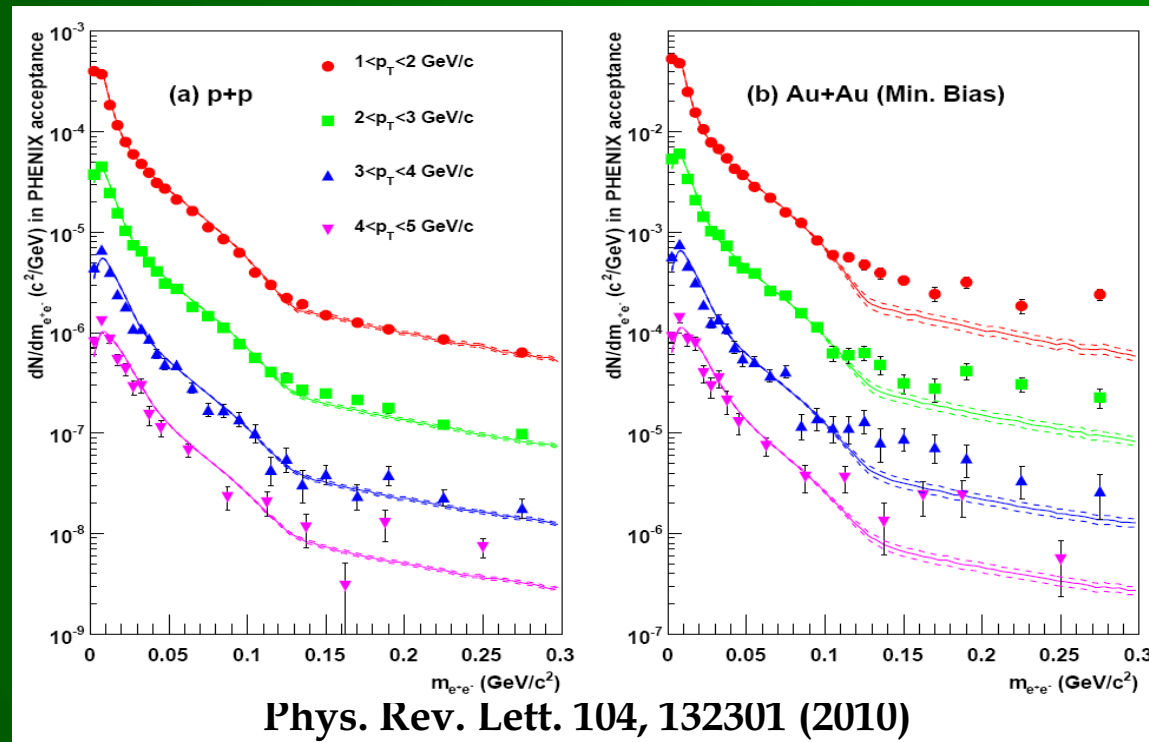
- S process dependent, dn_{γ^*}/dn_γ , for π_0 and η e.g.:

$$S(m_{ee}) = |F(m_{ee}^2)|^2 \left(1 - m_{ee}^2 / M_h^2 \right)^3, \quad S(m_{ee}) = 0 \text{ for } m_{ee} > M_h$$

- For $p_t \gg m_{ee} \gg m_e$: $L, S \rightarrow 1$

Dielectron spectrum measurement

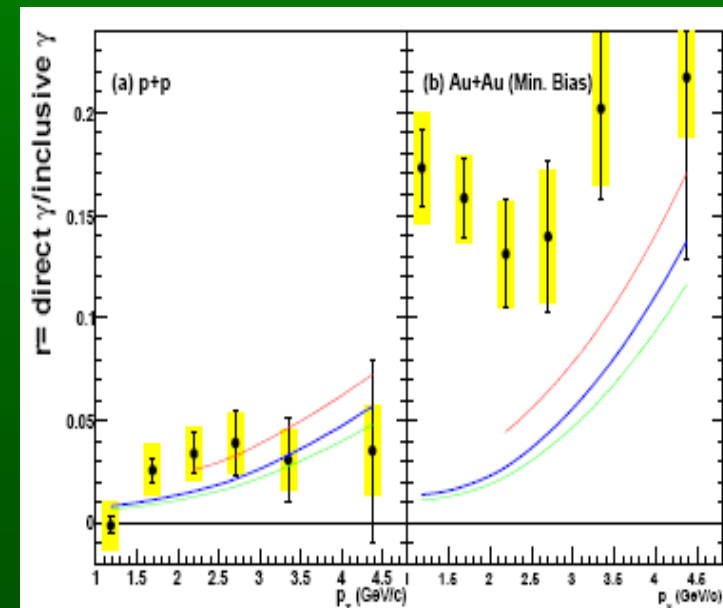
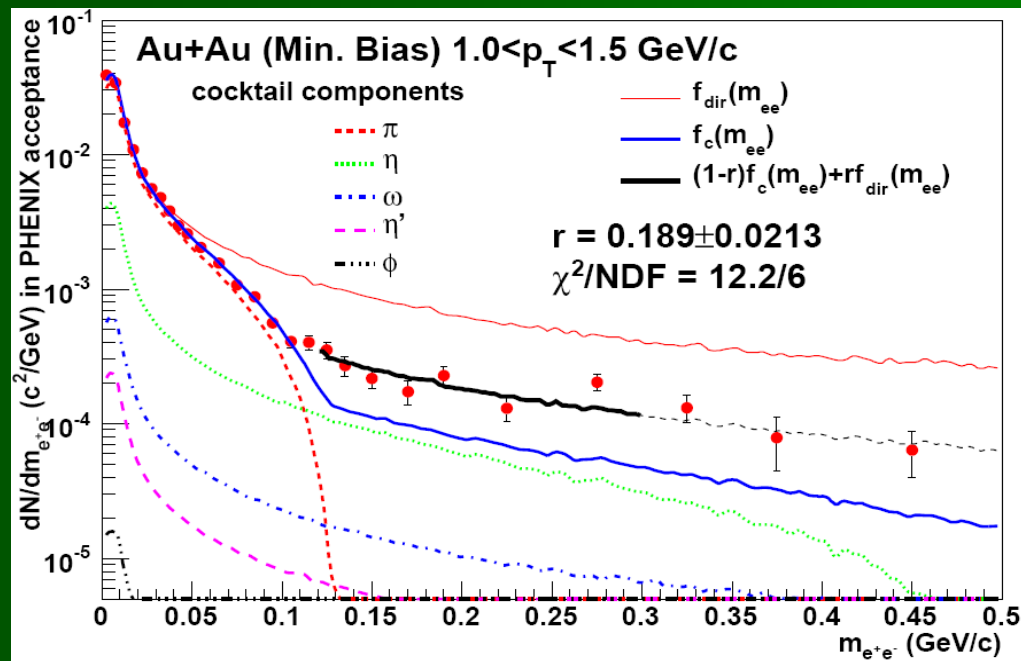
- Measured electron pairs with p_t of 1-5 GeV
- Easy via electron ID capabilities
- Compare to dielectrons from hadronic cocktail



- Excess seen above pion mass due to virtual γ

Direct photons versus decay photons

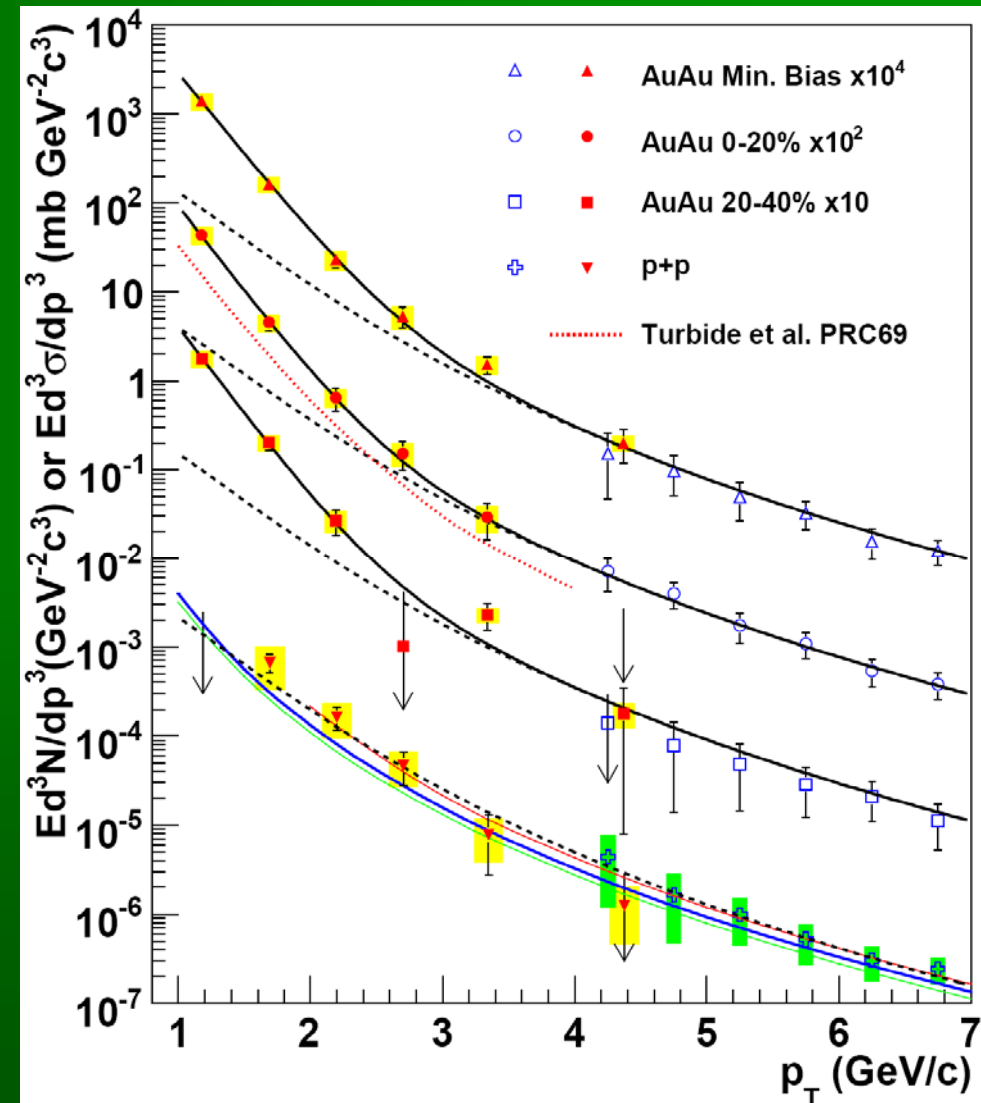
- Excess: virtual direct photons (decaying into e^+e^- pairs)
- Inclusive e^+e^- : hadronic + dir. virtual photon components
 - Hadronic electron pairs (f_c), calculated from cocktail: $\pi, \eta, \omega, \eta', \phi$
 - Electron pairs from direct virtual photons (f_{dir}) calculated from f_c via previous formula
- Determine ratio r by fit for separate p_t bins
- Use r to scale inclusive photon spectra



Direct photon spectra measured

- Inclusive photons: from the $m_{ee} < 0.03$ GeV/ c^2 inclusive e^+e^- yield ($S=1$ here)
- Direct photon spect.:

$$N_{\text{dir}}(p_t) = r \times N_{\text{incl}}(p_t)$$
- Thermal below roughly 3 GeV
- What temperature does this correspond to?



Initial temperature

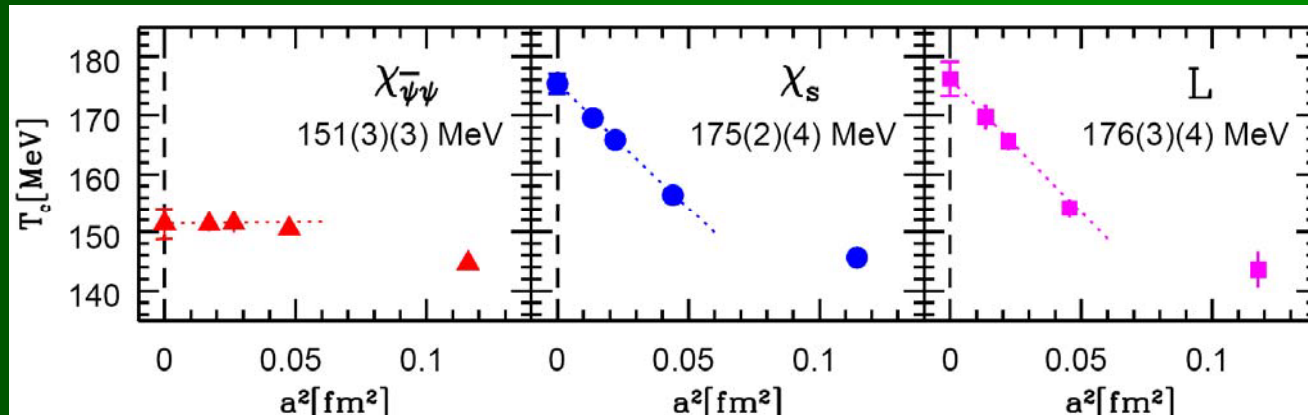
- Direct photon spectra can be fitted with a single average temperature
- $233 \pm 14 \pm 19$ MeV for minimum bias
- Initial temperature definitely higher
- More sophisticated method: using hydro
- Photon yield would then be:

$$\int \frac{p_\mu d^3\Sigma_\mu}{e^{\frac{p_\mu u^\mu}{T}} - 1} dt$$

- Integration over the whole time evolution

Transition temperature

- Budapest-Wuppertal group (arXiv:0908.3341) :
 - Chiral susceptibility: 151(3)(3) MeV
 - Strange quark number susceptibility: 175(2)(4) MeV
 - Polyakov loops: 176(3)(4) MeV



- MILC group, PRD71, 034504 (2005):
 - chiral susceptibility: 169(12)(4) MeV
- RBC-Bielefeld, PRD74, 054507 (2006):
 - chiral susceptibility and Polyakov loops: 192(7)(4) MeV
- Some disagreement, but well below 200 MeV
- Experimental data: T_i at least $2-3\sigma$ above T_c

Summary

- Hydro working at RHIC
- Indirect signs of quark DoF seen
- T_c calculated via lattice QCD
- Hadronic data at RHIC + hydro: $T_{\text{final}} \cong T_c$
- Photon data at RHIC + hydro: $T_{\text{initial}} \gg T_c$

**Thank you for your
attention**

Why we use hydrodynamics?

- Hadron & parton cascade models
 - Nonequilibrium models
 - Don't describe most of particle creation at low p_t
- Flow + hadronic cascade and resonance corrections
 - Small correction at RHIC energies ([arXiv:0903.1863](#))
- Most of particle creation according to hydro
 - Pion and kaon HBT radii m_t scaling ([Acta Phys.Polon.Supp.1:521-524,2008](#))
- After ~ 10 fm/c nonequilibrium expanding hadron gas
 - Anomal diffusion, no correction for spectra or flow ([Braz.J.Phys.37:1002-1013,2007](#))
- American Institute of Physics: success of hydro is the physics story 2005

Little vocabulary of hydrodynamics

- Exact/parametric solution
 - Solution of hydro equations analytically, without approximation
 - Usually has free parameters
- Hydro inspired parameterization
 - Distribution determined at freeze-out only, their time dependence is not considered
- Numerical solution
 - Solution of hydro equations numerically
 - Start from arbitrary initial state

Nonrelativistic solutions

Solution	Symmetry	Density prof.	EoS	Observables
Csizmadia et al. Phys. Lett. B443:21-25, 1998	Sphere	Gaussian	$\varepsilon = \frac{3}{2}nT$	Calculated
Csörgő Central Eur.J.Phys.2: 556-565,2004	Sphere	Arbitrary	$\varepsilon = \frac{3}{2}nT$	Not calculated
Akkelin et al. Phys.Rev. C67,2003	Ellipsoid	Gaussian ($T=T(t)$)	$\varepsilon = \kappa(T)nT$	Calculated
Csörgő Acta Phys.Polon. B37:483-494,2006	Ellipsoid	Arbitrary ($T=T(r,t)$)	$\varepsilon = \kappa nT$	Not calculated
Csörgő, Zimányi Heavy Ion Phys.17:281-293,2003	Ellipsoid	Gaussian	$\varepsilon = \kappa_e nT - B$	Calculated

Equations of relativistic hydro

- Assuming local thermal equilibrium
- For a perfect fluid: $T^{\mu\nu} = wu^\mu u^\nu - pg^{\mu\nu}$, $w = \varepsilon + p \Rightarrow \partial_\nu T^{\mu\nu} = 0$
- Equations in four-vector form and nonrelativistic notation

- Euler equation:

$$wu^\nu \partial_\nu u^\mu = (g^{\mu\eta} - u^\mu u^\eta) \partial_\eta p \quad \frac{w}{1-v^2} \frac{d\mathbf{v}}{dt} = - \left(\nabla p + \mathbf{v} \frac{\partial p}{\partial t} \right)$$

- Energy conservation:

$$w \partial_\nu u^\nu + u^\nu \partial_\nu \varepsilon = 0 \quad \frac{1}{w} \frac{d\varepsilon}{dt} = -\nabla \cdot \mathbf{v} - \frac{1}{1-v^2} \frac{d}{dt} \frac{v^2}{2}$$

- Charge conservation:

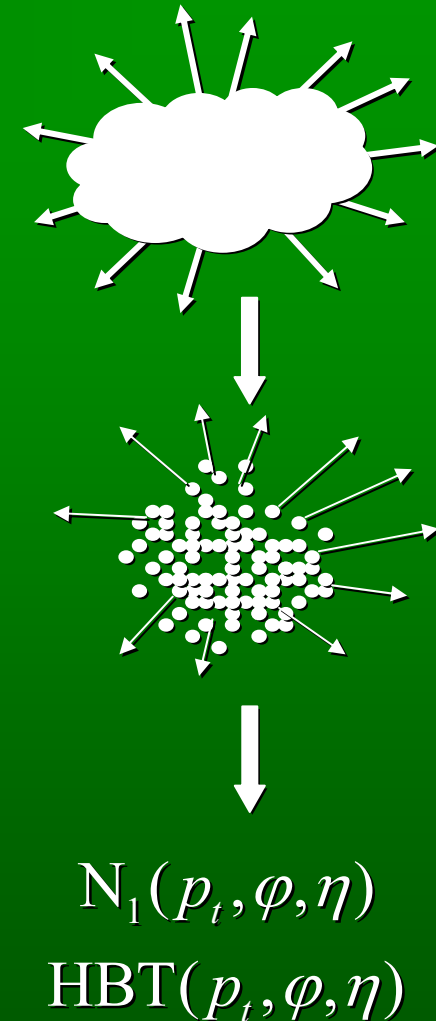
$$\partial_\mu (nu^\mu) = 0 \quad \frac{d}{dt} \ln \frac{n}{\sqrt{1-v^2}} = -\nabla \cdot \mathbf{v}$$

$$u^\mu \partial_\mu = \frac{d}{d\tau} \quad \text{comoving proper-time derivative}$$

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla \quad \text{comoving derivative}$$

How analytic hydro works

- Take hydro equations and EoS
- Find a solution
 - Will contain parameters (like Friedmann, Schwarzschild etc.)
 - Will use a possible set of initial conditions
- Use a freeze-out condition
 - Eg fixed proper time or fixed temperature
 - Generally a hyper-surface
- Calculate the hadron source function
- Calculate observables
 - E.g. spectra, flow, correlations
 - Straightforward calculation
- Hydrodynamics: Initial conditions \otimes dynamical equations \otimes freeze-out conditions

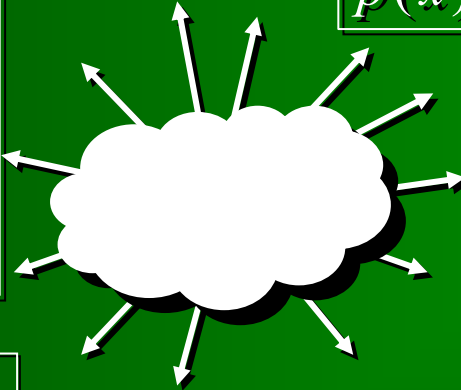


How analytic hydro works

Hydro equations + EoS PLB505:64-70,2001
hep-ph/0012127 → Self-similar solution:

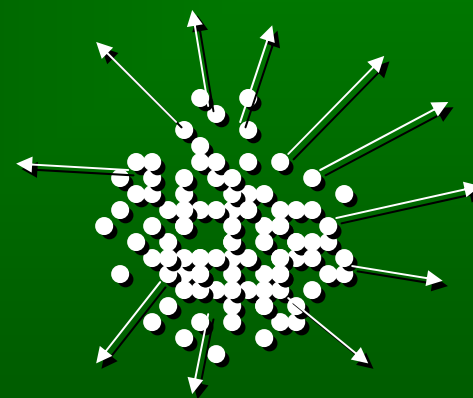
$$\begin{aligned}
 (\partial_t + \nabla \vec{v})n &= 0 \\
 (\partial_t + \nabla \vec{v})T &= -p \nabla \vec{v} \\
 (\partial_t + \vec{v} \nabla) \vec{v} &= -(1/n) \nabla p \\
 T &= \frac{3}{2} p
 \end{aligned}$$

$$p(x), n(x), T(x), \vec{v}(x)$$



Phase-space distribution
Boltzmann-equation

$$\begin{aligned}
 f(x, p) &\sim n(x) \exp \left\{ -\frac{(p - mv)^2}{kT} \right\} \\
 (\partial_t + \vec{v} \nabla) f(x, p) &= S(x, p)
 \end{aligned}$$



PRC67:034904,2003
hep-ph/0108067 → Source
 $S(x, p)$

PRC54:1390-1403,1996
hep-ph/9509213

Observables
 $N_1(p), C_2(p_1, p_2), v_2(p)$

Scheme works also backwards

The source function

- Source function: probability of a particle created at x with p
- Maxwell-Boltzmann distribution + extra terms

$$S(x, p)d^4x = \mathcal{N}n(x) \exp\left[-\frac{p_\mu u^\mu(x)}{T(x)}\right] \frac{p_\mu u^\mu}{u^0} H(\tau)d^4x$$

– \mathcal{N} normalization

– $H(\tau)d\tau$ freeze-out distribution

if sudden: $H(\tau) = \delta(\tau - \tau_0)$

– $\frac{p_\mu u^\mu}{u^0} d^3x$ Cooper-Fry prefactor (flux term)

– Validity: $\tau_0 > R_{\text{HBT}}, m_t > T_0$

Single particle spectrum

- Source function: spatial origin and momentum
- Momentum distribution

→ integrate on spatial coordinates:

$$N_1(p) = \int_{R^4} S(x, p) d^4x$$

- Second order Gaussian approximation around emission maximum
- After integration:

$$N_1(p) = \bar{N} \cdot \bar{E} \cdot \bar{V} \cdot \exp \left[\frac{p^2}{2ET_0} - \frac{E}{T_0} - \frac{p_x^2}{2ET_x} - \frac{p_y^2}{2ET_y} - \frac{p_z^2}{2ET_z} \right]$$

- Directional slope parameter:

$$T_x = T_0 + \frac{ET_0 \dot{X}_0^2}{b(T_0 - E)}$$

Transverse momentum spectrum

- Go to mid-rapidity ($y=0$)
- Integrate on transverse angle ϕ

$$N_1(p_t) = \bar{N} \bar{V} \left(m_t - \frac{p_t^2 (T_{\text{eff}} - T_0)}{m_t T_{\text{eff}}} \right) \exp \left[-\frac{p_t^2}{2m_t T_{\text{eff}}} + \frac{p_t^2}{2m_t T_0} - \frac{m_t}{T_0} \right]$$

- The effective temperature is from the slopes:

$$T_x = T_0 + m_t \dot{X}^2 \frac{T_0}{b(T_0 - E)}, \quad T_y = T_0 + m_t \dot{Y}^2 \frac{T_0}{b(T_0 - E)},$$

$$\frac{1}{T_{\text{eff}}} = \frac{1}{2} \left(\frac{1}{T_x} + \frac{1}{T_y} \right)$$

The elliptic flow

- The elliptic flow can be calculated as:

$$v_2 = \frac{\int_0^{2\pi} d\phi N_1(p_t, \phi) \cos(2\phi)}{\int_0^{2\pi} d\phi N_1(p_t, \phi)}$$

- Result (similar to other models): $v_2 = \frac{I_1(w)}{I_0(w)}$

- $I_n(w)$: modified Bessel functions

$$I_n(w) = \frac{1}{2\pi} \int_0^{2\pi} e^{w \cos(2\phi)} \cos(2n\phi) d\phi$$

- Where w is: $w = \frac{p_t^2}{4m_t} \left(\frac{1}{T_y} - \frac{1}{T_x} \right) \sim E_K \frac{\varepsilon}{T_{eff}}$

Two-particle correlation radii

- Definition:

$$C_2(p_1, p_2) = \frac{N_2(p_1, p_2)}{N_1(p_1)N_1(p_2)}$$

- From the source function: $C_2(q, K) = 1 + \lambda \left| \frac{\tilde{S}(q, K)}{\tilde{S}(0, K)} \right|^2$
- Changing coordinates

$$q = p_1 - p_2, K = 0.5(p_1 + p_2) \Rightarrow C_2(q, K) = 1 + \lambda \exp\left(-R_x^2 q_x^2 - R_y^2 q_y^2 - R_z^2 q_z^2\right)$$

- Result: $R_i = \frac{T_0 \tau_0 (T_i - T_0)}{m_t T_i}$

- The usual scaling (same for kaons!): $R_i^2 \sim \frac{1}{m_t}$

- Bertsch-Pratt coordinates: $R_{out} = R_{side} = 0.5 \left(R_x^2 + R_y^2 \right)$

- Freeze-out: $\tau = \text{const.} \leftrightarrow \Delta\tau = 0 \rightarrow R_{out} = R_{side}$

Famous solutions

- Landau's solution (1D, developed for p+p):
 - Accelerating, implicit, complicated, 1D
 - L.D. Landau, *Izv. Acad. Nauk SSSR* 81 (1953) 51
 - I.M. Khalatnikov, *Zhur. Eksp. Teor. Fiz.* 27 (1954) 529
 - L.D. Landau and S.Z. Belenkij, *Usp. Fiz. Nauk* 56 (1955) 309
- Hwa-Bjorken solution:
 - Non-accelerating, explicit, simple, 1D, boost-invariant
 - R.C. Hwa, *Phys. Rev. D* 10, 2260 (1974)
 - J.D. Bjorken, *Phys. Rev. D* 27, 40 (1983)
- Others
 - Chiu, Sudarshan and Wang
 - Baym, Friman, Blaizot, Soyeur and Czyz
 - Srivastava, Alam, Chakrabarty, Raha and Sinha

Virtual to real photon ratio

The yields of virtual photons dN_{γ^*} and electron pairs dN_{ee} are related:

$$\frac{d^2 N_{ee}}{dM^2} = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} dN_{\gamma^*}, \quad (\text{B1})$$

$$L(M) = \sqrt{1 - \frac{4m_e^2}{M^2}} \left(1 + \frac{2m_e^2}{M^2} \right), \quad (\text{B2})$$

$$\frac{d^2 N_{ee}}{dM^2} = \frac{\alpha}{3\pi} \frac{L(M)}{M^2} S(M, q) dN_{\gamma}. \quad (\text{B3})$$

Here we have introduced $S(M, q) = dN_{\gamma^*}(M)/dN_{\gamma}$ to factor out the difference between real photon emission and virtual photon emission. The factor $S(M, q)$ is pro-

$$\frac{d^2 n_{ee}}{dm_{ee}} = \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} \sqrt{1 - \frac{4m_e^2}{m_{ee}^2}} \left(1 + \frac{2m_e^2}{m_{ee}^2} \right) S dn_{\gamma} \quad (1)$$

tors, phase space, and spectral functions. $S(M, q)$ approaches 1 for small M , $S(M, q) \rightarrow 1$ for $M \rightarrow 0$. Additionally, since $L(M) \simeq 1 - 6m_e^2/M^2$ for $m_e \ll M$, $L(M) = 1$ is a very good approximation. Thus the relationship between the electron pair yield and the direct photon yield simplifies to

$$\frac{d^2 N_{ee}}{dM^2} \simeq \frac{\alpha}{3\pi} \frac{1}{M^2} dN_{\gamma}, \quad (\text{B4})$$

$$\frac{d^2 N_{ee}}{dM} \simeq \frac{2\alpha}{3\pi} \frac{1}{M} dN_{\gamma}. \quad (\text{B5})$$