



**The Abdus Salam
International Centre for Theoretical Physics**



2167-25

Advanced School on Direct and Inverse Problems of Seismology

27 September - 8 October, 2010

Understanding the Earth's Interior from Relaxation Normal Modes

1 Basic theory

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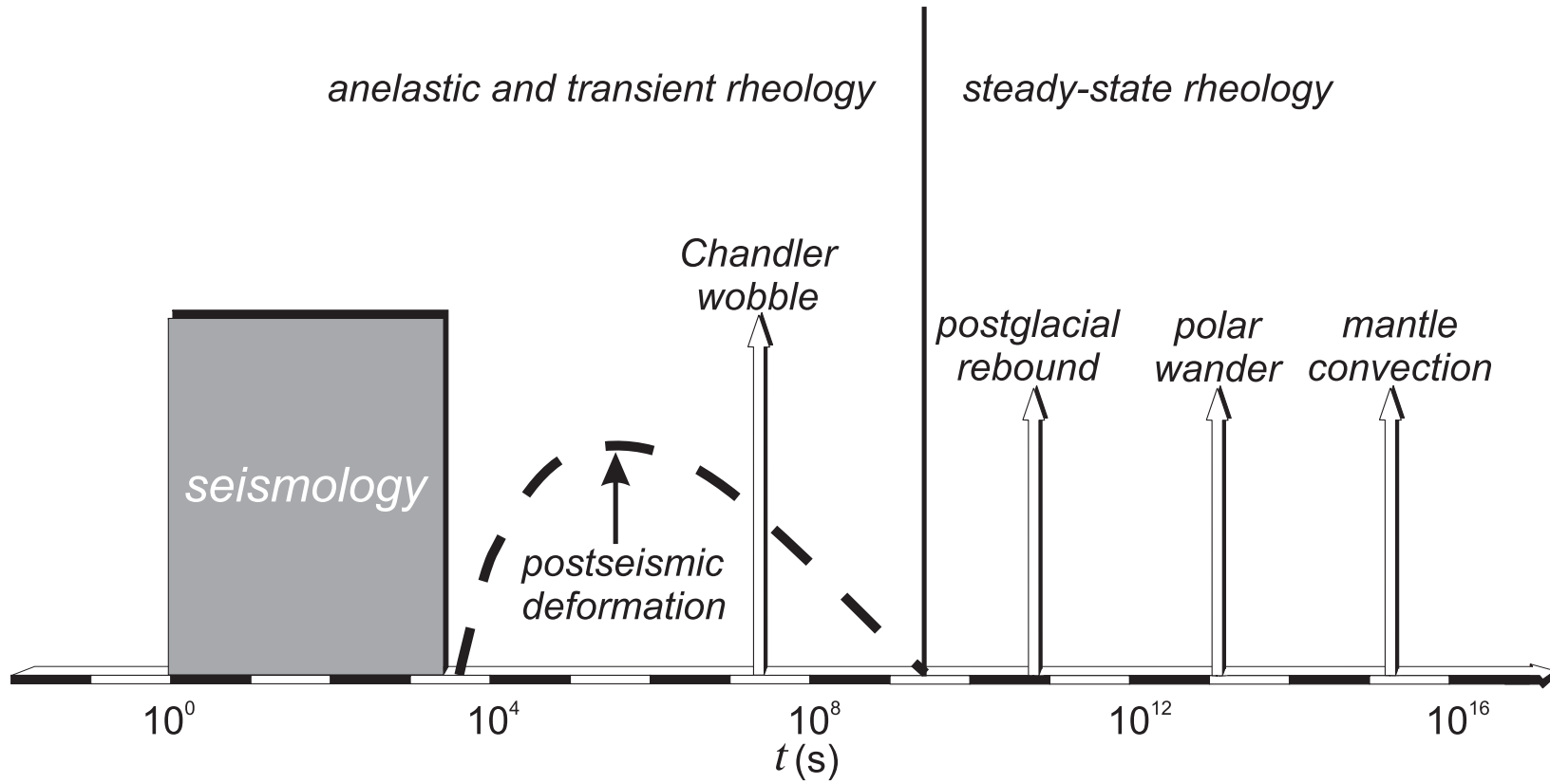
Understanding the Earth's Interior from Relaxation Normal Modes 1 – Basic theory

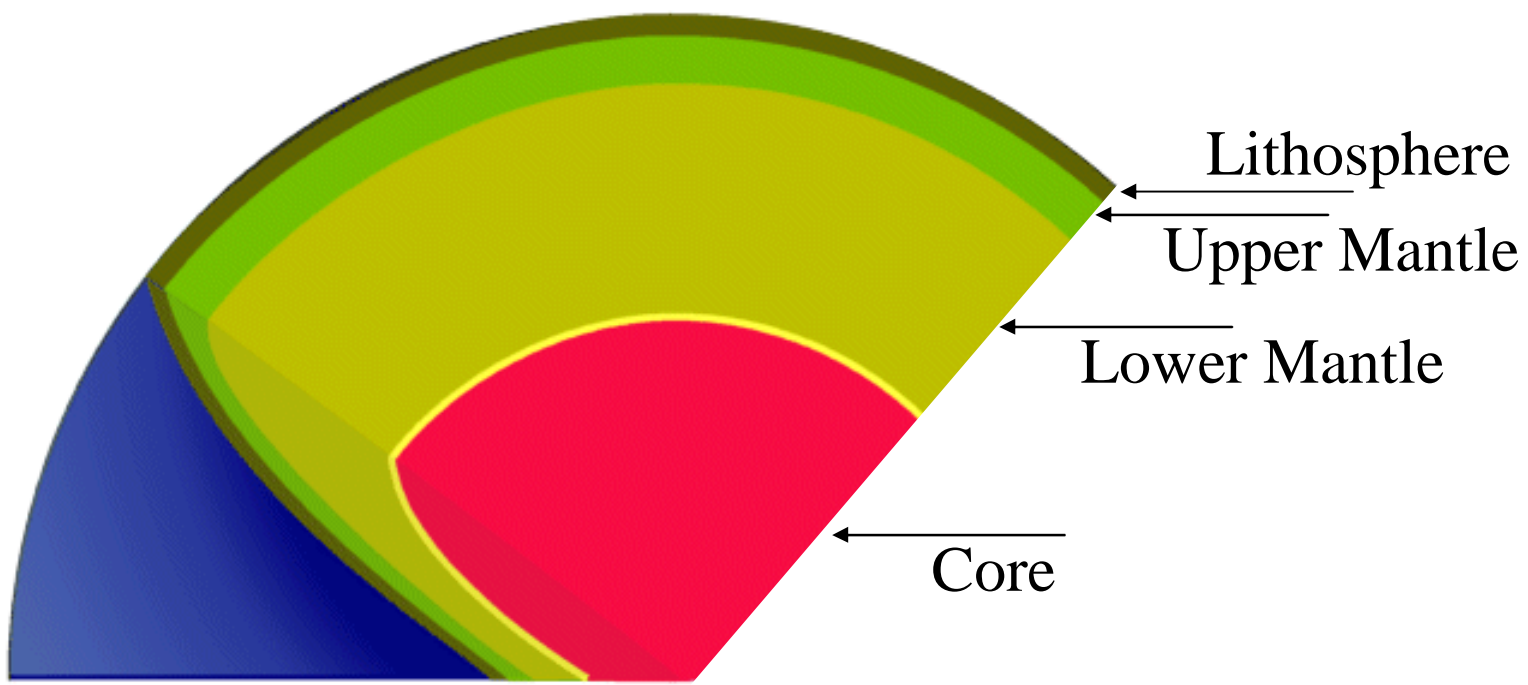
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$$\text{Maxwell time} = \nu_1 / \mu_1 \sim 0 (10^2 \text{ y})$$



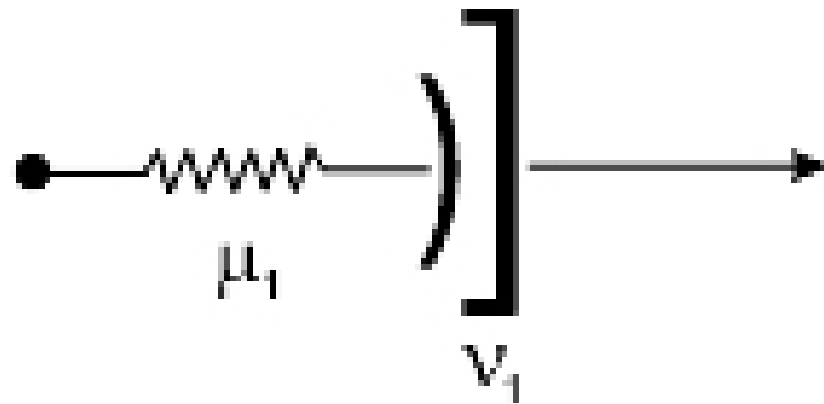


Lithosphere

Upper Mantle

Lower Mantle

Core



$$\left\{ \begin{array}{l} \nabla \cdot \boldsymbol{\sigma}' - \nabla(\rho g \mathbf{u} \cdot \hat{\mathbf{r}}) - \rho \nabla \phi' - \rho' g \hat{\mathbf{r}} + \mathbf{f} = 0 \\ \nabla^2 \phi' = 4\pi G (\rho' + \rho_f) \end{array} \right.$$

$$\dot{\sigma}_{ij} + \frac{\mu}{\nu} \left(\sigma_{ij} - \frac{1}{3} \sigma_{kk} \delta_{ij} \right) = 2 \mu \dot{\epsilon}_{ij} + \lambda \dot{\epsilon}_{kk} \delta_{ij}$$

$$\mathcal{L}[\sigma_{ij}] = 2 \hat{\mu}(s) \mathcal{L}[\epsilon_{ij}] + \hat{\lambda}(s) \mathcal{L}[\epsilon_{kk}] \delta_{ij}$$

$$\hat{\mu}(s) = \frac{\mu s}{s + \tau} \quad \hat{\lambda}(s) = \frac{\lambda s + \kappa \tau}{s + \tau} \quad \tau = \frac{\mu}{\nu} \quad \kappa = \lambda + \frac{2}{3} \mu$$

$$u(\mathbf{r}) = \sum_{n=2}^{\infty} U_n(r) P_n(\cos \theta)$$

$$v(\mathbf{r}) = \sum_{n=2}^{\infty} V_n(r) \partial_{\theta} P_n(\cos \theta)$$

$$\phi'(\mathbf{r}) = - \sum_{n=2}^{\infty} \phi_n(r) P_n(\cos \theta)$$

$$P_n(x) = \frac{1}{2^n n!} \frac{d^n (x^2 - 1)^n}{dx^n}$$

$$y(r, n, s) = \begin{pmatrix} \tilde{U}_n \\ \tilde{V}_n \\ \hat{\lambda} \tilde{\chi}_n + 2 \hat{\mu} \partial_r \tilde{U}_n \\ \hat{\mu} \left(\partial_r \tilde{V}_n + \frac{1}{r} \tilde{U}_n - \frac{1}{r} \tilde{V}_n \right) \\ -\tilde{\phi}_n \\ -\partial_r \tilde{\phi}_n - \frac{n+1}{r} \tilde{\phi}_n + 4 \pi G \rho \tilde{U}_n \end{pmatrix}$$

$$\nabla \cdot \mathbf{u} = \sum_{n=2}^{\infty} \chi_n(r) P_n(\cos \theta)$$

$$\chi_n(r) = \partial_r U_n + \frac{2}{r} U_n - \frac{n(n+1)}{r} V_n$$

$$\partial_r \mathbf{y}(r, s, n) = \mathbf{A}(r, s, n) \mathbf{y}(r, s, n) + \delta(r - r_S) \mathbf{f}(n)$$

$$\mathbf{A}(r, s, n) = \begin{pmatrix} \frac{-2\hat{\lambda}}{r\beta} & \frac{N\hat{\lambda}}{r\beta} & \frac{1}{\beta} & 0 & 0 & 0 \\ -\frac{1}{r} & \frac{1}{r} & 0 & \frac{1}{\hat{\mu}} & 0 & 0 \\ -\frac{4g\rho}{r} + \frac{4\gamma}{r^2\beta} & \frac{Ng\rho}{r} - \frac{2N\gamma}{r^2\beta} & -\frac{4\hat{\mu}}{r\beta} & \frac{N}{r} & -\frac{(1+n)\rho}{r} & \rho \\ \frac{g\rho}{r} - \frac{2\gamma}{r^2\beta} & \frac{4N\hat{\mu}(\hat{\lambda}+\hat{\mu})}{r^2\beta} - \frac{2\hat{\mu}}{r^2} & -\frac{\hat{\lambda}}{r\beta} & -\frac{3}{r} & \frac{\rho}{r} & 0 \\ -4\pi G\rho & 0 & 0 & 0 & -\frac{n+1}{r} & 1 \\ -\frac{4\pi G(n+1)\rho}{r} & \frac{4\pi GN\rho}{r} & 0 & 0 & 0 & \frac{n-1}{r} \end{pmatrix}$$

$$N = n(n+1)$$

Green functions - Incompressible

$$\mathbf{Y}(r, s, n) = [\mathbf{Y}_R \ \mathbf{Y}_I]$$

$$\mathbf{Y}_R(r, s, n) = r^n \begin{pmatrix} \frac{nr}{2(2n+3)} & \frac{1}{r} & 0 \\ \frac{(n+3)r}{2(2n^2+5n+3)} & \frac{1}{nr} & 0 \\ \frac{2((n-1)n-3)\mu+gnr\rho}{2(2n+3)} & \frac{2(n-1)\mu+gr\rho}{r^2} & -\rho \\ \frac{n(n+2)\mu}{(n+1)(2n+3)} & \frac{2(n-1)\mu}{nr^2} & 0 \\ 0 & 0 & -1 \\ \frac{2Gn\pi r\rho}{3+2n} & \frac{4G\pi\rho}{r} & -\frac{1+2n}{r} \end{pmatrix} \quad \mathbf{Y}_I(r, s, n) = \frac{1}{r^n} \begin{pmatrix} \frac{n+1}{2(2n-1)} & r^{-2} & 0 \\ -\frac{n-2}{2n(2n-1)} & -\frac{1}{(n+1)r^2} & 0 \\ \frac{g(n+1)r\rho-2(n(3+n)-1)\mu}{2(2n-1)r} & \frac{gr\rho-2(2+n)\mu}{r^3} & -\frac{\rho}{r} \\ \frac{\mu(n^2-1)}{rn(2n-1)} & \frac{2(2+n)\mu}{(n+1)r^3} & 0 \\ 0 & 0 & -\frac{1}{r} \\ \frac{2G\pi(n+1)\rho}{2n-1} & \frac{4G\pi\rho}{r^2} & 0 \end{pmatrix}$$

$$\mathbf{Y}_j(R_{j+1}, s, n)\mathbf{C}_j = \mathbf{Y}_{j+1}(R_{j+1}, s, n)\mathbf{C}_{j+1}$$

$$y_{omo}(r, s, n) = \mathbf{D}(r, s, n)y_C$$

$$\tilde{\mathbf{X}}(r, s, n) = \frac{[\mathbf{P}_2\mathbf{D}(a, s, n)\mathbf{I}_C(n)] [\mathbf{P}_1\mathbf{D}(a, s, n)\mathbf{I}_C(n)]^\dagger \mathbf{b}(s, n)}{\Delta_{sec}(s, n)}$$

$$\mathbf{P}_2\mathbf{y}(r, s, n) = \begin{pmatrix} \tilde{U}(r, s, n) \\ \tilde{V}(r, s, n) \\ -\tilde{\Phi}(r, s, n) \end{pmatrix} = \tilde{\mathbf{X}}(r, s, n)$$

$$\mathbf{I}_C(n) = \begin{pmatrix} -\frac{3}{4\pi G\rho_C}r_C^{n-1} & 0 & 1 \\ 0 & 1 & 0 \\ 0 & 0 & \frac{4\pi G\rho_C^2}{3}r_C \\ 0 & 0 & 0 \\ r_C^n & 0 & 0 \\ 2(n-1)r_C^{n-1} & 0 & 4\pi G\rho_C \end{pmatrix}$$

$$\Delta_{sec}(s, n) = \det [\mathbf{P}_1\mathbf{D}(a, s, n)\mathbf{I}_C(n)]$$

$$\begin{pmatrix} U(r, t, n) \\ V(r, t, n) \\ -\Phi(r, t, n) \end{pmatrix} = \mathbf{X}(r, t, n) = \int_{s_0 - i\infty}^{s_0 + i\infty} \tilde{\mathbf{X}}(r, s, n) e^{st} ds = \mathbf{k}_E \delta(t) + \sum \mathbf{k}_j e^{s_j t}$$

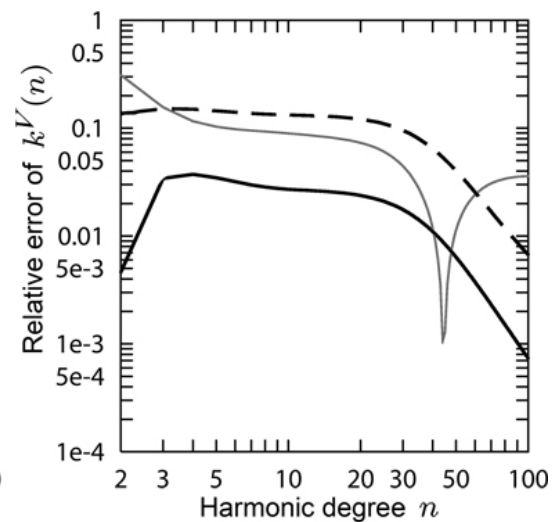
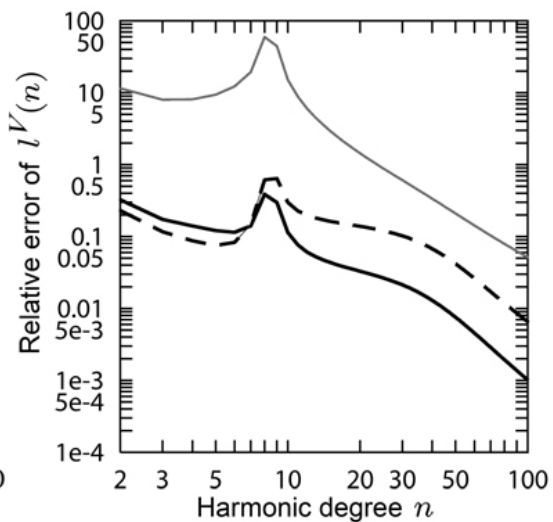
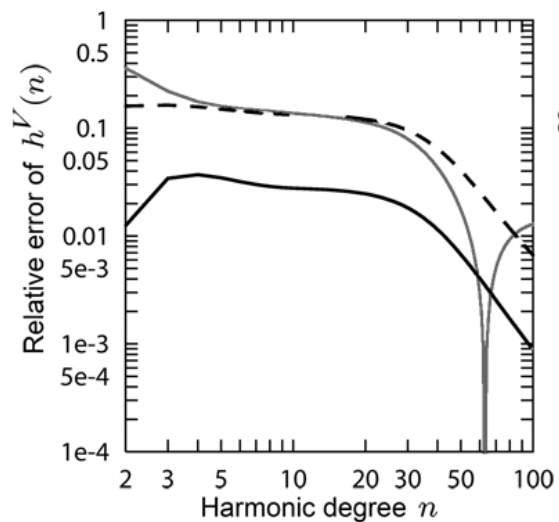
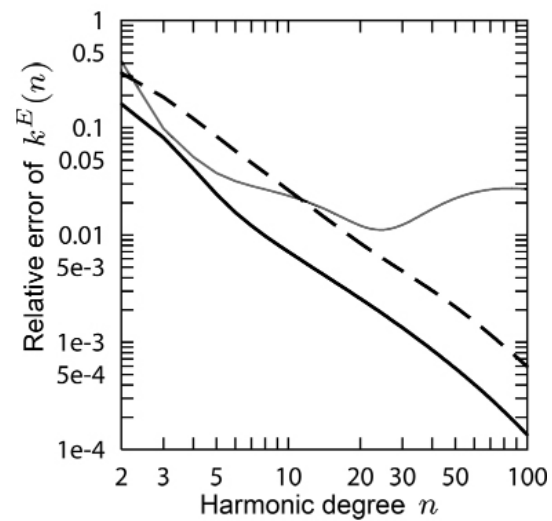
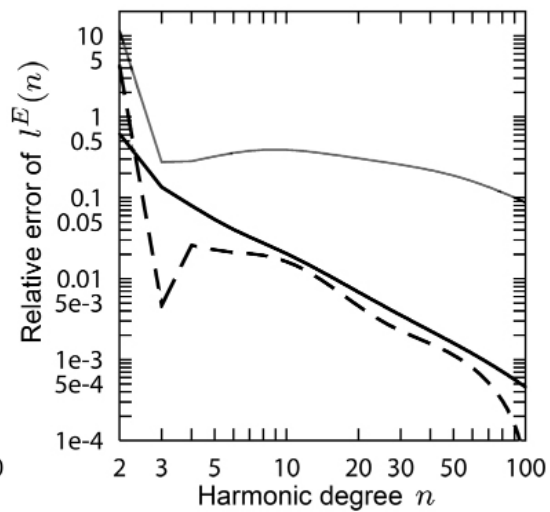
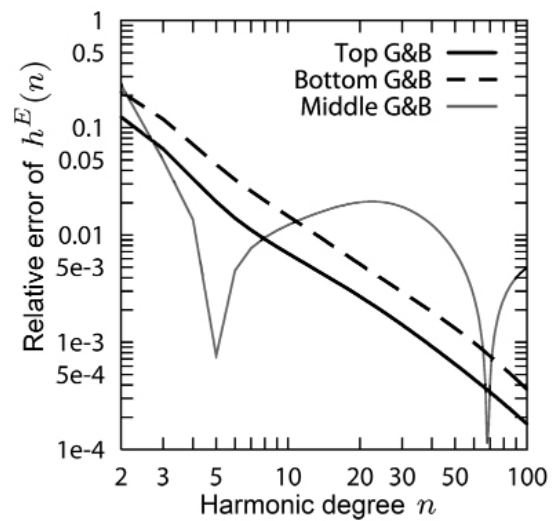
$$\mathbf{k}_E = \lim_{s \rightarrow -\infty} \tilde{\mathbf{X}}(r, s, n)$$

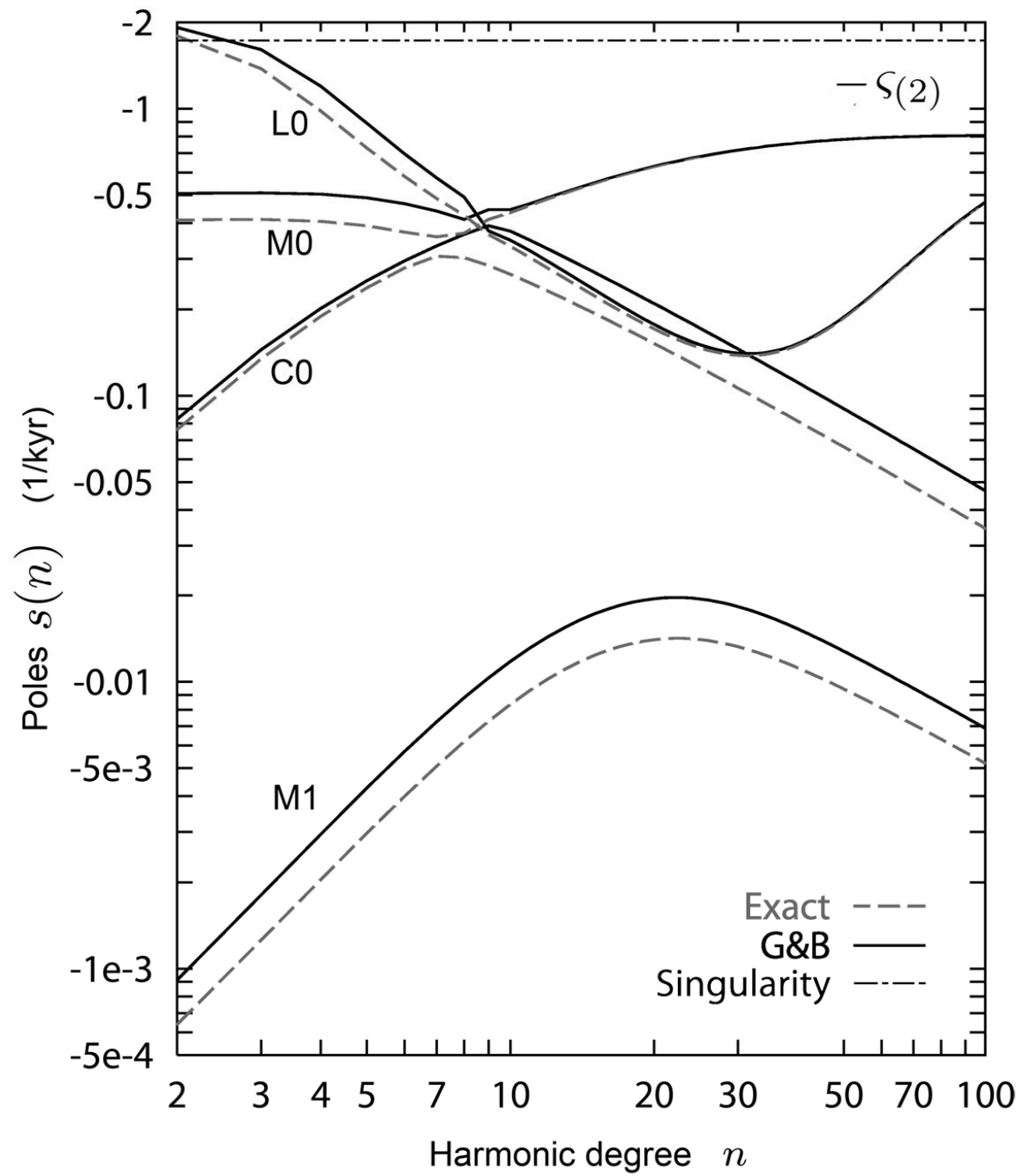
$$\mathbf{k}_j = \lim_{s \rightarrow s_j} (s - s_j) \tilde{\mathbf{X}}(r, s, n)$$

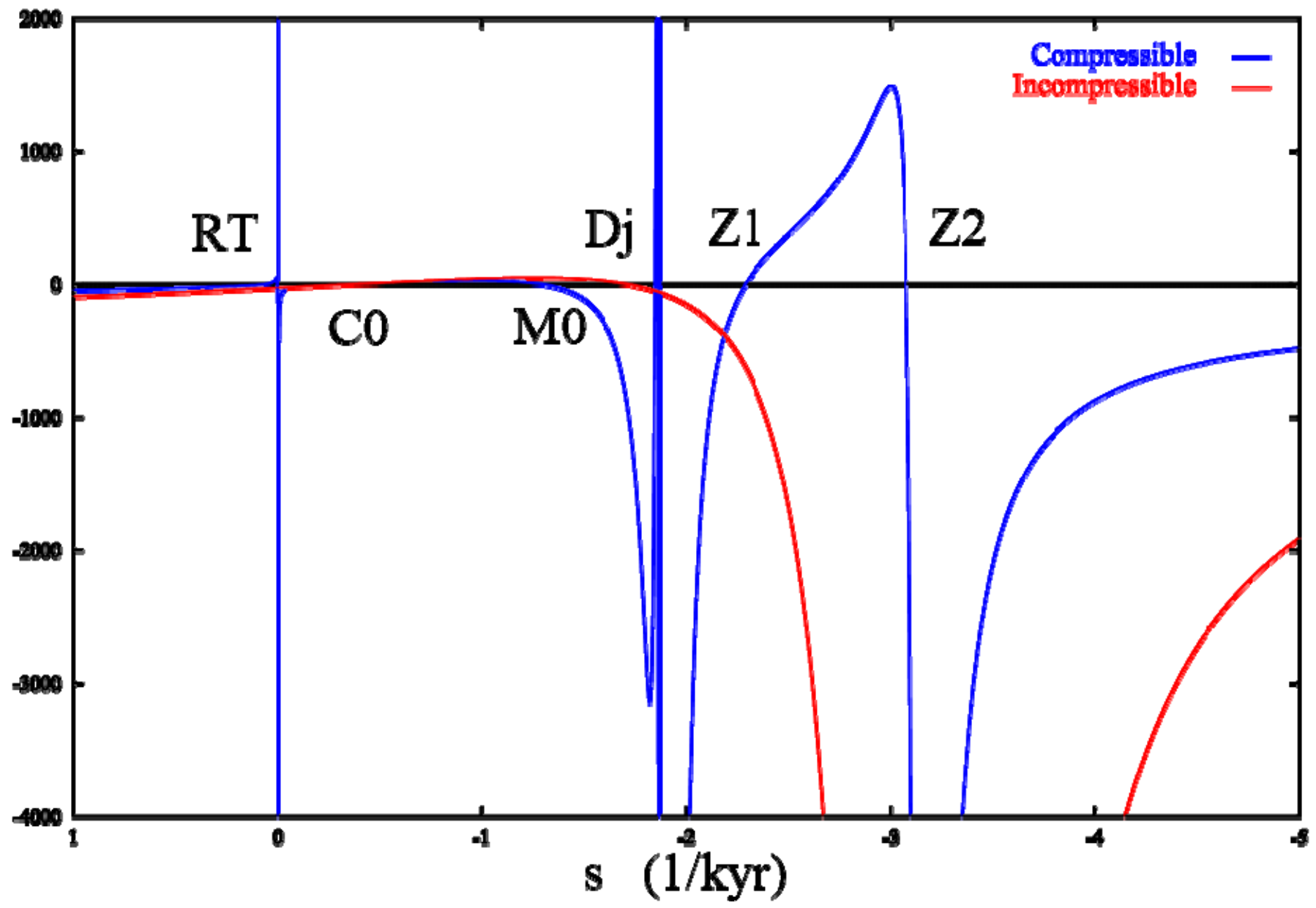
Compressible (approximated) model Helmholtz equation

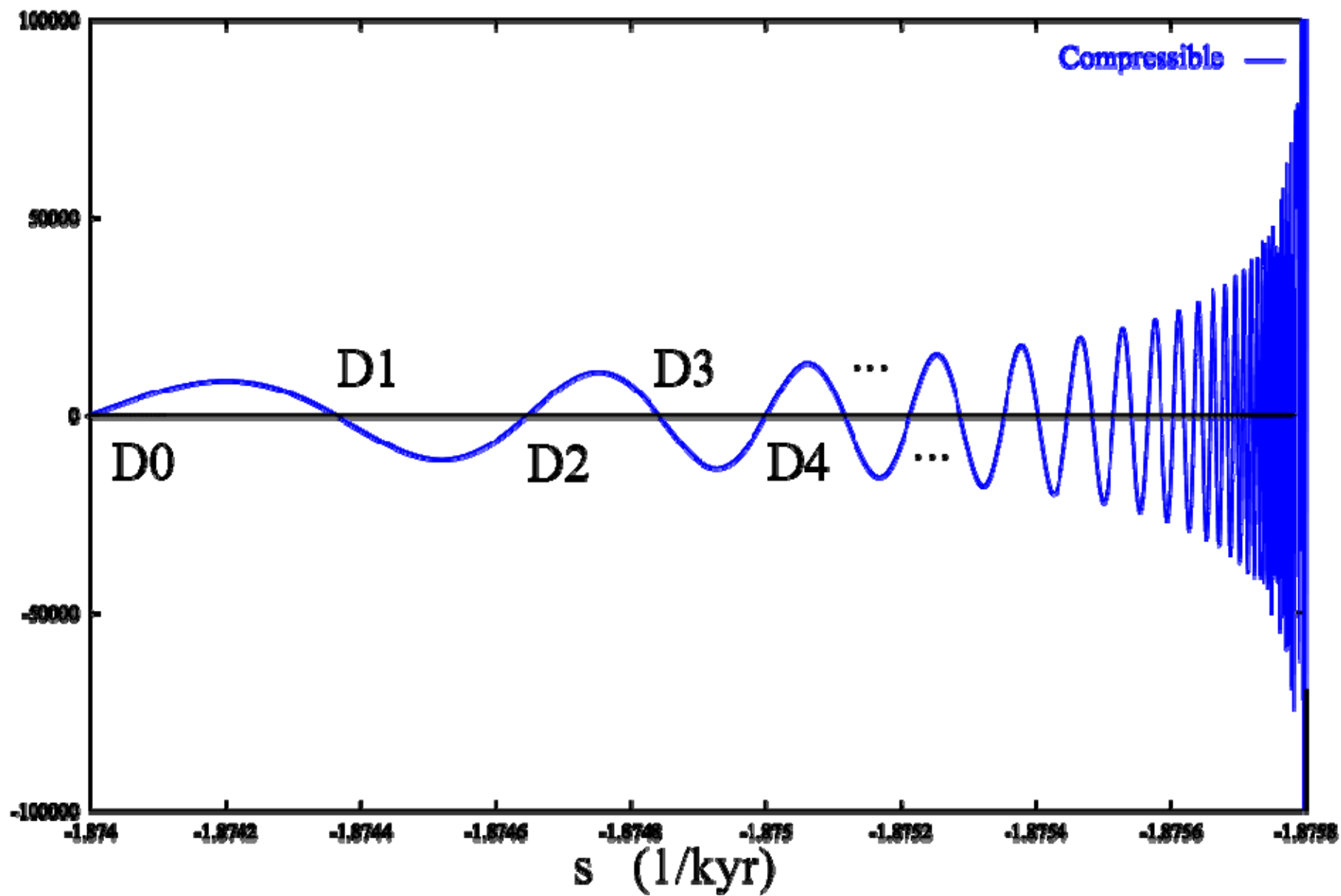
$$\mathbf{y}_k(r, s, n) = \begin{pmatrix} -\frac{NC}{k^2 r} J(kr) - \partial_r J(kr) \\ -\frac{1+C}{k^2 r} J(kr) - C \partial_r J(kr) \\ \hat{\mu} \frac{2N(1+C) + k^2 r^2 \beta}{k^2 r^2} J(kr) - \hat{\mu} \frac{2(NC-2)}{r} \partial_r J(kr) \\ \hat{\mu} \frac{2 + (k^2 r^2 - 2N + 2)C}{k^2 r^2} J(kr) + \hat{\mu} \frac{2(C-1)}{r} \partial_r J(kr) \\ \frac{\zeta}{k^2} J(kr) \\ -\frac{(n+1)\zeta(nC-1)}{k^2 r} J(kr) \end{pmatrix}$$

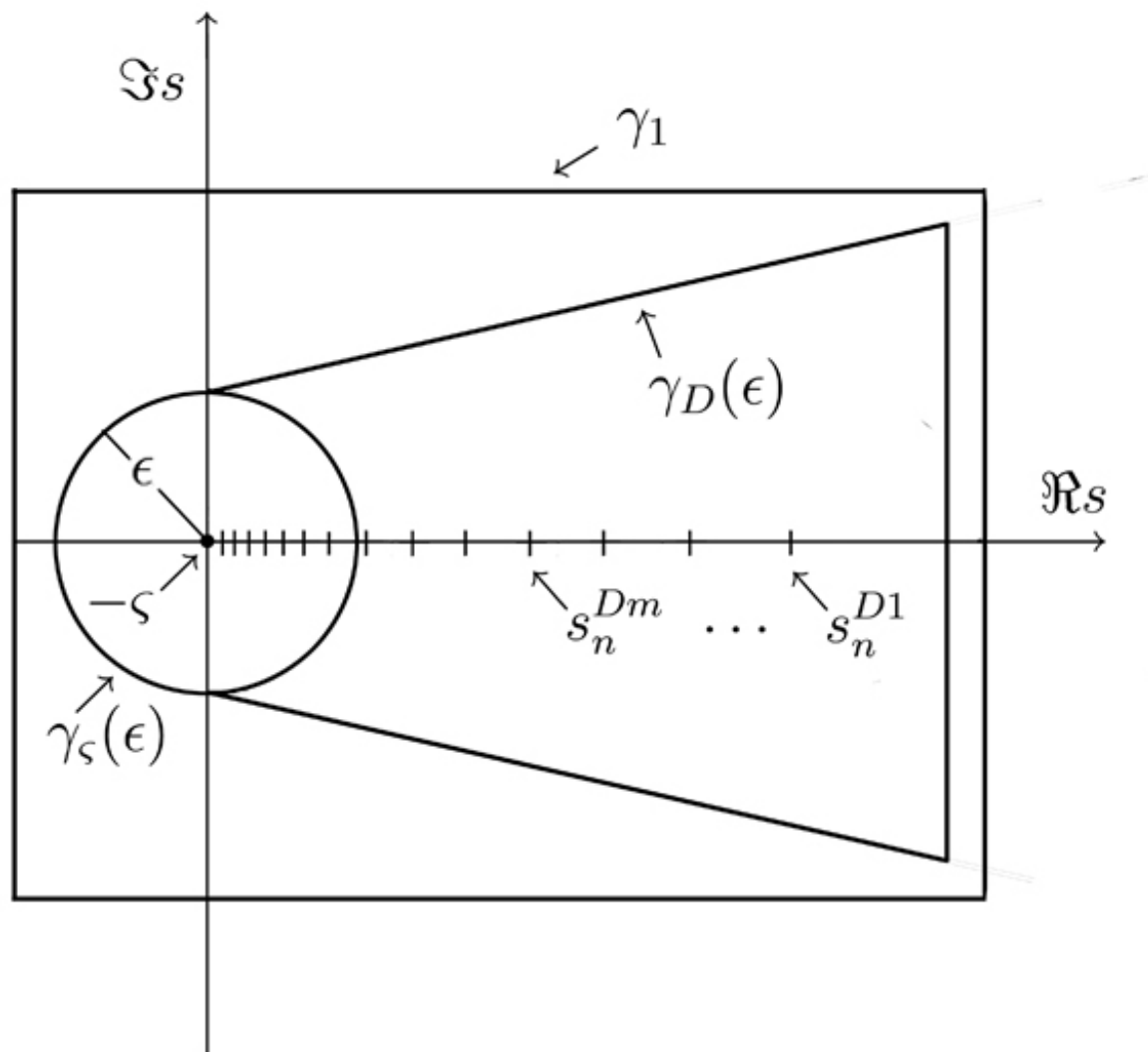
$$\xi(r) = \frac{g(r)}{r} \quad \rightarrow \quad \bar{\xi}(r) = \xi_j$$

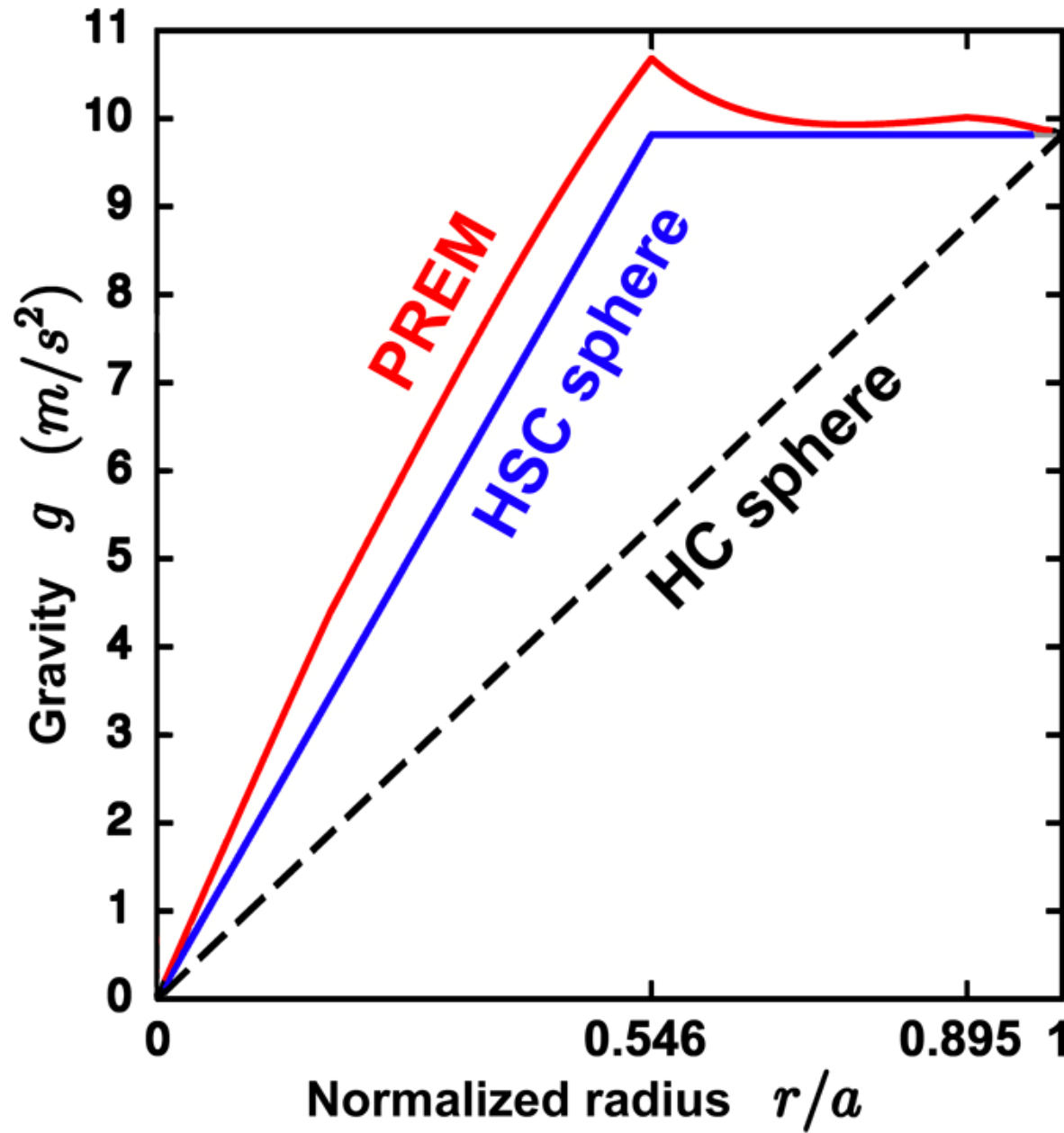












*Williamson-Adams
equation*

$$d_r \rho + \frac{g \rho^2}{\kappa} - \lambda = 0$$

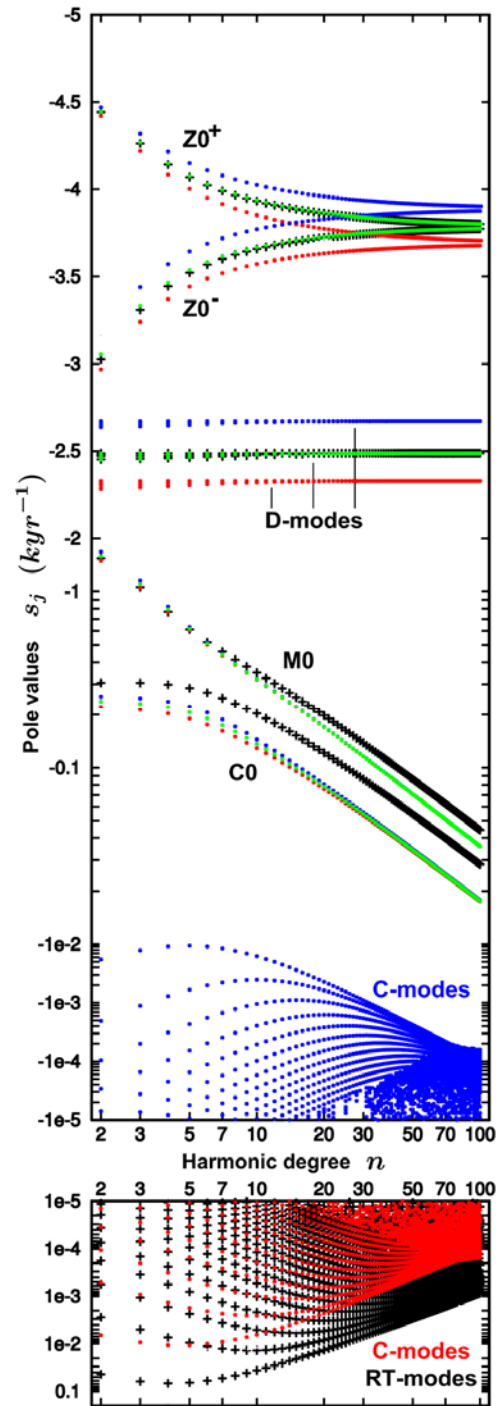
$$\bar{\lambda} \approx \frac{g \rho^2}{\kappa} > 0$$

*Continuous
density profile*

$$\rho = \frac{\alpha}{r}$$

$$g = 2\pi G \alpha$$

$$\kappa_{SC} = 2\pi G \alpha^2$$



***A new class of modes
Compositional C-modes***

Fluid limit

$$\begin{pmatrix} U(r, t, n) \\ V(r, t, n) \\ -\Phi(r, t, n) \end{pmatrix} = \mathbf{X}(r, t, n) = \int_{s_0 - i\infty}^{s_0 + i\infty} \tilde{\mathbf{X}}(r, s, n) e^{st} ds = \mathbf{k}_E \delta(t) + \sum \mathbf{k}_j e^{s_j t}$$

$$\bar{k}_n^\infty = \lim_{t \rightarrow \infty} \bar{k}_n(t) = \bar{k}_E - \sum \frac{\bar{k}_j}{s_j} = \bar{k}_n^{ISO}$$

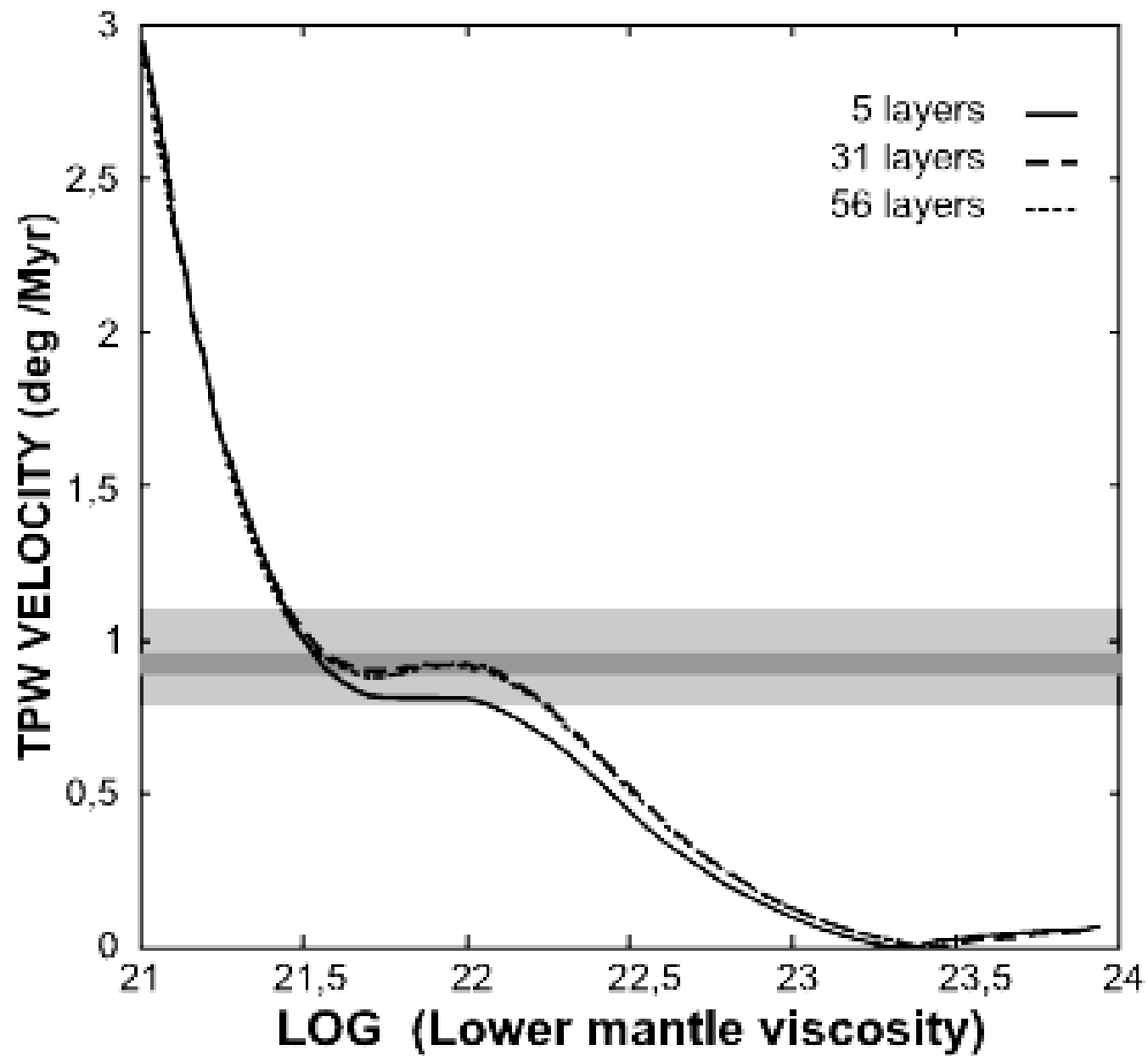


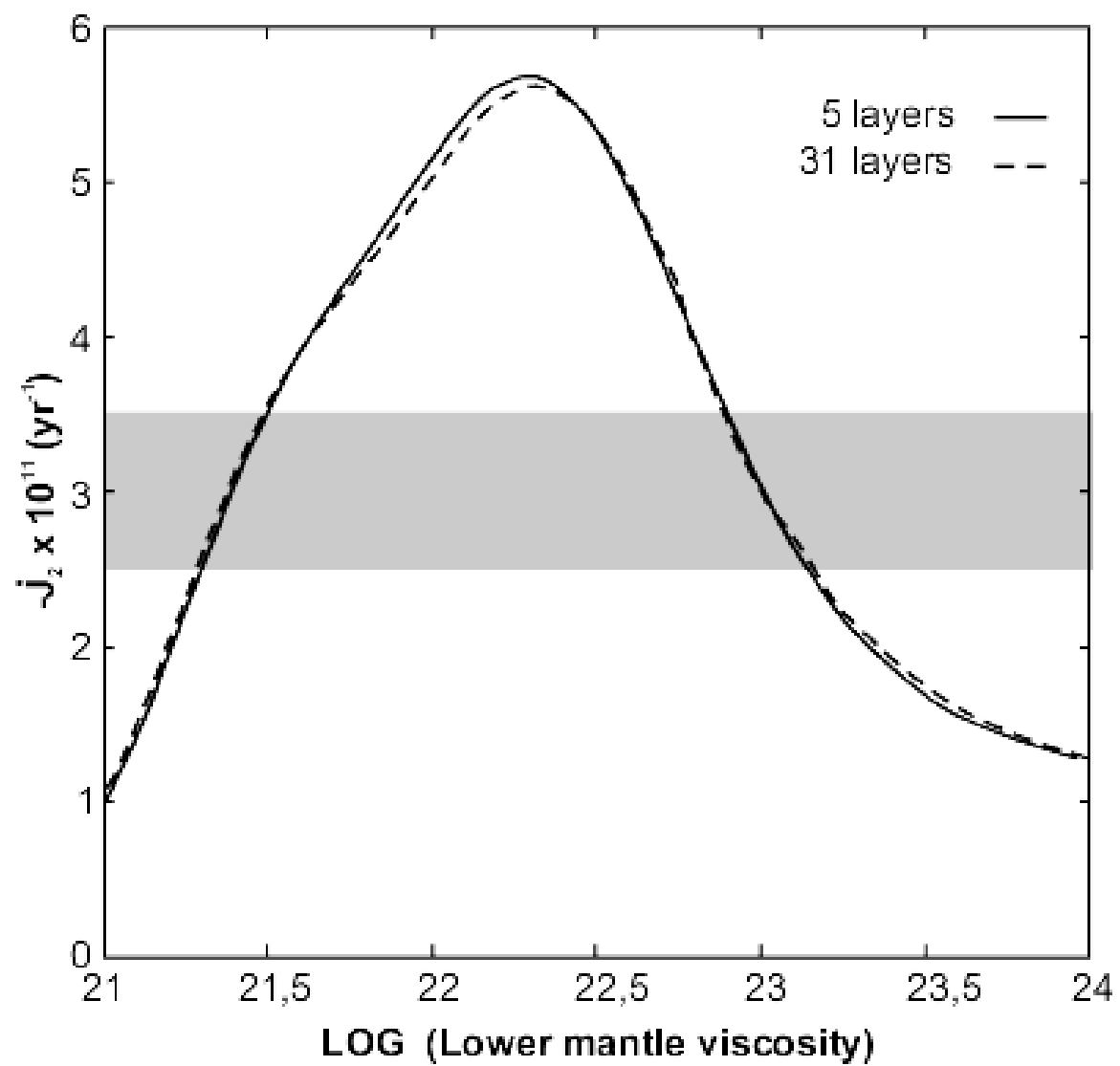
Understanding the Earth's Interior from Relaxation Normal Modes 2- Long term Earth's rotation

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$$\mathbf{J} = \mathbf{J}^\omega + \mathbf{J}^\delta$$

$$J_{ij}^\omega = I \delta_{ij} + \frac{a^5}{3G} k^T \star \left(\omega_i \omega_j - \frac{1}{3} \omega^2 \delta_{ij} \right)$$

$$k_F^T = \lim_{t \rightarrow \infty} k^T(t) \star H(t)$$

$$\mathbf{J}^\omega(t=0) = \text{Diag}[A, A, C]$$

$$C = \frac{2}{3} \frac{a^5 \Omega^2}{3G} k_F^T$$

$$A = -\frac{1}{3} \frac{a^5 \Omega^2}{3G} k_F^T$$

$$k_F^T = \frac{3G(C - A)}{a^5 \Omega^2}$$

$$\mathbf{J}^\omega = \text{Diag}[A, A, C] + \Delta \mathbf{I}^\omega$$

$$\Delta I_{j3}^\omega = \frac{a^5 \Omega^2}{3G} k^T(t) \star m_j(t) (1 + m_3(t)) \quad j = 1, 2$$

$$\frac{is}{\sigma_r} \tilde{\mathbf{m}}(s) + \left(1 - \frac{\tilde{\mathbf{k}}^T(s)}{k_F^T}\right) \tilde{\mathbf{m}}(s) = \left(1 + \tilde{k}_L(s)\right) \tilde{\phi}(s)$$

$$\tilde{k}(s) = k_E + \sum \frac{k_j}{s - s_j}$$

$$k_F = k_E - \sum \frac{k_j}{s_j}$$

$$\tilde{\mathbf{m}}(s) = \frac{1 + k_F^L + s \sum \frac{k_j^L}{s_j(s - s_j)}}{s \left(\frac{i}{\sigma_r} - \frac{1}{k_F^T} \sum \frac{k_j^T}{s_j(s - s_j)} \right)} \tilde{\phi}(s)$$

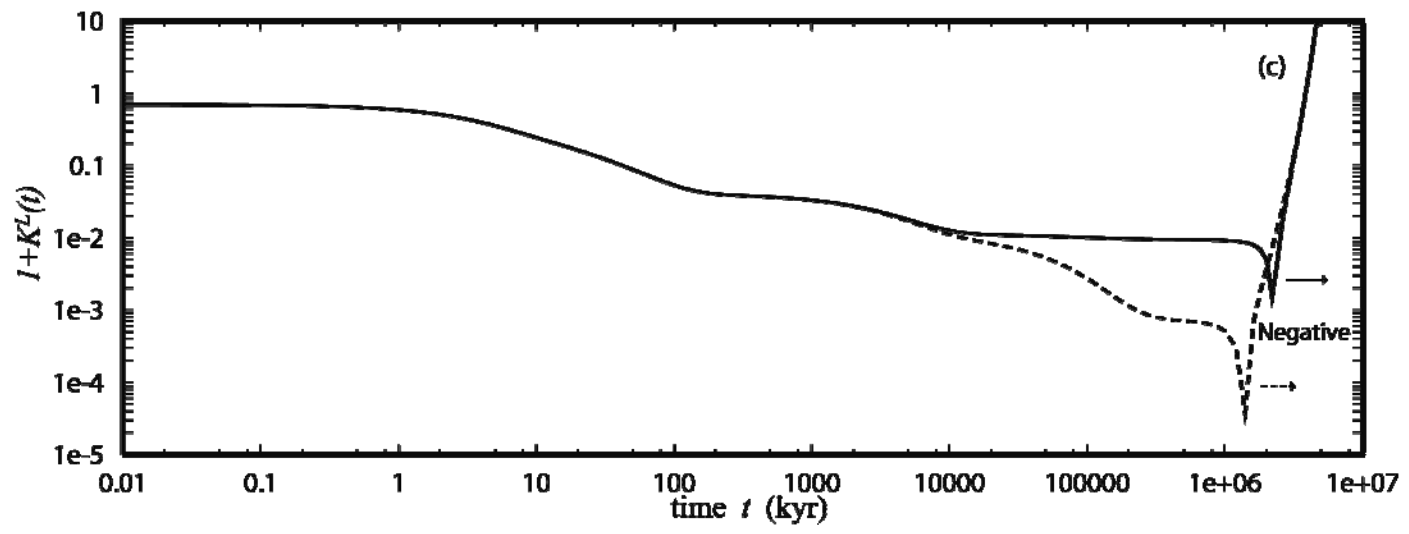
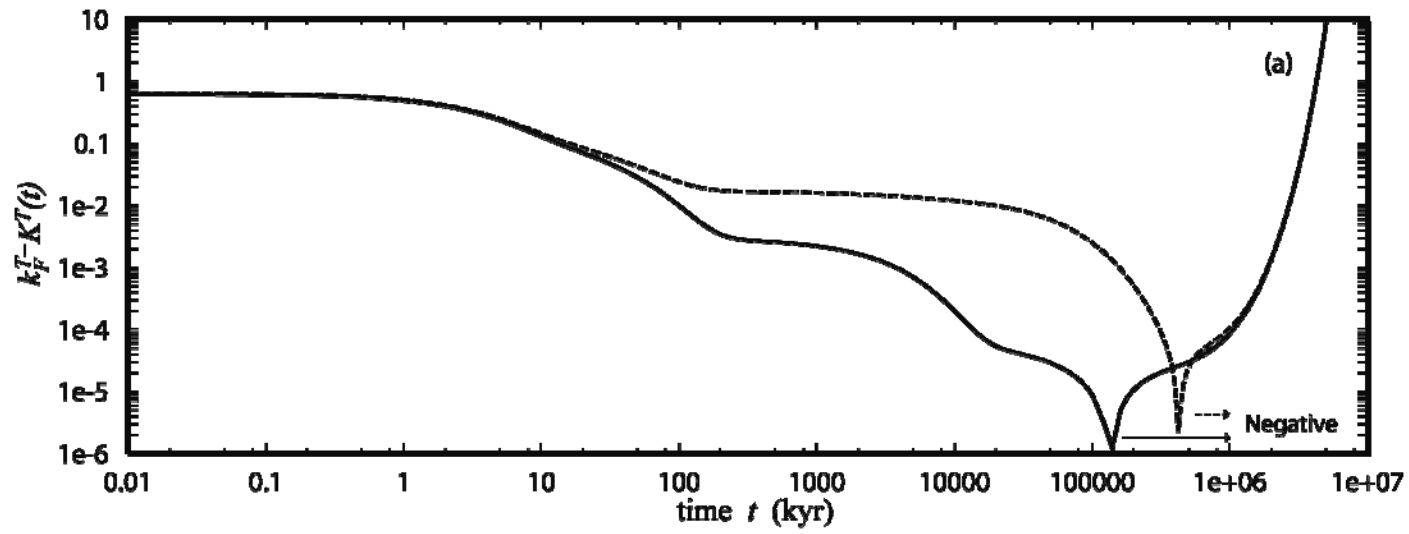
$$C \rightarrow C + \Delta I_{33}^C \qquad A \rightarrow A + \frac{\Delta I_{11}^C + \Delta I_{22}^C}{2}$$

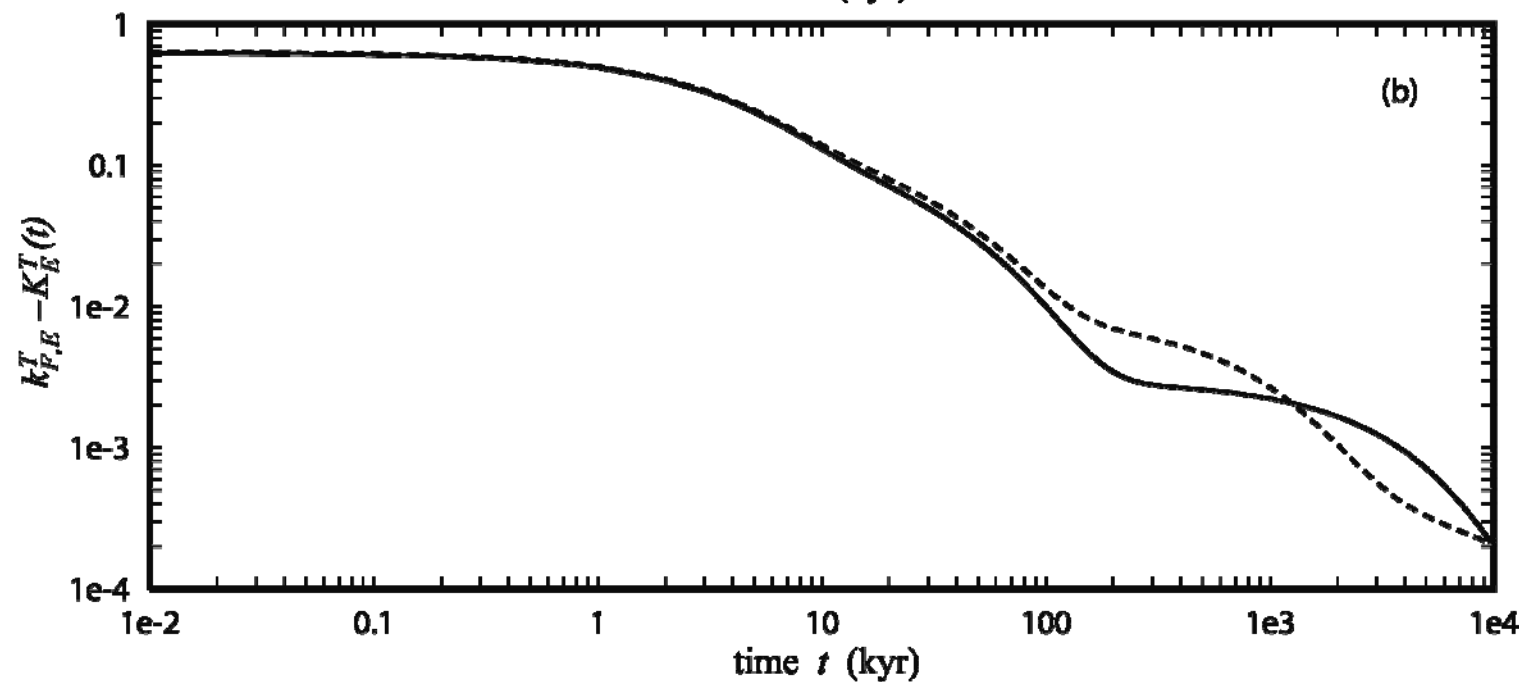
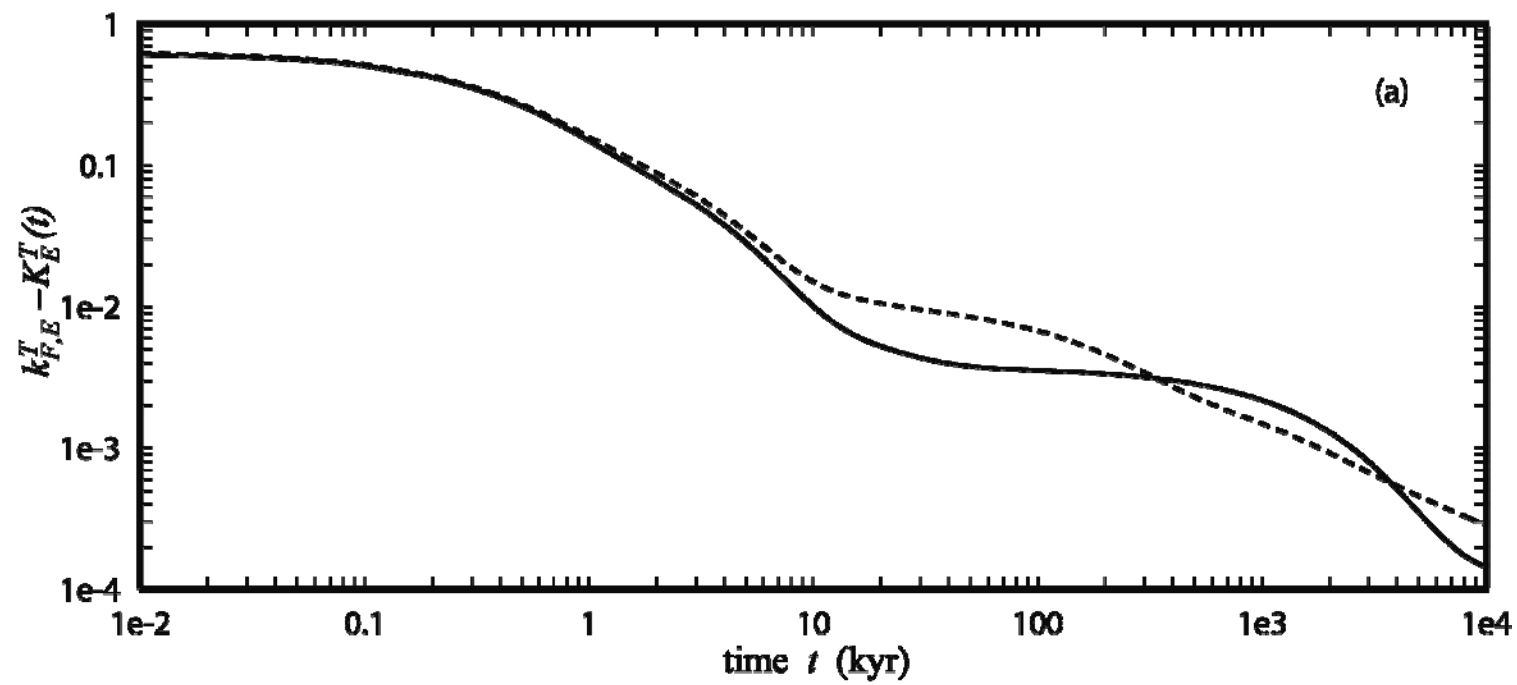
$$k_{F,obs}^T = k_F^T + \beta = \frac{3G(C - A)}{a^5 \Omega^2}$$

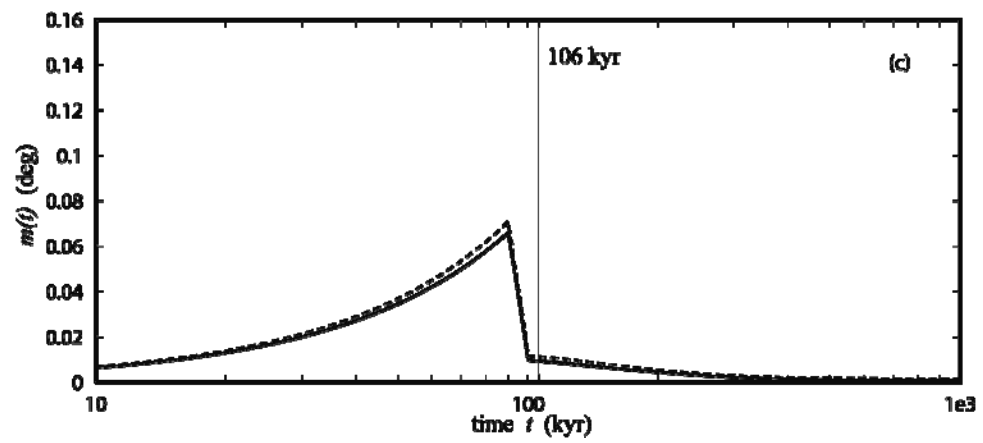
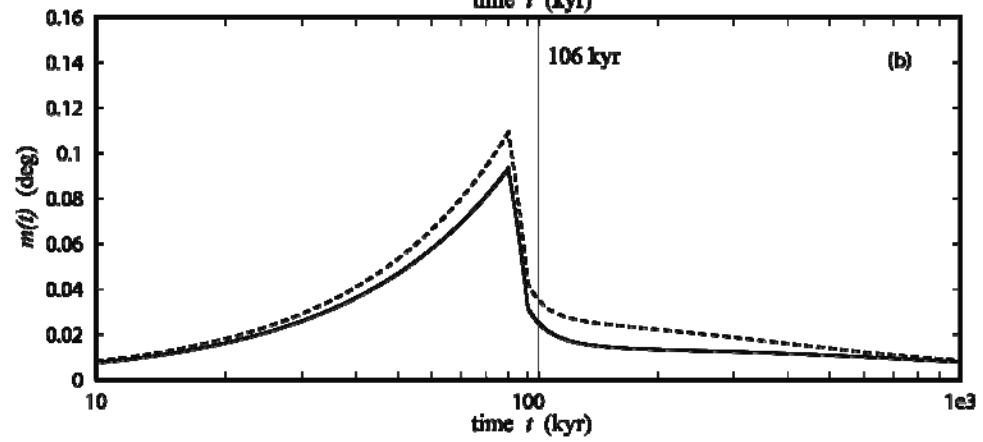
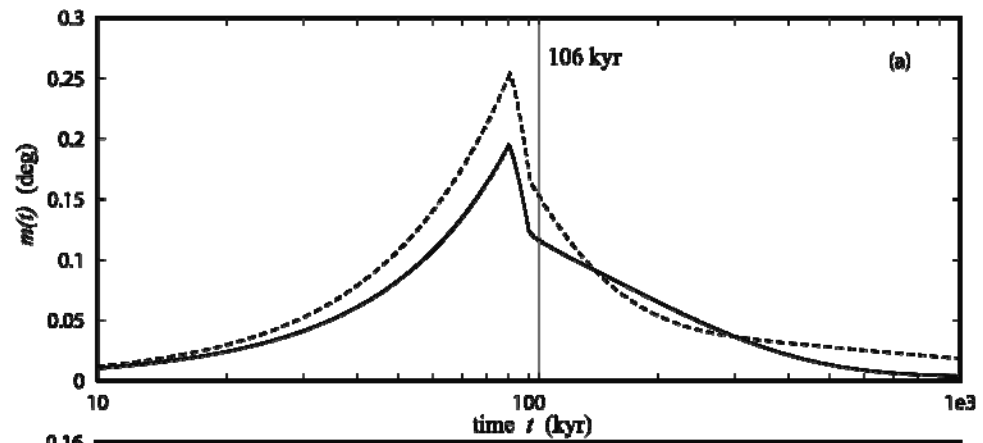
$$\tilde{\mathbf{m}}(s) = \frac{1 + k_F^L + s \sum \frac{k_j^L}{s_j(s-s_j)}}{\frac{\beta}{k_F^T + \beta} + s \left(\frac{i}{\sigma_r} - \frac{1}{k_F^T + \beta} \sum \frac{k_j^T}{s_j(s-s_j)} \right)} \tilde{\phi}(s)$$

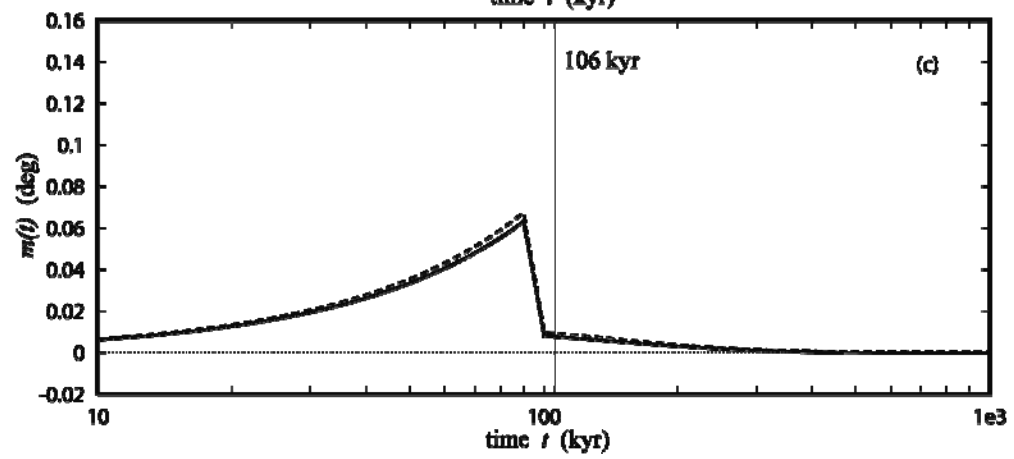
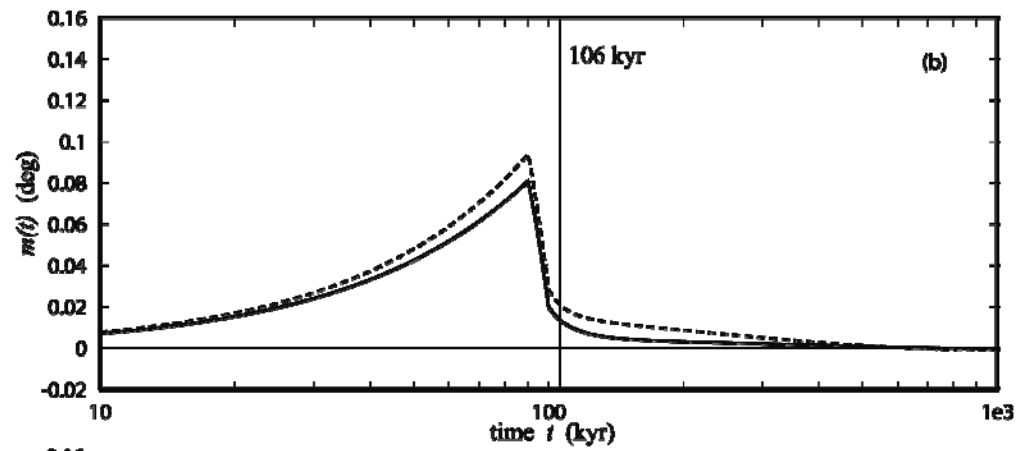
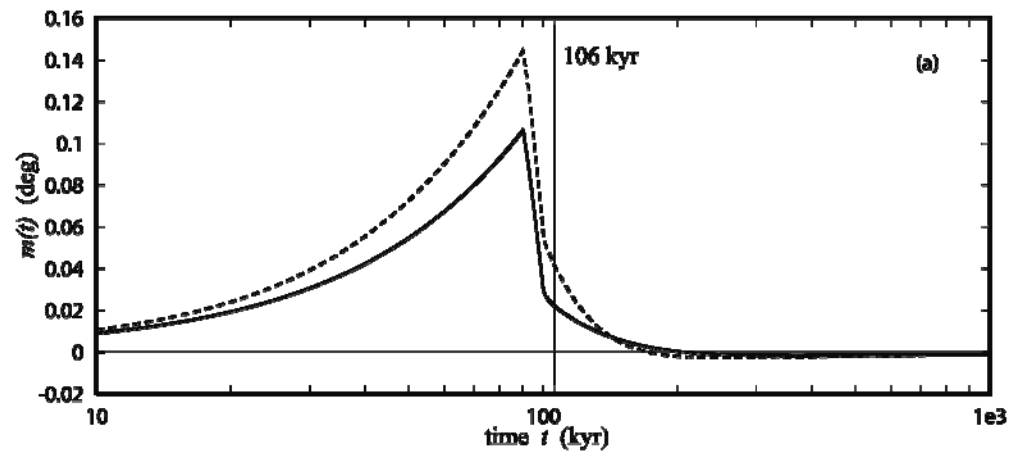
$$\mathbf{K}(f; n, t) = \mathcal{L}^{-1} \left[\tilde{\mathbf{k}}(n, s) f(s) \right] = \int_{\gamma} \tilde{\mathbf{k}}(n, s) \tilde{f}(s) e^{st} ds$$

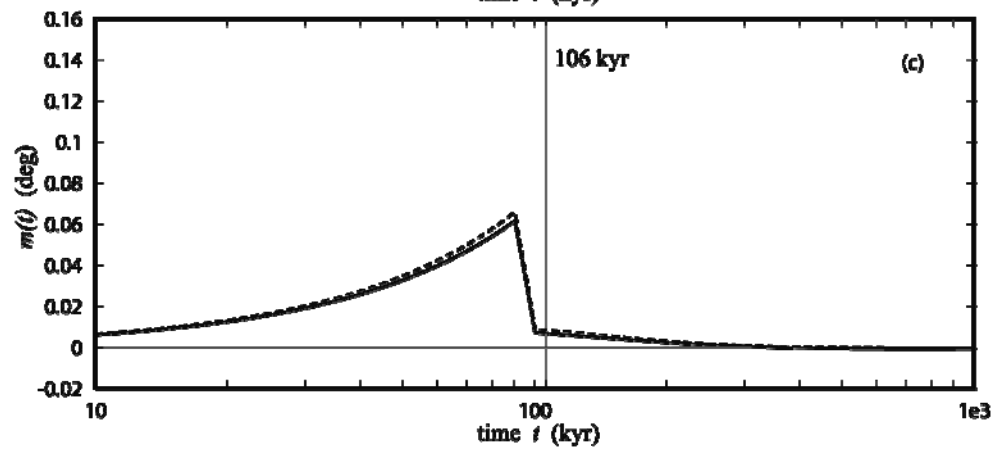
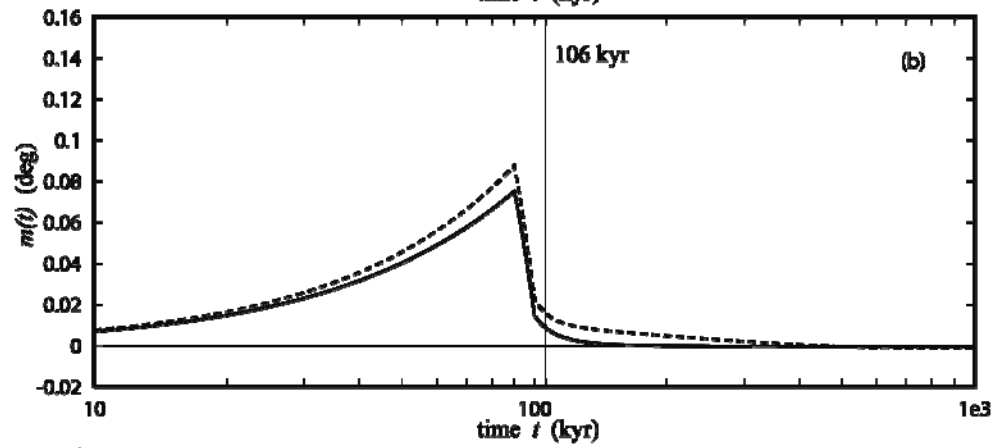
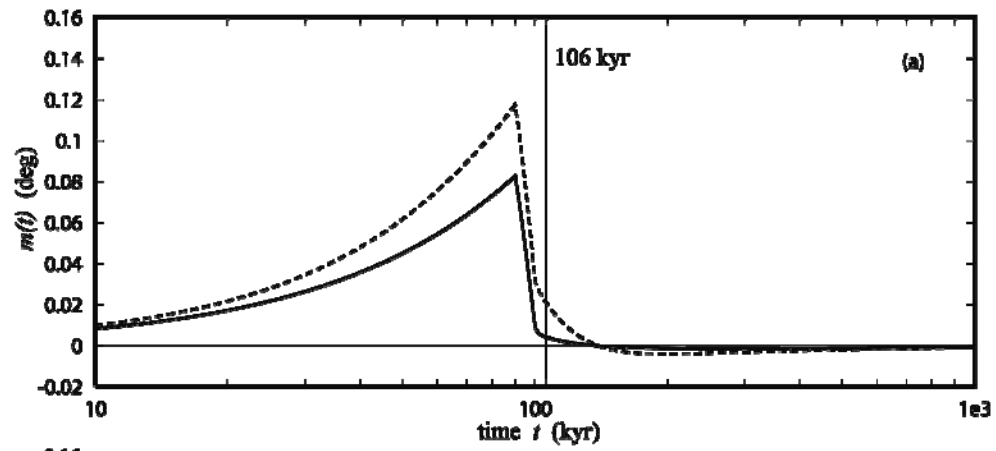
$$\mathbf{m}(t) = \int_{\gamma} \frac{1 + \tilde{k}_L(s)}{1 - \frac{\tilde{k}^T(s)}{k_F^T + \beta}} \tilde{\phi}(s) e^{st} ds$$

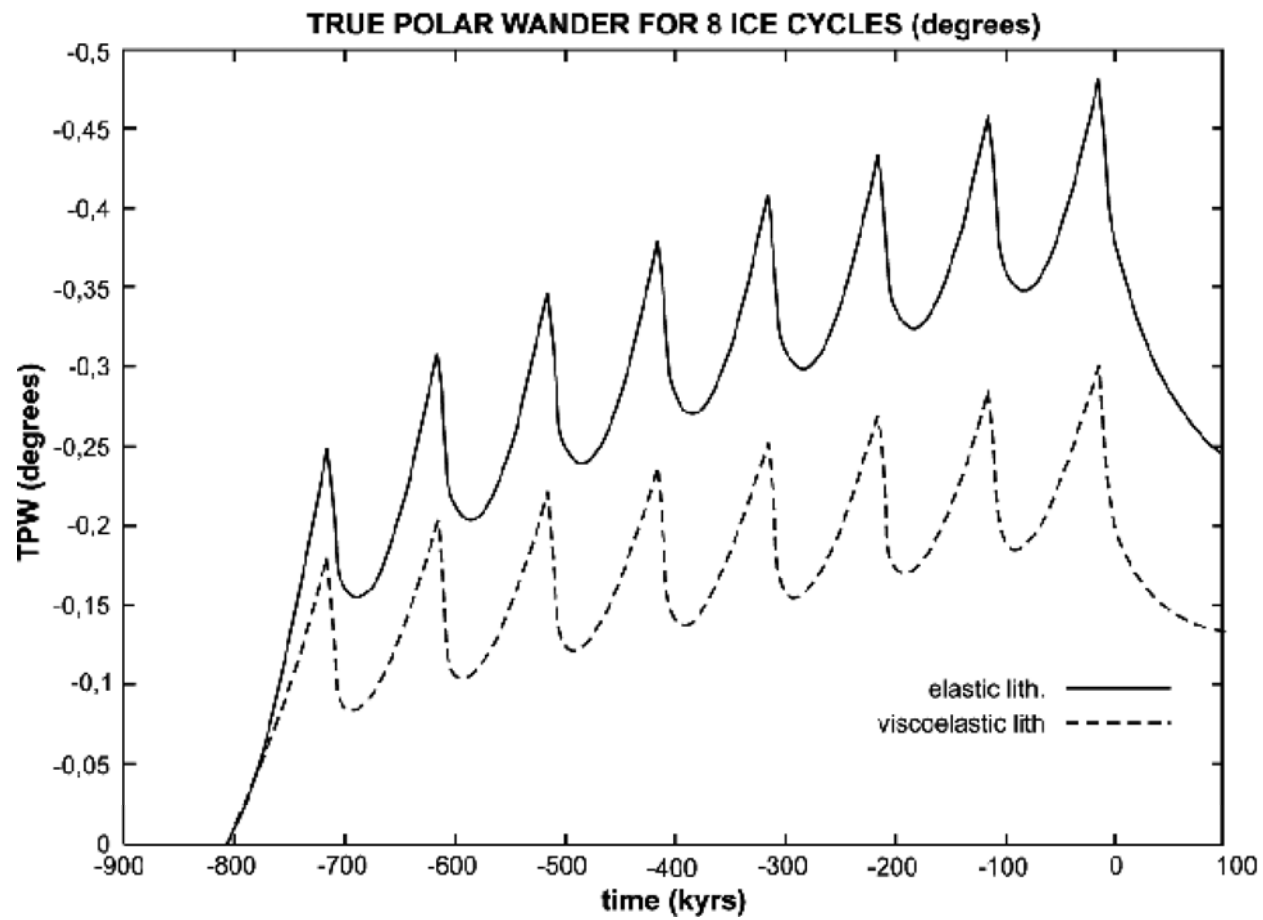


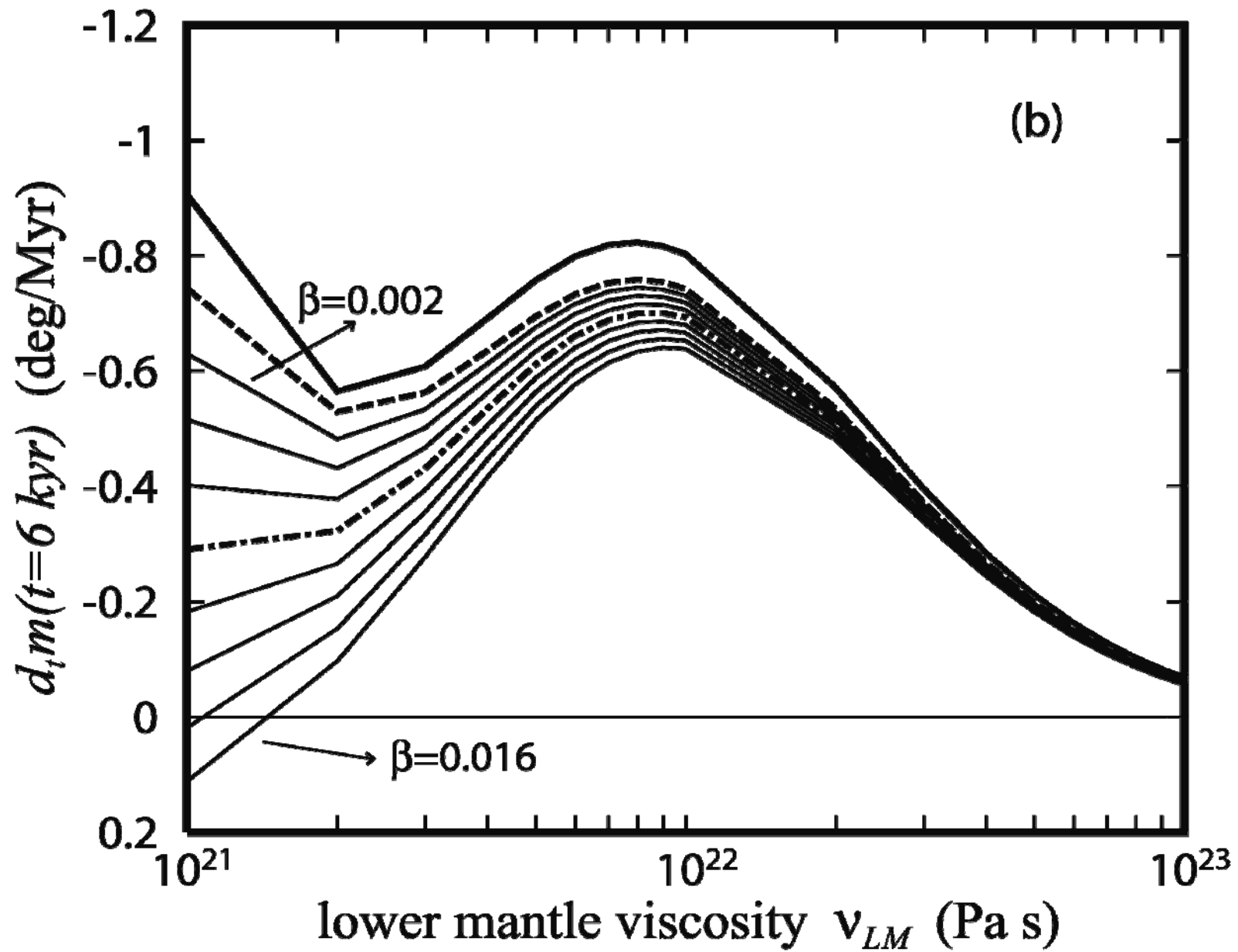


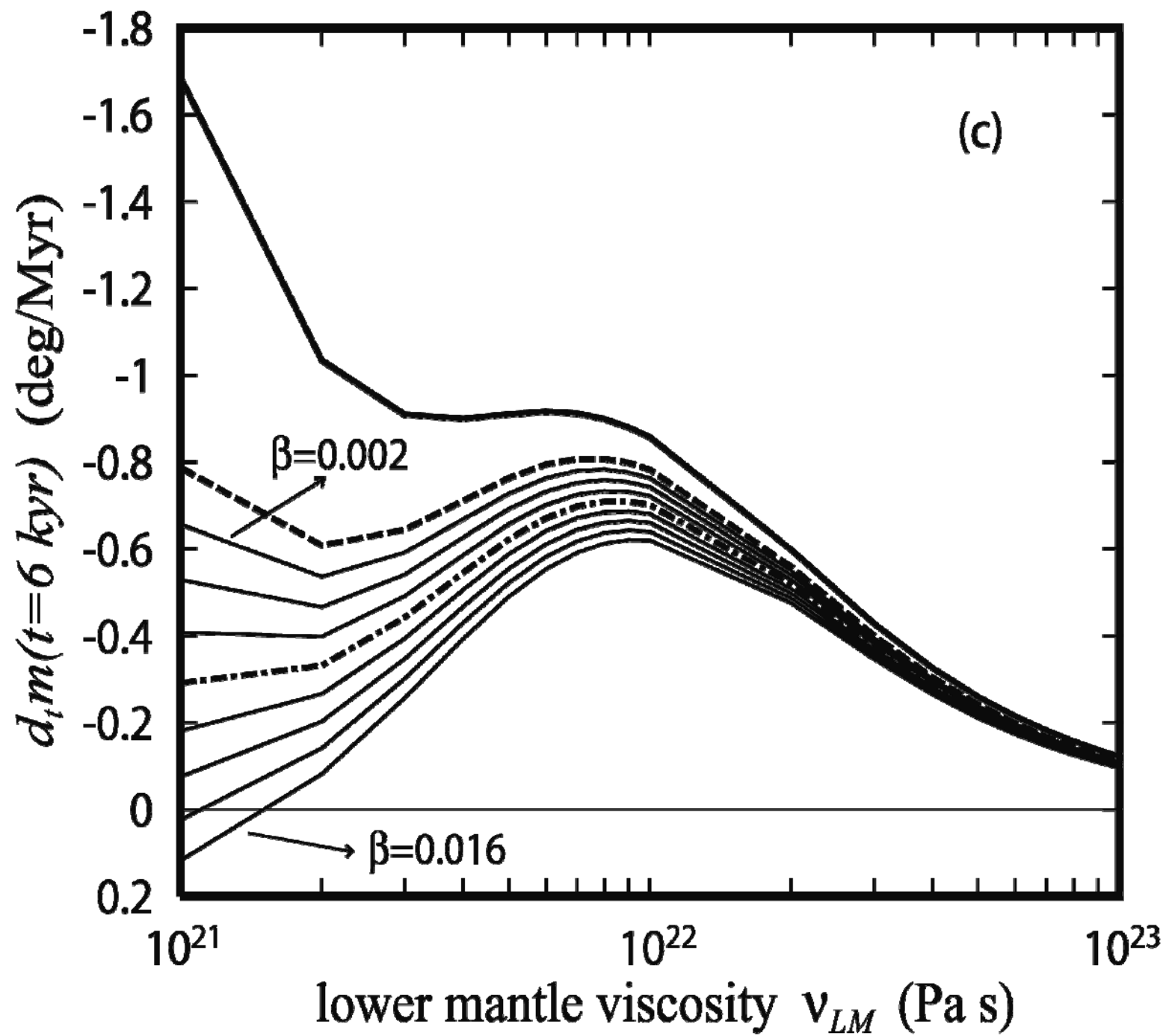


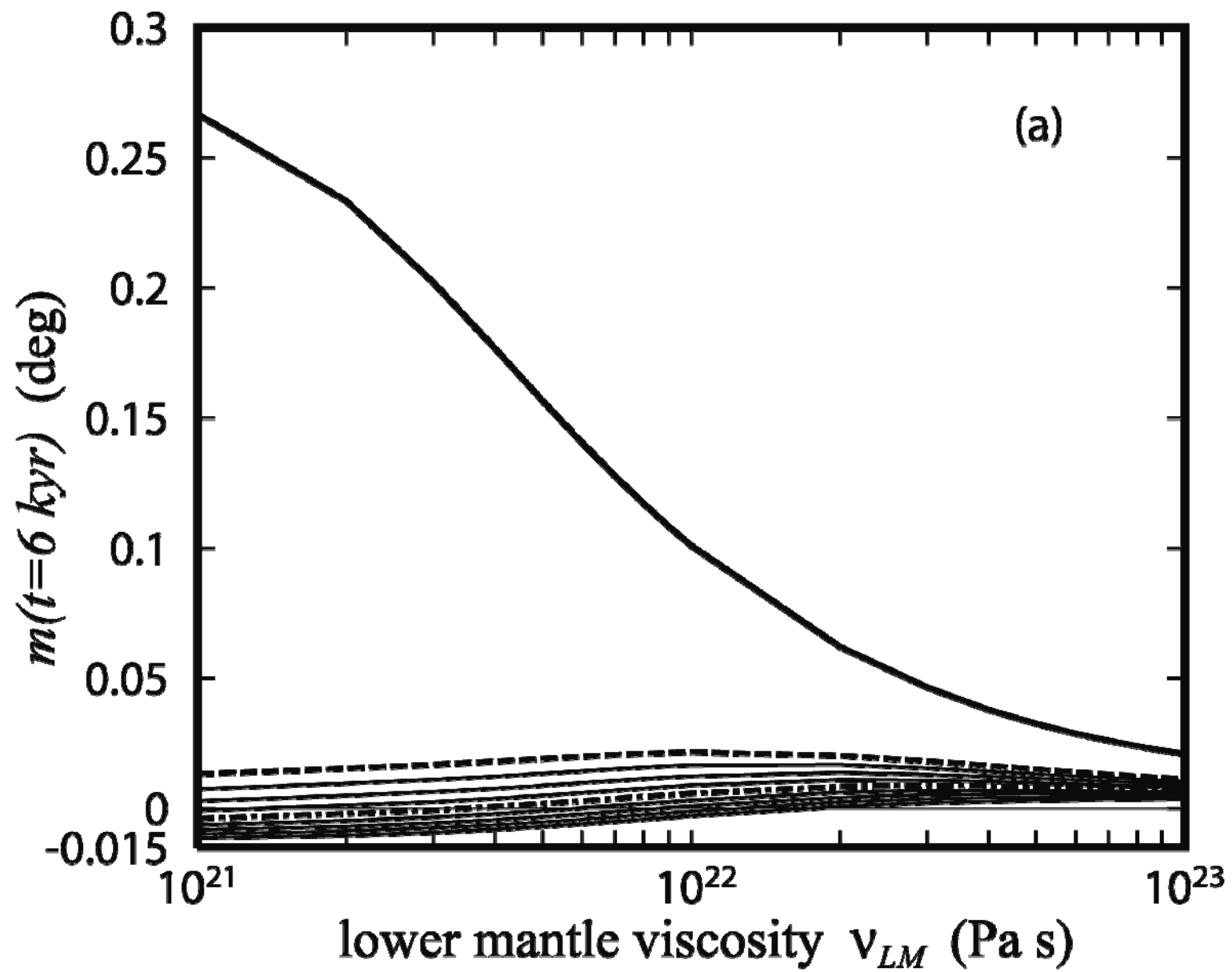


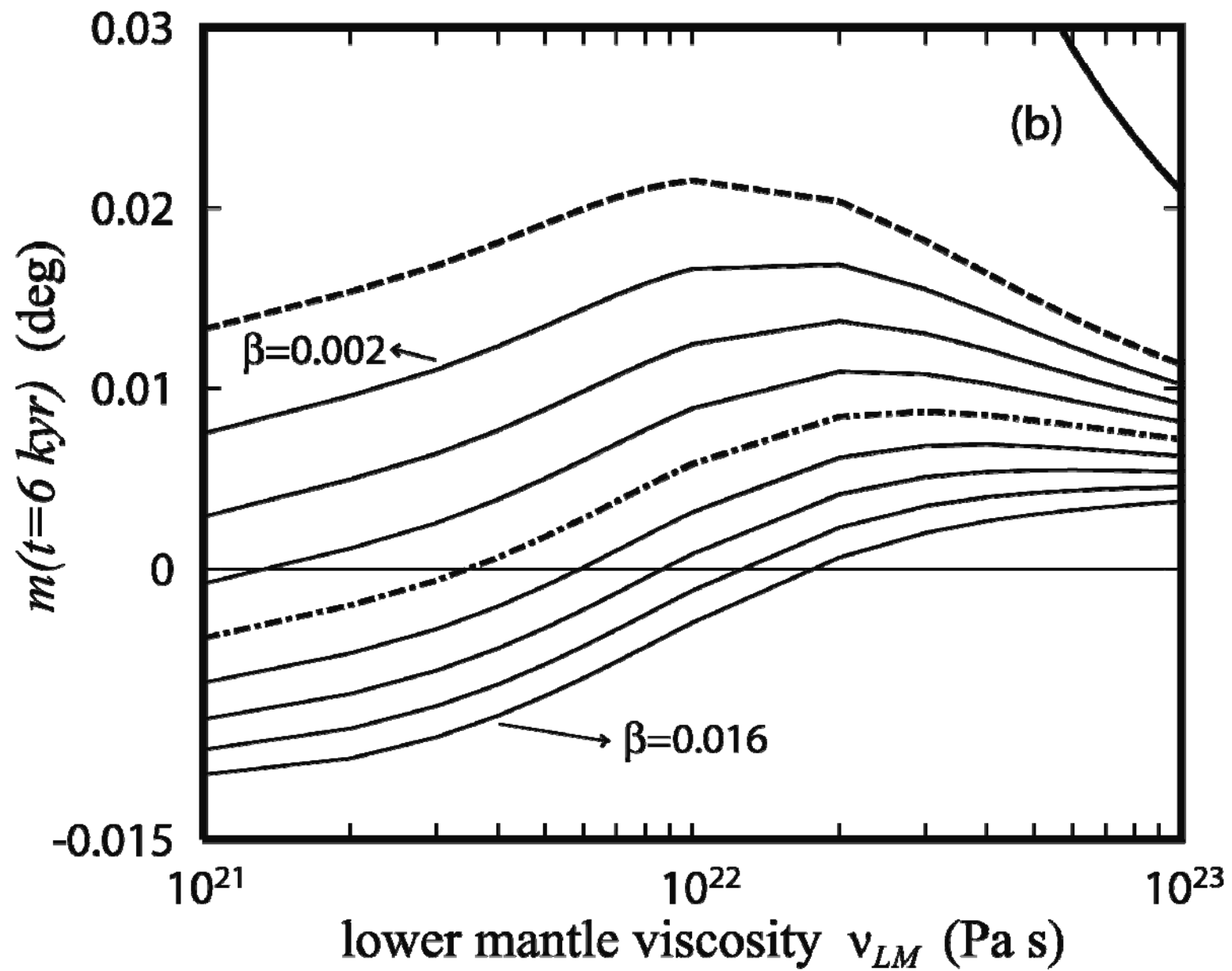












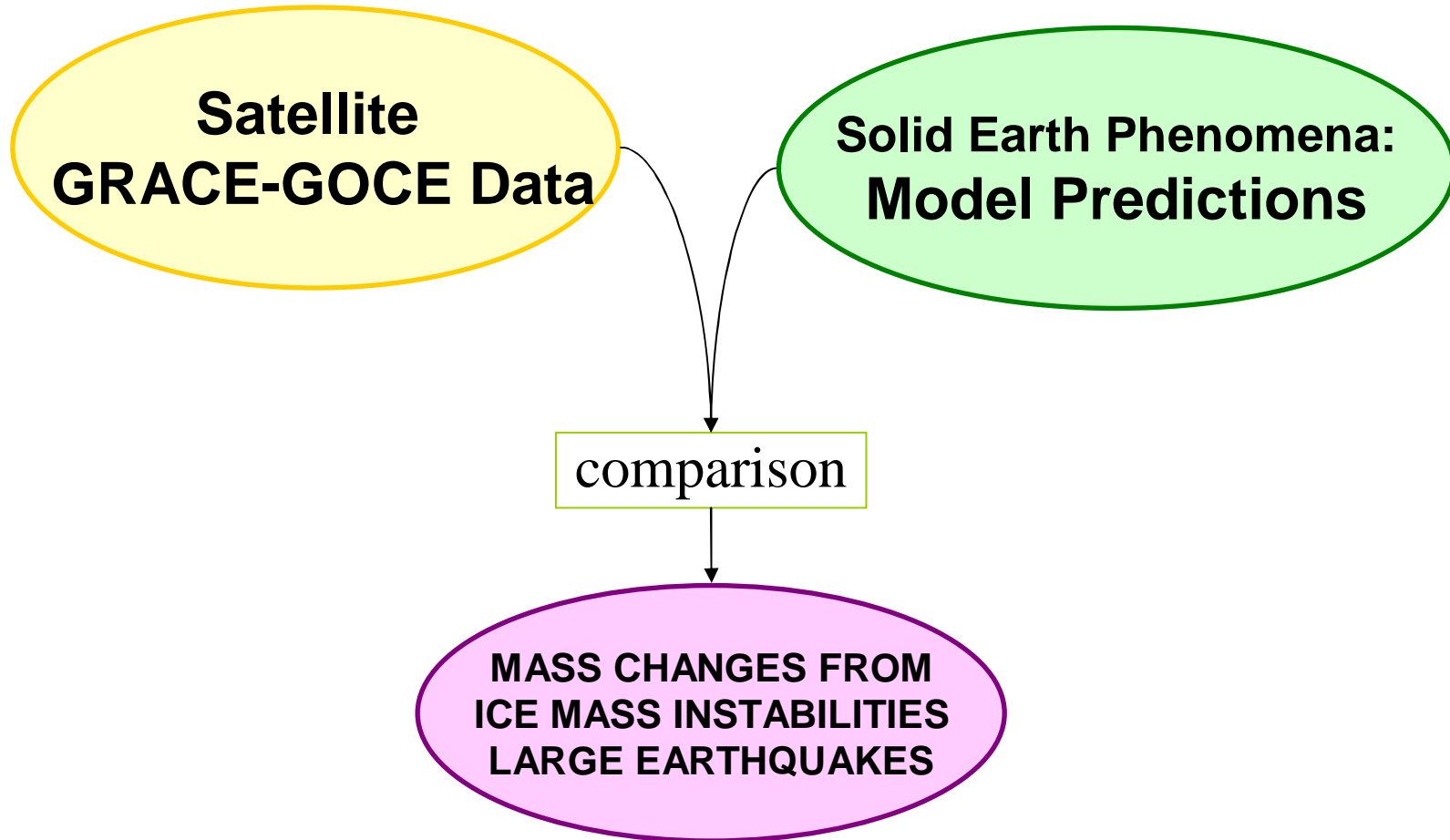


**Understanding the Earth's Interior from
Relaxation Normal Modes
3 – Ice mass balance, Sumatran
earthquake from gravity
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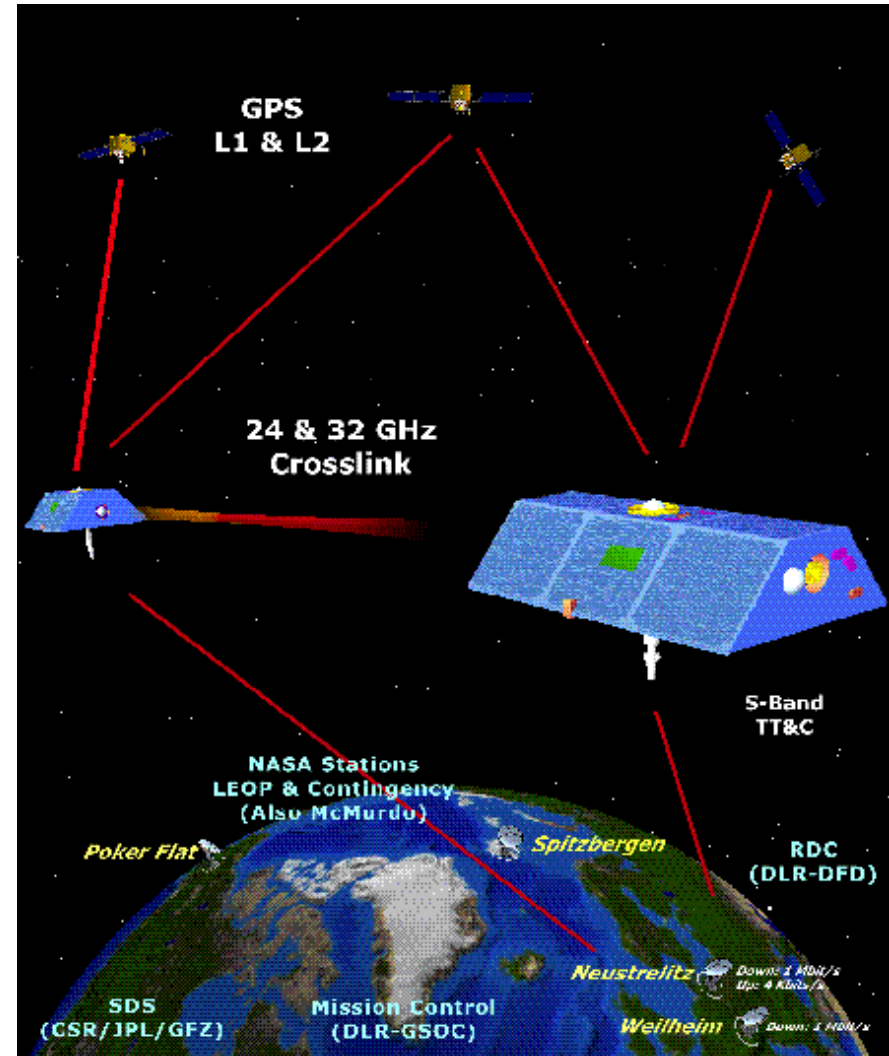
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General Scheme



SLR and GRACE

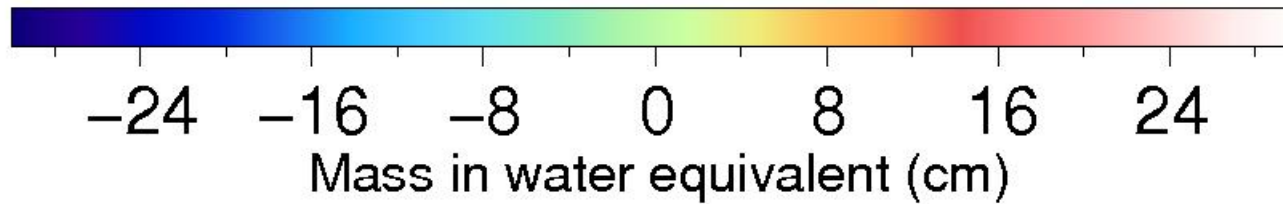
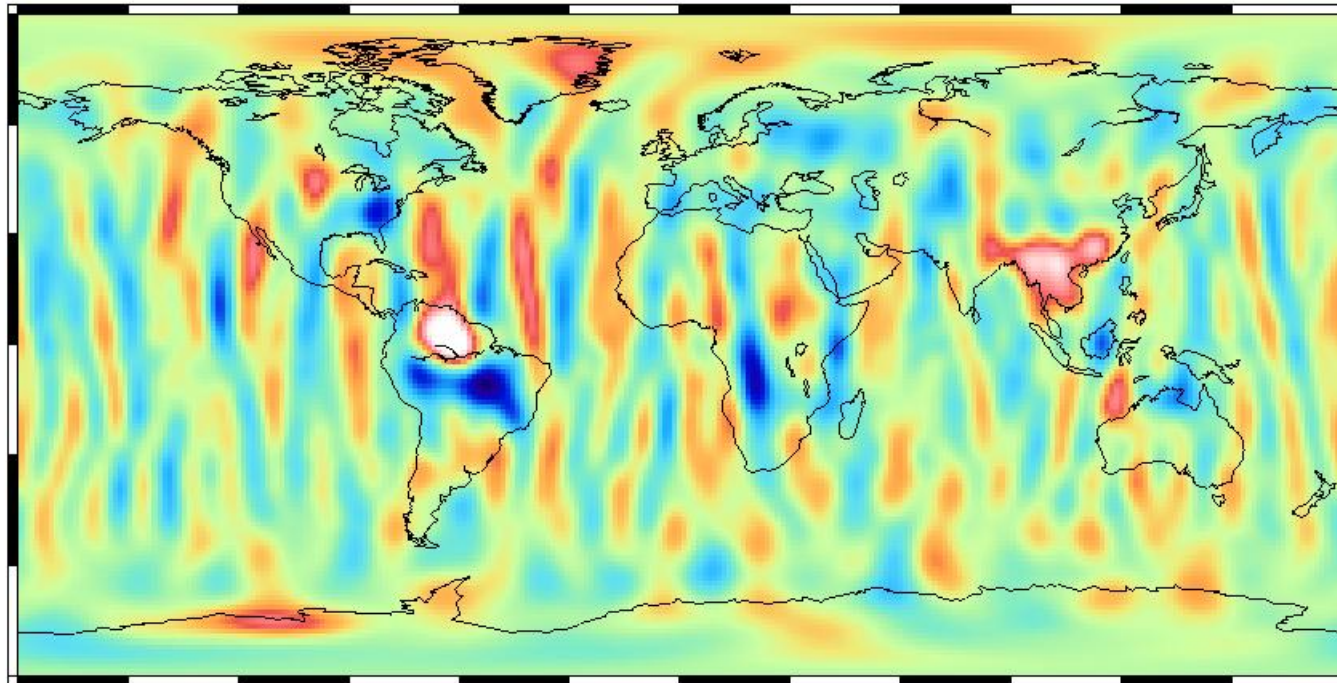
Satellite Laser Ranging

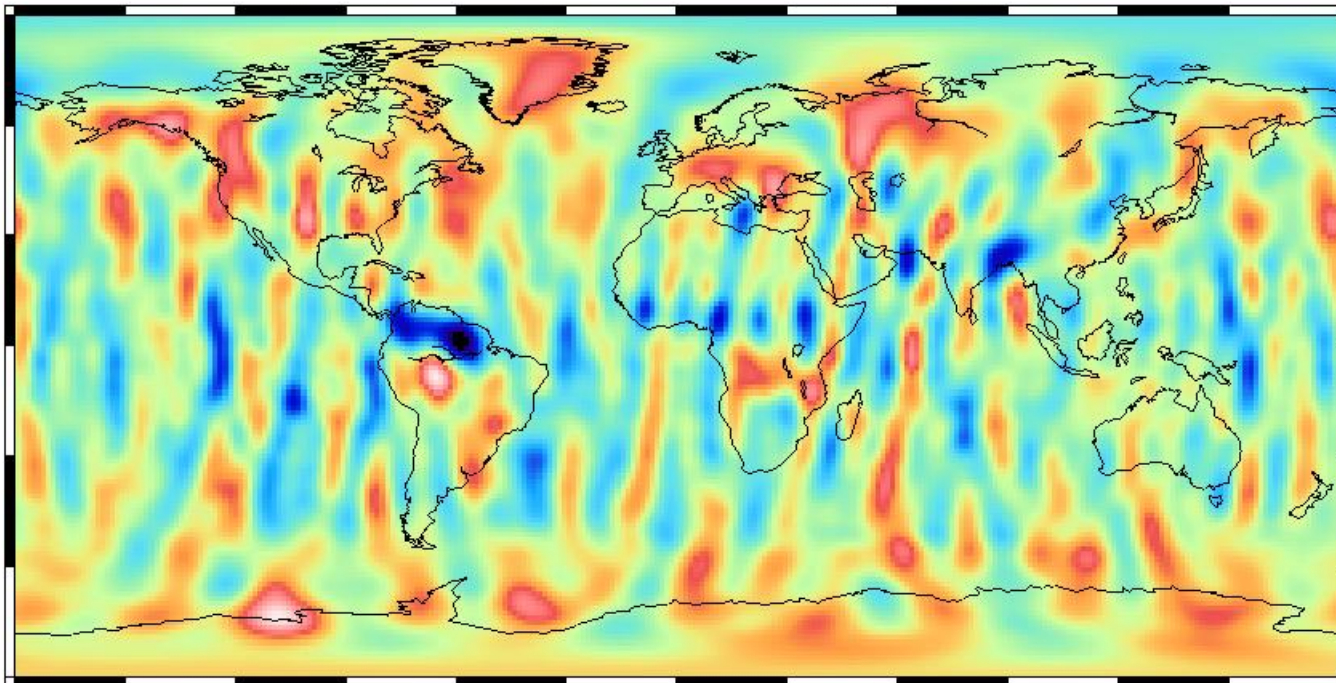




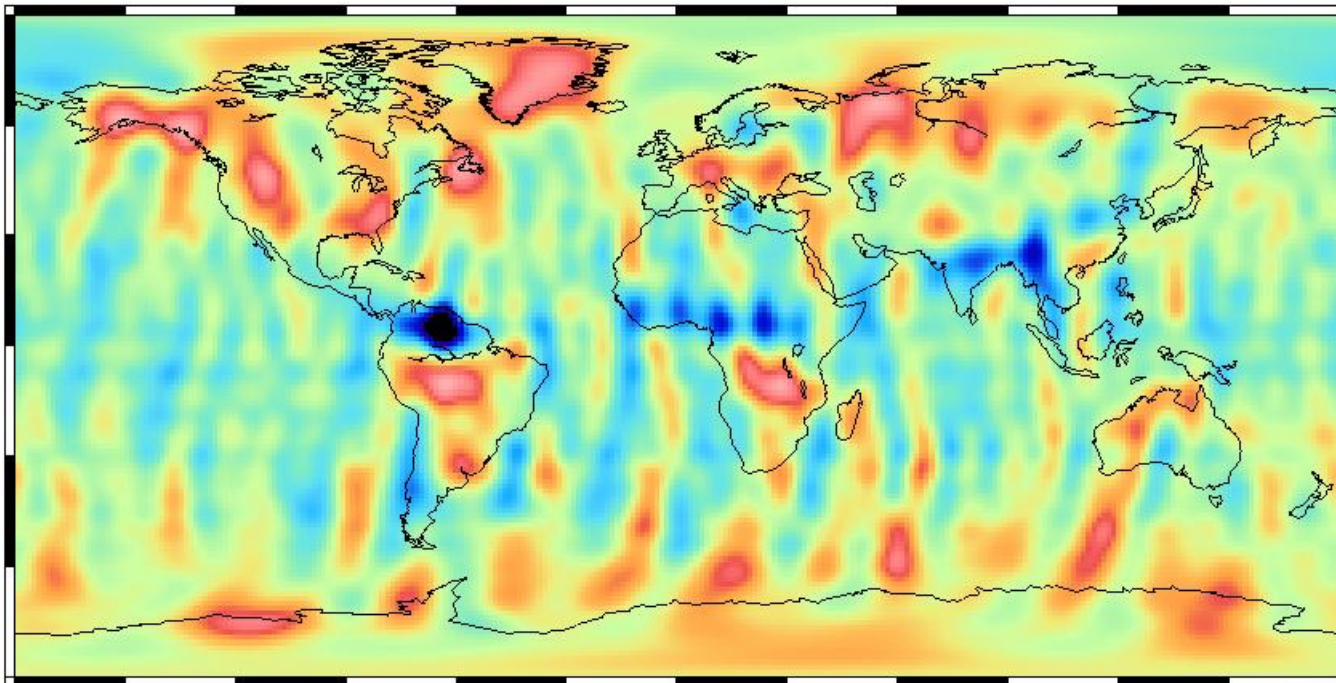
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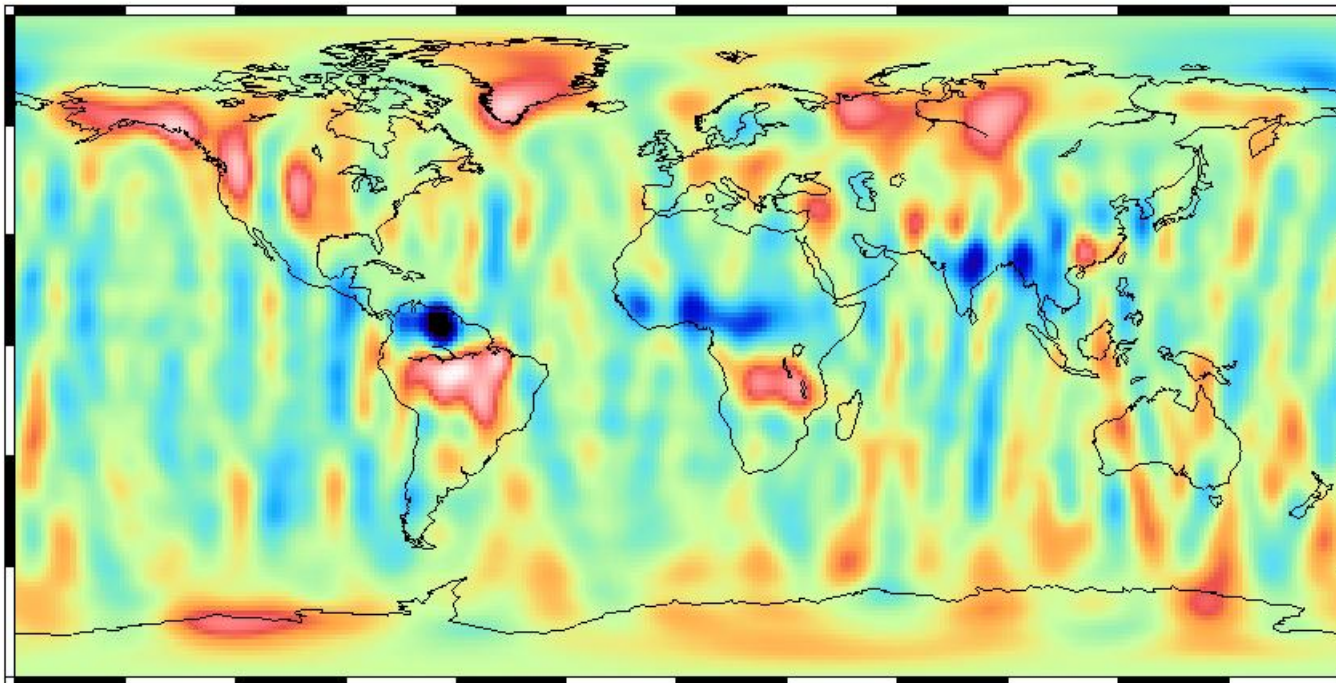




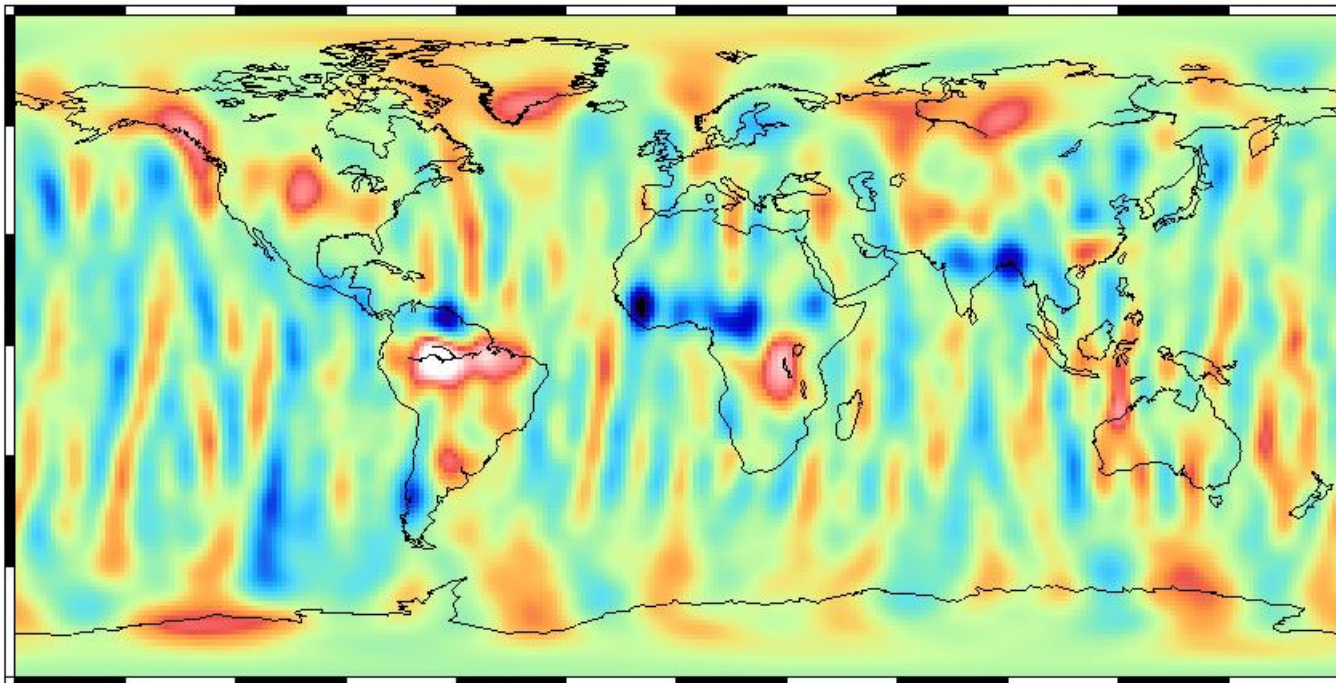
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Mass in water equivalent (cm)



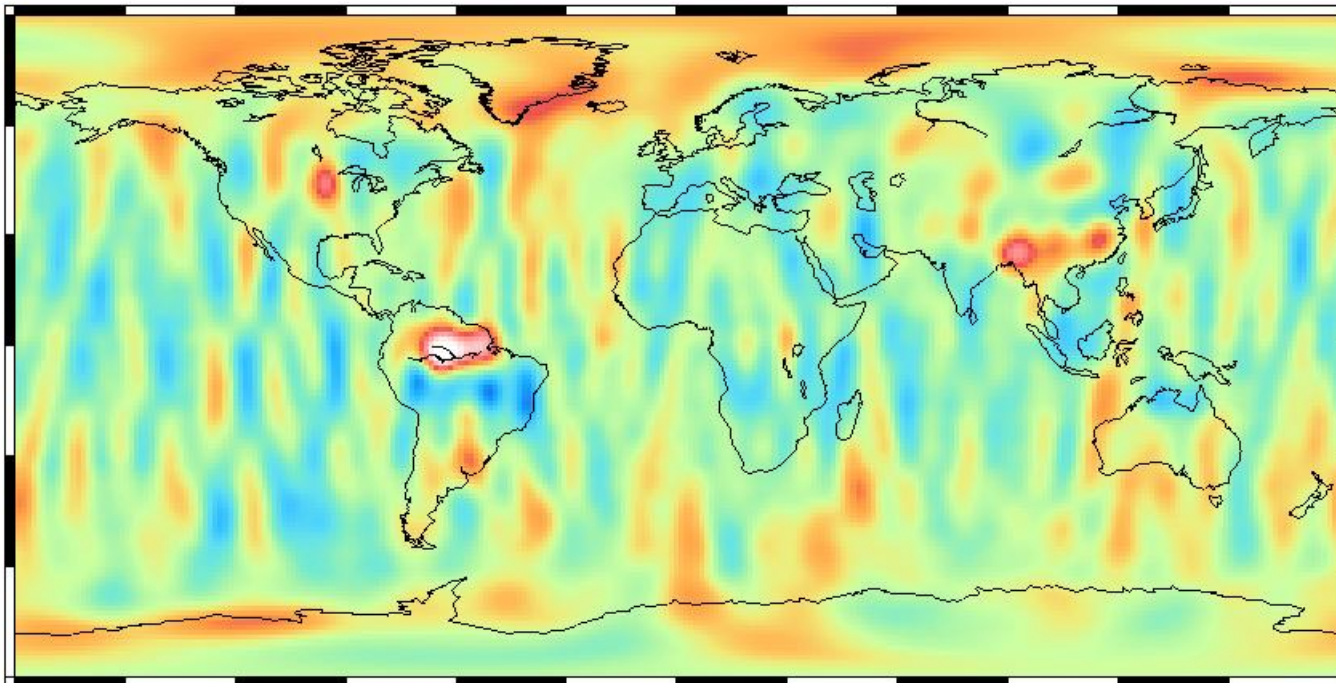
-24 -16 -8 0 8 16 24
Mass in water equivalent (cm)



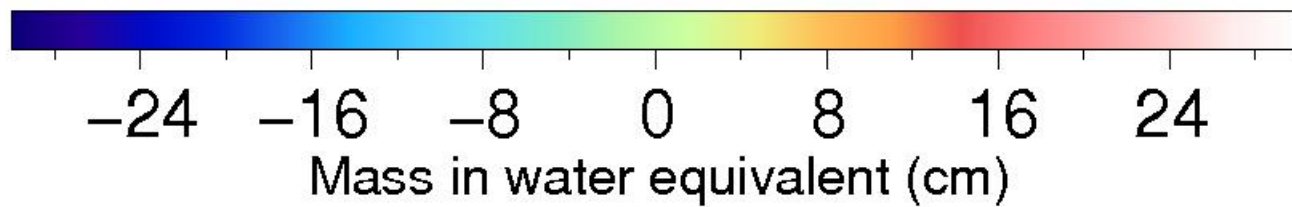
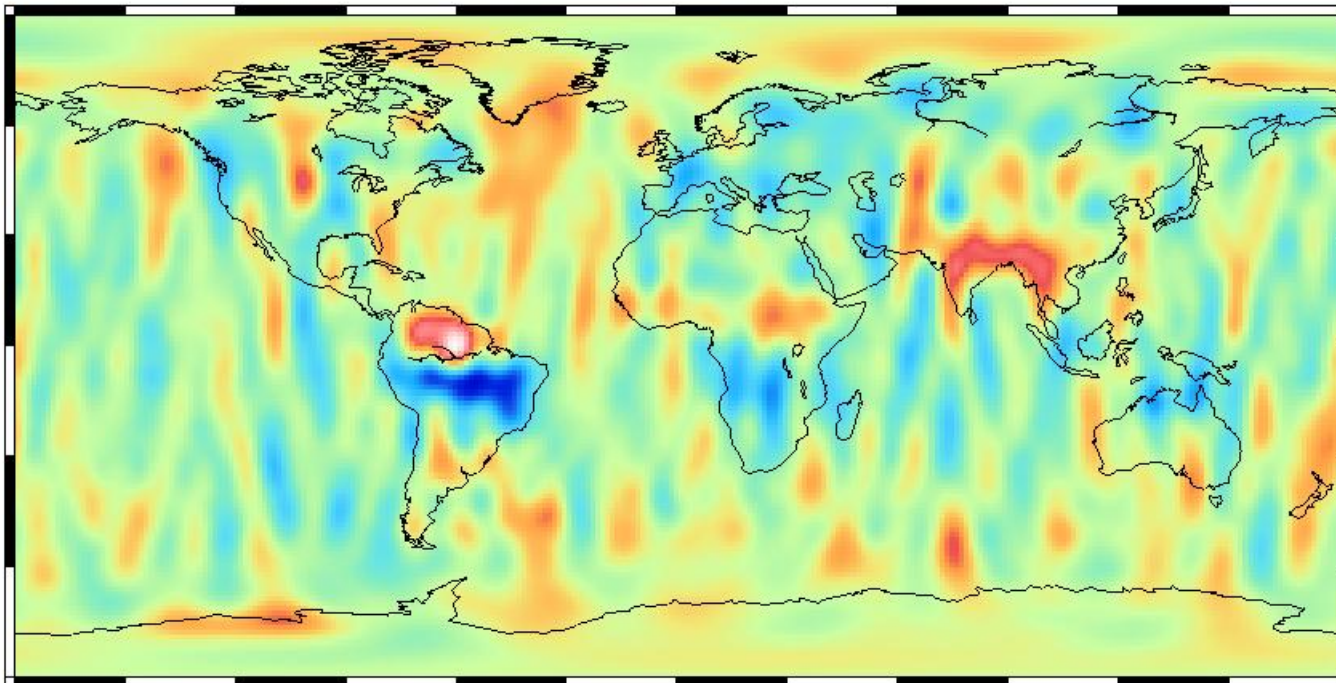
-24 -16 -8 0 8 16 24
Mass in water equivalent (cm)

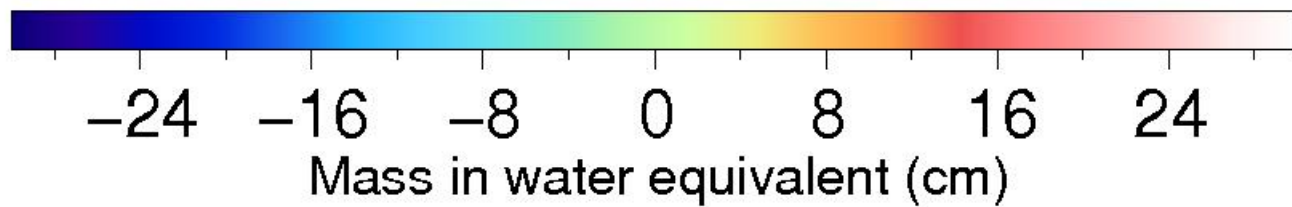
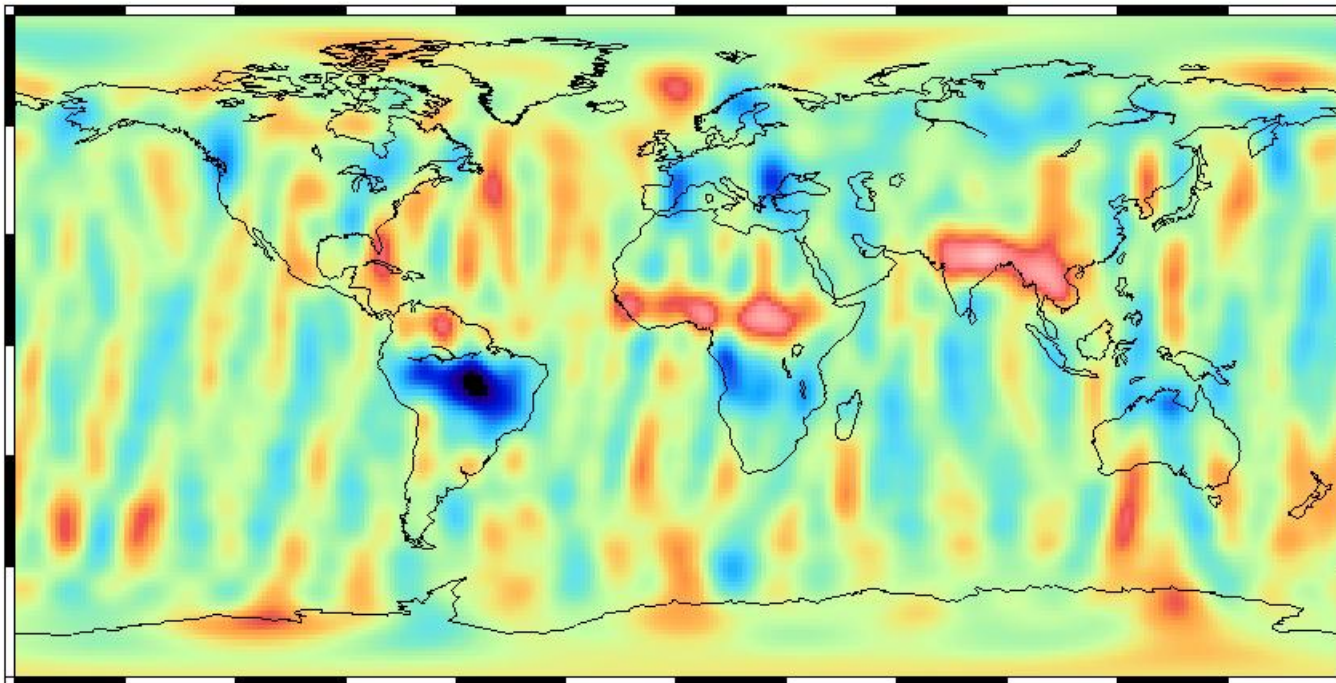


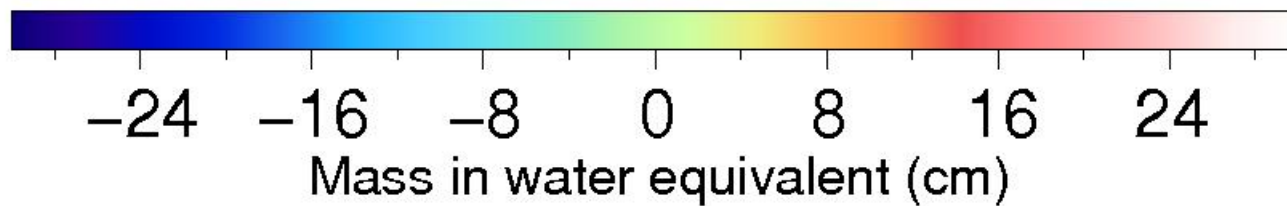
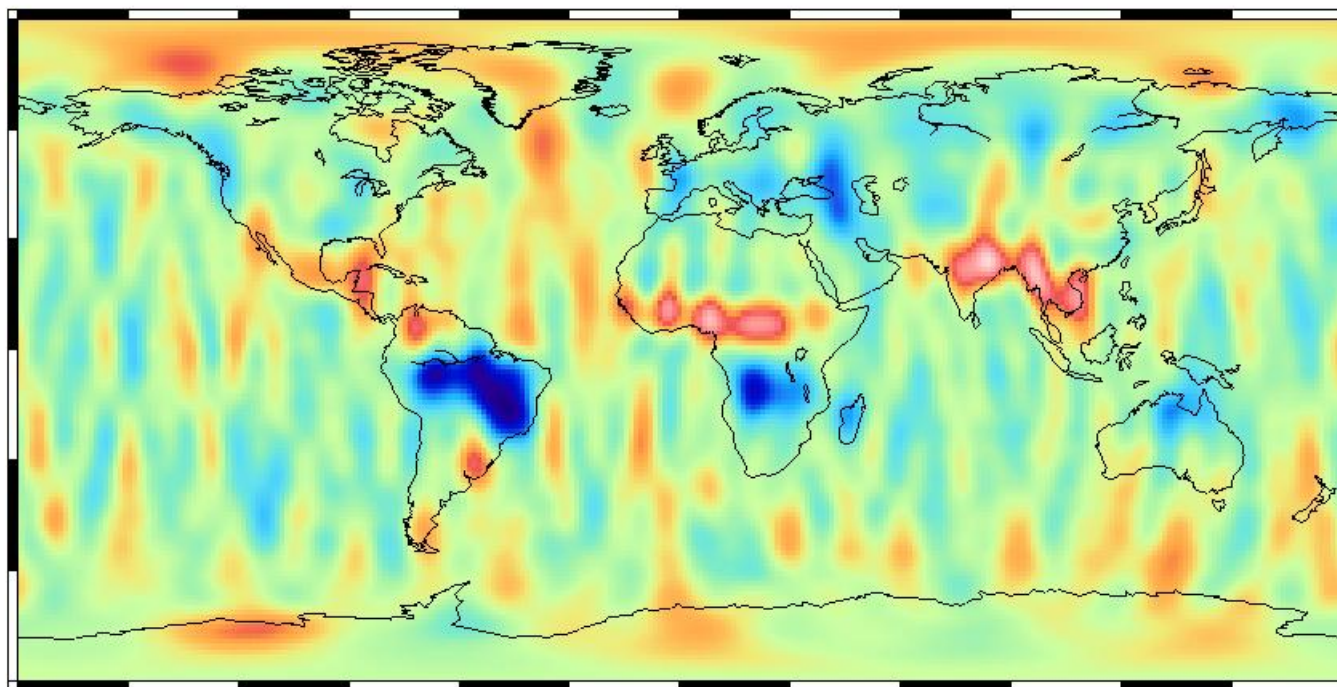
-24 -16 -8 0 8 16 24
Mass in water equivalent (cm)

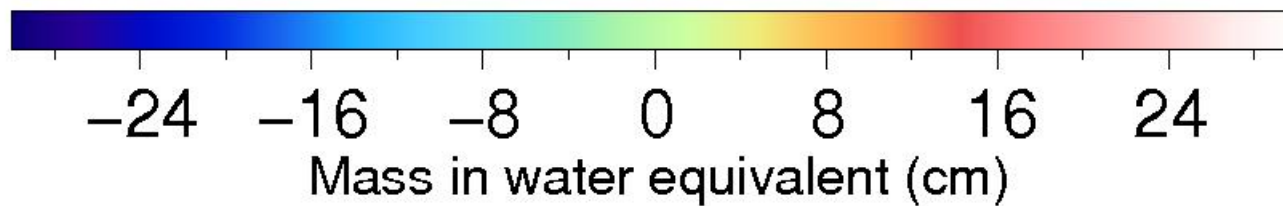
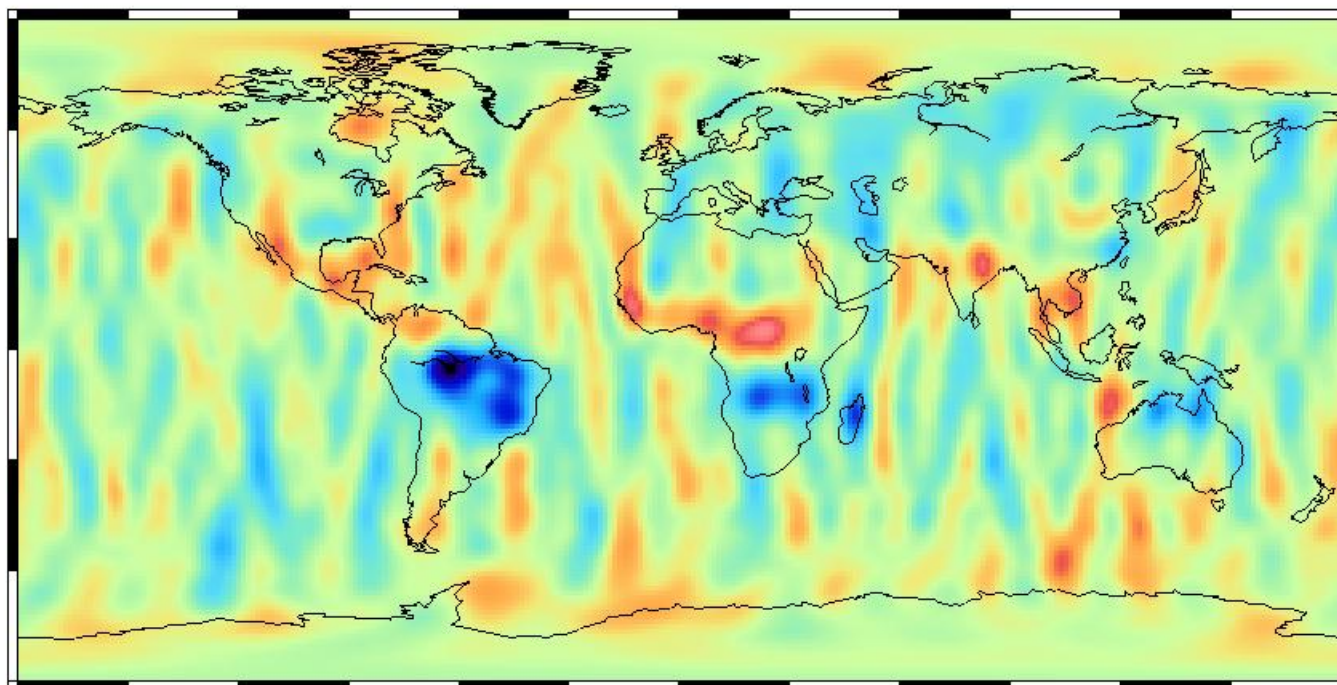


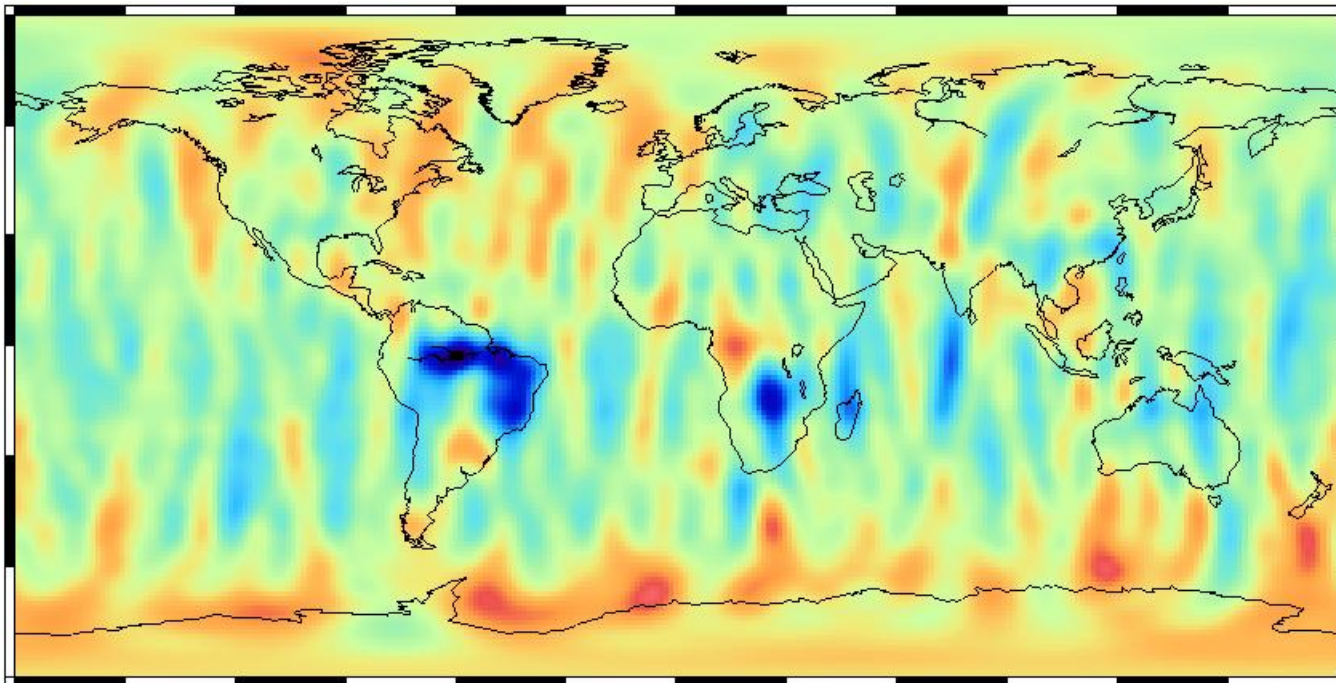
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Mass in water equivalent (cm)



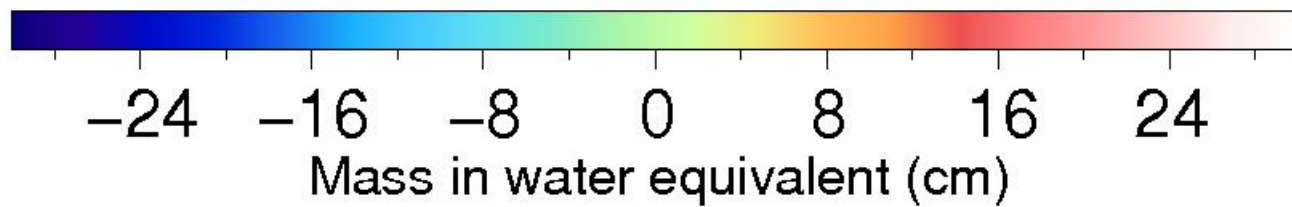
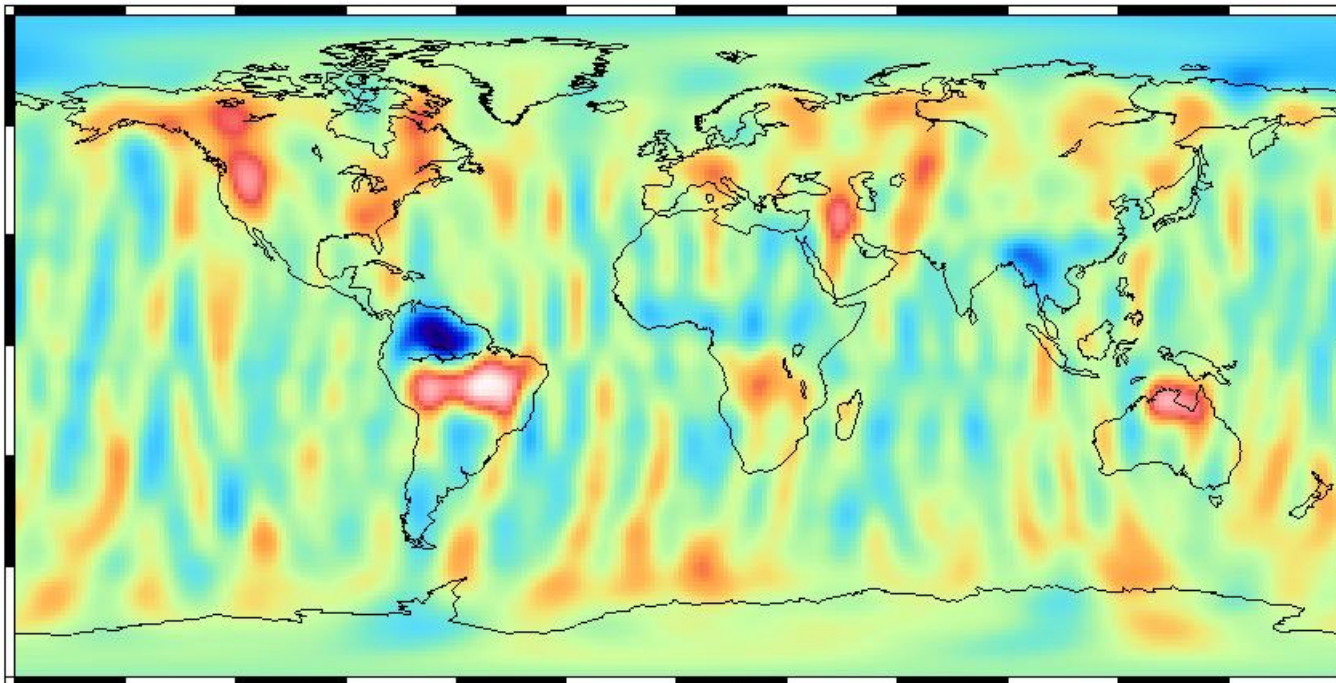


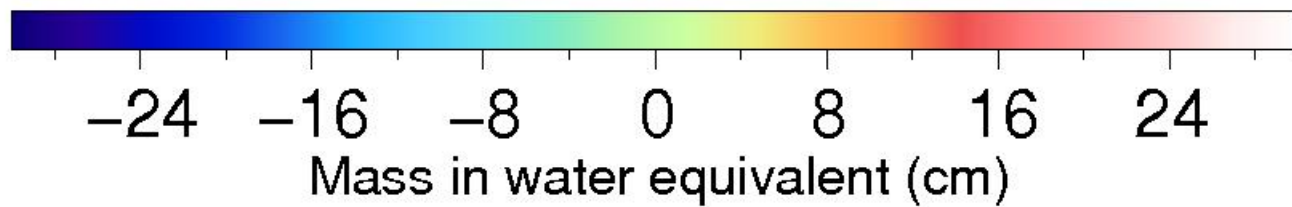
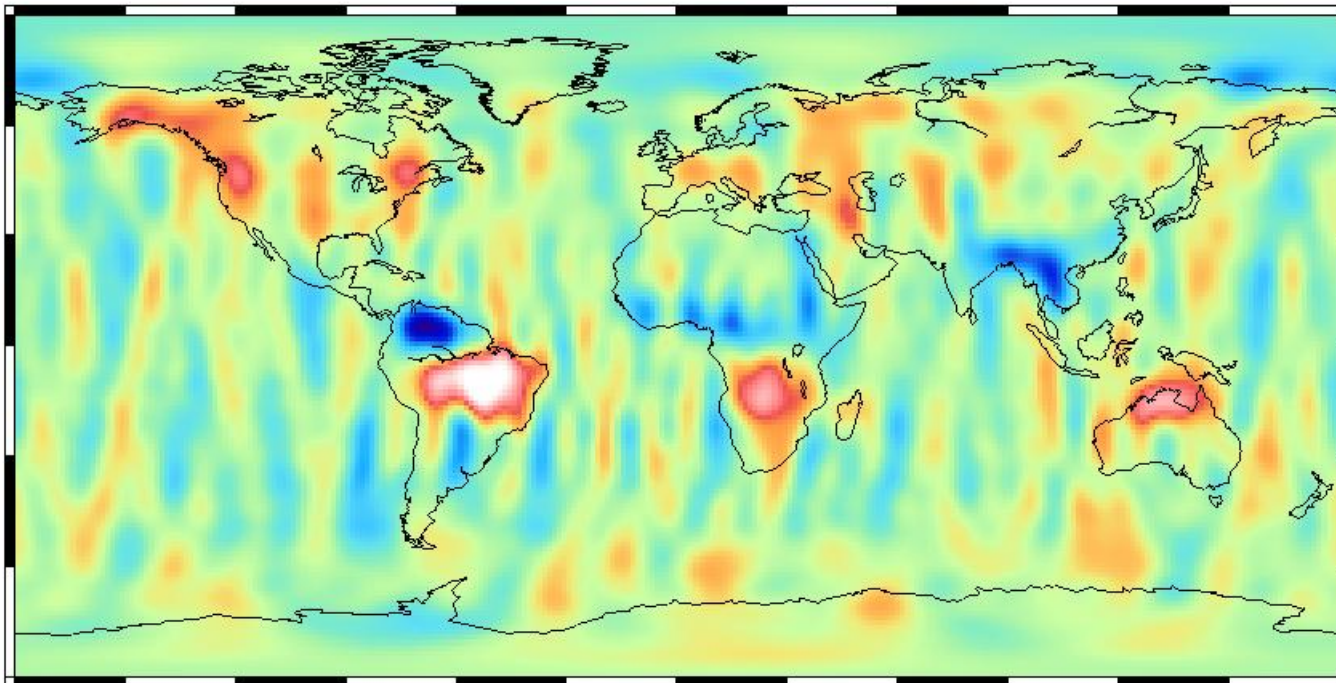


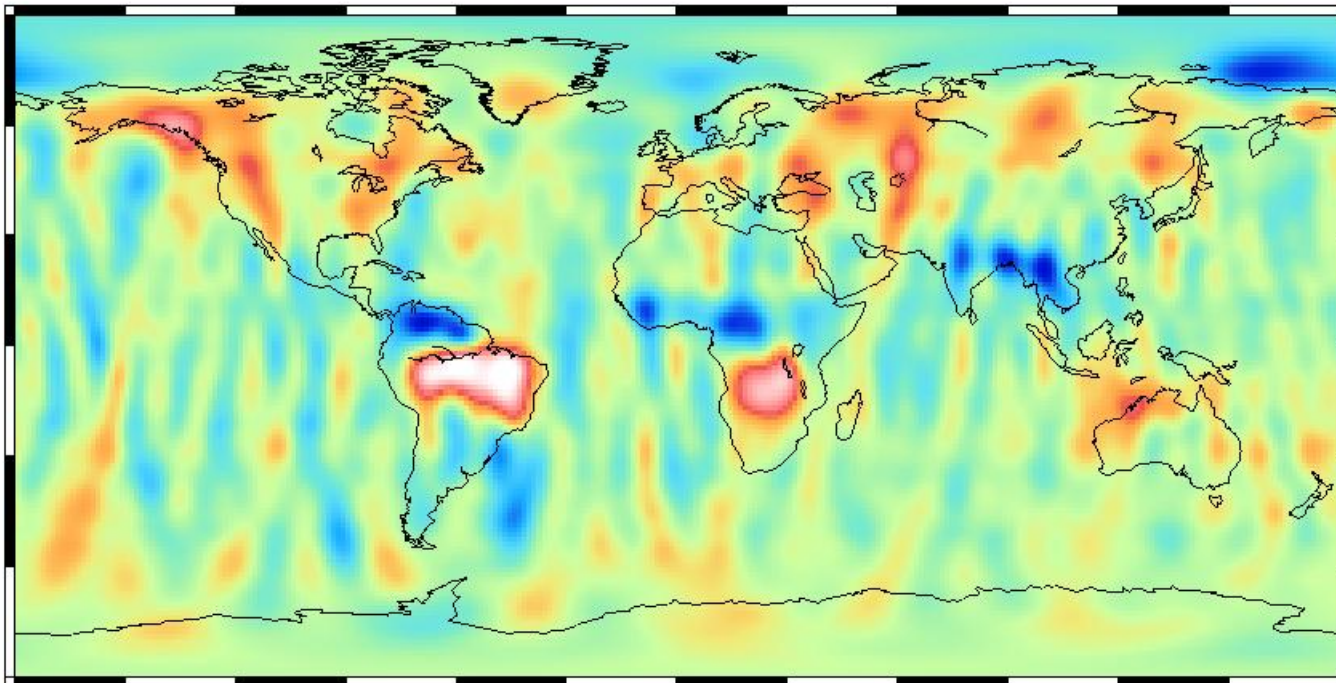




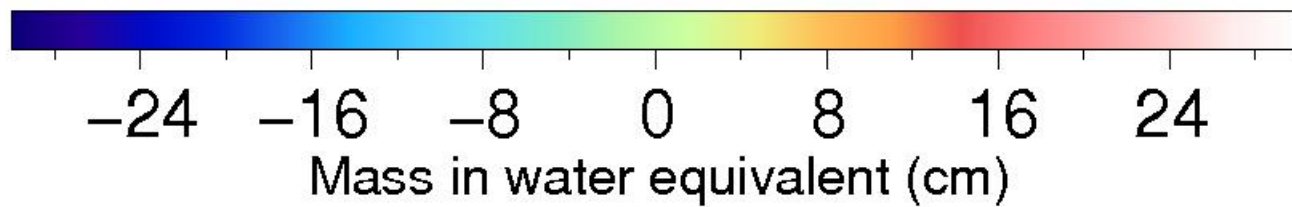
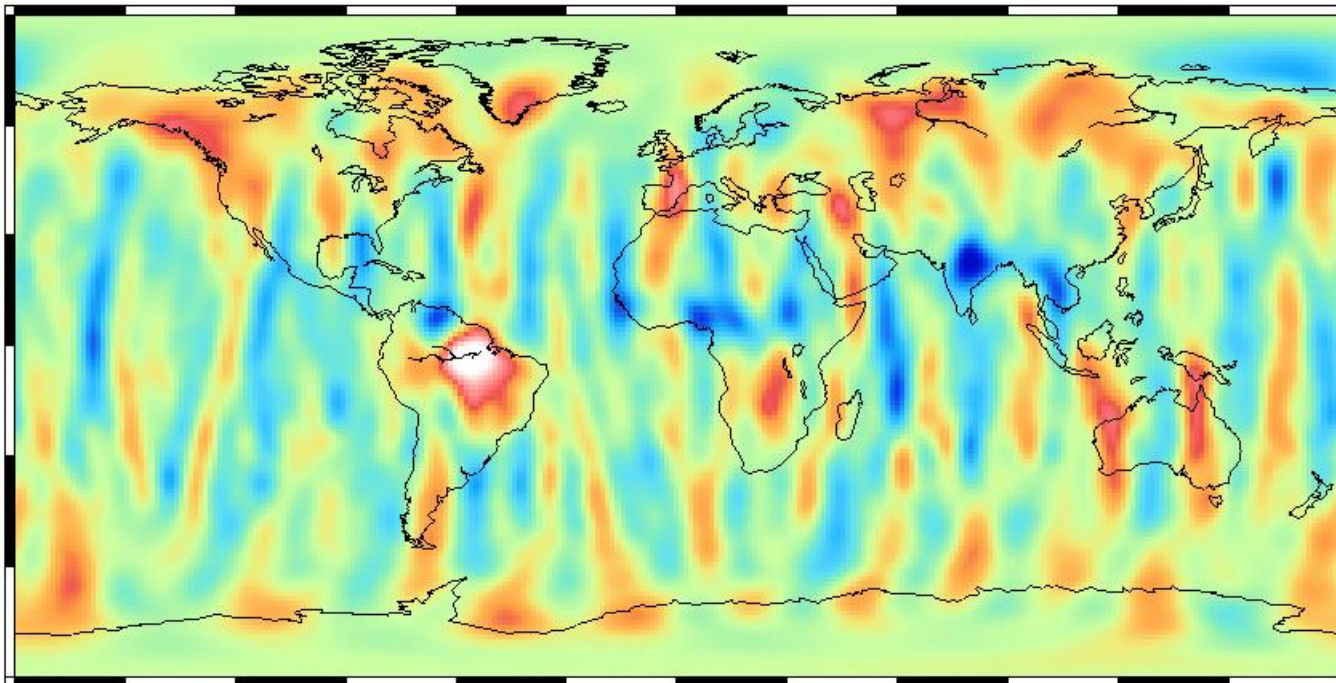
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Mass in water equivalent (cm)

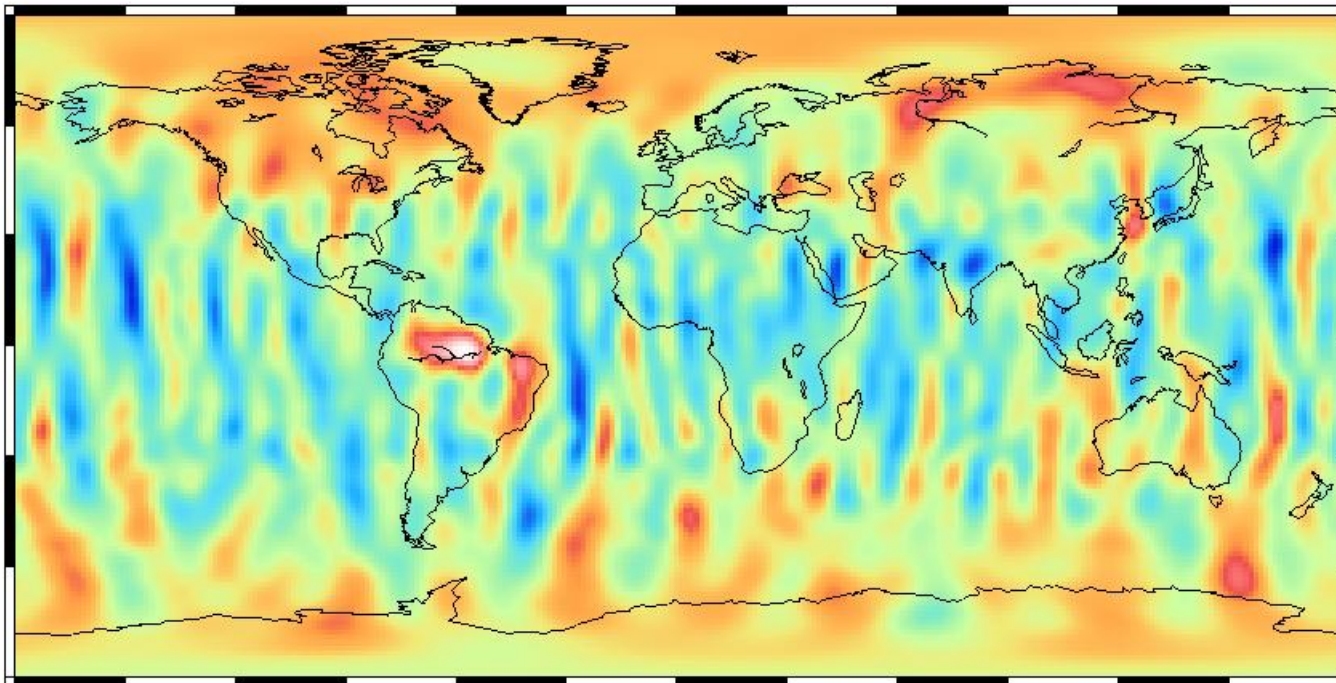




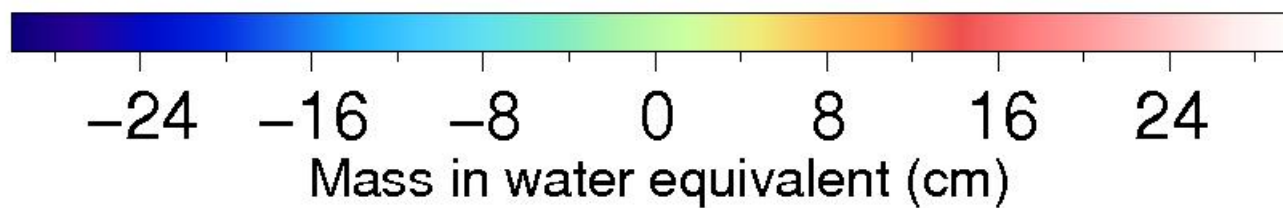
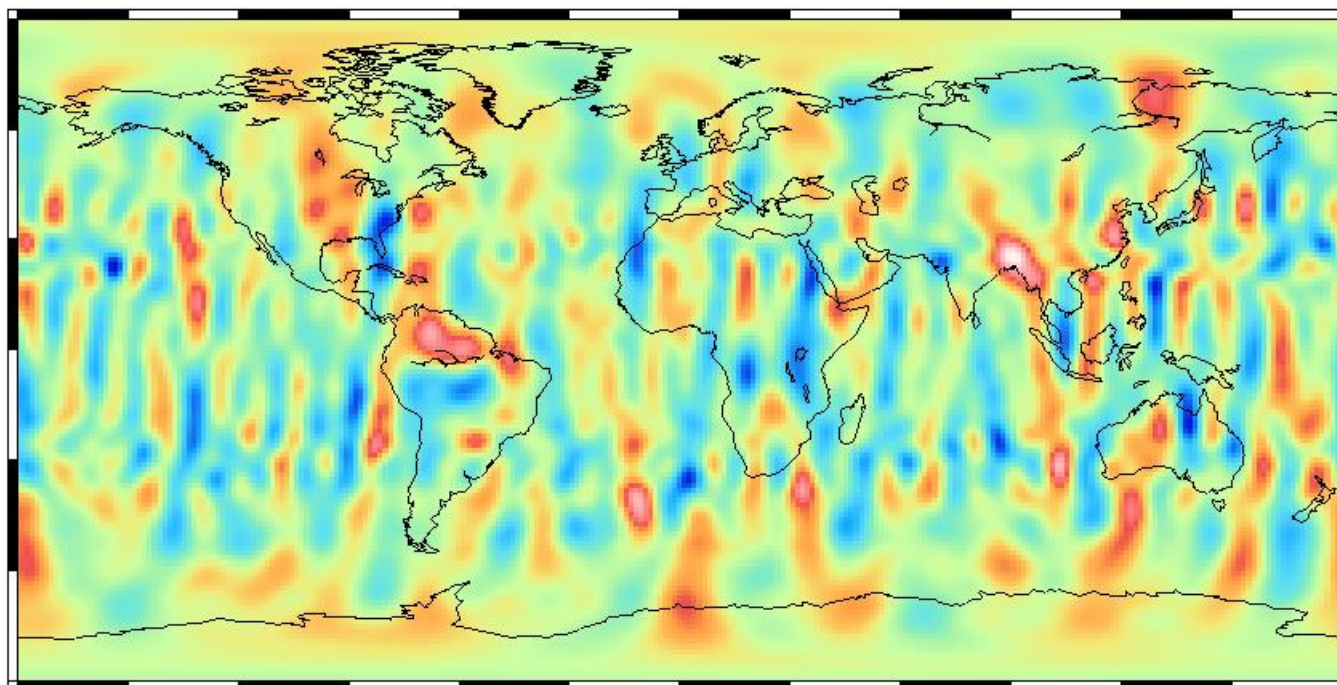


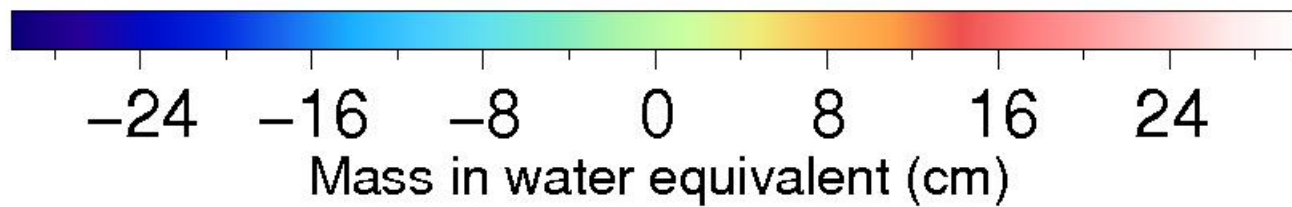
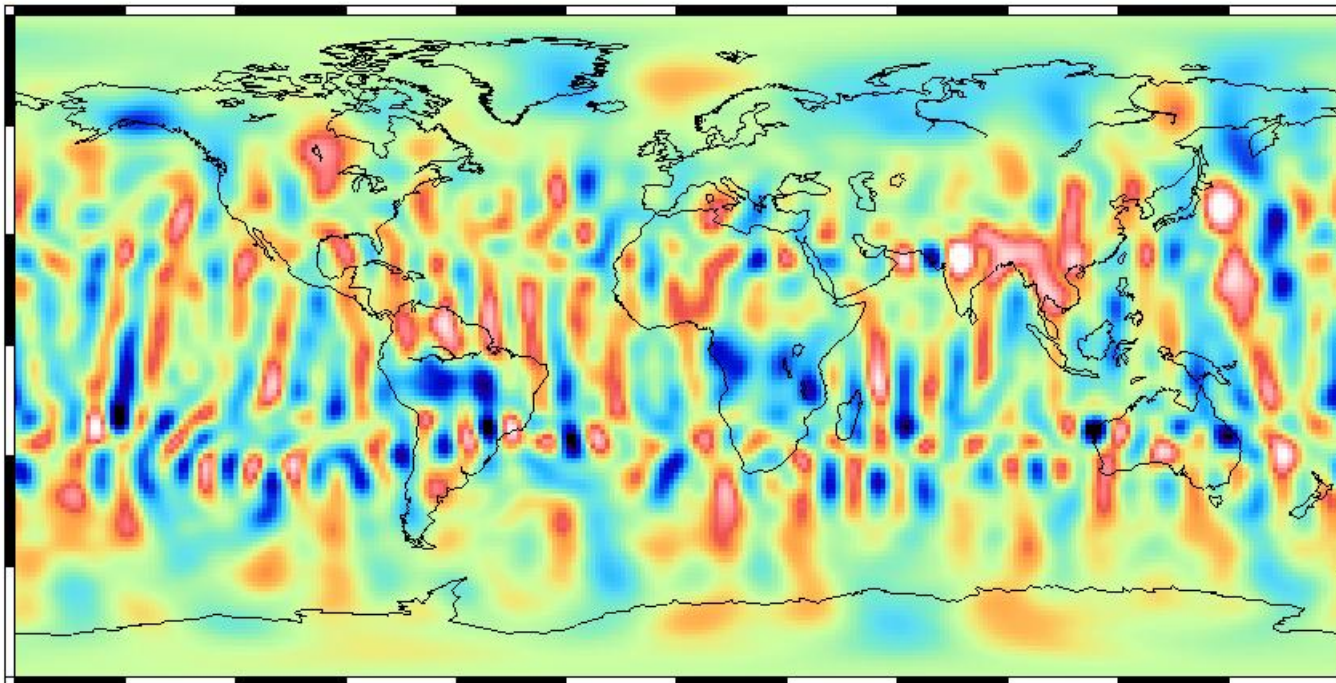
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Mass in water equivalent (cm)

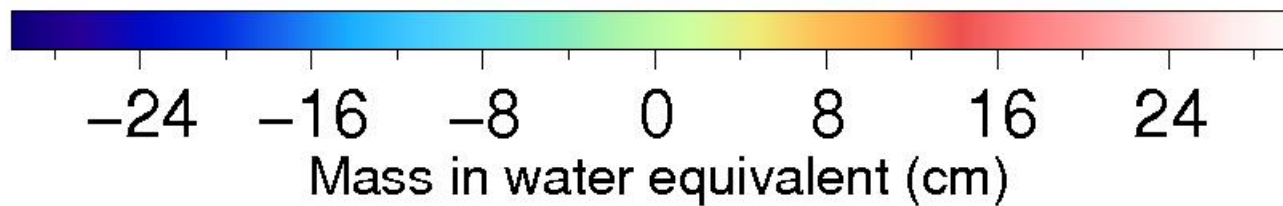
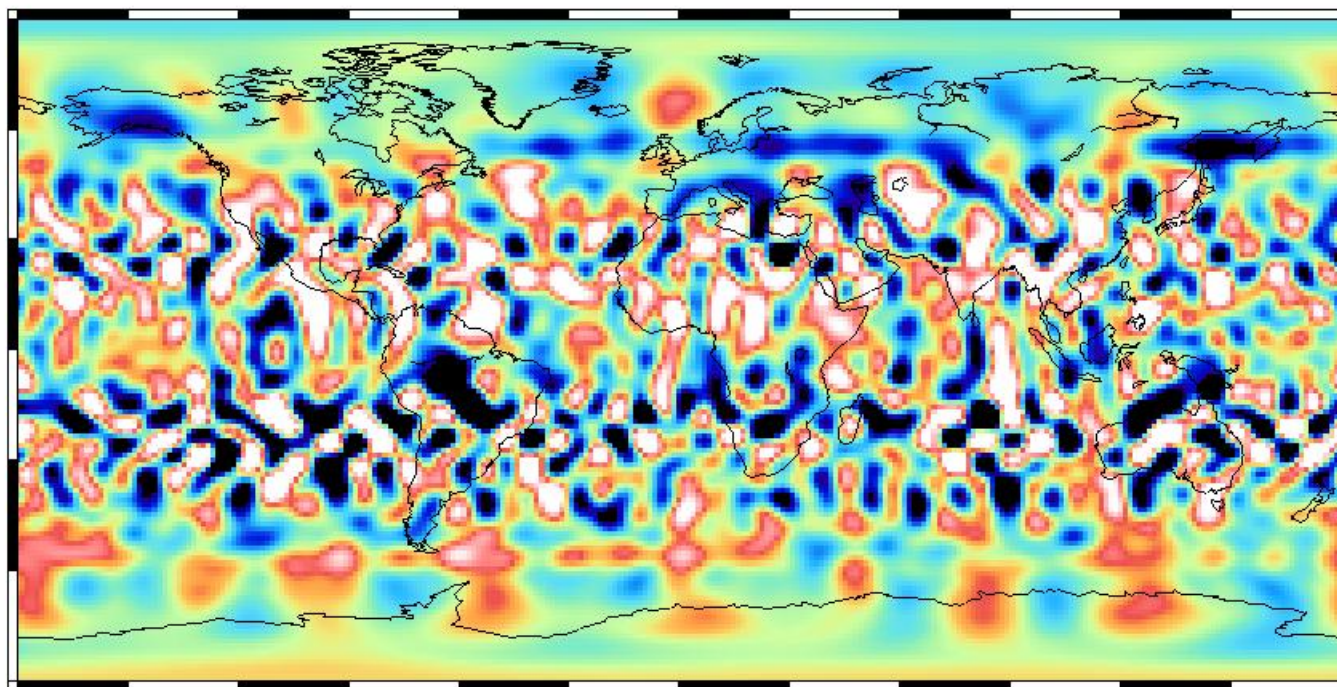


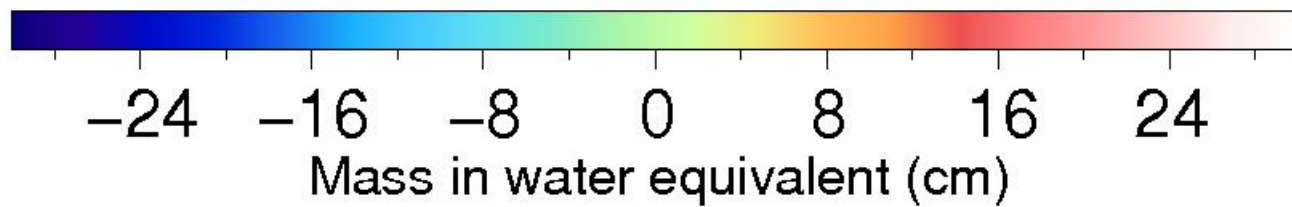
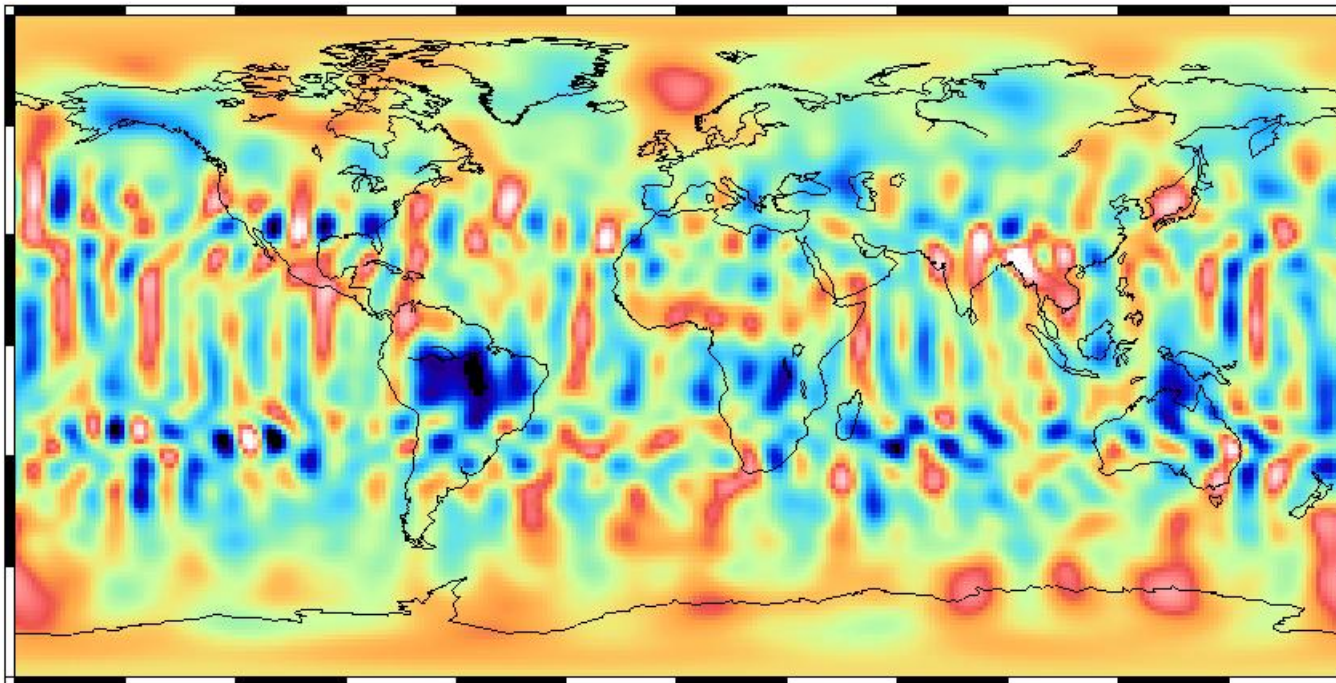


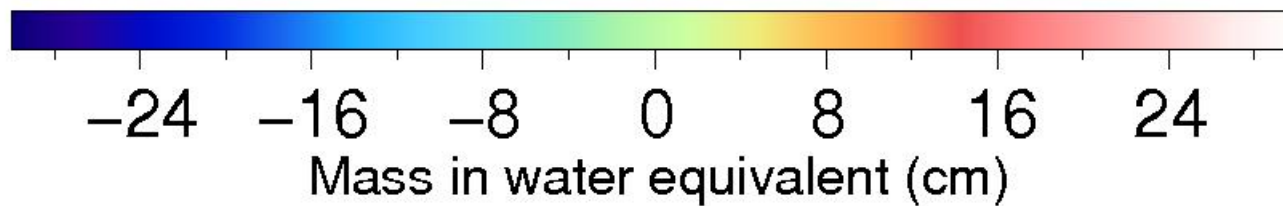
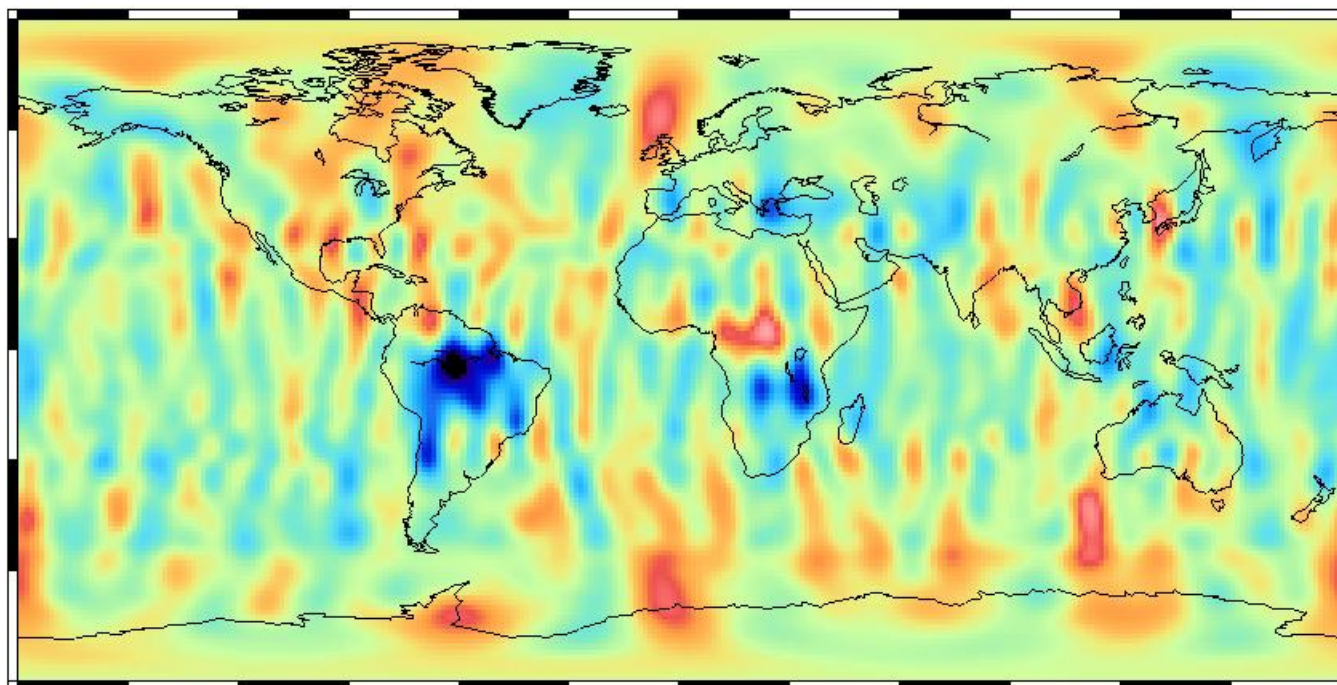
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Mass in water equivalent (cm)

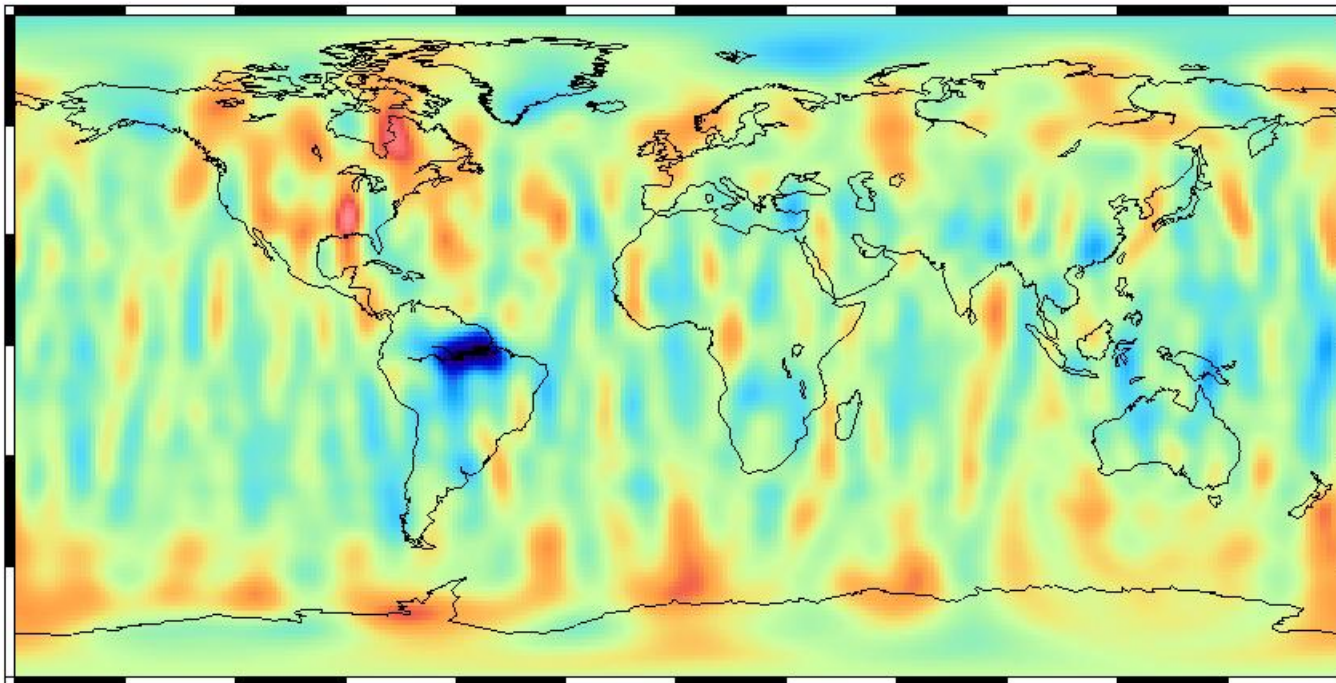




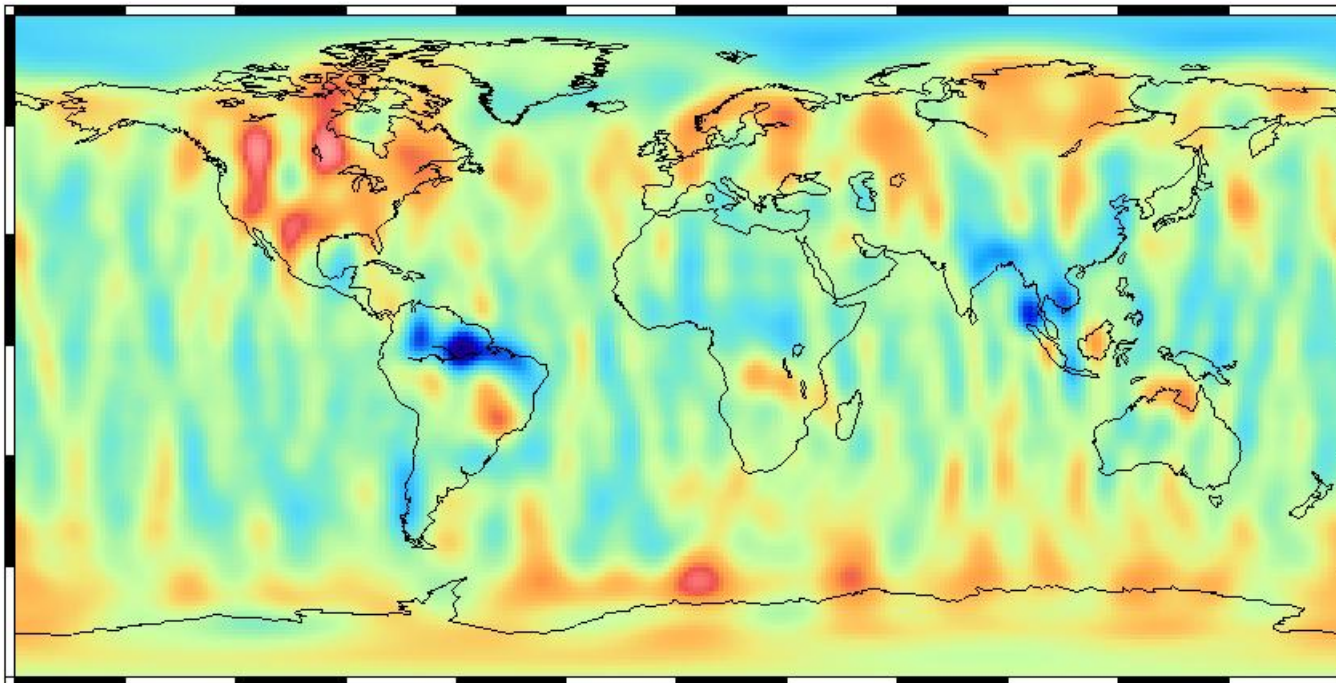




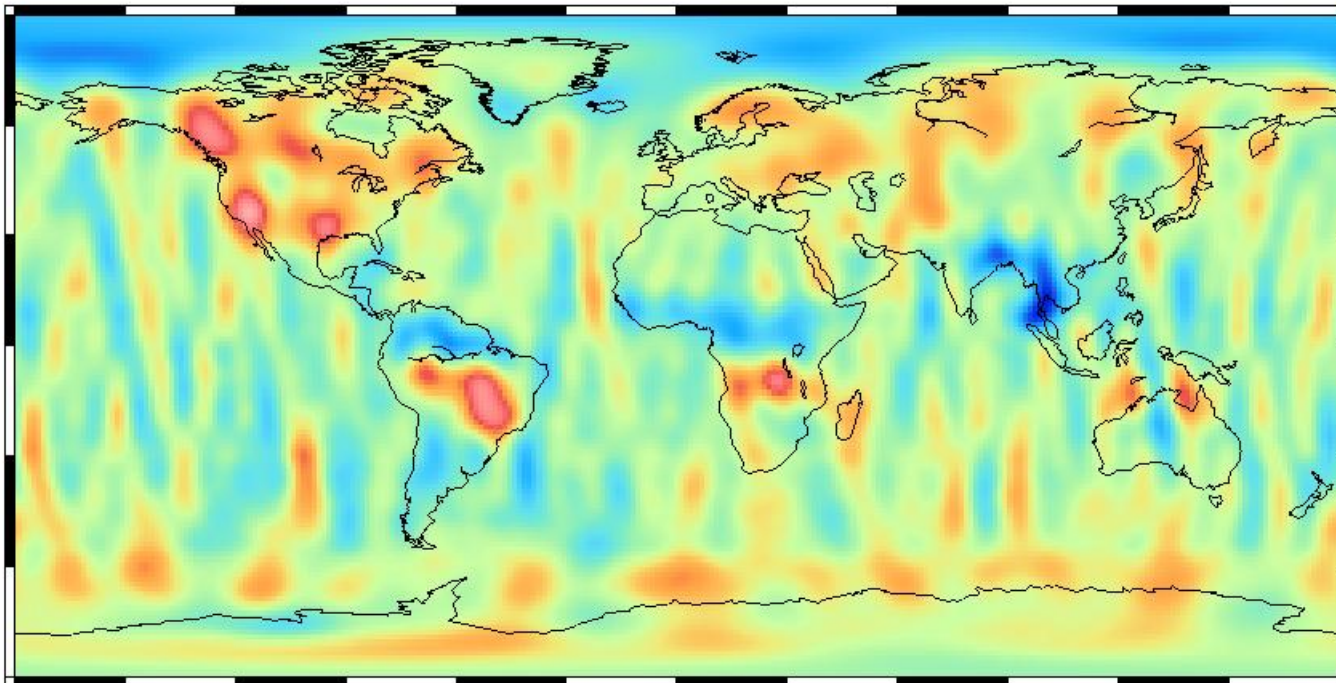




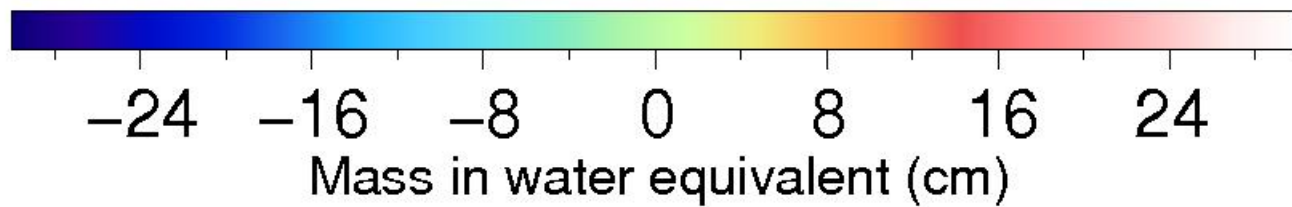
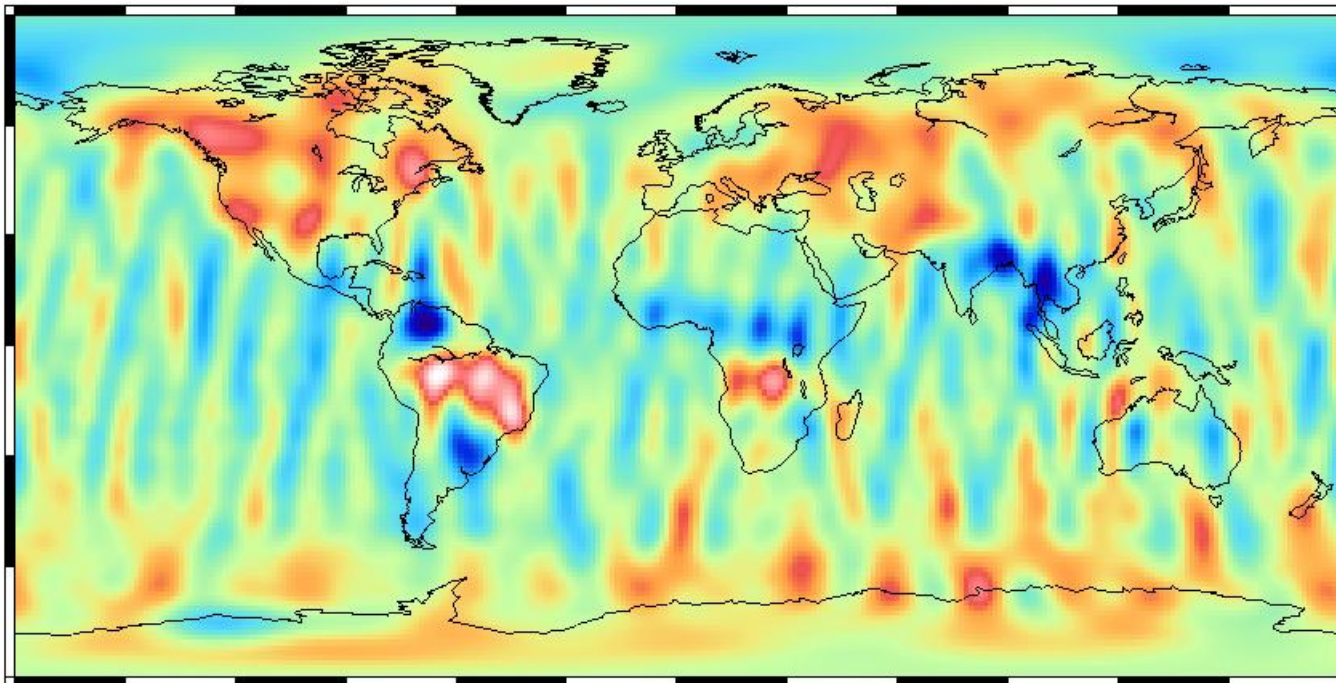
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Mass in water equivalent (cm)

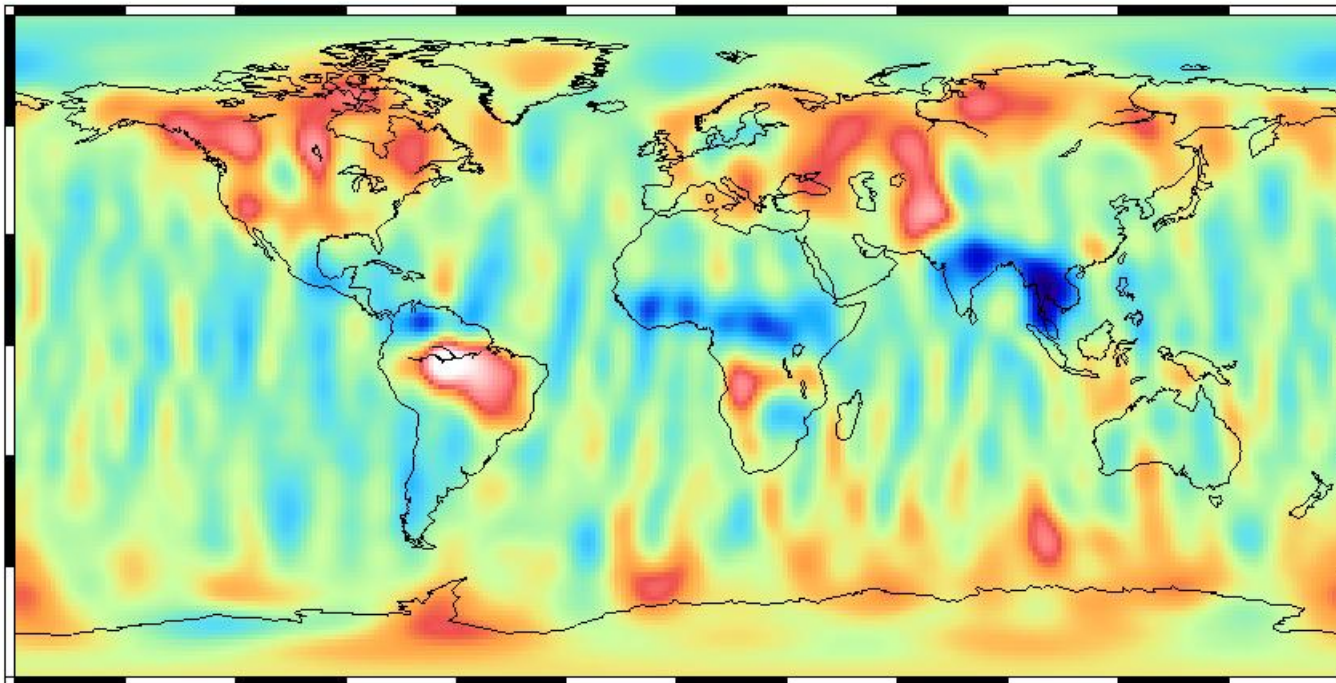


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Mass in water equivalent (cm)

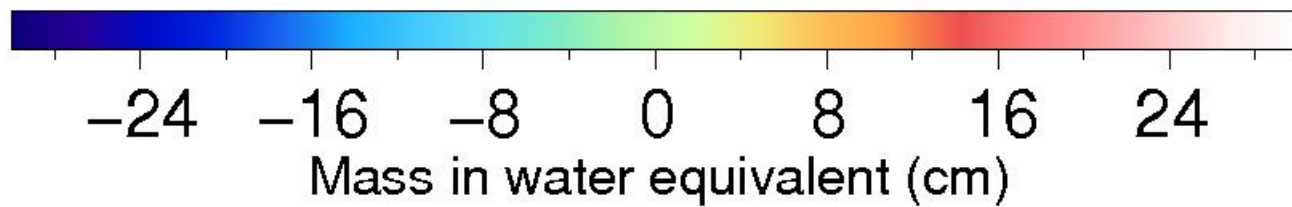
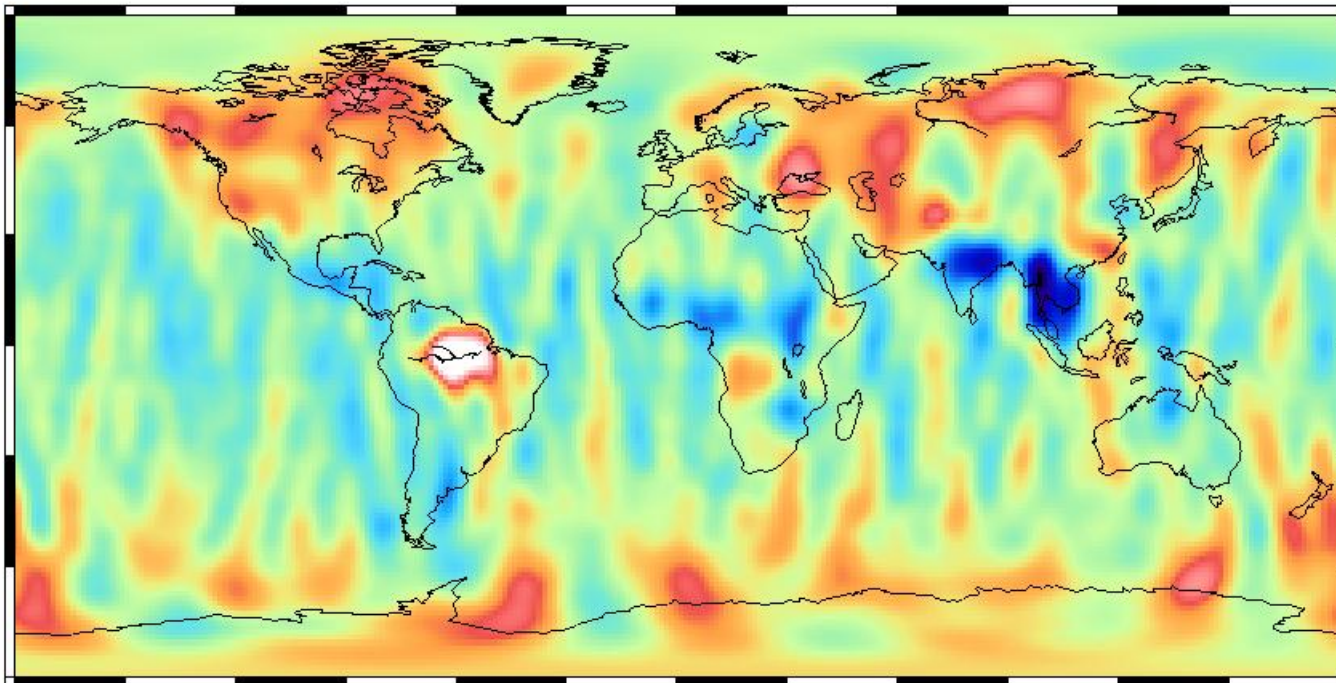


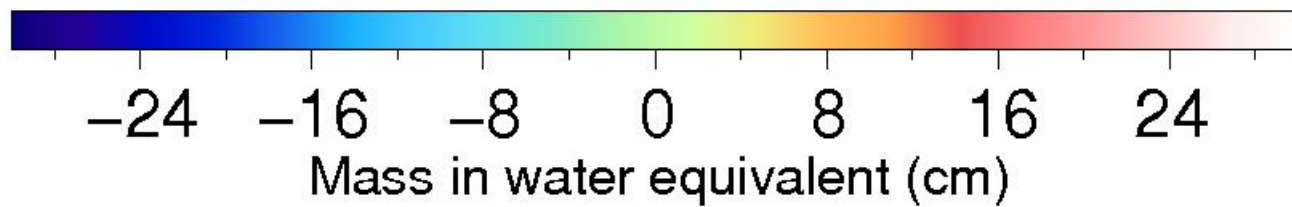
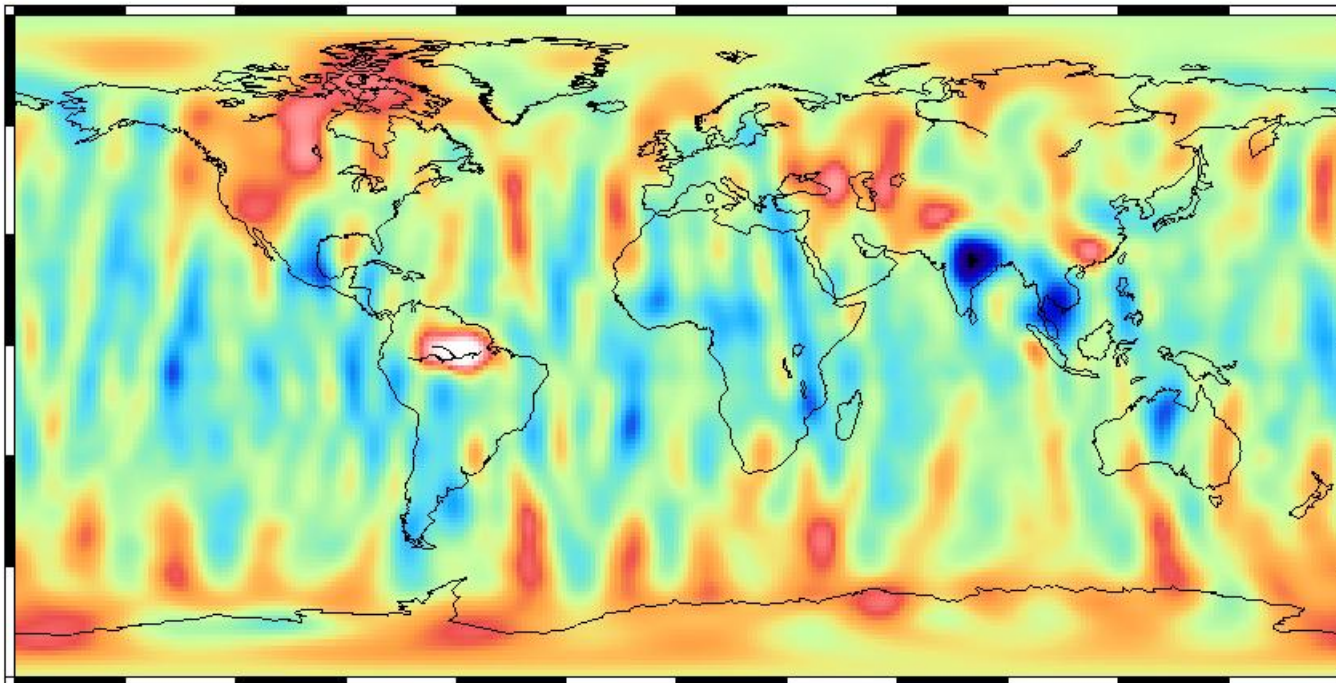
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Mass in water equivalent (cm)

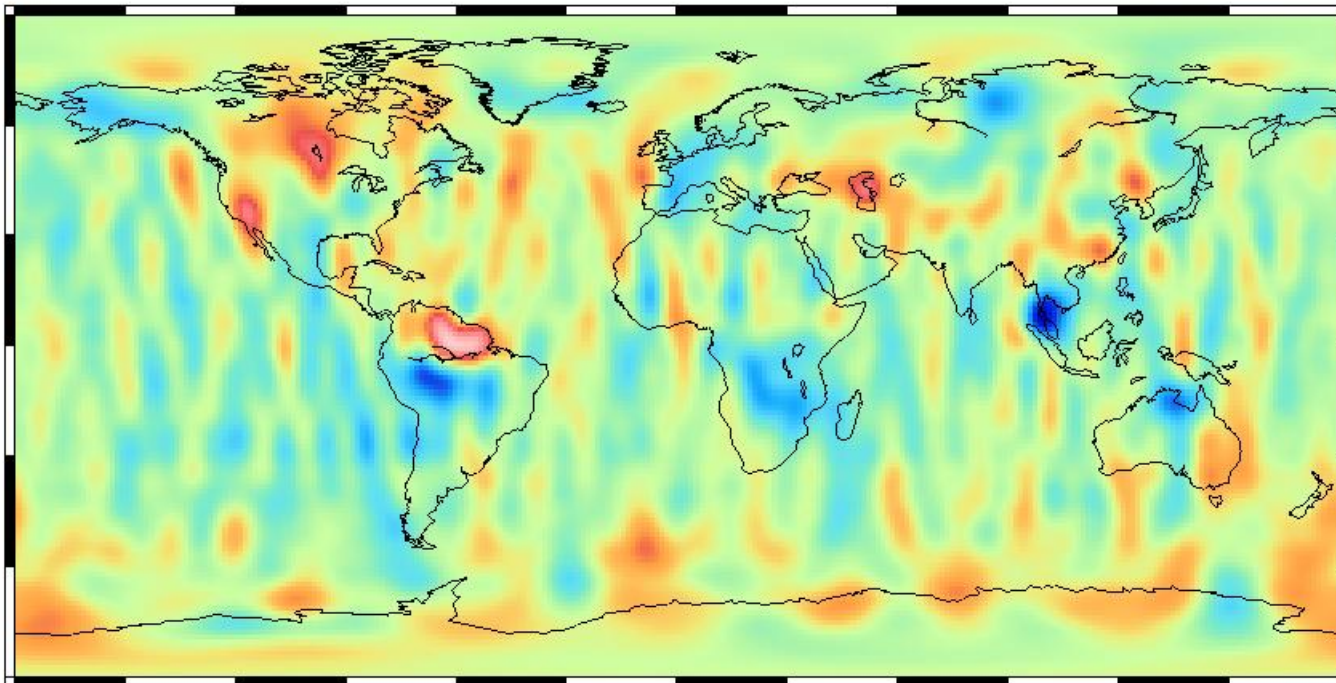




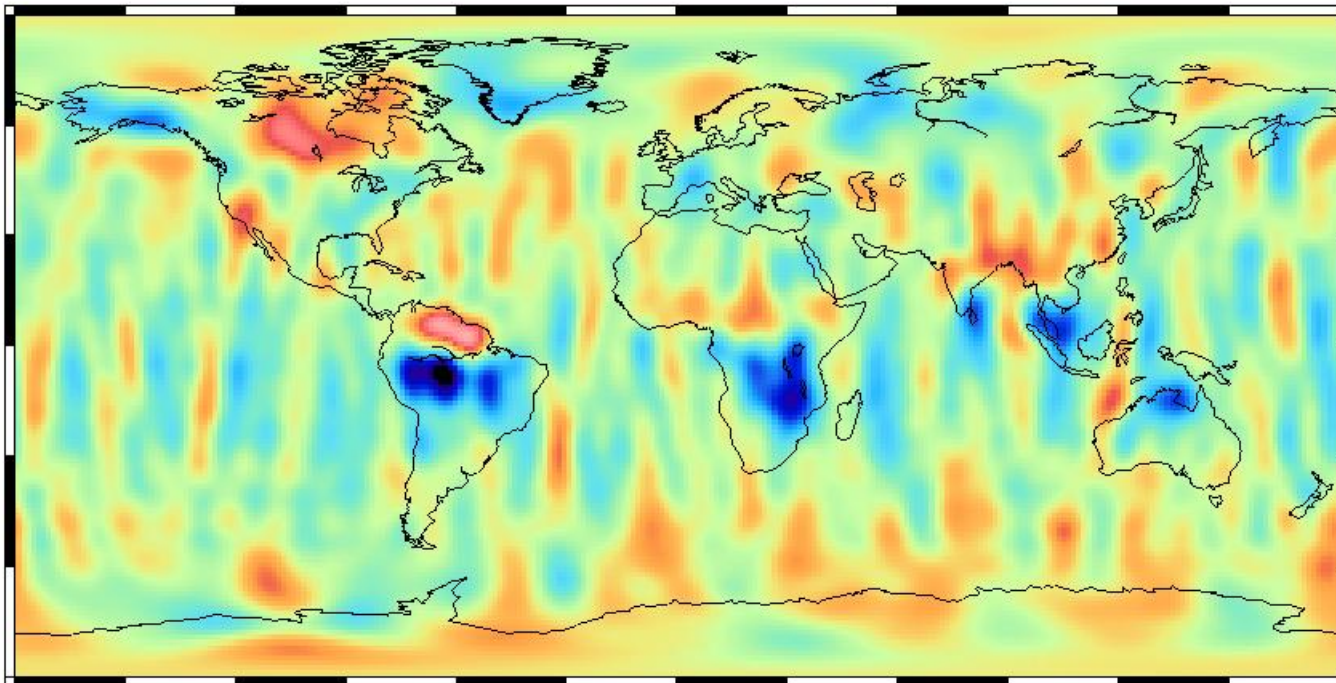
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Mass in water equivalent (cm)



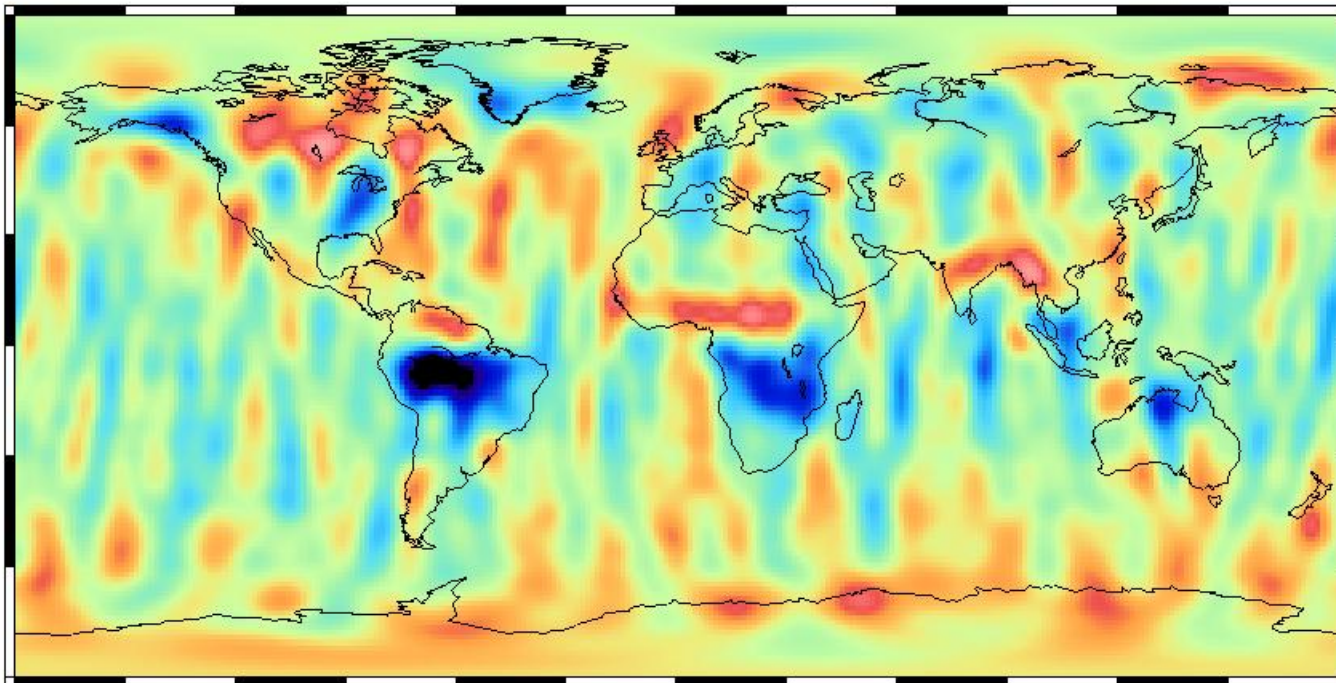




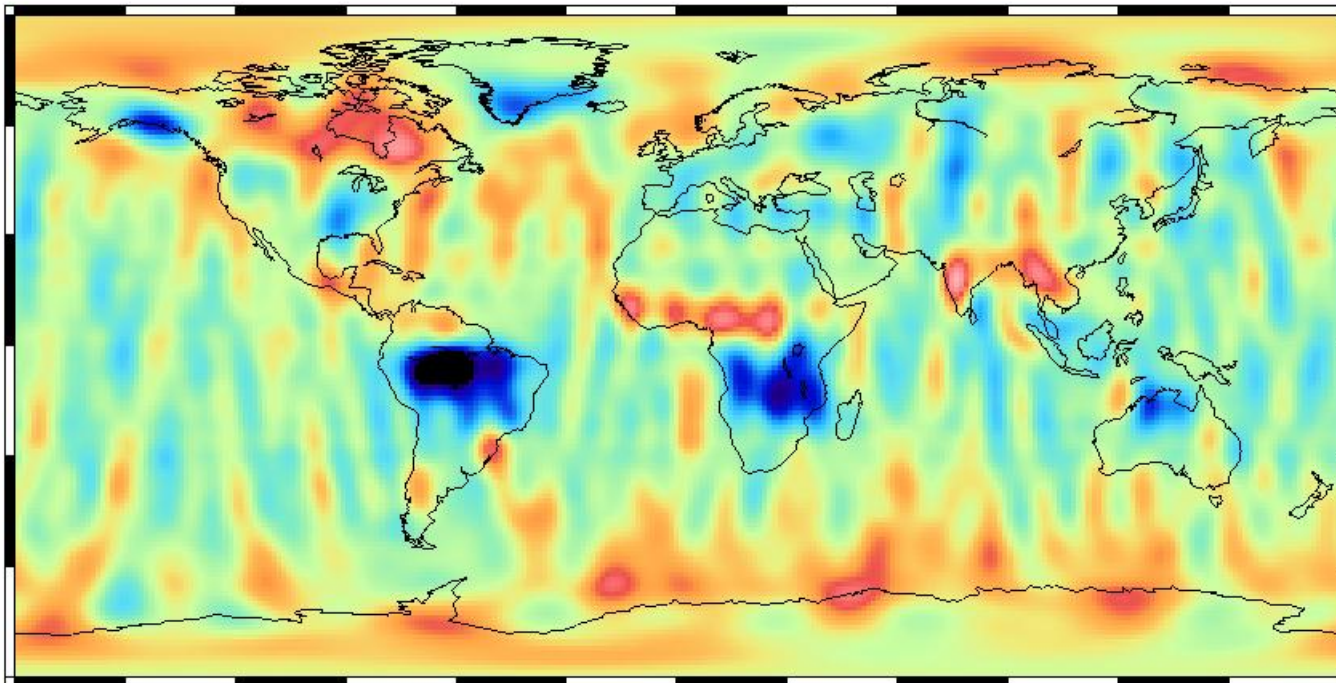
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Mass in water equivalent (cm)



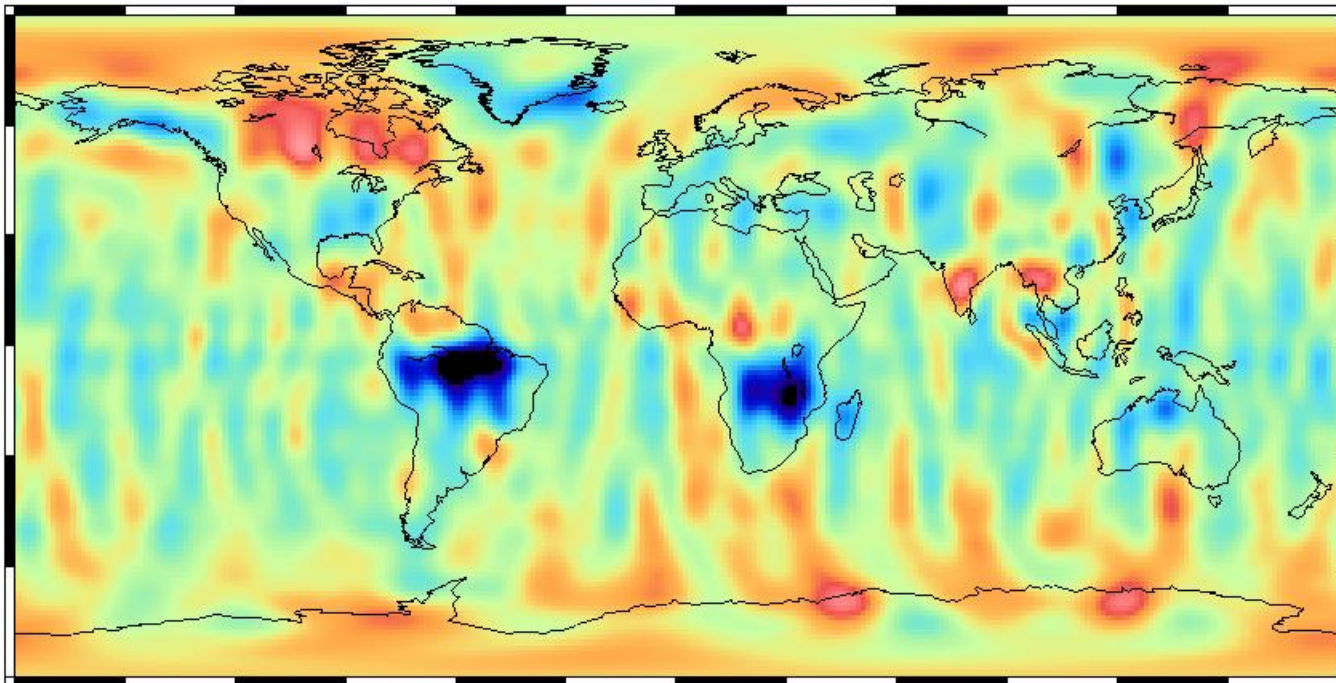
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Mass in water equivalent (cm)



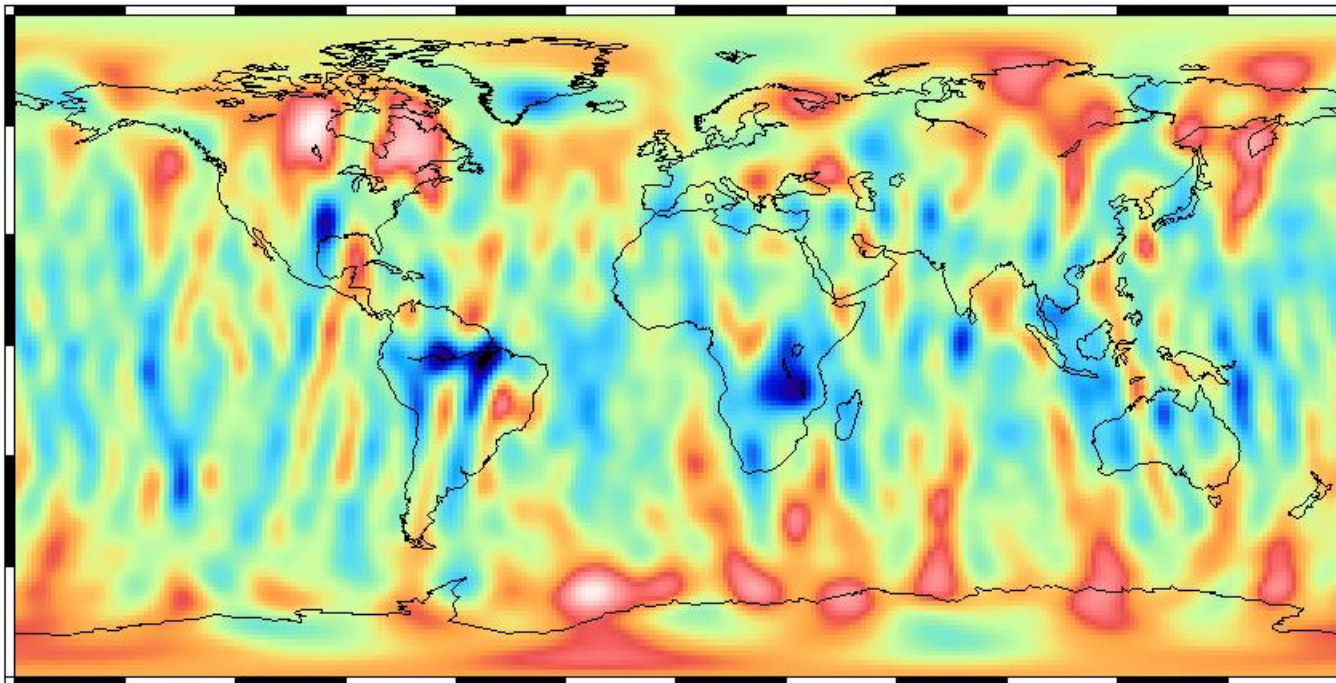
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Mass in water equivalent (cm)



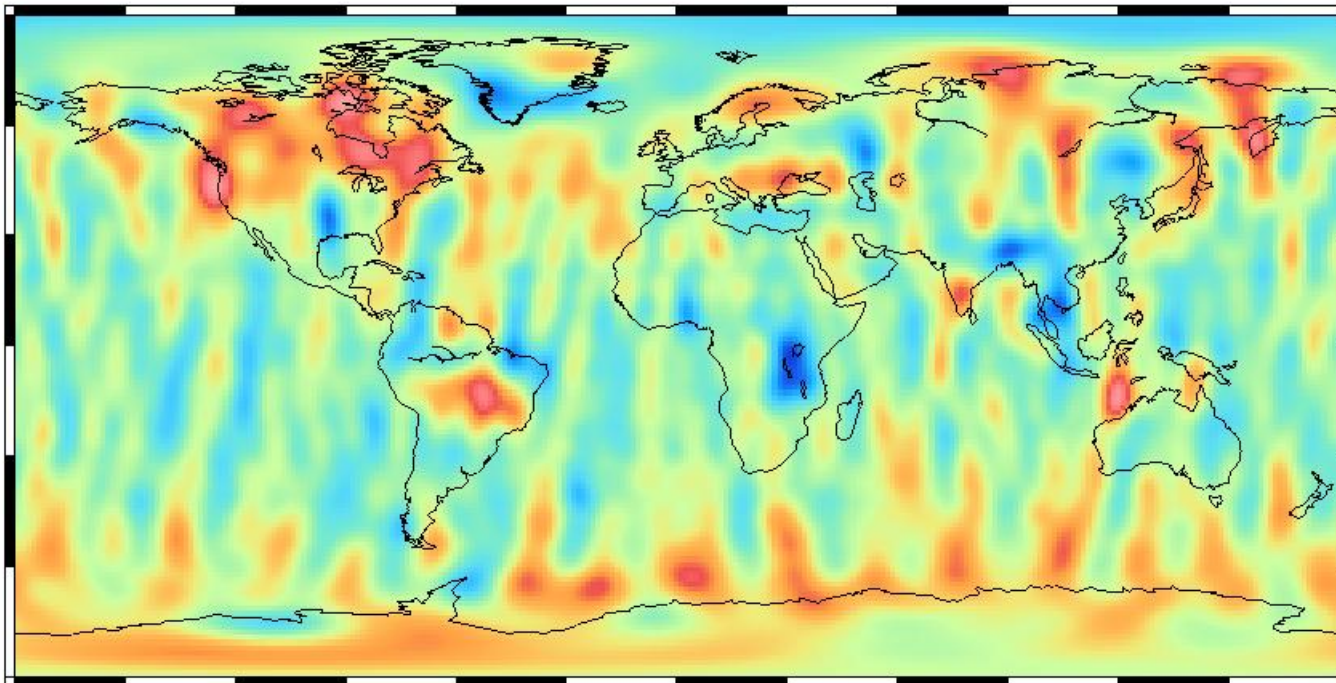
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Mass in water equivalent (cm)



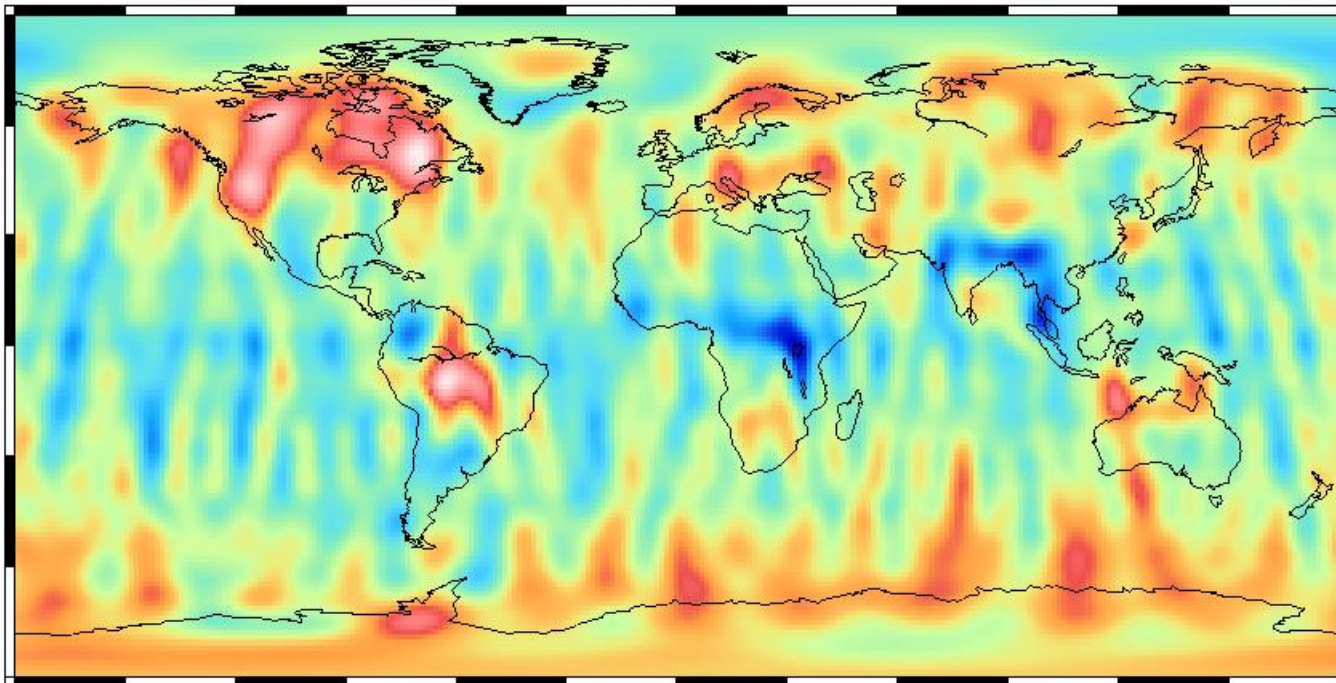
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Mass in water equivalent (cm)



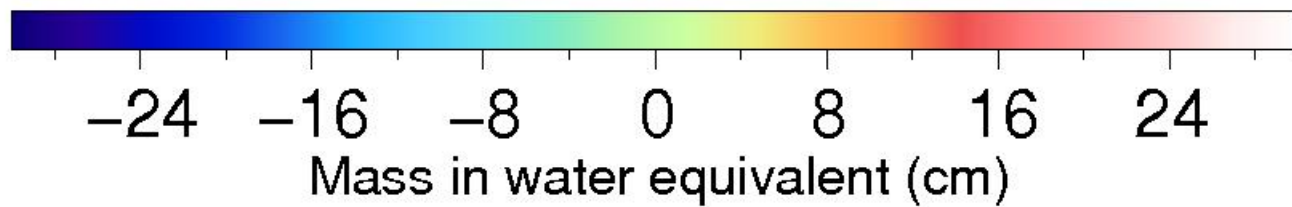
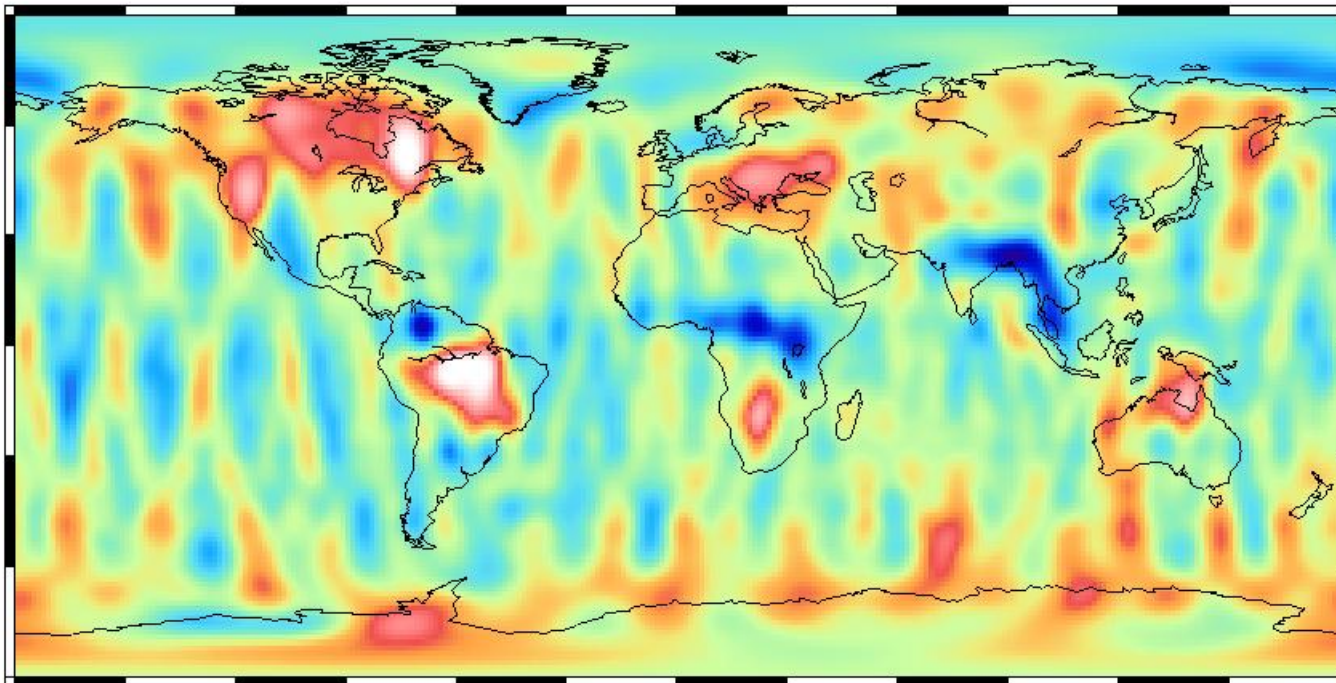
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Mass in water equivalent (cm)

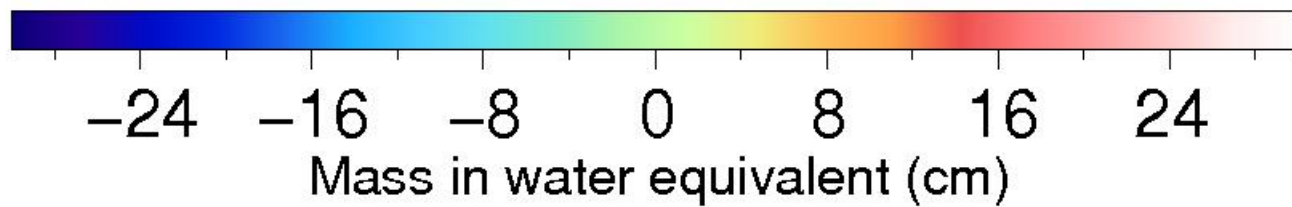
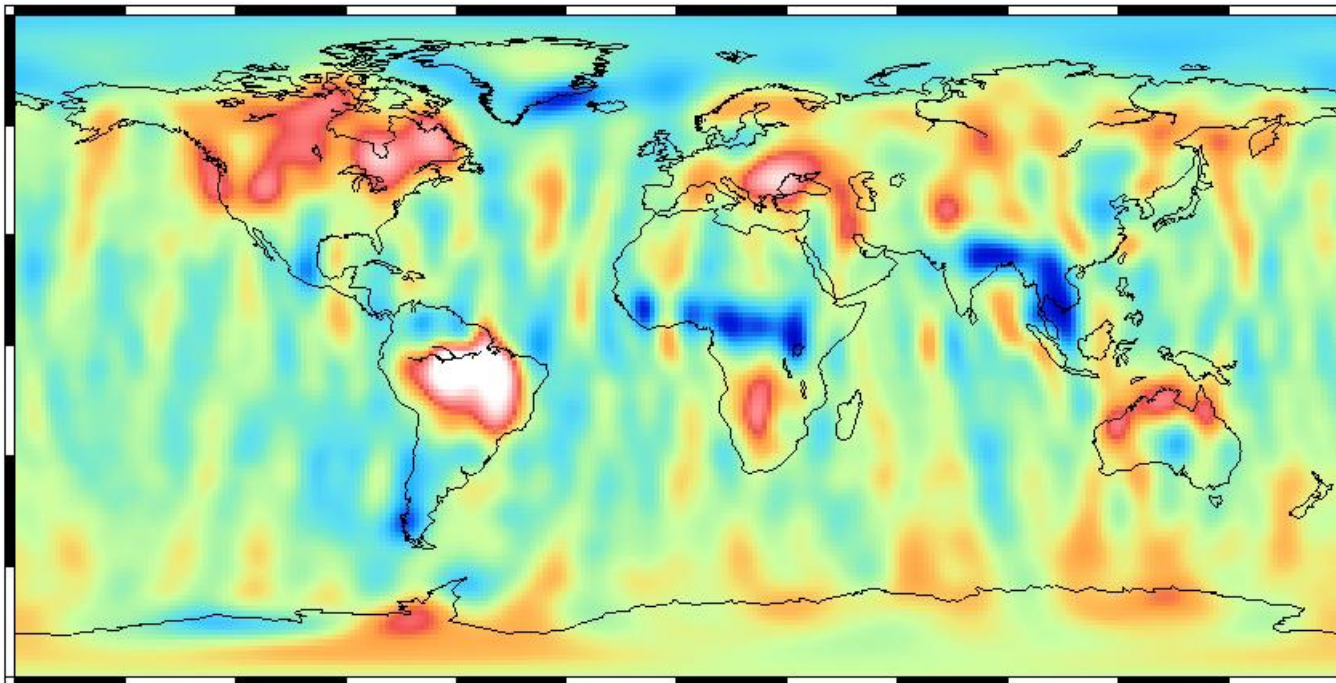


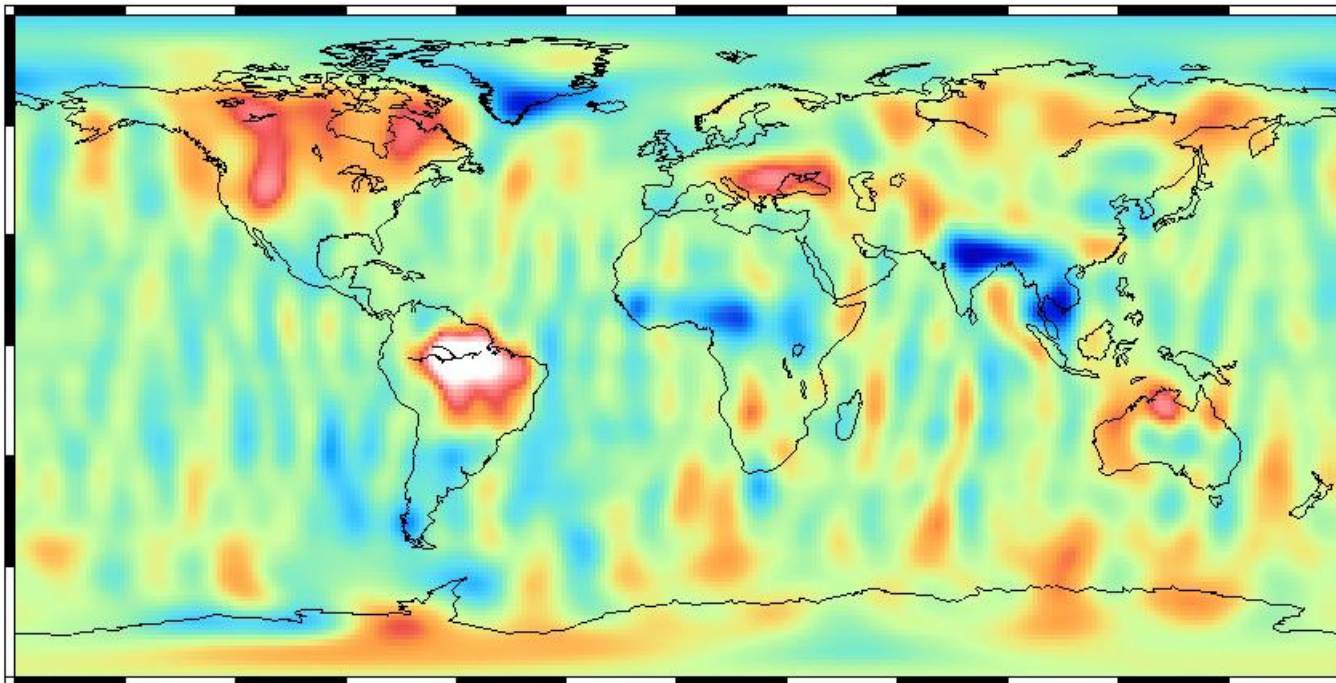
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Mass in water equivalent (cm)



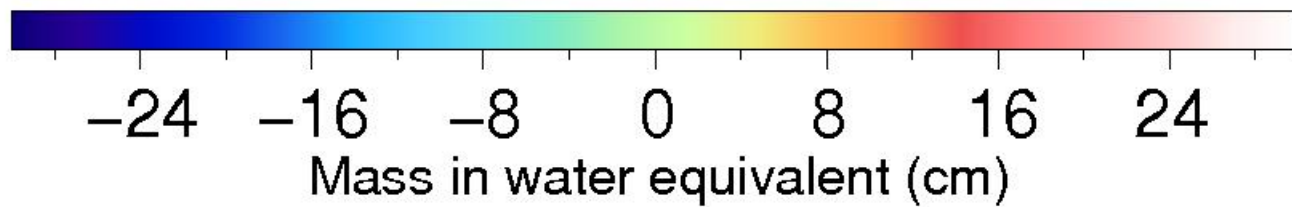
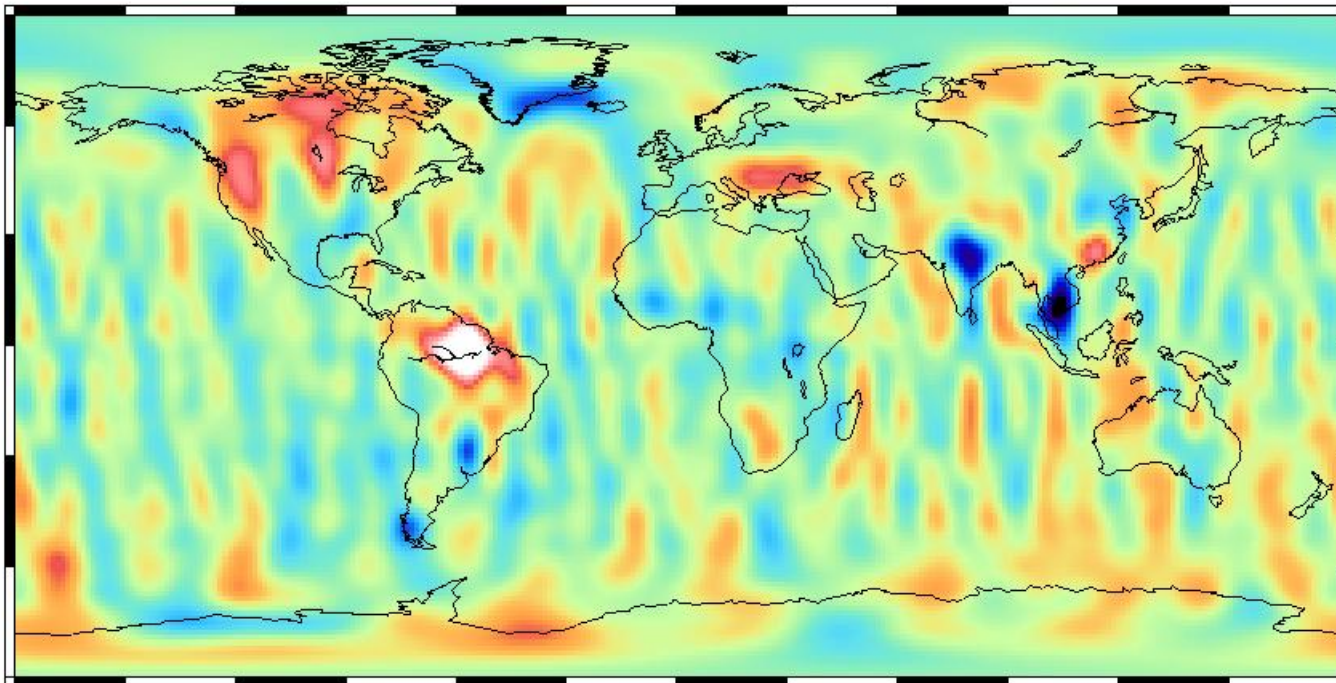
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Mass in water equivalent (cm)

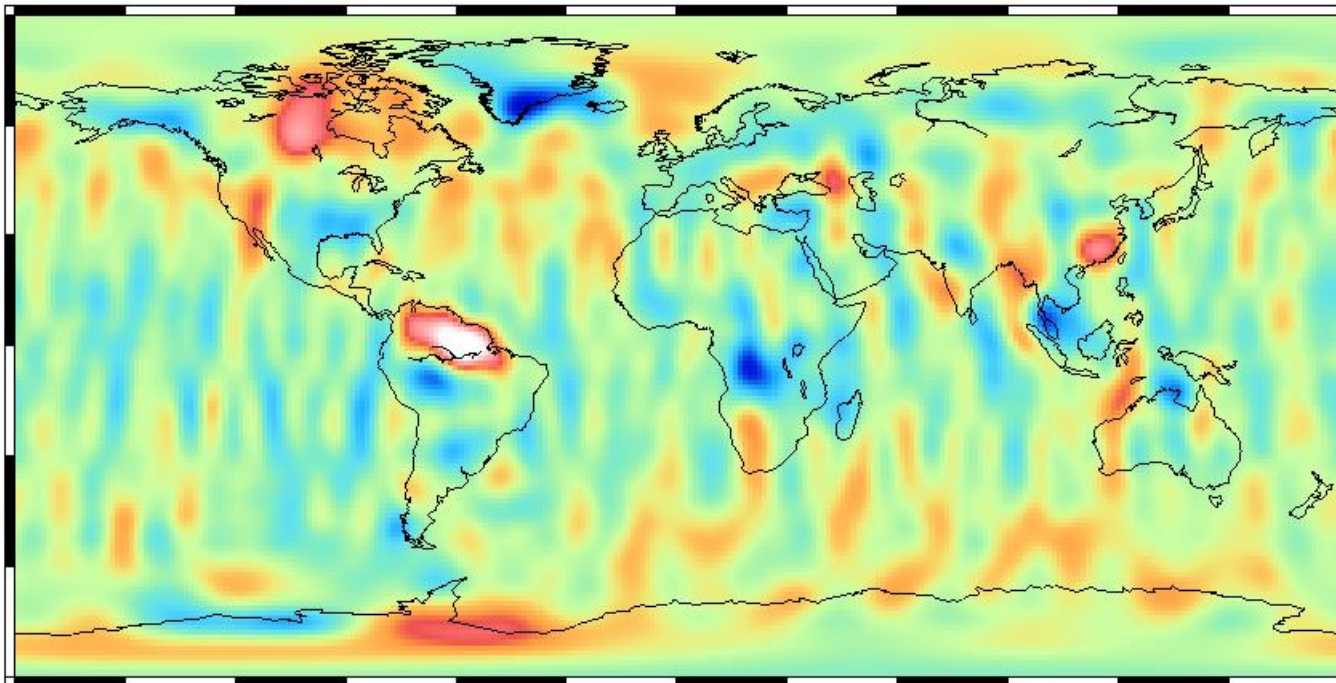




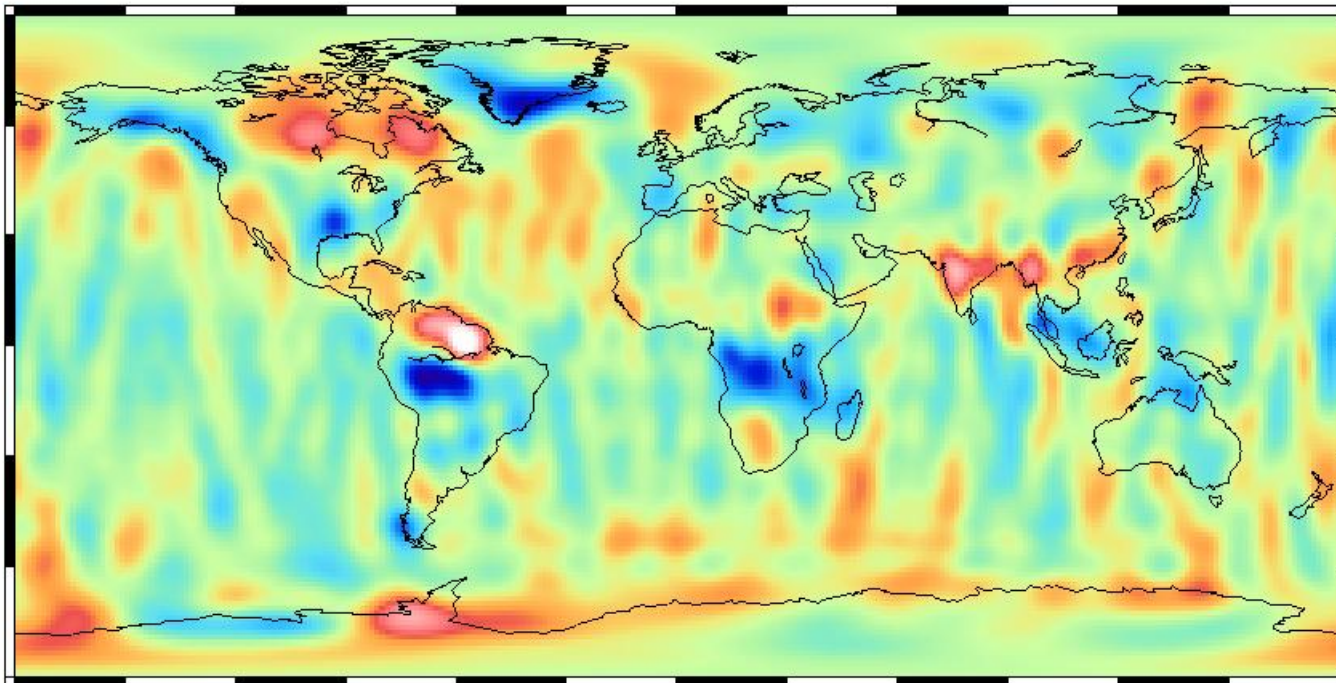


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Mass in water equivalent (cm)

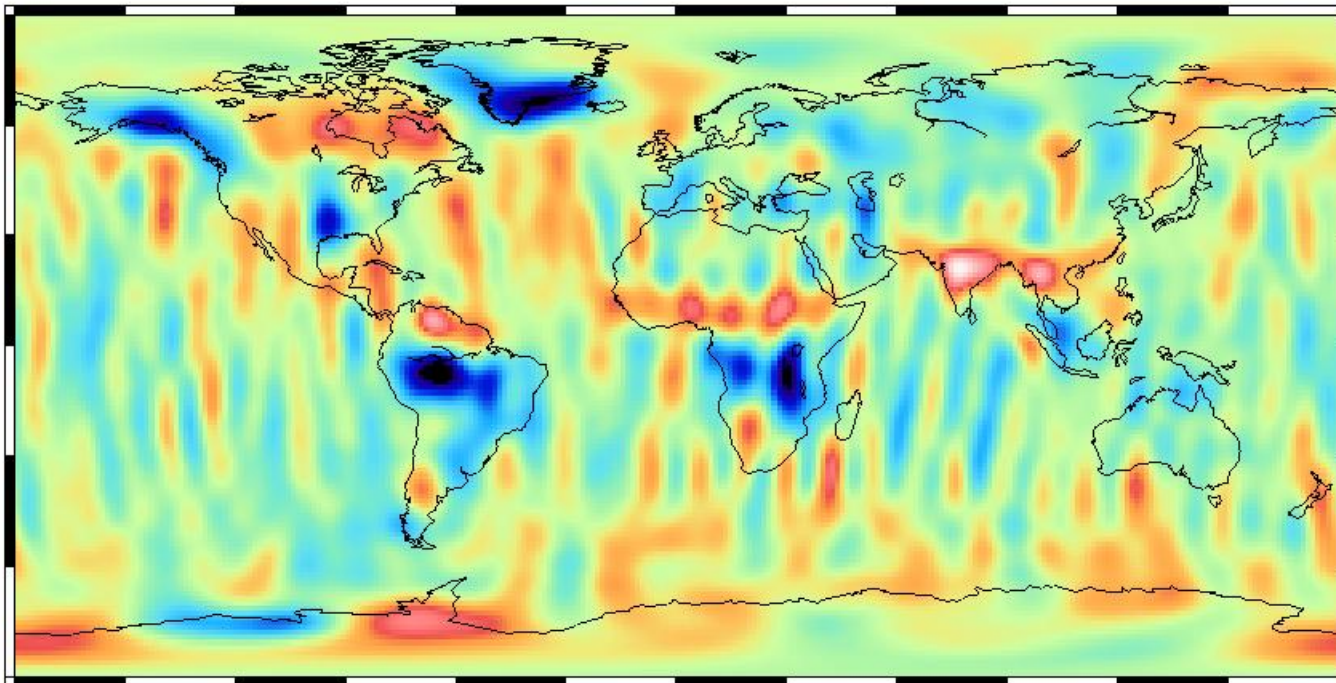




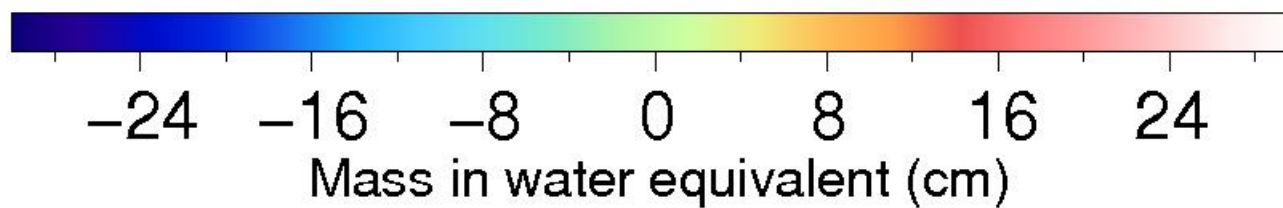
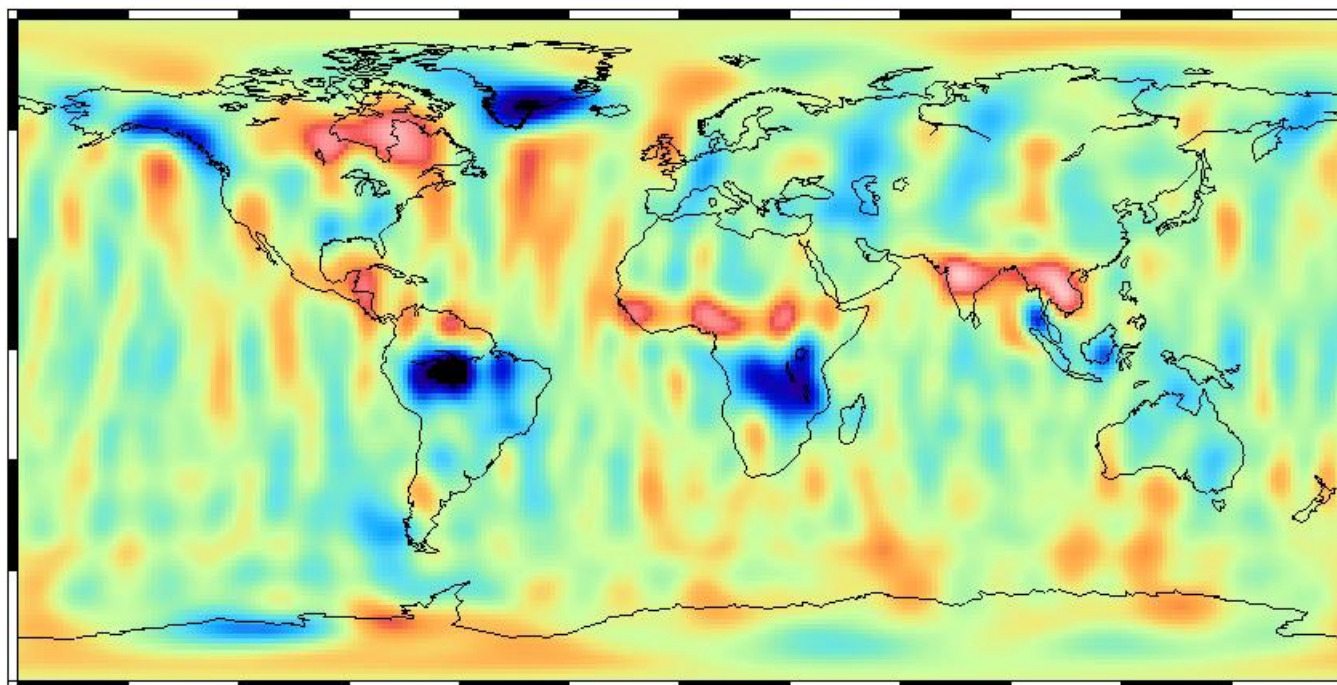
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Mass in water equivalent (cm)

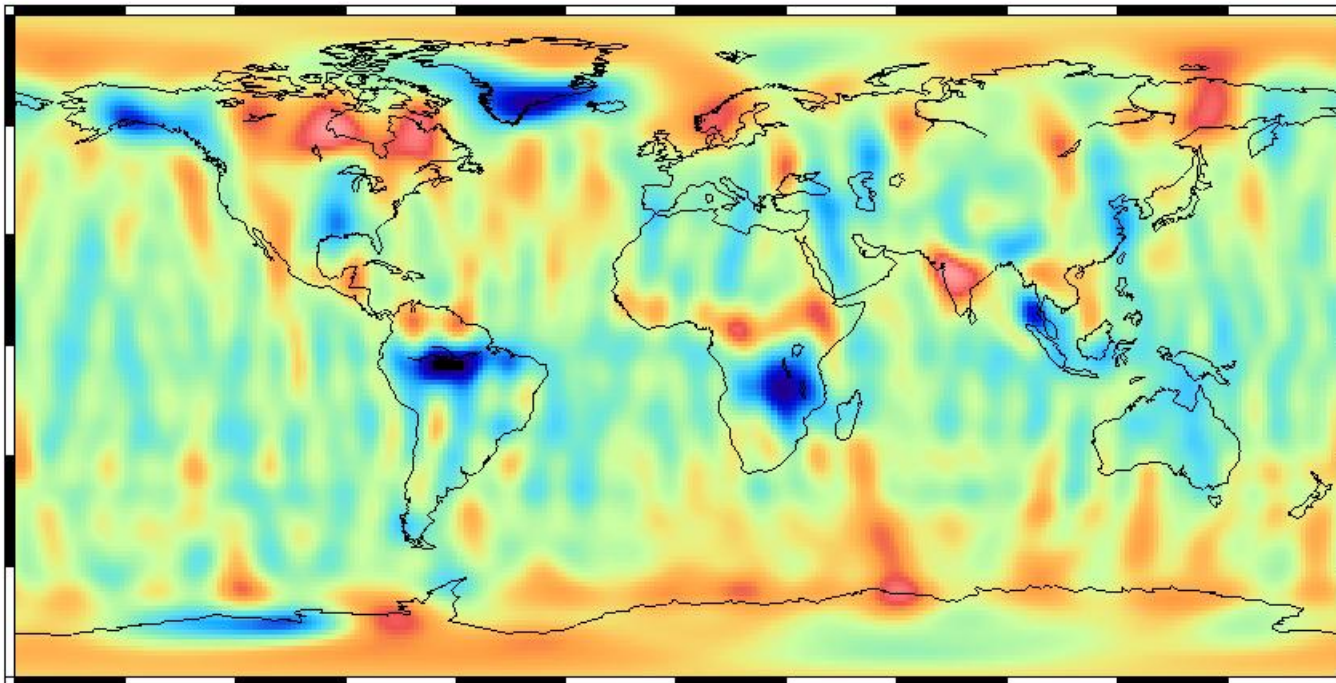


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Mass in water equivalent (cm)

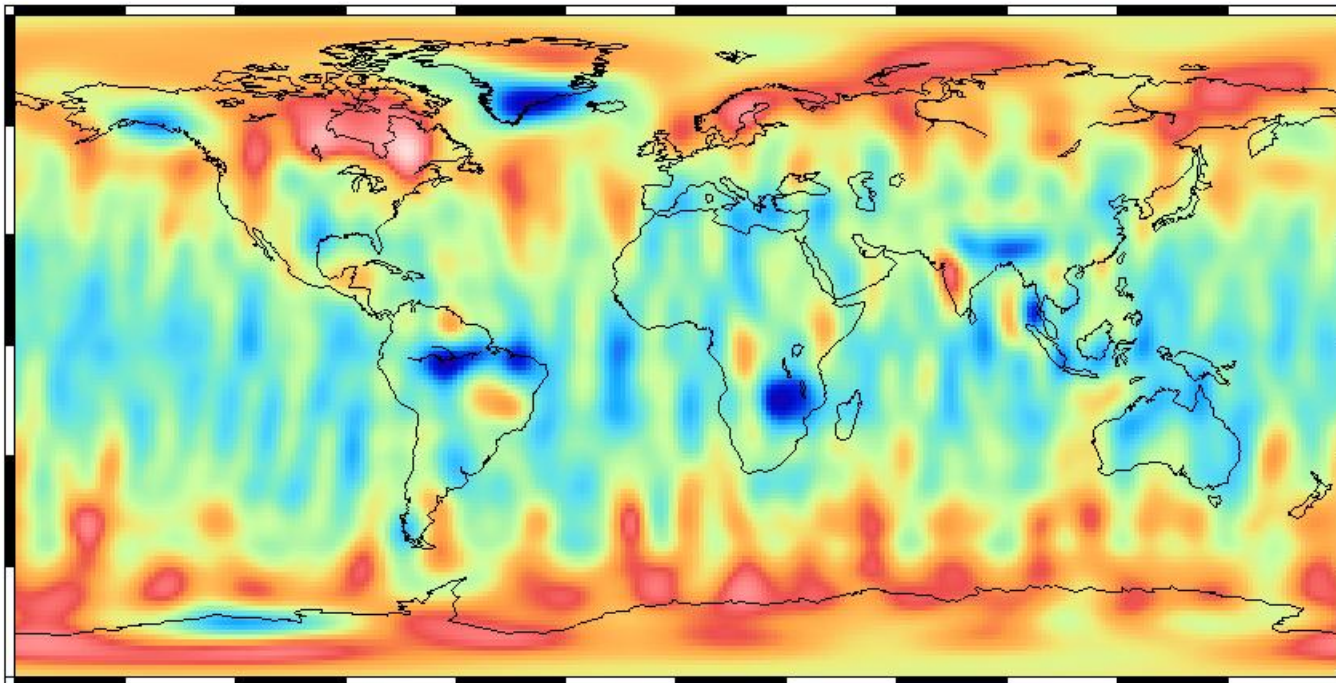


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Mass in water equivalent (cm)

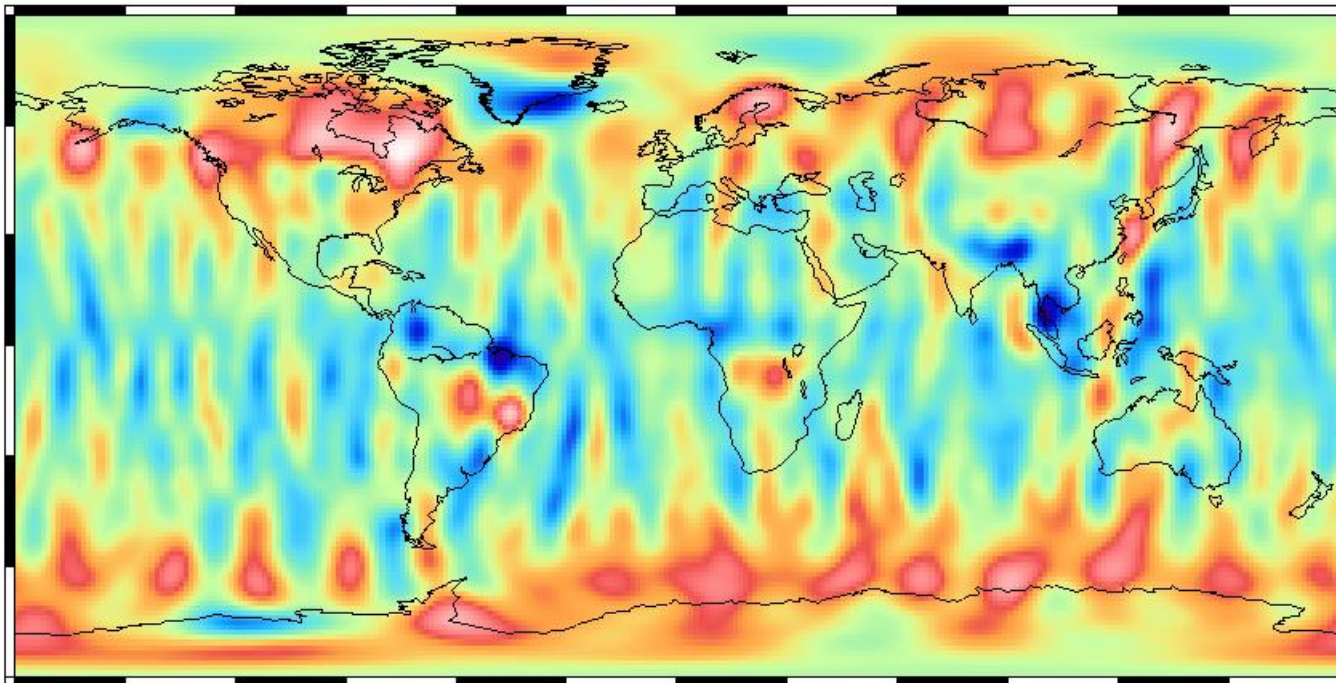




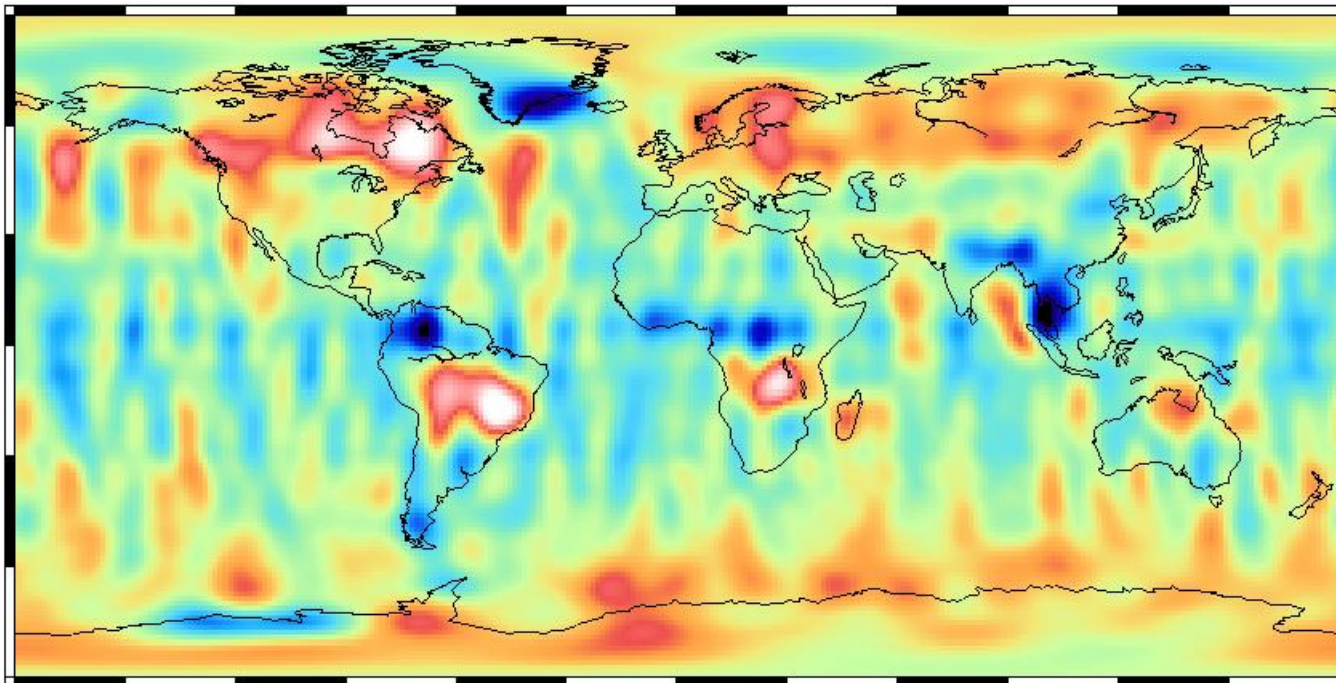
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Mass in water equivalent (cm)



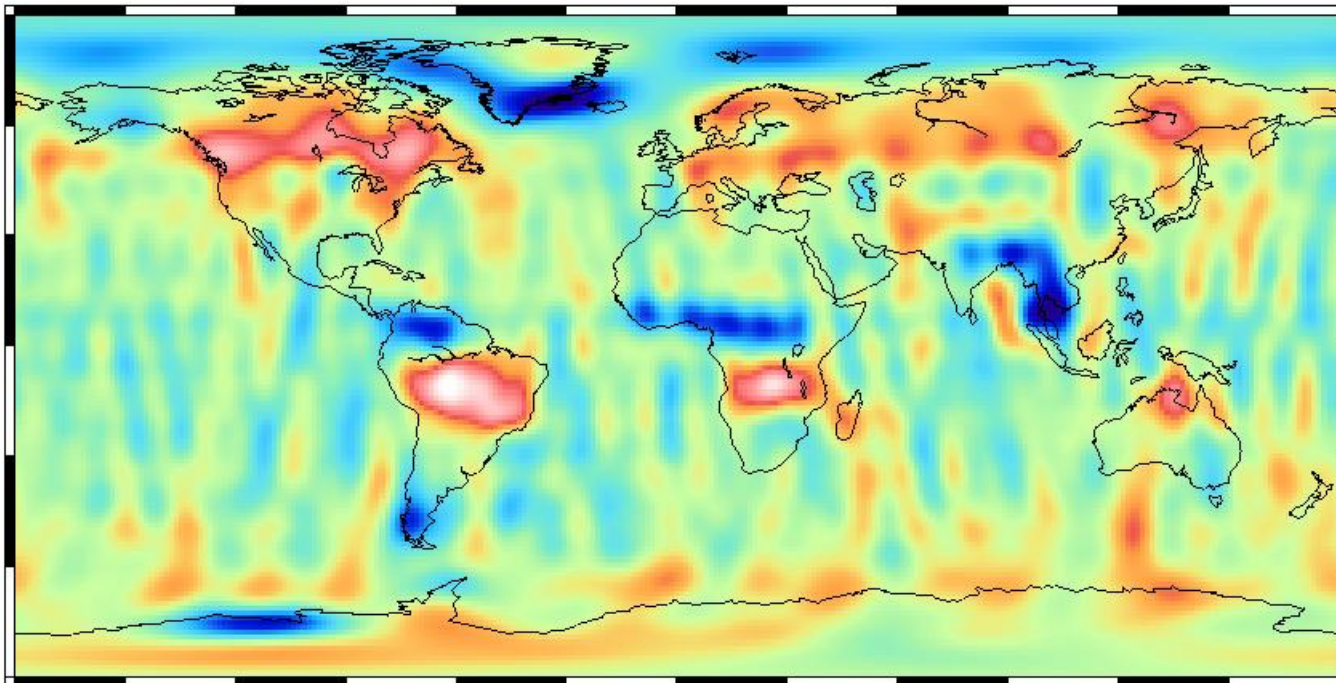
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Mass in water equivalent (cm)



-24 -16 -8 0 8 16 24
Mass in water equivalent (cm)

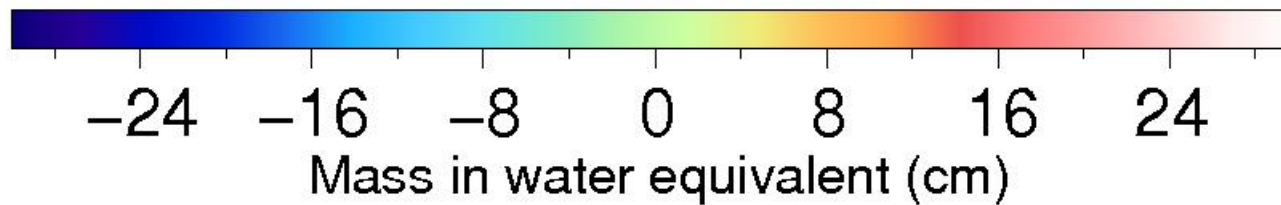
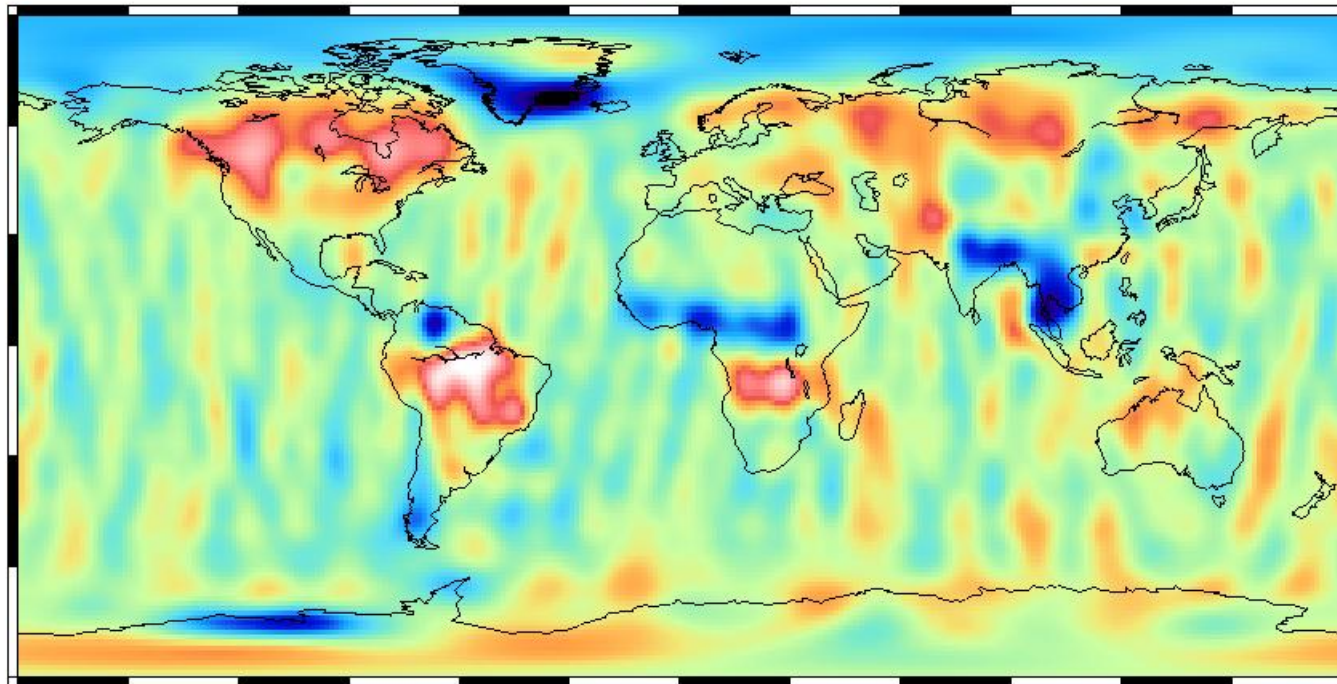


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Mass in water equivalent (cm)

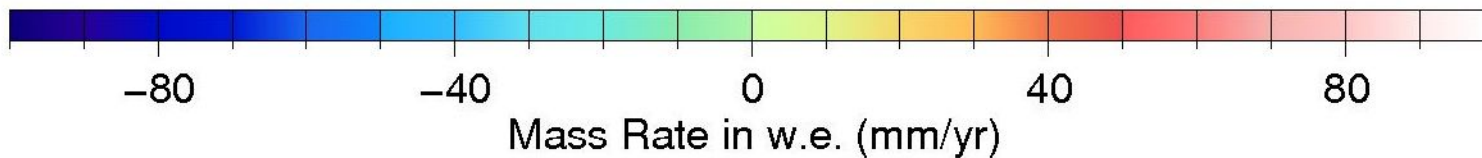
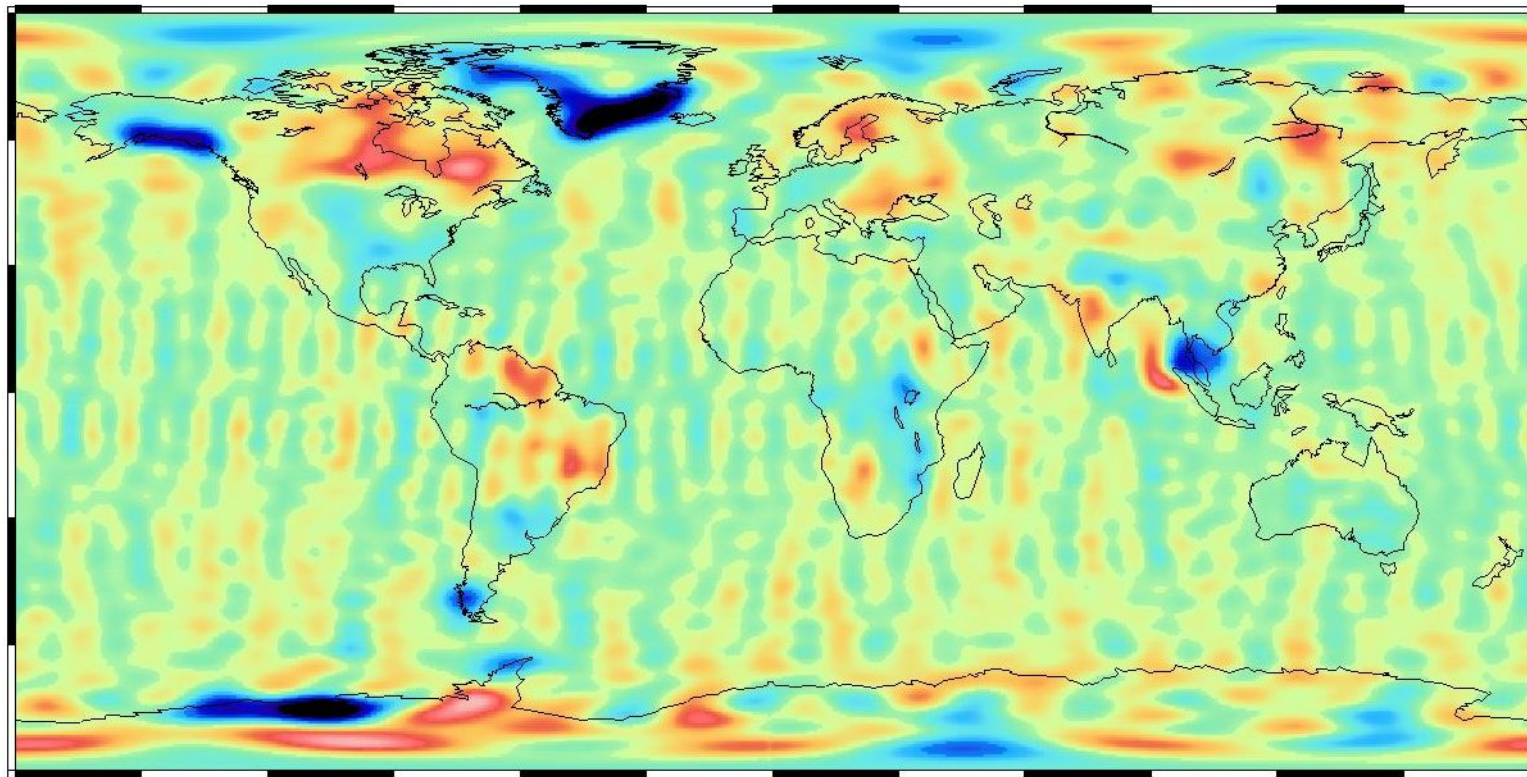


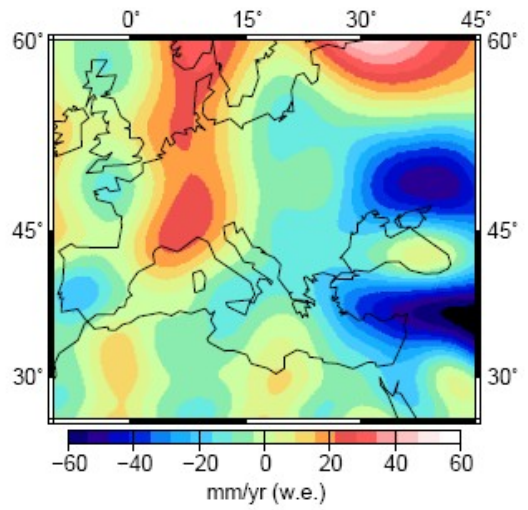
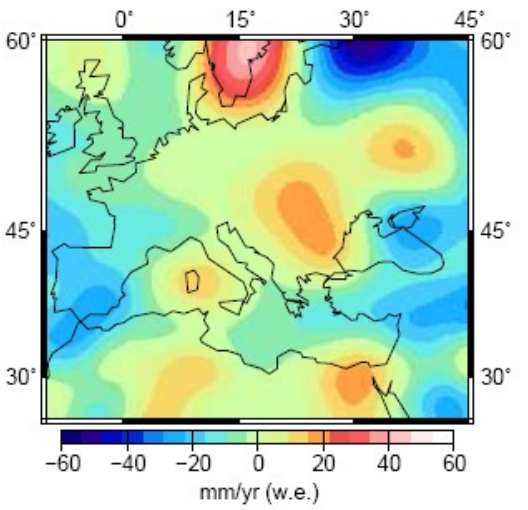
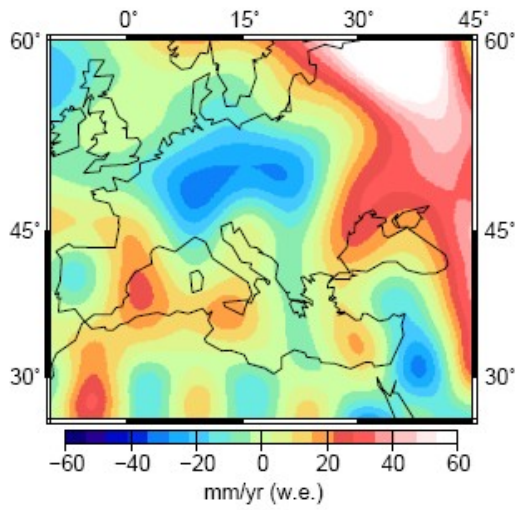
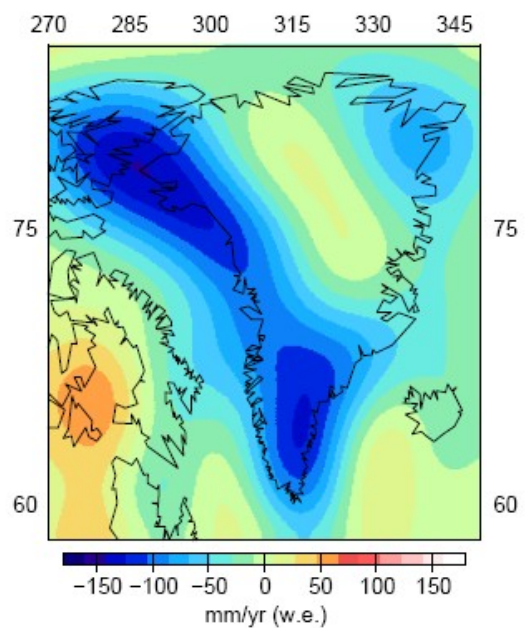
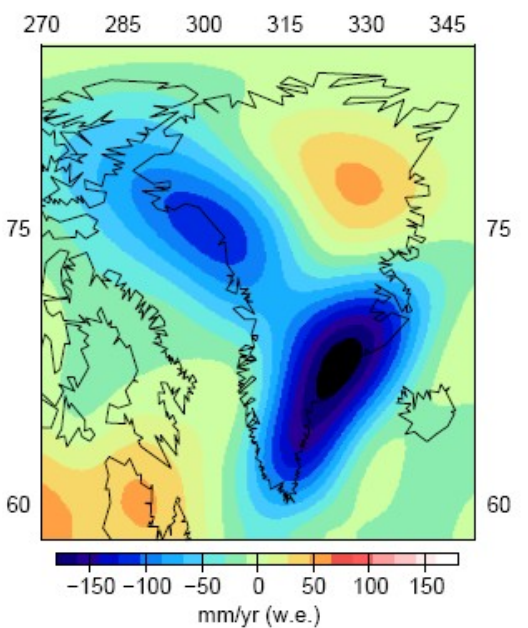
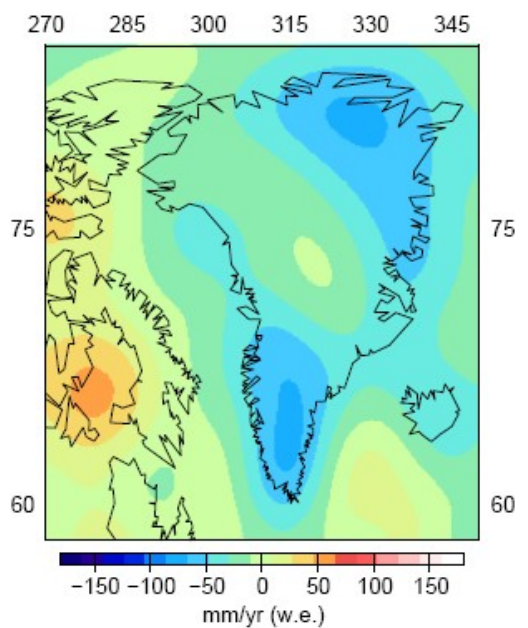
-24 -16 -8 0 8 16 24
Mass in water equivalent (cm)

April 2007

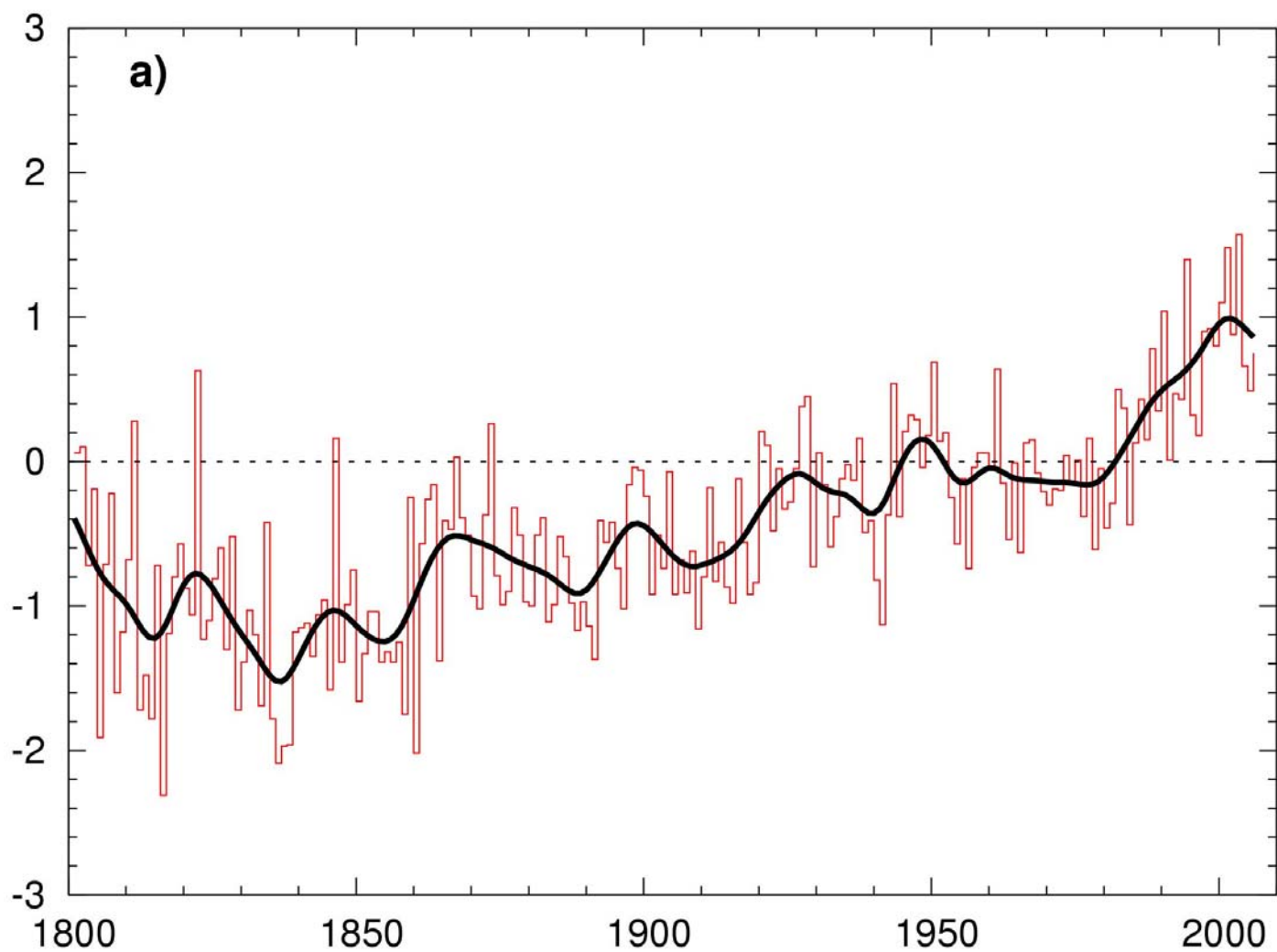


The Map of Mass Variation Trend - Filtered

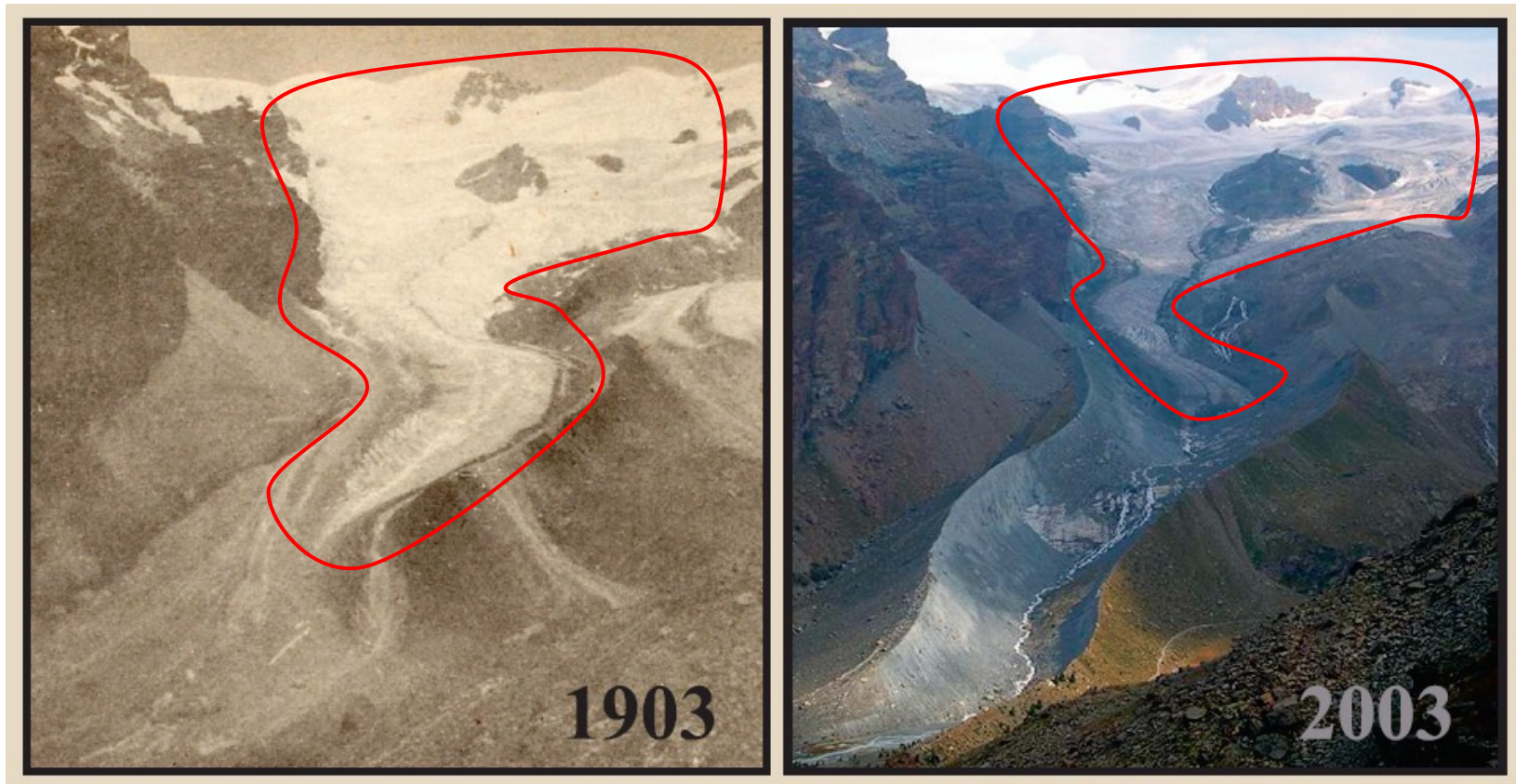


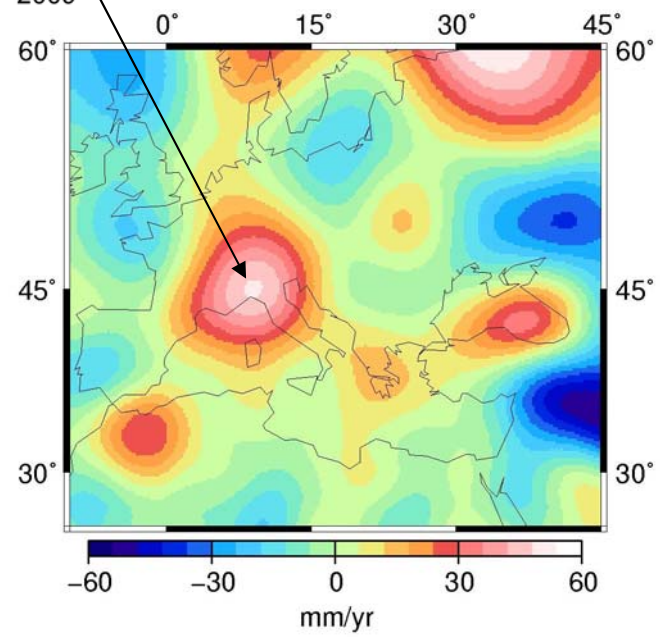
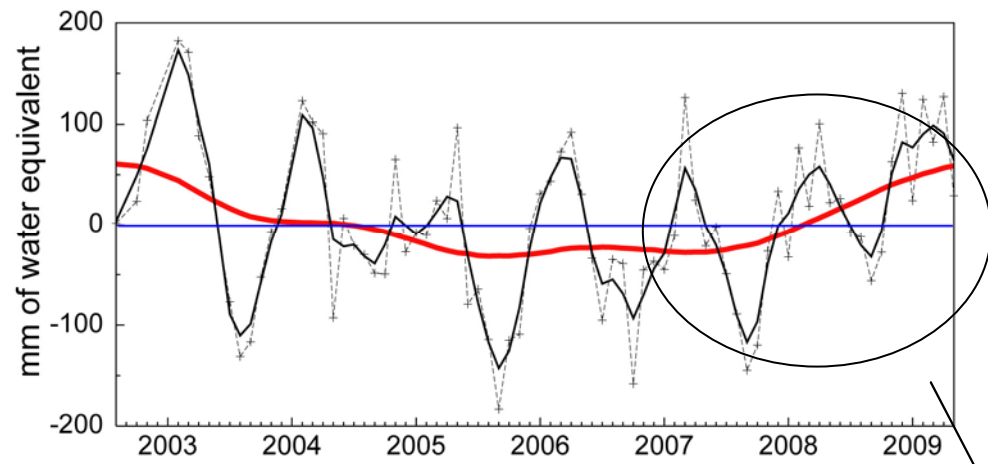


La temperatura media in Italia (scarti, in °C, rispetto ai valori medi del periodo di riferimento 1961-1990) dall'inizio del XIX secolo



Ghiacciaio Verra (Monte Rosa)



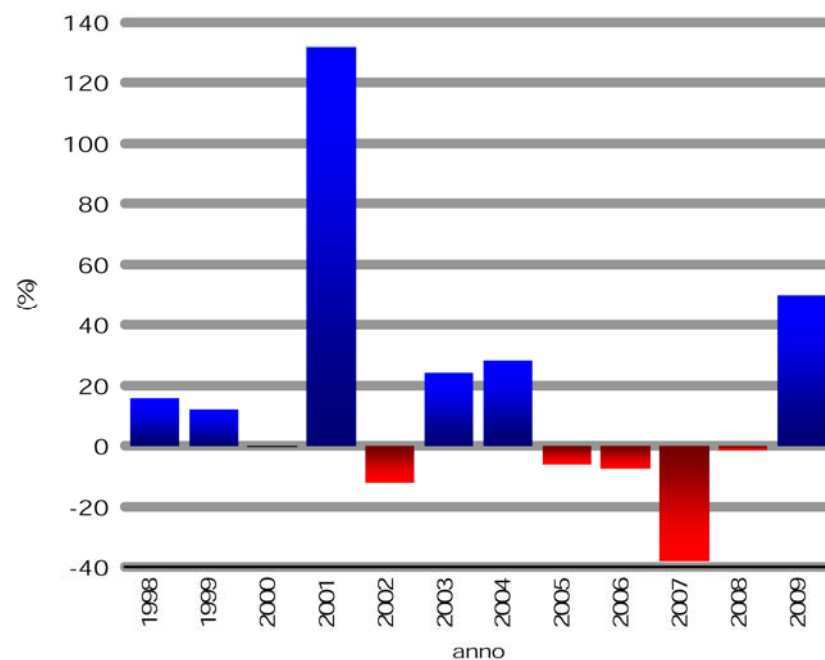




BILANCI DI ACCUMULO 2008/2009 IN LOMBARDIA

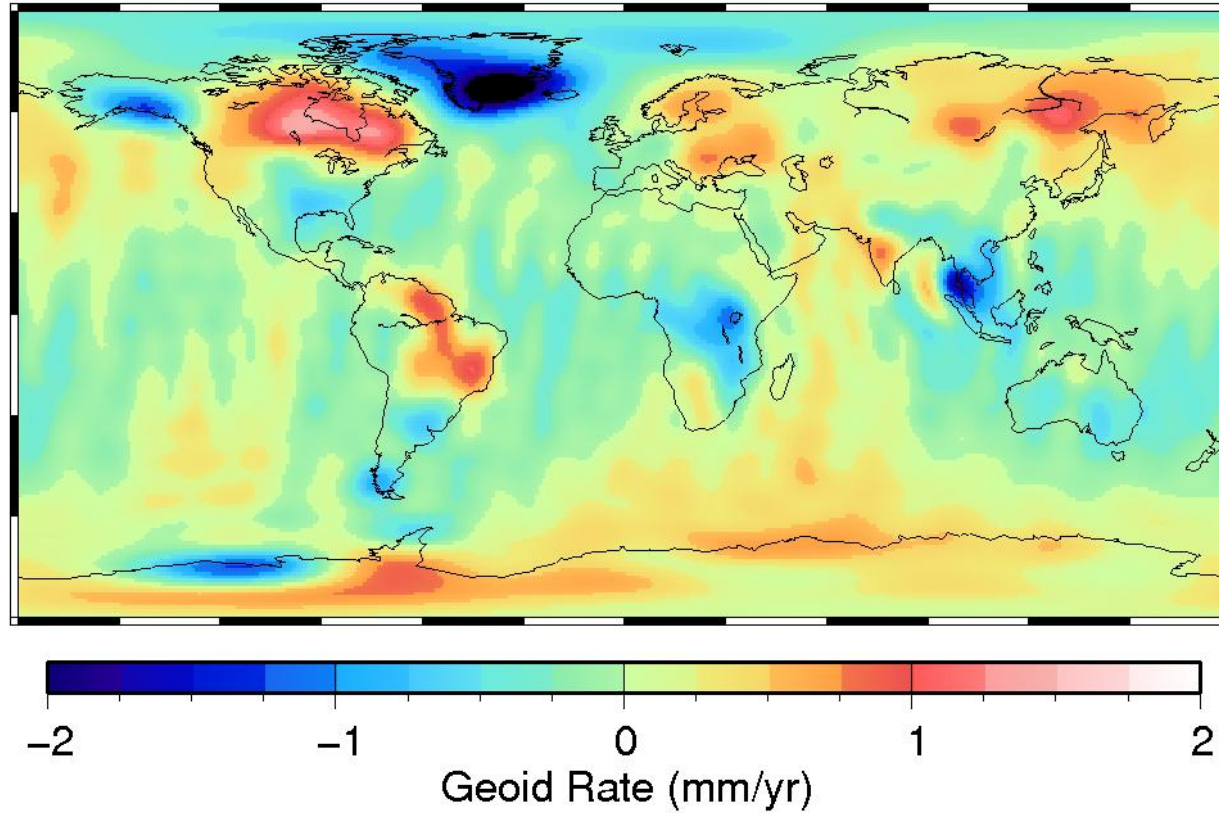
Dati dai siti nivologici del Servizio Glaciologico Lombardo

Risultati preliminari

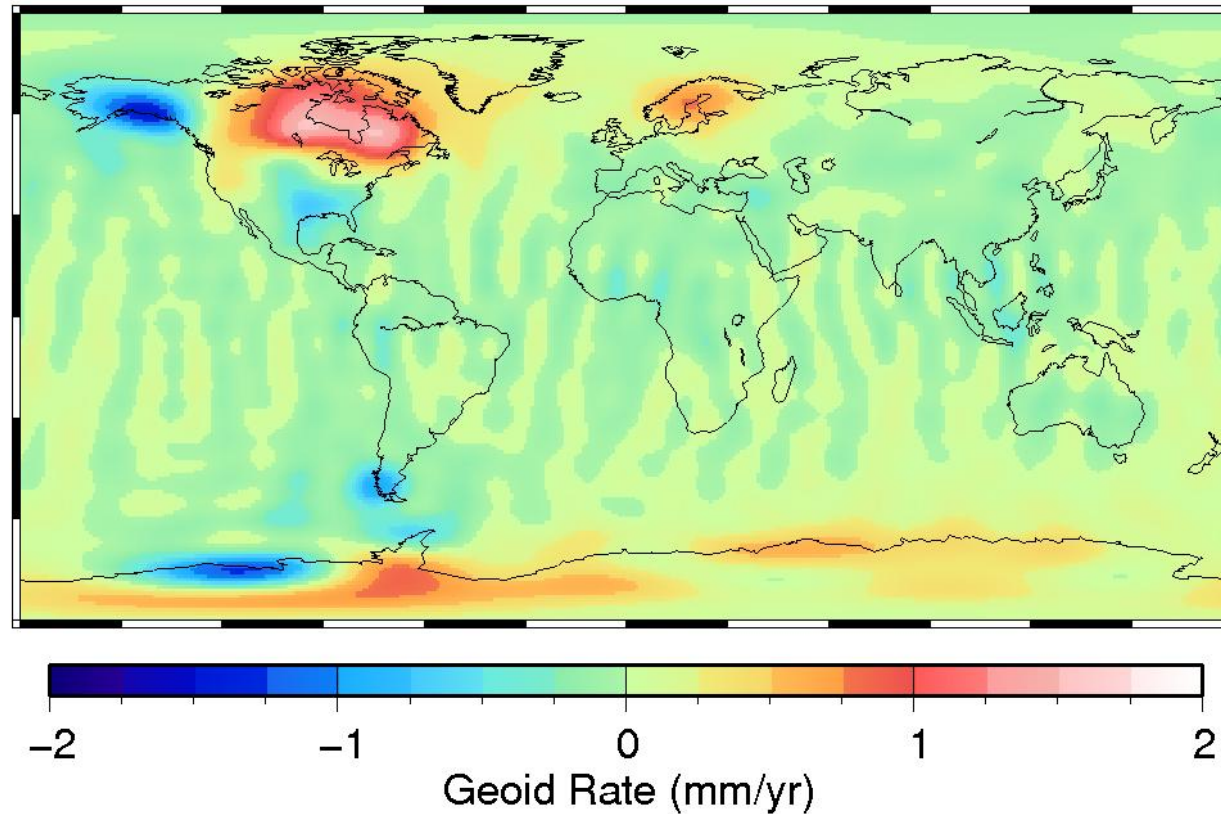


**Variazione percentuale dell'altezza neve (HN)
presso i siti nivologici SGL rispetto alla media 2003/2008**

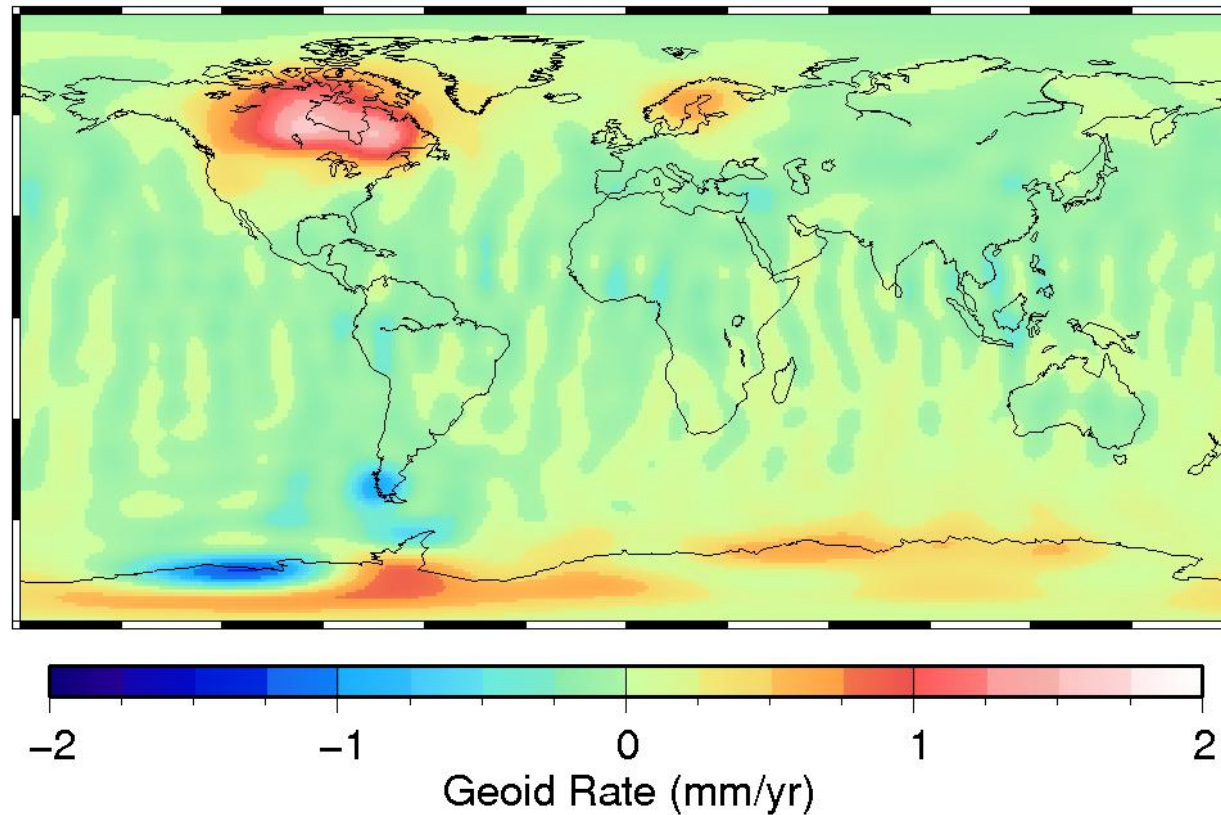
Geoid from GRACE



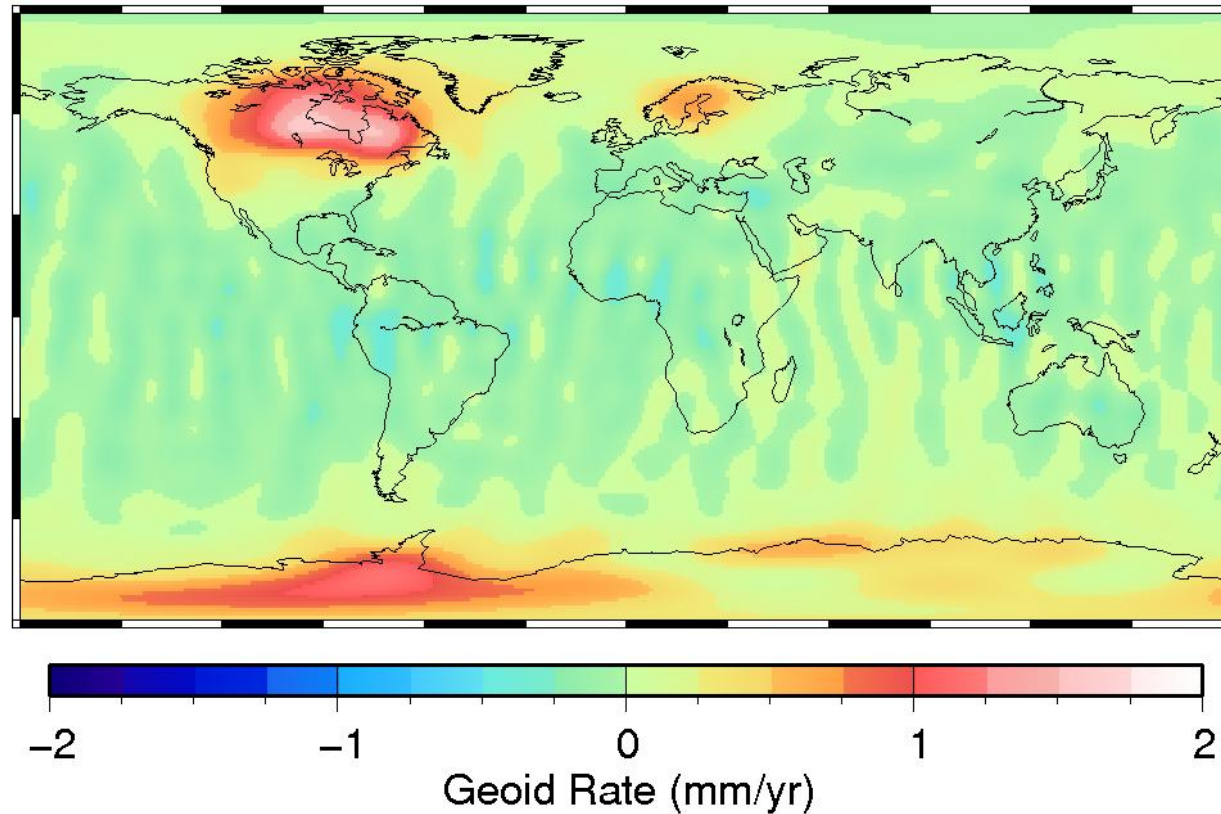
GRACE up 30 - Nearby Fennoscandia *Removed*



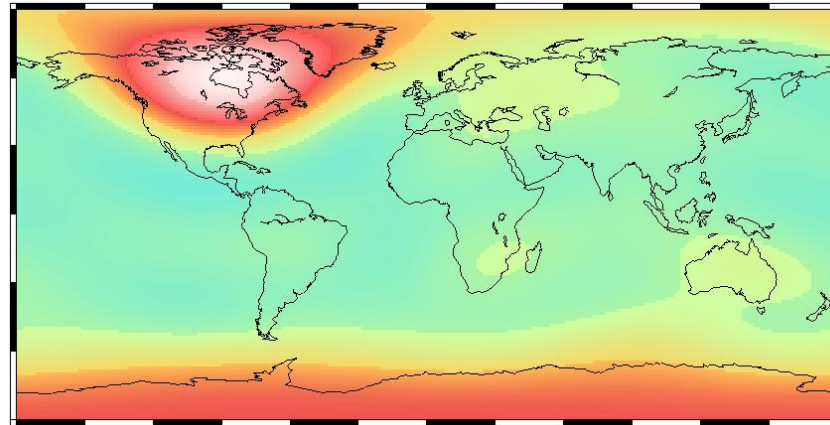
GRACE up 30 - Nearby Hudson Bay *Bay Removed*



GRACE up 30 - West Antarctica *Removed*

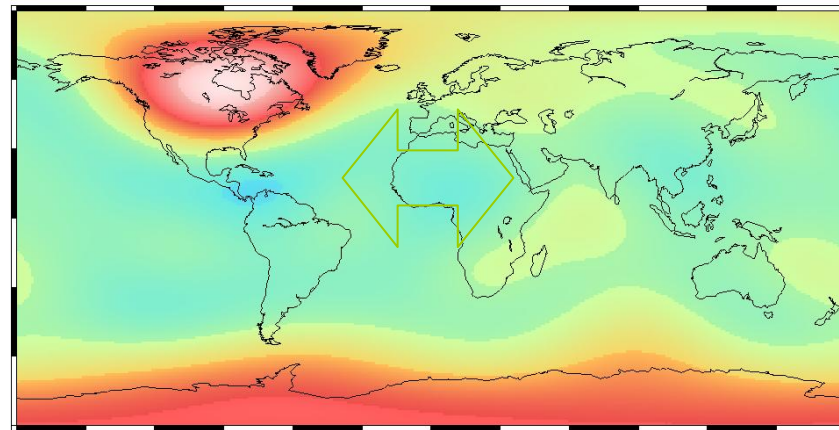
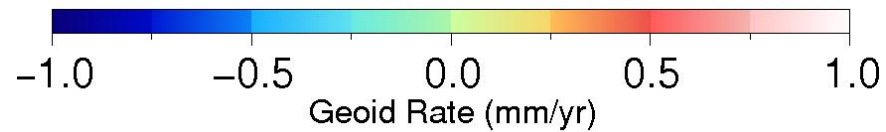


Global Problem - Search for best viscosity

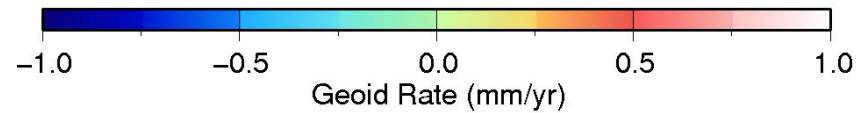


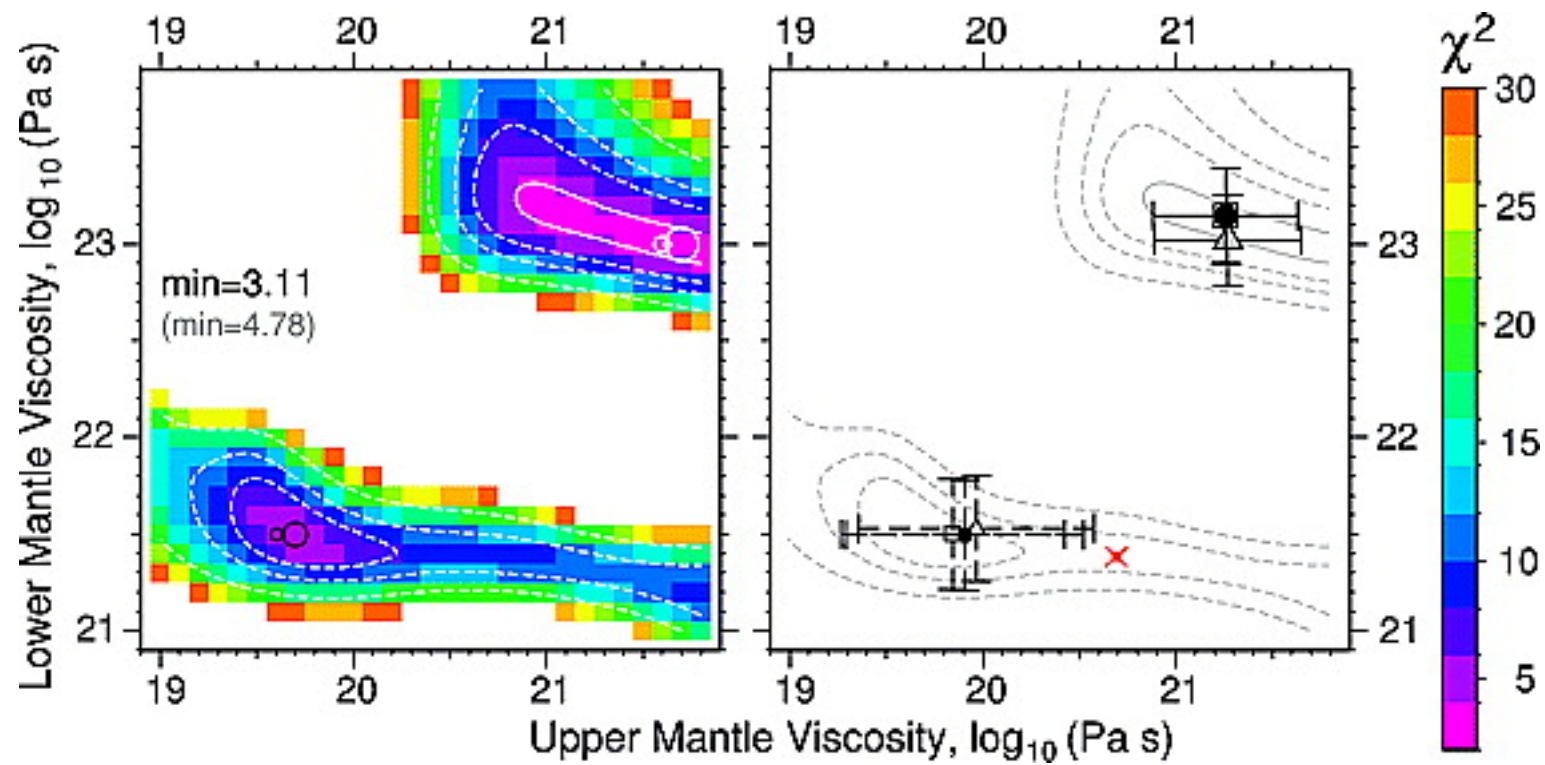
$\nu_{UP} =$
 $3.1 \times 10^{20} \text{ Pa s}$

$\nu_{LW} =$
 $1.5 \times 10^{23} \text{ Pa s}$

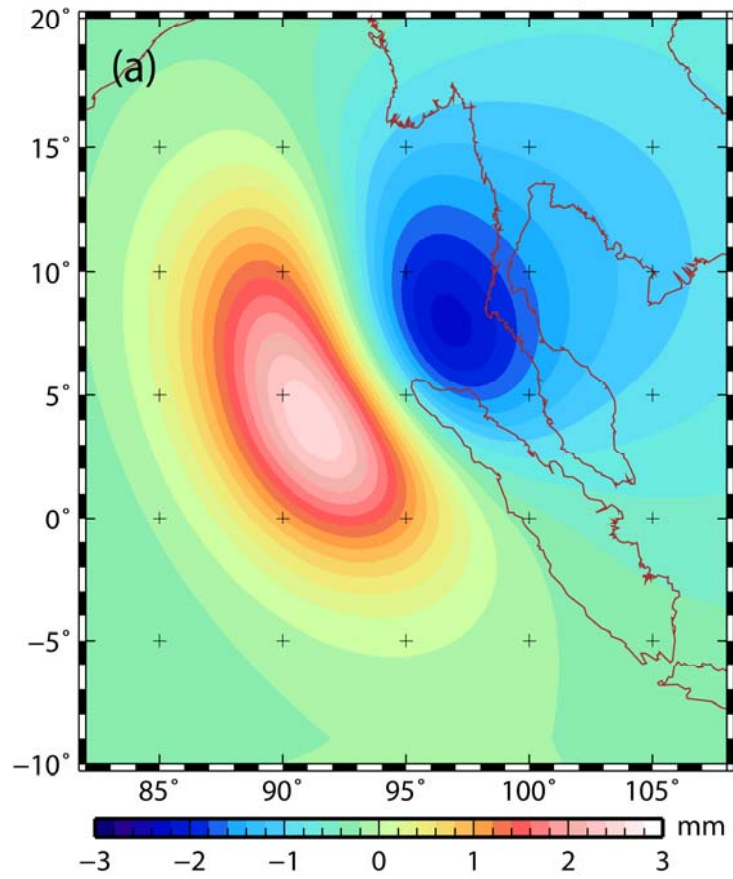


Cleaned **GRACE**

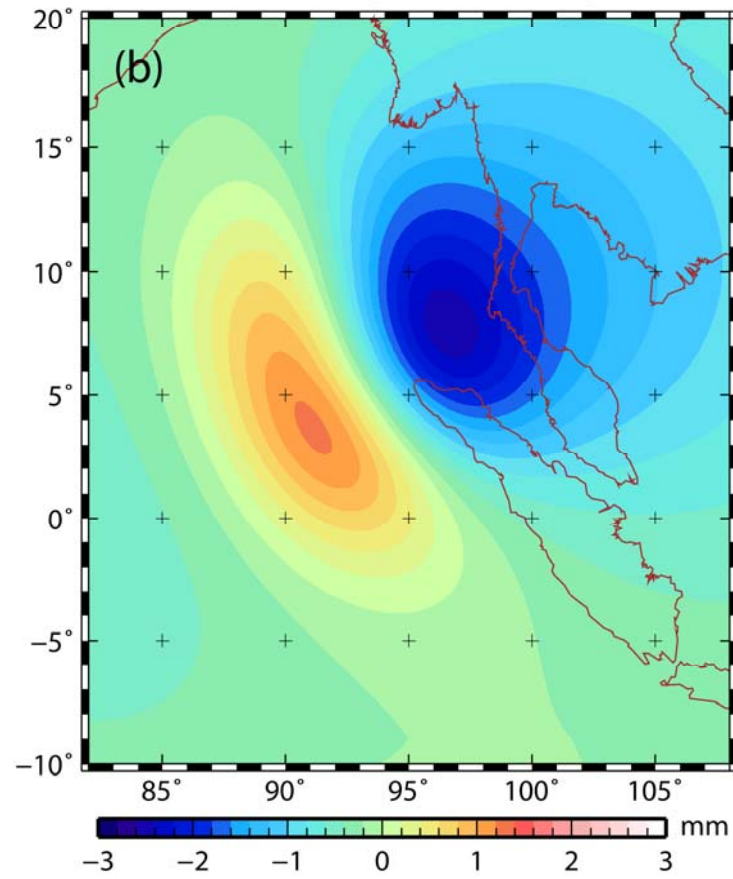


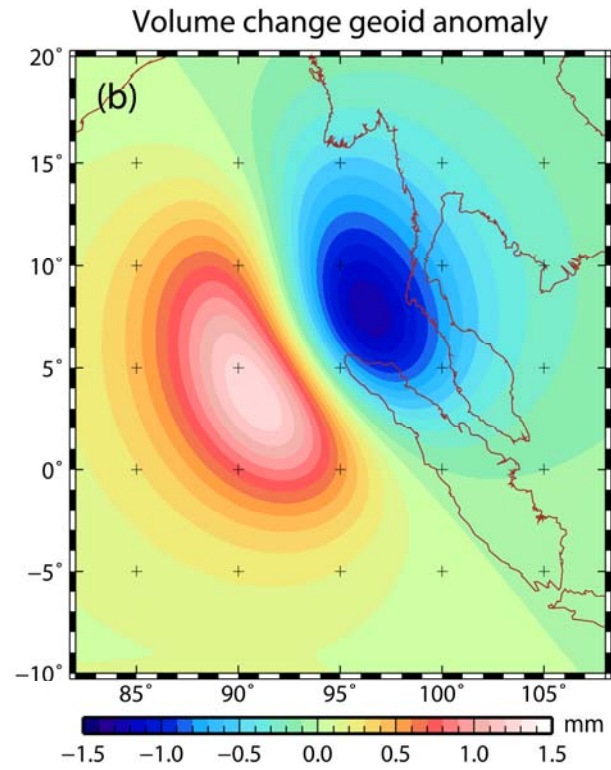
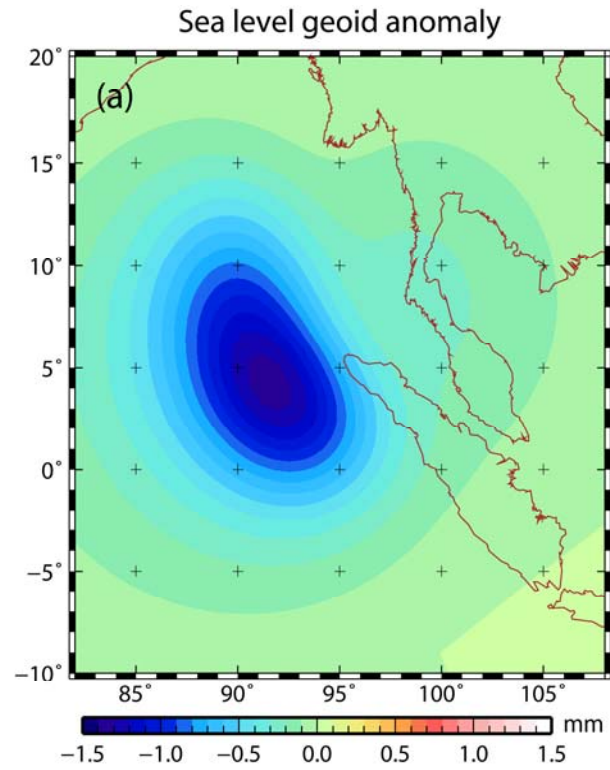


S-PREM geoid anomaly

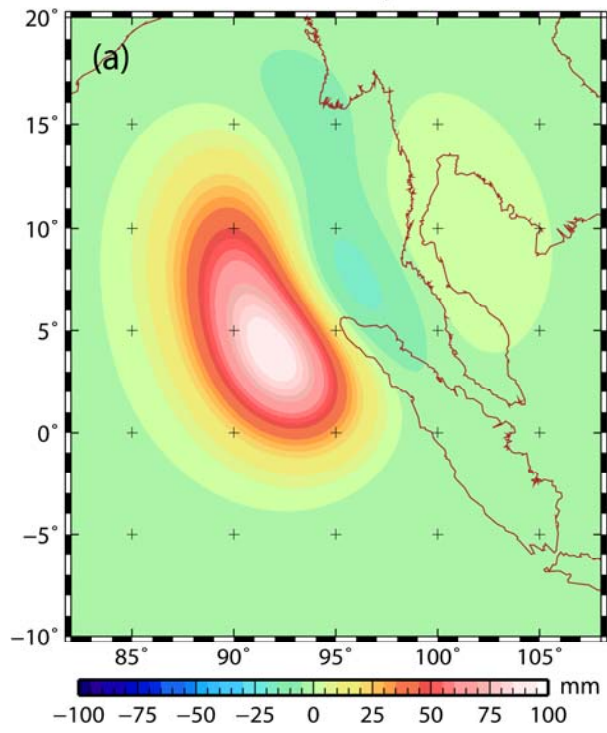


O-PREM geoid anomaly

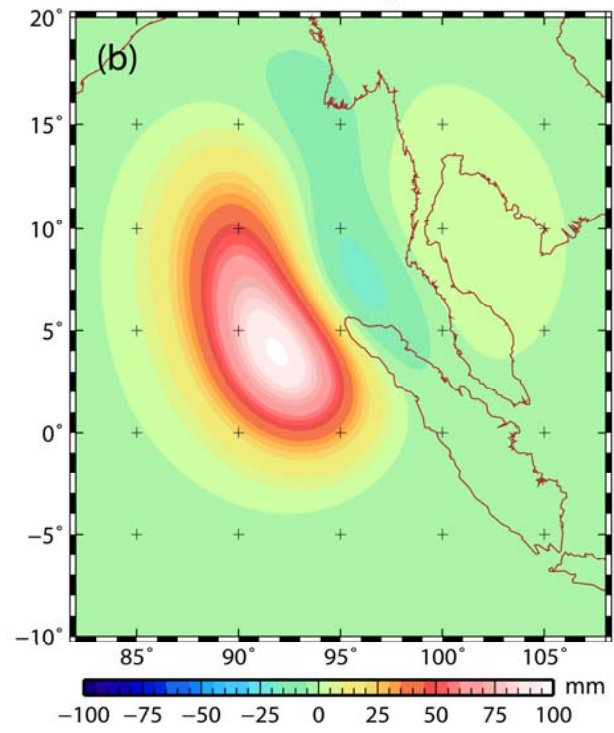


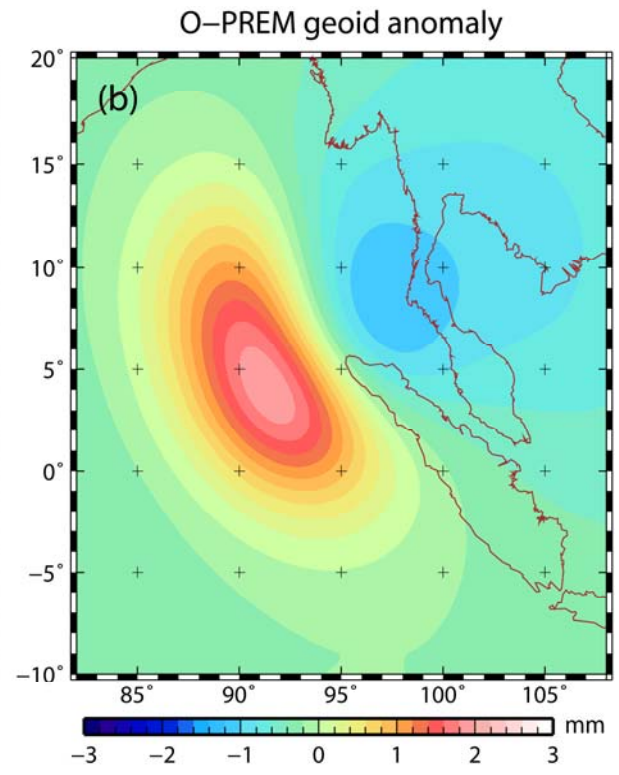
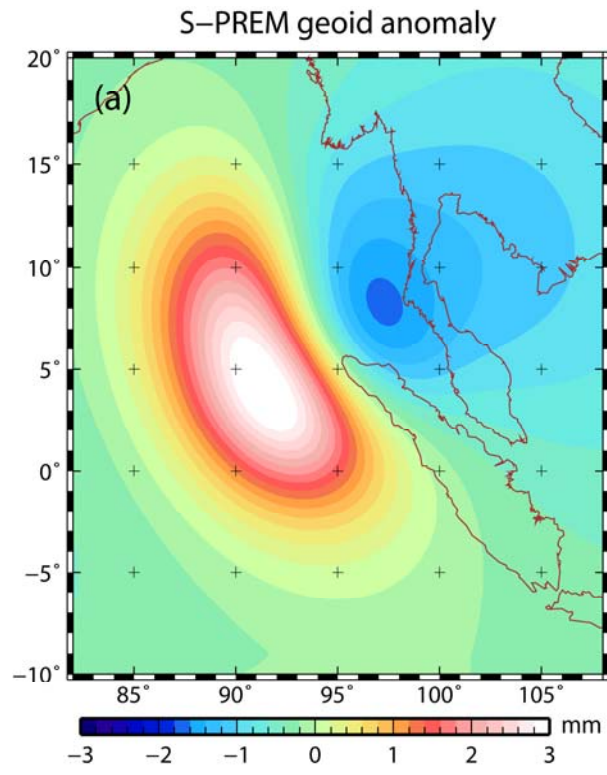


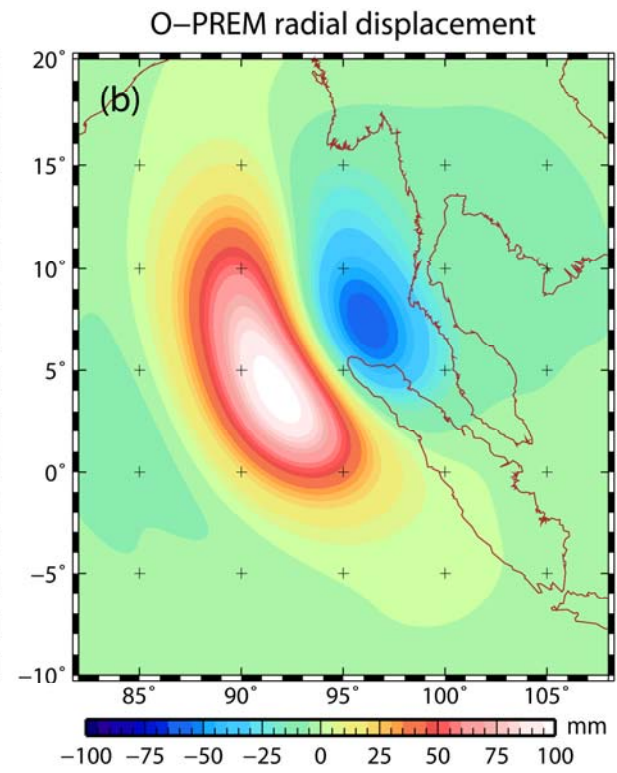
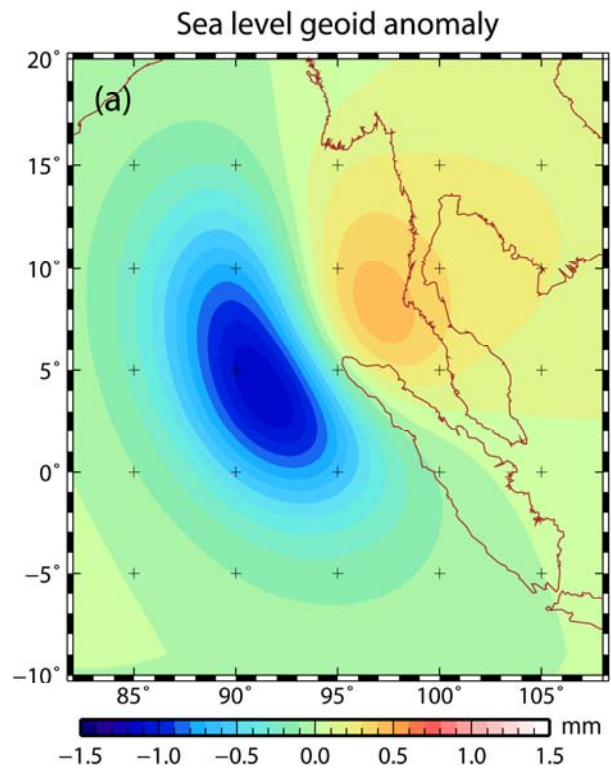
S-PREM radial displacement

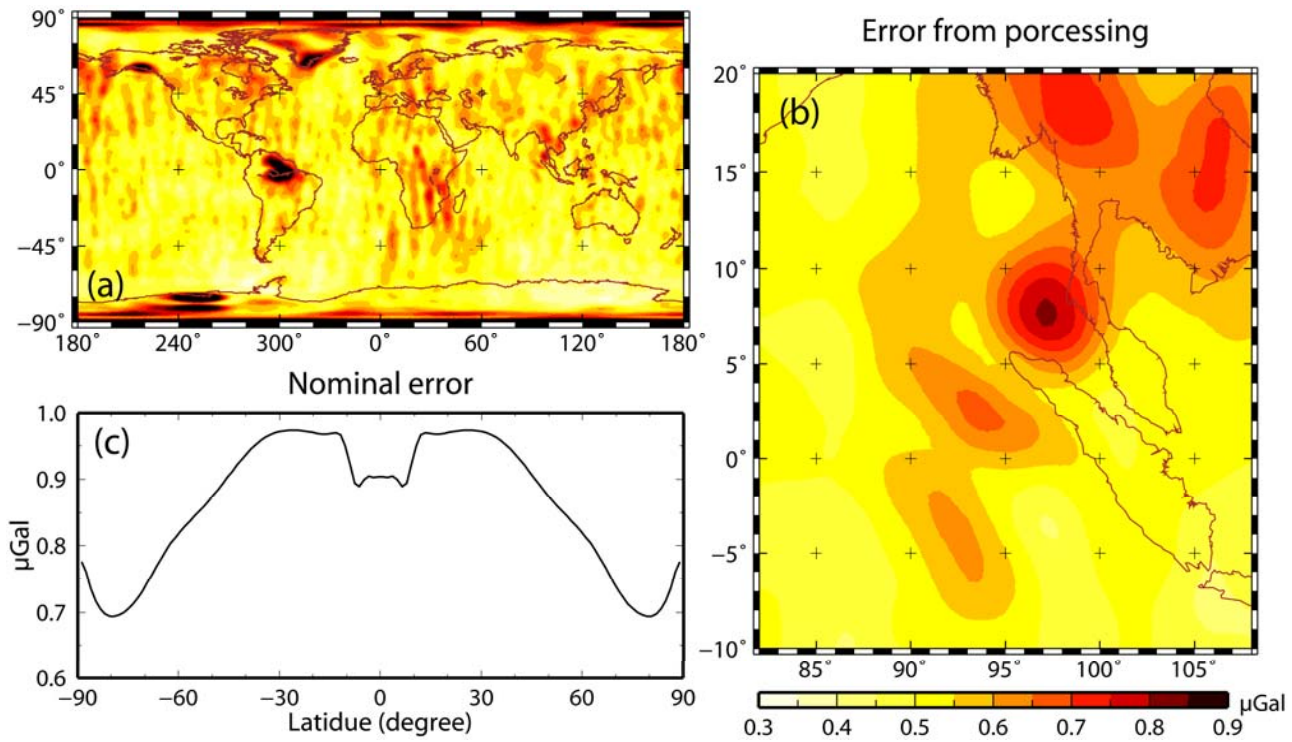


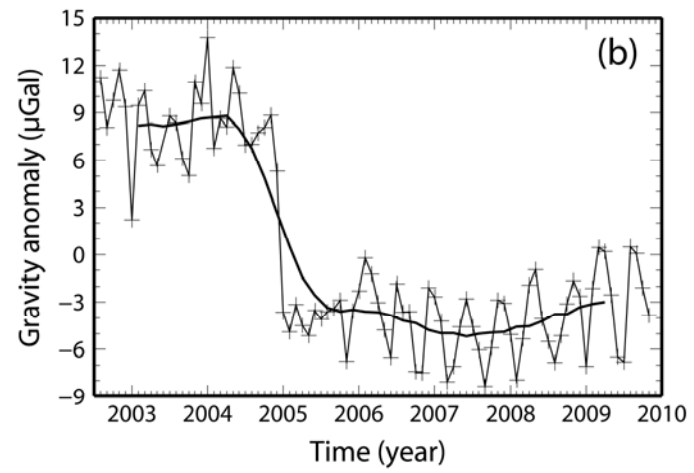
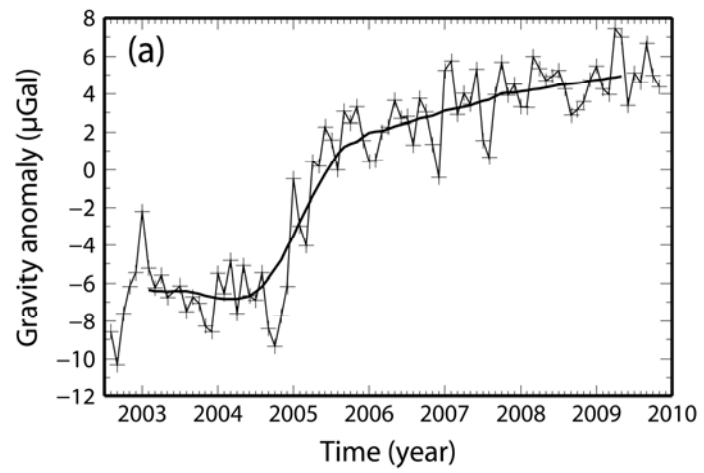
O-PREM radial displacement

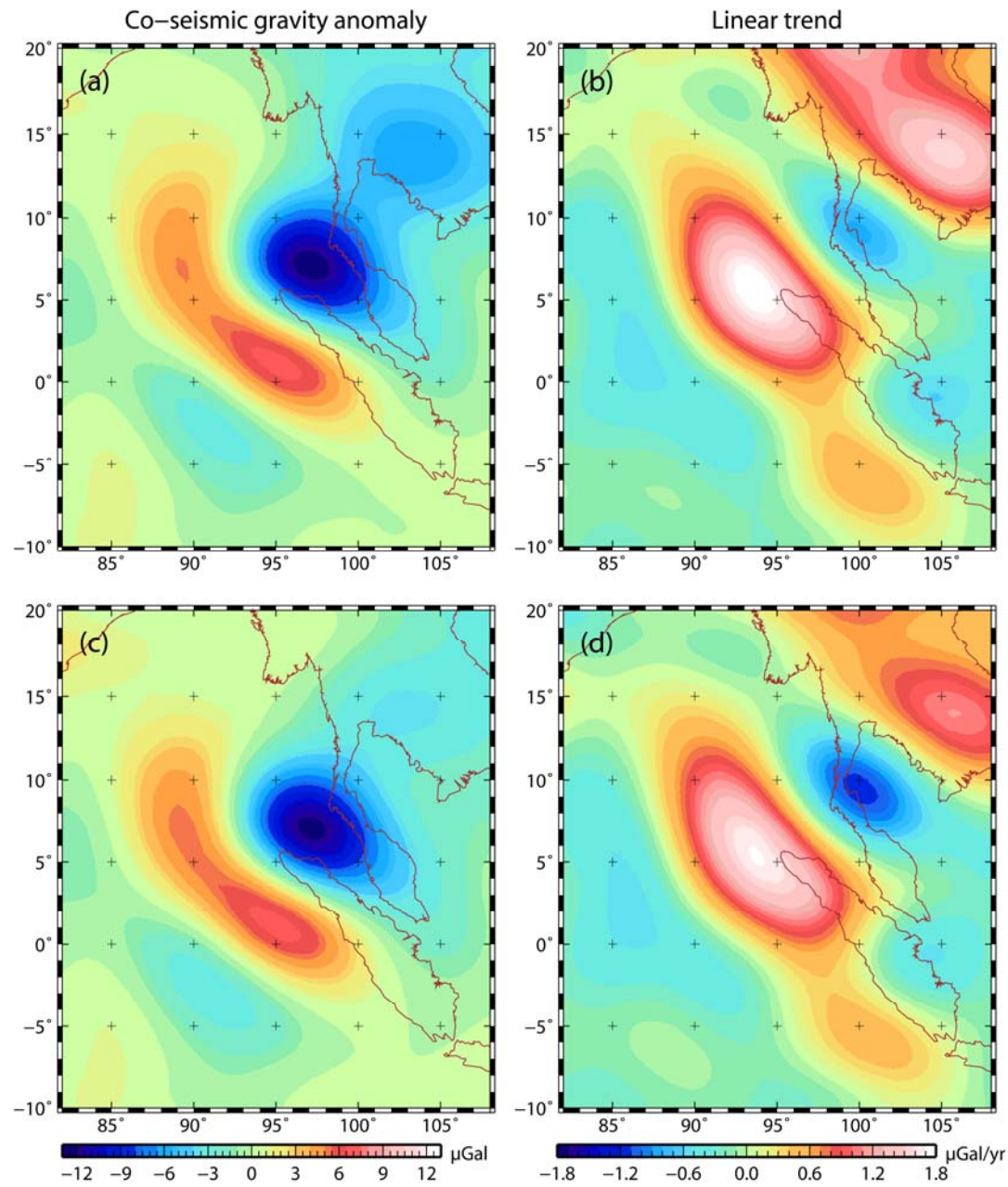


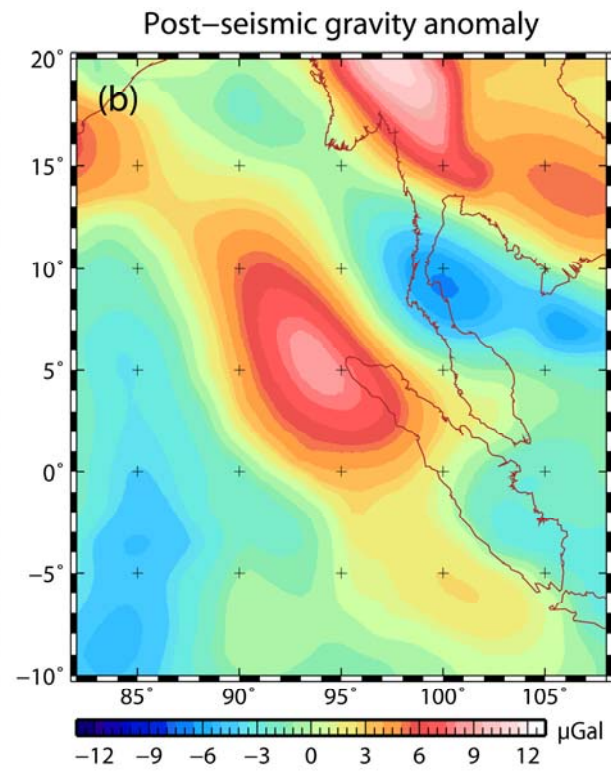
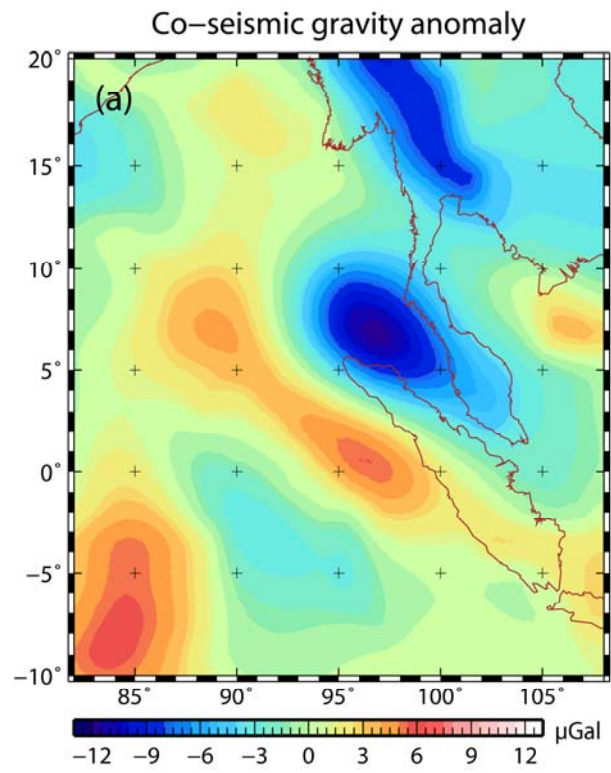


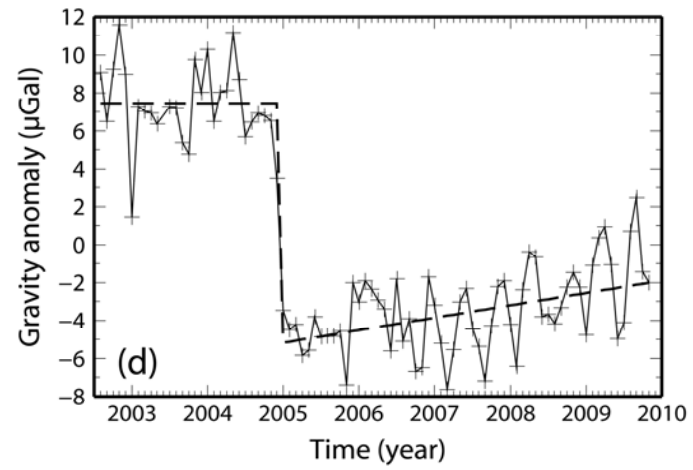
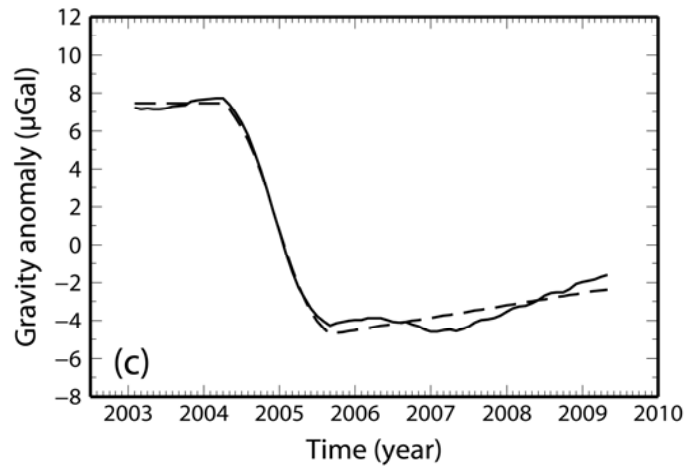
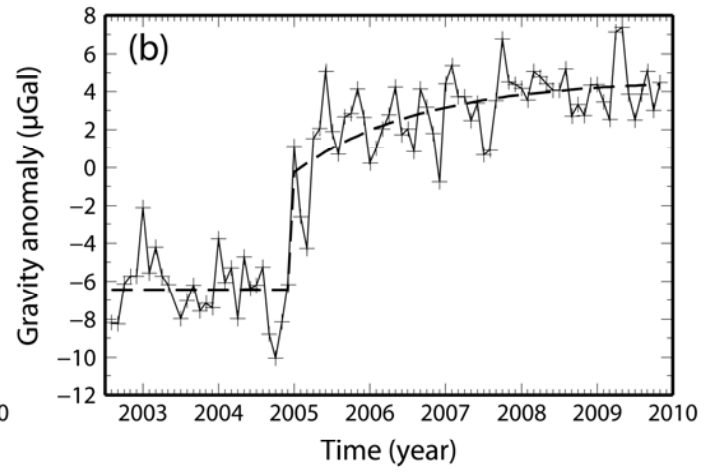
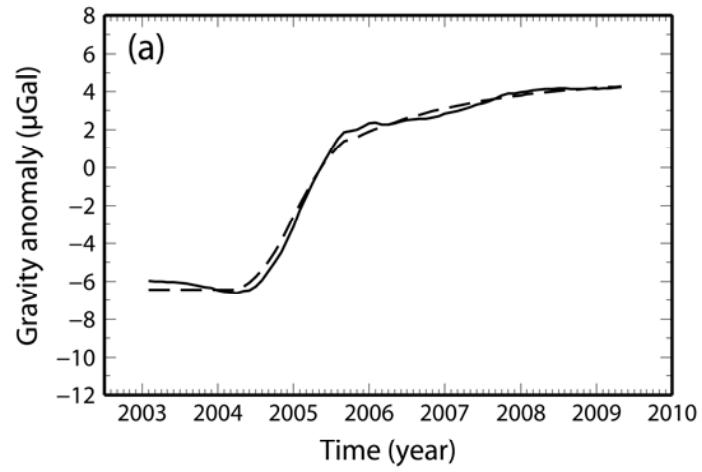


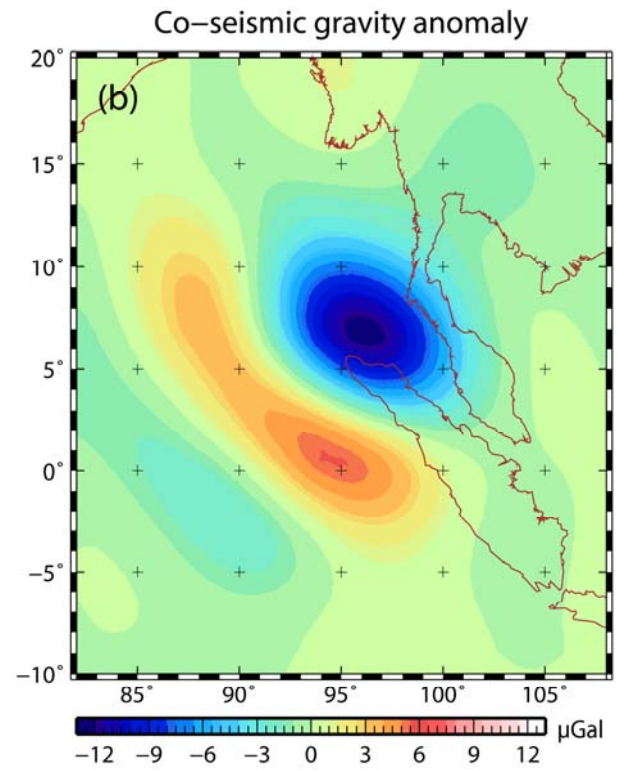
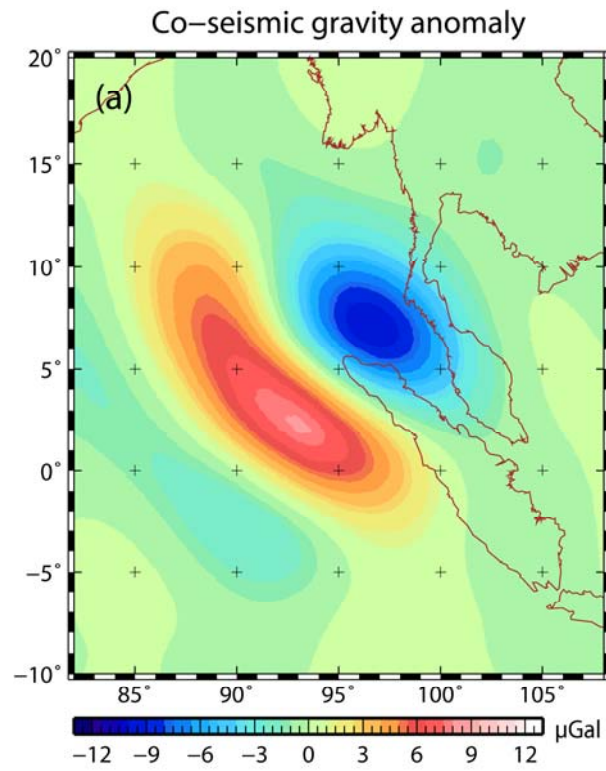















*Normal mode relaxation theory
and new generation space missions:
revealing the
Physics of Earth's interior*