



The Abdus Salam
International Centre for Theoretical Physics



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**Advanced Workshop on Anderson Localization, Nonlinearity and
Turbulence: a Cross-Fertilization**

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Inertial particle clustering and random walks in random environments

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Inertial particle clustering and random walks in random environments

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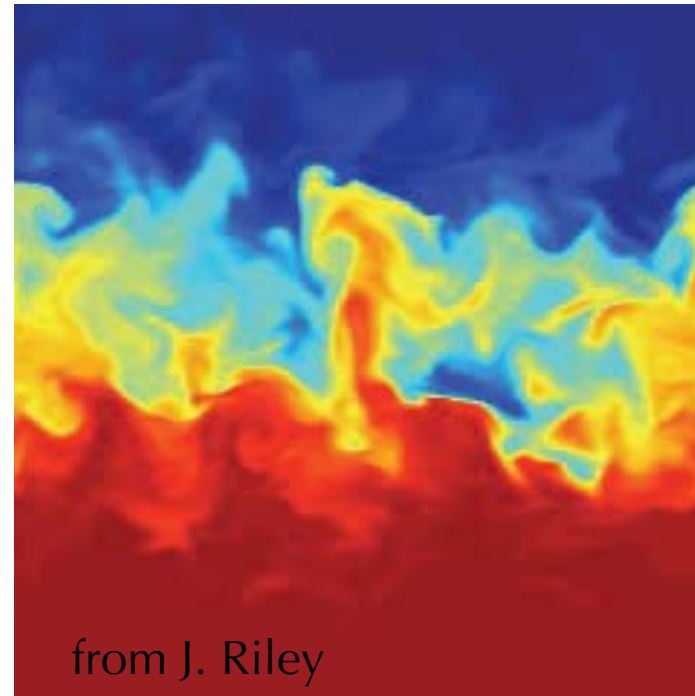
in collaboration with

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B. Merlet (Paris), and **S. Musacchio** (Nice)

Turbulent Transport/Mixing

- ▶ **Industrial/Natural problems:** passive or active transport of species by a turbulent flow
- ▶ **Effect of turbulence:** enhance mixing/dispersion (w.r.t. molecular diffusion)

Eddy diffusivity
(~mean-field effect)



- ▶ **Quantifying fluctuations?** What are the mechanisms leading to the presence of very high concentrations?

Fluctuations in turbulent transport

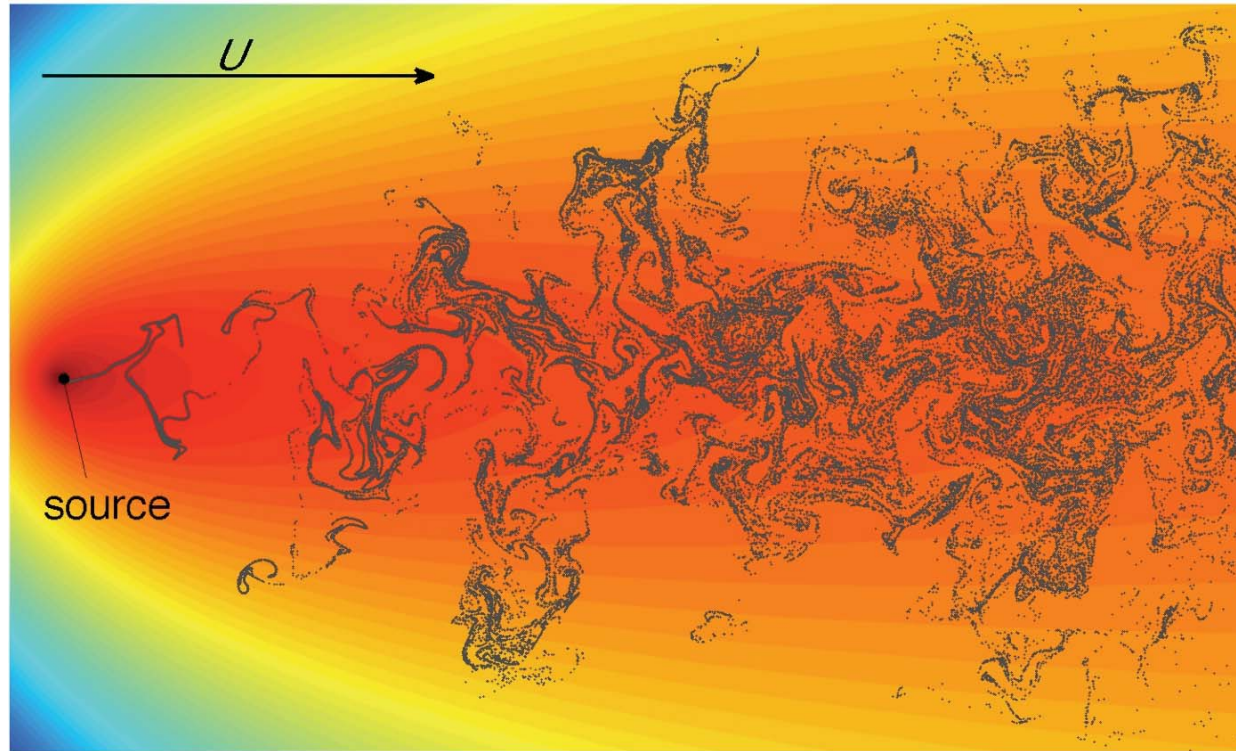
- ▶ Fluctuations are important for risk assessments



- ▶ **Models/Observations:** space and/or time averages

Mean vs. meandering plumes

- ▶ One source of fluctuations = the turbulent transport itself



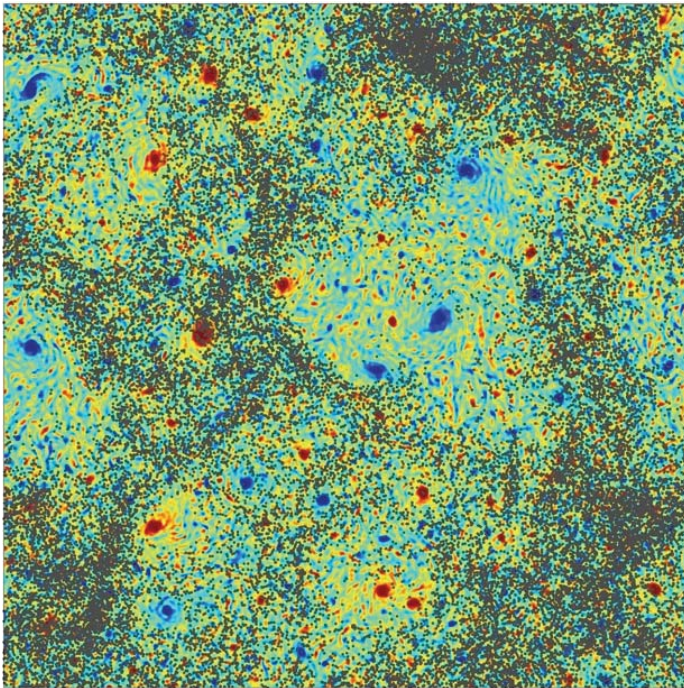
- ▶ Concentration PDFs have tails rather far from Gaussian (trapping events for the random walk of particles in the random environment of subjacent turbulent fluctuations?)

Universality??

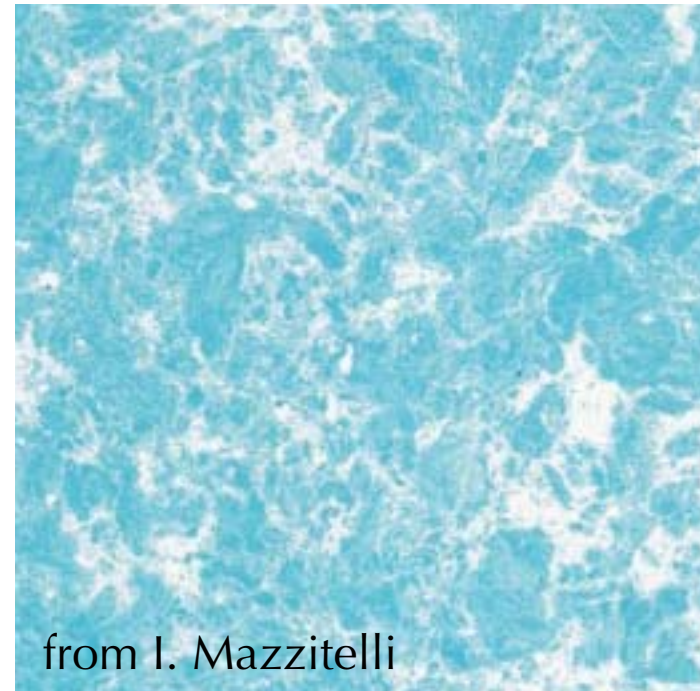
Particles finite mass

- ▶ **Most particles are not tracers but have inertia**

Heavy particles are ejected from eddies



Light particles cluster in their cores



from I. Mazzitelli

Preferential concentration

Very heavy particles

- ▶ Spherical particles much much heavier than the fluid, feeling no gravity, evolving with moderate velocities: **one of the simplest model**

$$\ddot{\mathbf{X}} = -\frac{1}{\tau} \left(\dot{\mathbf{X}} - \mathbf{u}(\mathbf{X}, t) \right)$$

↑
Prescribed velocity field
(random or solution to NS)

Stokes time:

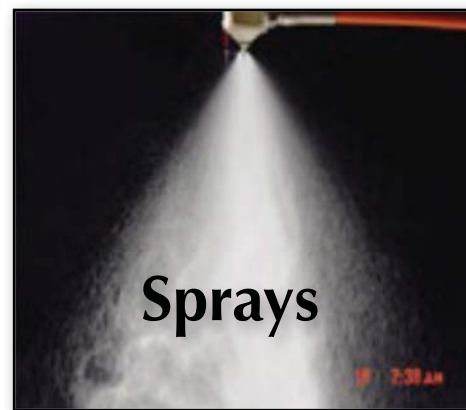
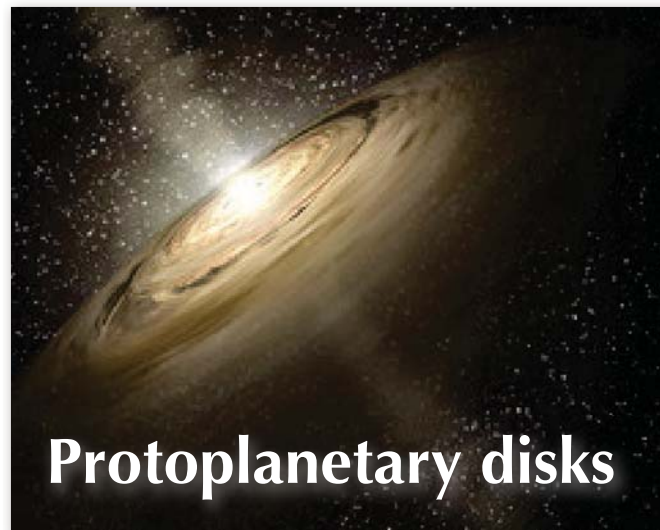
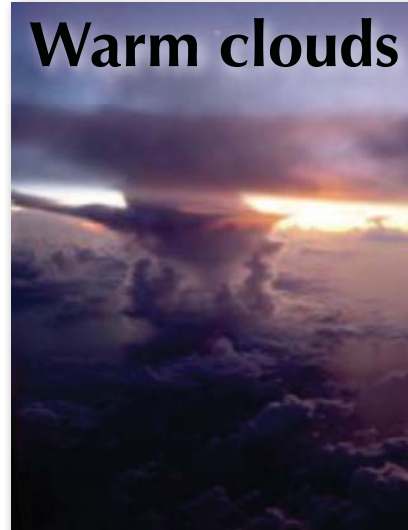
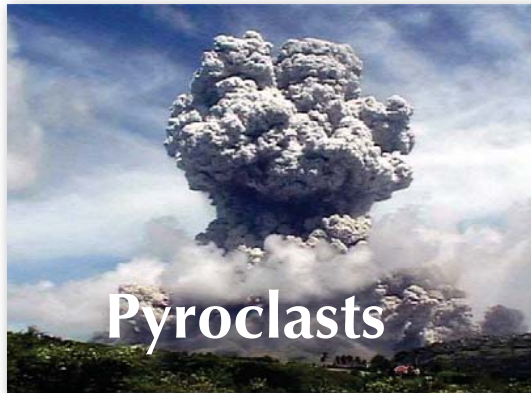
$$\tau = \frac{2 \rho_p a^2}{9 \rho_f \nu}$$

Inertia measured by the Stokes number

$$St = \tau / \tau_f$$

- ▶ **Dissipative dynamics** (even if $\mathbf{u}(\mathbf{x}, t)$ is incompressible)
Lagrangian averages correspond to an SRB measure (with support on the attractor) that depends on time and on the realization of the fluid velocity field.

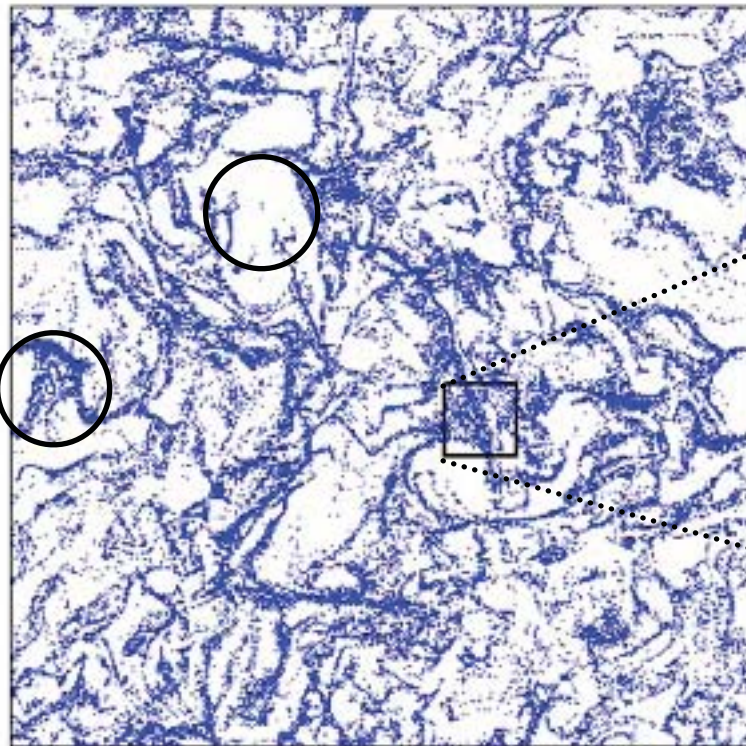
Very heavy particles



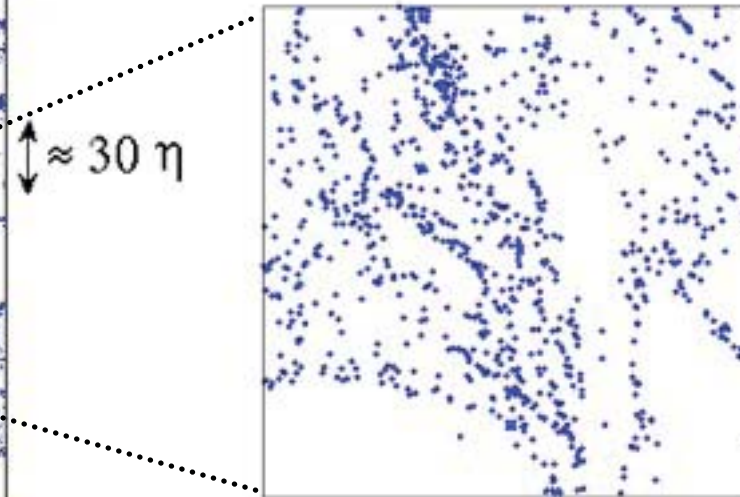
- ▶ Need for quantifying inertia effects on concentration fluctuations, which in turn affect global quantities

Preferential concentrations

- ▶ Observed for a long time in experiments
Eaton & Fessler (1994); Douady, Couder, & Brachet (1991)
- ▶ **Quantifying them is important for**
 - * the rates at which particles interact
 - * the fluctuations in the concentration of a pollutant
 - * the possible feedback of the particles on the fluid



Inertial-range clusters and voids

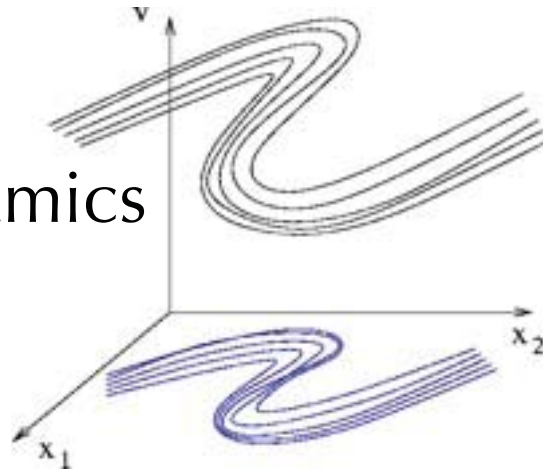


Multifractal distribution
at dissipative scales

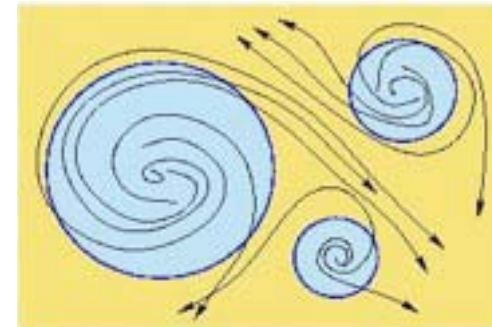
Phenomenology of clustering

▶ Different mechanisms:

Dissipative dynamics
⇒ **attractor**



Ejection from **eddies**
by centrifugal forces

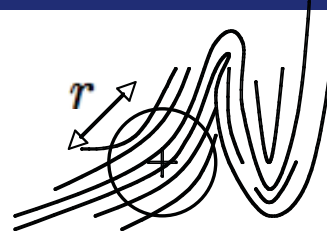


- ▶ **Theory:** requires elaborating models to disentangle these two effects. For instance:
 - ▶ flows with no structures (uncorrelated in time) to isolate the effects of a dissipative dynamics
 - ▶ coarse-grained closures to understand ejection from eddies
- ▶ Numerics show that these effects act at different scales

Small-scale clustering

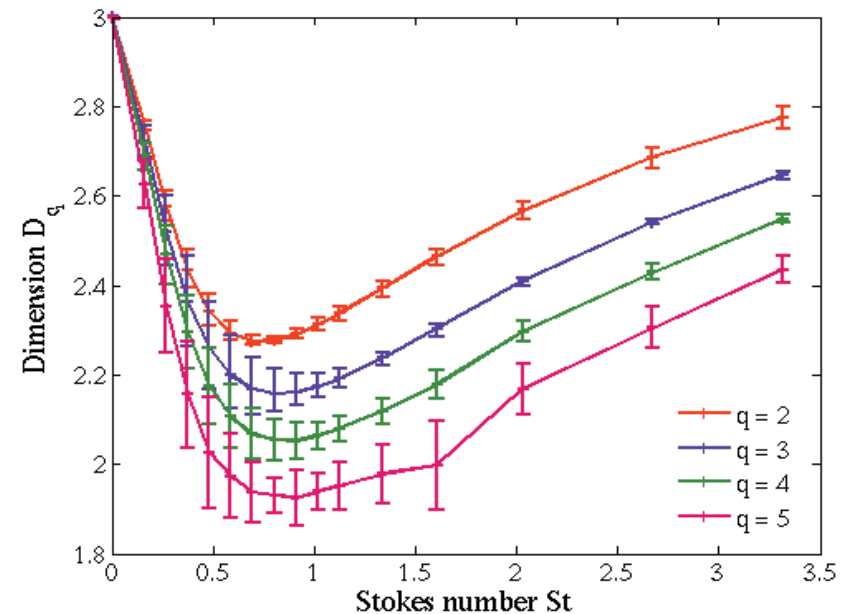
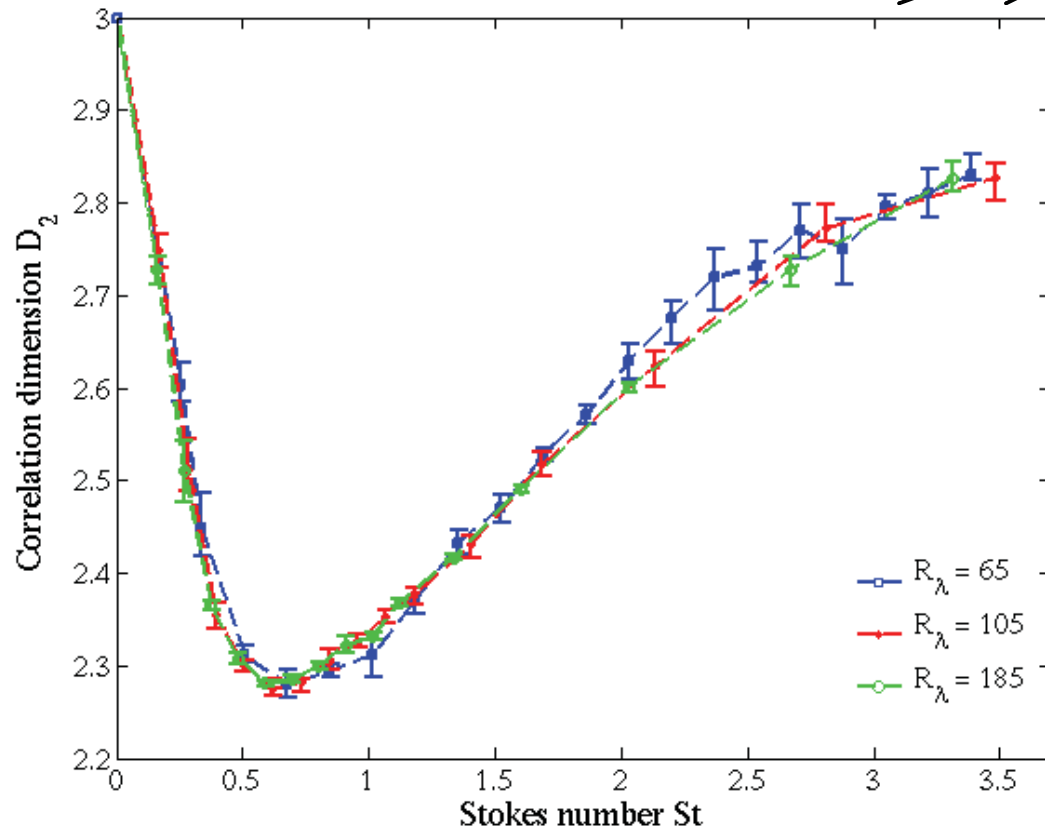
► Fractal dimensions $r \ll \eta$

Coarse-grained density $\bar{\rho}_r$



$$\langle \bar{\rho}_r^p \rangle \sim r^{p(\mathcal{D}_{p+1} - d)}$$

Hentschel–Procaccia (1983)
spectrum for dimensions



► Spectrum \mathcal{D}_p is a function of St but almost not depends on Re

PDF local dimension $\delta_r = \frac{\ln \bar{\rho}_r}{\ln r}$

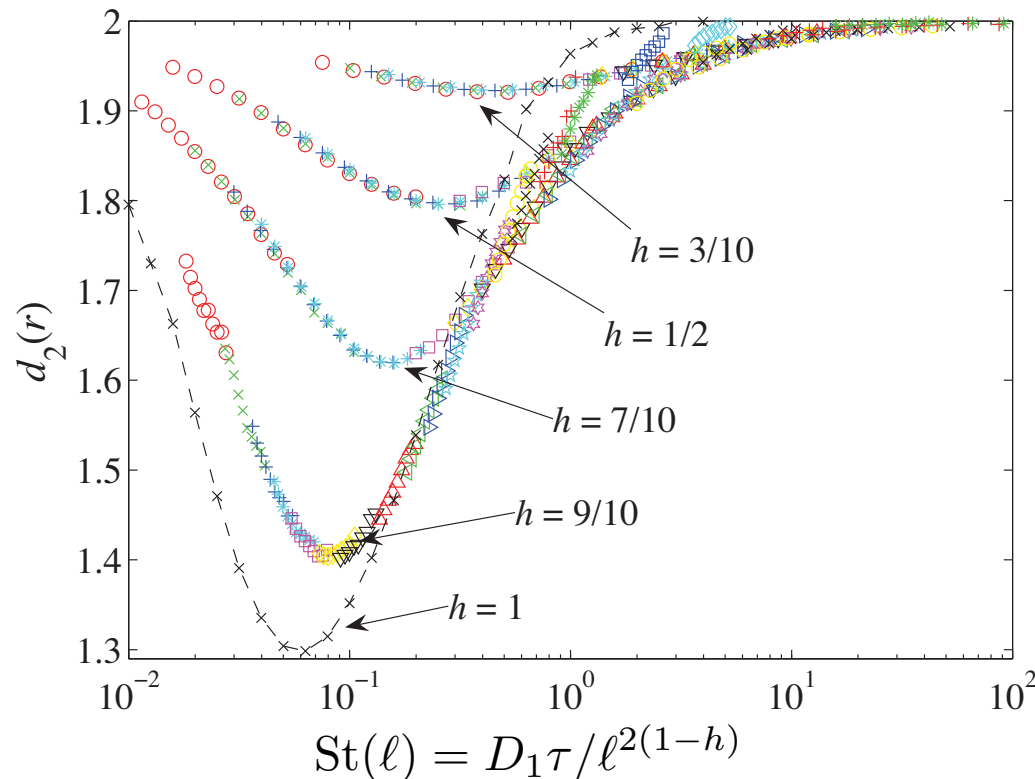
$$p_r(\delta) \propto r^{\mathcal{S}(\delta, St)}$$

Inertial-range clustering ?

- Case of *non-differentiable* Kraichnan: particle dynamics at scale ℓ depends on a **local (scale-dependent) Stokes number**

$$\text{St}(\ell) = \tau / \tau_\ell = \varepsilon^{1/3} \tau / \ell^{2/3}$$

Falkovich, Fouxon, Stepanov (2003)
JB, Cencini, Hillerbrand (2007)



Both the scale-invariance of the fluid flow and that of the particle distribution are broken

$$\ell \rightarrow \infty \quad \text{St}(\ell) \rightarrow 0$$

inertia becomes negligible

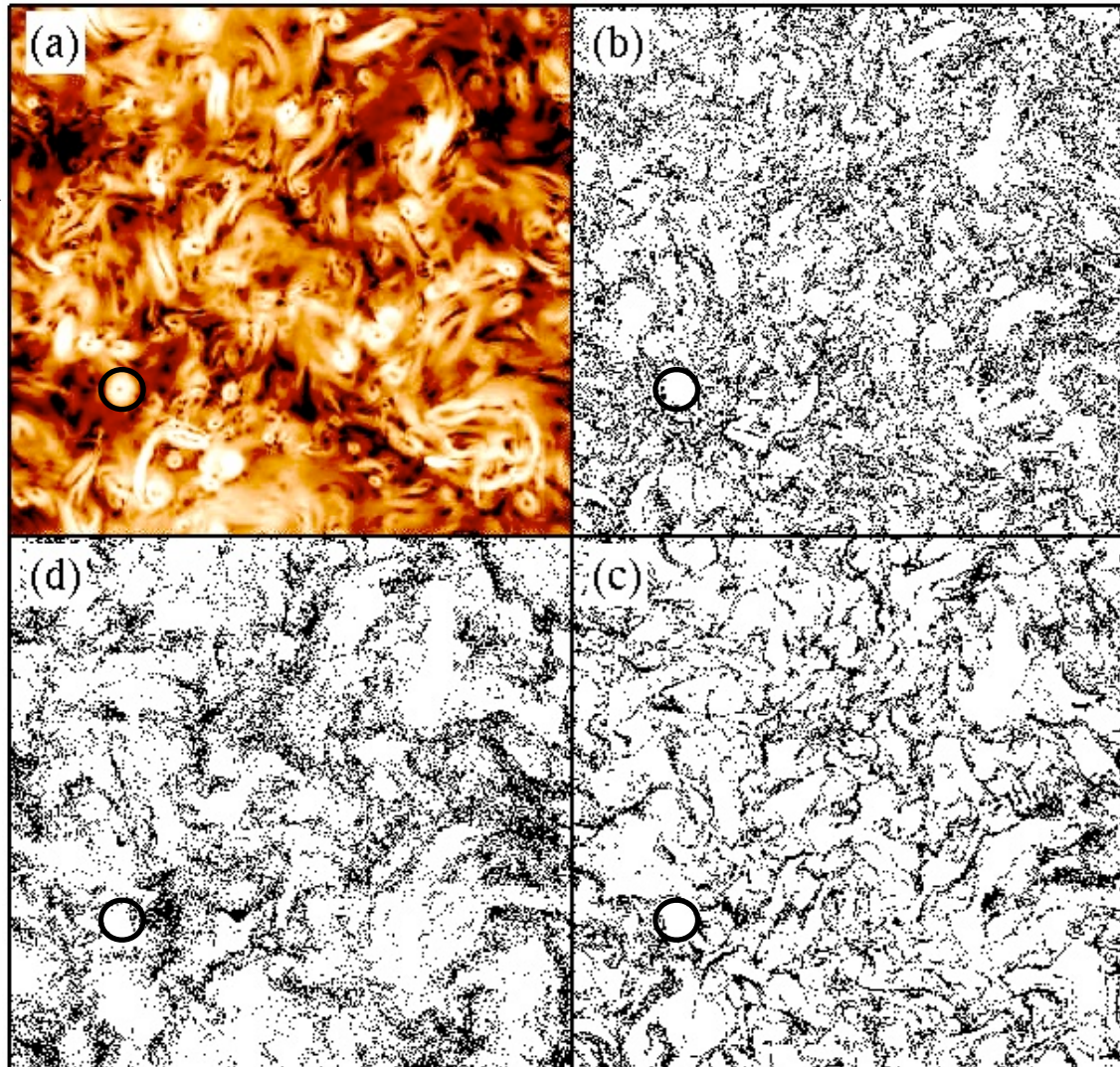
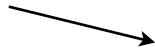
$$\ell \rightarrow 0 \quad \text{St}(\ell) \rightarrow \infty$$

particles move almost ballistically

Particles in turbulent flow

Real flow have structure and particle distribution correlates with the acceleration field

Modulus of acceleration



$St = 0.16$

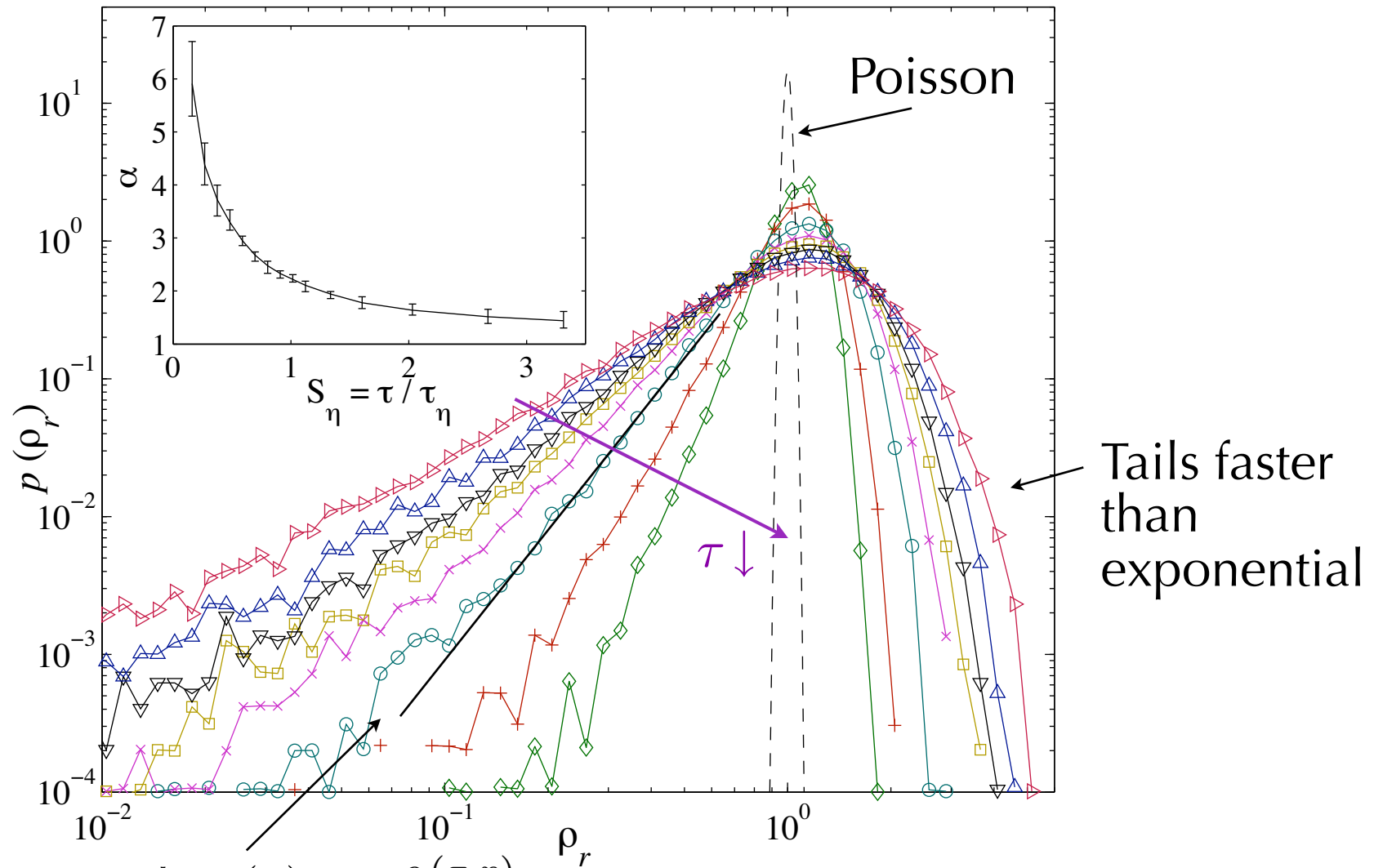
JB, Biferale
et al. (2007)

$R_\lambda = 185$

$St = 3.3$

$St = 0.8$

Coarse-grained density



Algebraic tails $p(\rho) \propto \rho^{\alpha(\tau,r)}$
(signature of voids)

Time scales of clustering

- ▶ The local Stokes number $St(\ell) = \varepsilon^{1/3} \tau / \ell^{2/3}$ is not relevant
- ▶ **Non dimensional contraction rate**

When inertia is very weak: $\tau \ll \tau_f$

$$\ddot{\mathbf{X}} = -\frac{1}{\tau} \left(\dot{\mathbf{X}} - \mathbf{u}(\mathbf{X}, t) \right) \Rightarrow \dot{\mathbf{X}} = \mathbf{u}(\mathbf{X}, t) - \tau \ddot{\mathbf{X}}$$

$$\dot{\mathbf{X}} \approx \mathbf{v}(\mathbf{X}, t) = \mathbf{u}(\mathbf{X}, t) - \tau [\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u}] \quad \text{Maxey's approximation}$$

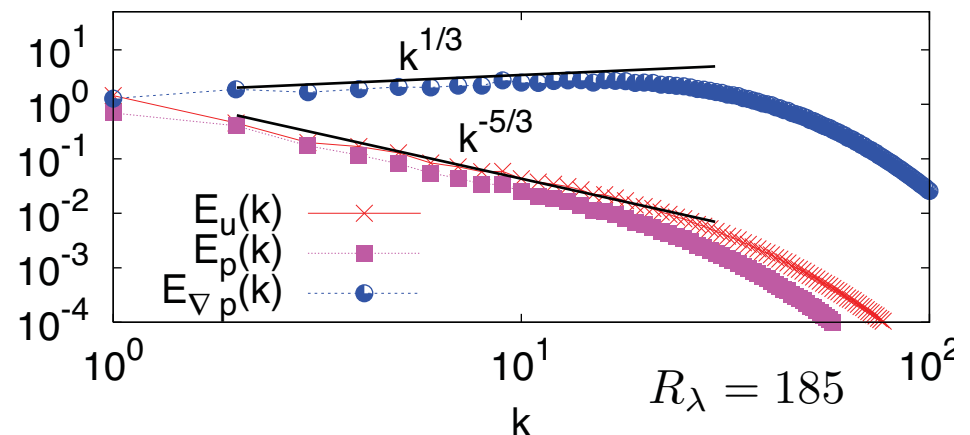
Rate at which a particle blob with size r is contracted

$$\Gamma_{r,\tau} = \frac{1}{r^3} \int_{|\mathbf{x}| < r} \nabla \cdot \mathbf{v} \, d^3x \simeq \frac{\tau}{r^2} \delta_r p$$

Pressure scaling

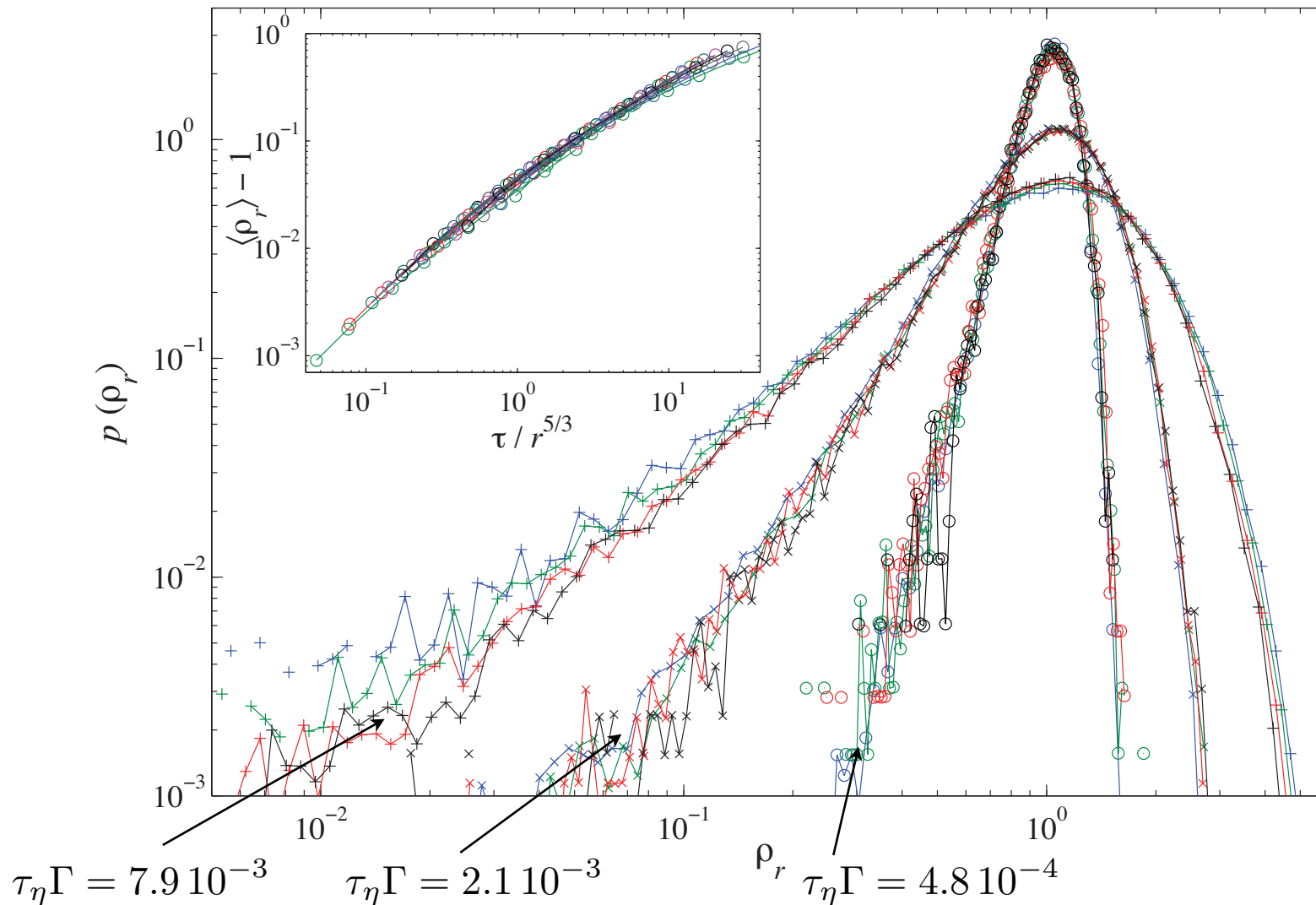
$$\text{Sweeping} \quad \delta_r p \sim U(\varepsilon r)^{1/3}$$

$$\Gamma_{r,\tau} \sim \tau / r^{5/3}$$



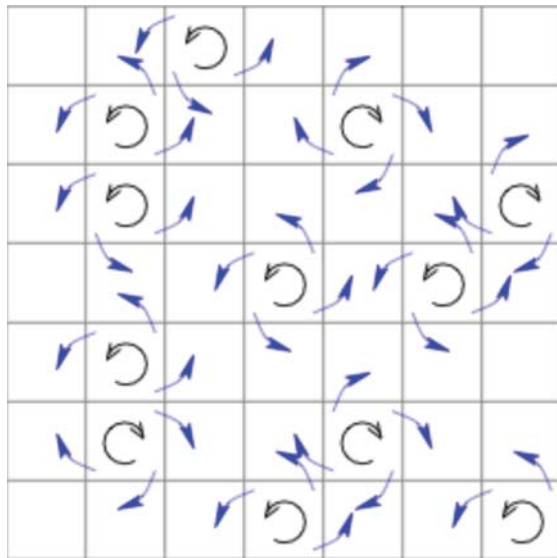
Scalable deviations from uniformity

- ▶ Mass distribution depends only on $\tau_\eta \Gamma_{r,\tau} \sim \text{Re}^{1/4} \text{St}(r/\eta)^{5/3} \sim \text{Re}^{-1} \text{St}(r/L)^{5/3}$

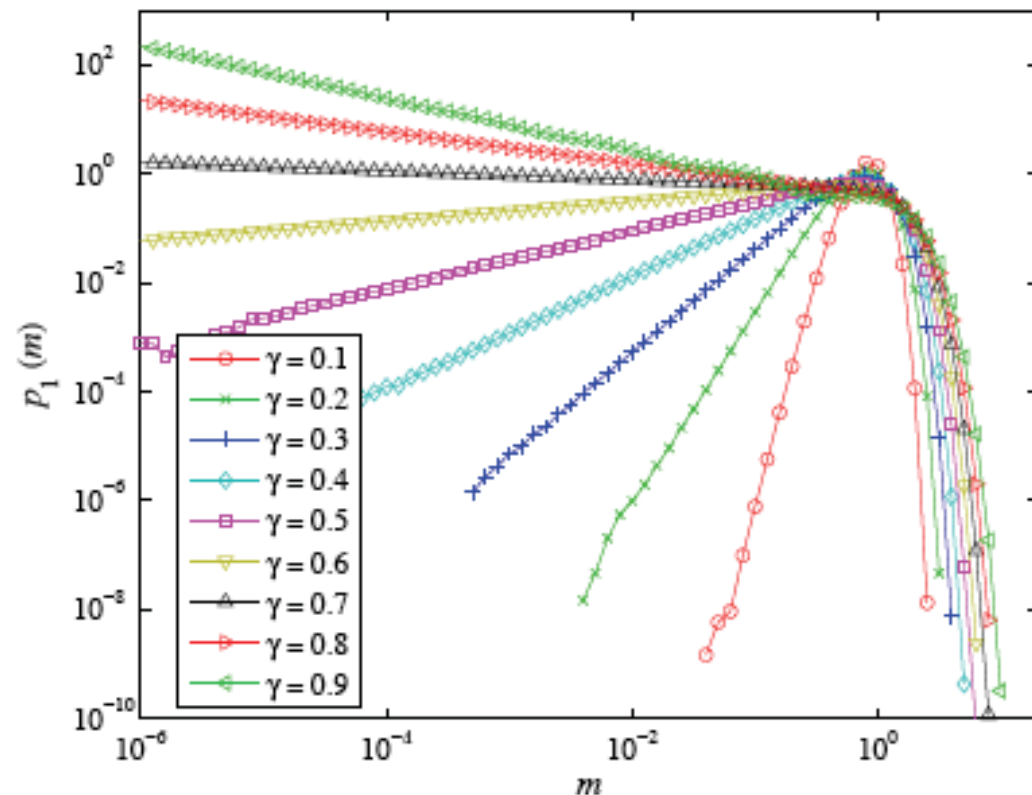


Mass transport model

- ▶ Find models belonging to the same universality class
- ▶ Discreteness in time and space
- ▶ At each time step some (randomly chosen with probability p) cells eject a fraction of their mass to their neighbors
- ▶ Parameter = γ ejection rate (related to particle inertia)



JB, R. Chérite (2007)



Tails

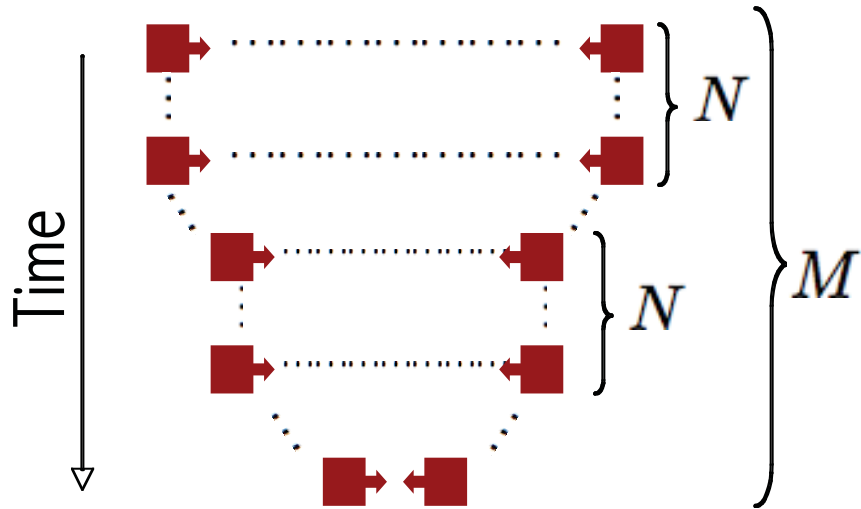
▶ Right tail = algebraic $p(m) \propto m^{\alpha(\gamma)}$

$$N \text{ times } \left\{ \begin{array}{l} \left[\rightarrow \left[\left[\rightarrow \right] \right] m_0 \approx 1 \\ \vdots \\ \left[\left[\left[\left[\rightarrow \right] \right] \right] m_N \approx (1 - \gamma)^N \end{array} \right.$$

$$\text{Prob} = p^N (1 - p)^{2N}$$

$$\Rightarrow \alpha(\gamma) = \frac{\ln p(1 - p)}{\ln(1 - \gamma)}$$

▶ Left tail = super-exponential



$$\text{Prob} = [p^2(1 - p)]^{N M}$$

$$m_{N M} = \frac{1 - [1 - (1 - \gamma)]^N]^M}{(1 - \gamma)^N}$$

$$\Rightarrow p(m) \propto \exp(-C m \ln m)$$

More general model

► **Continuous limit** $dm_n = \gamma_{n+1}m_{n+1} + \gamma_{n-1}m_{n-1} - 2\gamma_n m_n$

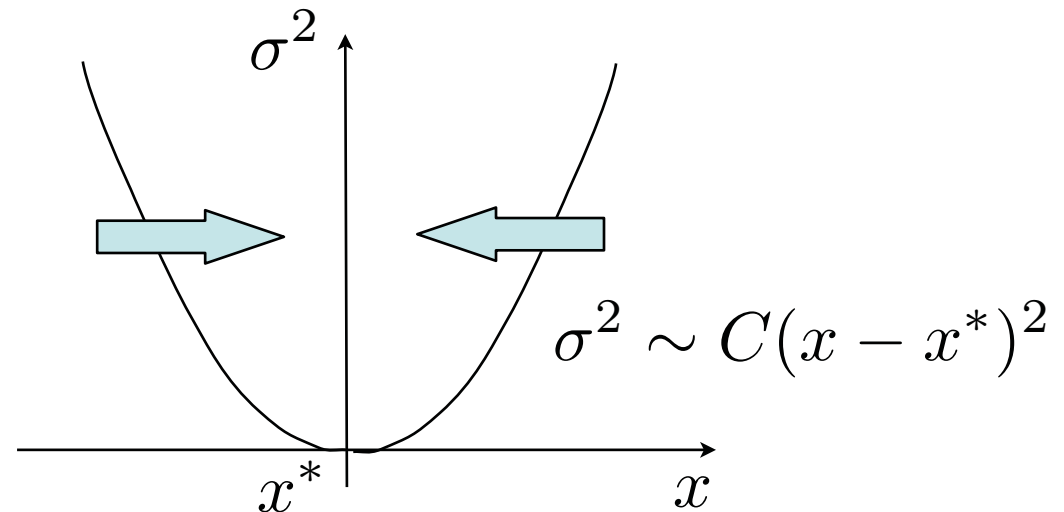
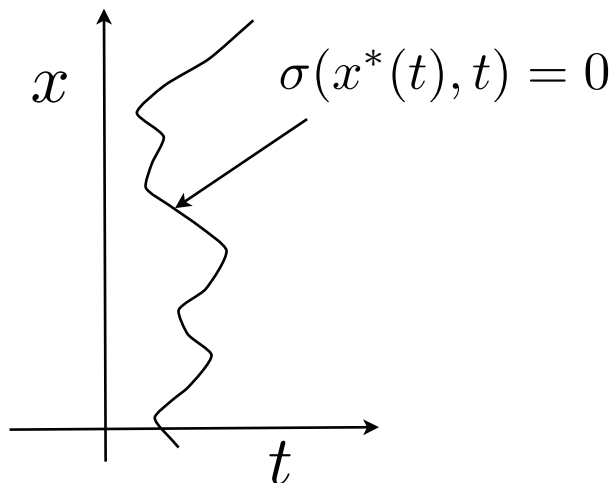
$$\left. \begin{array}{l} (dx^2/dt) \gamma_n \rightarrow \sigma^2(x, t)/2 \\ m_n/dx \rightarrow \rho(x, t) \end{array} \right\} \Rightarrow \partial_t \rho = \frac{1}{2} \partial_x^2 (\sigma^2 \rho)$$

σ random \Rightarrow diffusion in a random environment

Green function \leftrightarrow transition probability for $dX = \sigma(X, t) dW_t$

► **High densities:** near the zeros of the diffusion coefficient σ

Example: $\sigma(x, t)$ smooth, generic, Gaussian



Large density tail

$$\partial_t \rho = \partial_x^2 (x^2 \rho)$$

$$\rho(x, t) = \frac{1}{2} e^{2t} \operatorname{erfc} \left(\frac{\ln |x| + 3t}{2\sqrt{t}} \right)$$

Cumulative probability $P^>(\rho)$
 = fraction of space-time where
 the density is larger than ρ

$$\lambda = 2t / \ln \rho$$

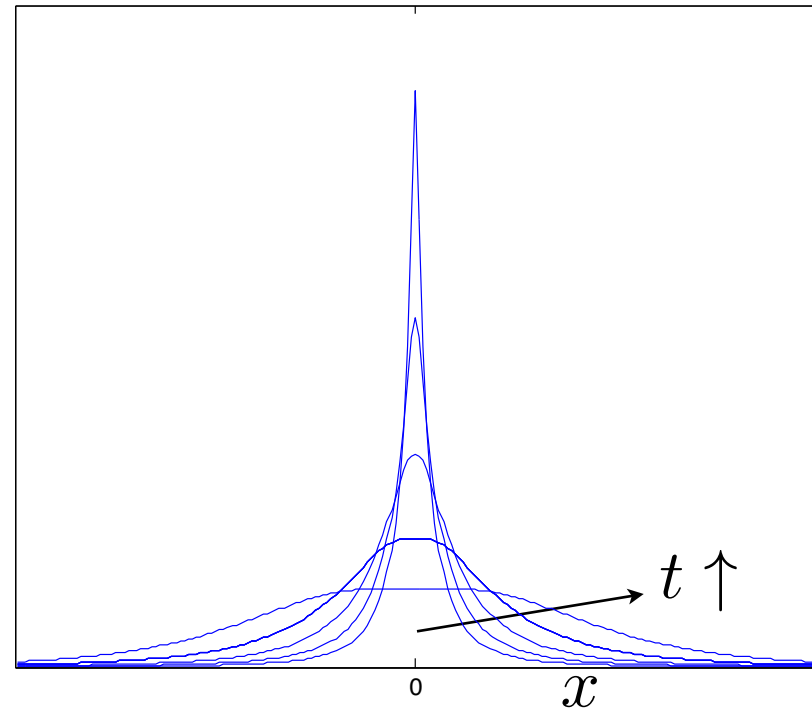
$$P^>(\rho) \simeq \ln \rho \int_1^\infty e^{-\mu(\lambda) \ln \rho} d\lambda$$

Saddle-point: $P^>(\rho) \propto \rho^{-1}$

Universal intermediate asymptotics where

$$p(\rho) = -\frac{dP^>}{d \ln \rho} \propto \rho^{-2}$$

+ non-universal cut-off (distribution of zeros lifetime)



$$\mu(\lambda) = \frac{3}{2} \lambda - \sqrt{2\lambda(\lambda - 1)}$$

Conclusions

▶ **Summary**

Clustering of heavy particles:

of two kinds, depending on the observation scale:

- * multifractal in the dissipative range
- * dependent on a rescaled contraction rate in the inertial range

Connection to random walks in random environments:

ejection models reproduce most features (in particular tails)

▶ **Open questions**

- * **Universality** of mass probability distribution?
- * **Spatial/temporal correlations** (e.g. scale invariance) of the ejection/diffusion rate?