



**The Abdus Salam
International Centre for Theoretical Physics**



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**Beyond the Standard Model: Results with the 7 TeV LHC Collision
Data**

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Drell-Yan Production of Metastable Sleptons at the LHC

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Metastable Staus at the LHC

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DER FORSCHUNG | DER LEHRE | DER BILDUNG

Based on Jan Heisig, JK, [arXiv:1106.0764](https://arxiv.org/abs/1106.0764) [hep-ph], and work in progress

Outline

- 1 Introduction
- 2 Drell-Yan Production
- 3 Cascade Production

1 Introduction

2 Drell-Yan Production

3 Cascade Production

LSP and Dark Matter Candidates

Assumption: LSP forms all dark matter

- ⇒ Electrically neutral
- ⇒ Uncolored
- ⇒ Lifetime $>$ age of universe
 \leadsto assume R parity conservation \Rightarrow stable
- ⇒ Massive
- ⇒ Non-relativistic during structure formation (cold dark matter)

Candidates in MSSM with supergravity

- Lightest neutralino
- Sneutrino (ruled out by direct searches)
- Gravitino (superpartner of graviton)

Expectations for Gravitino Mass

Depend on scenario for **mediation of SUSY breaking**

	Gravitino mass	Gravitino LSP?
Gravity mediation	$m_{3/2} \gtrsim 100 \text{ GeV}$	possible
Gaugino mediation	$m_{3/2} \gtrsim 10 \text{ GeV}$	
Gauge mediation	$m_{3/2} \ll 1 \text{ GeV}$	yes
Anomaly mediation	$m_{3/2} \gtrsim 10 \text{ TeV}$	no

Gravitino Problem

- Thermal production in early universe

$$\Omega_{3/2}^{\text{tp}} h^2 \propto \frac{T_R}{m_{3/2}}$$

Gravitino Problem

- Thermal production in early universe

$$\Omega_{3/2}^{\text{tp}} h^2 \simeq 0.11 \left(\frac{T_R}{2 \cdot 10^9 \text{ GeV}} \right) \left(\frac{67 \text{ GeV}}{m_{3/2}} \right) \left(\frac{M_{\tilde{g}}}{10^3 \text{ GeV}} \right)^2$$

- Observed dark matter abundance: $\Omega_{\text{DM}} h^2 \simeq 0.11$

- **Inflation, leptogenesis**

↪ large reheating temperature $T_R \sim 10^9 \text{ GeV}$ desirable

↪ Consider $m_{3/2} \gtrsim 1 \text{ GeV}$

- Gravitino interacts via **gravity** ↪ extremely weakly

↪ **lifetime** $\sim 10^{-2} \text{ s} \dots \text{ years}$ for **unstable** gravitino

- Energetic decay products destroy nuclei produced in **Big Bang Nucleosynthesis (BBN)**

↪ $T_R \lesssim 10^7 \text{ GeV}$ or $m_{3/2} \gg 1 \text{ TeV}$

↪ Consider **gravitino LSP**

Gravitino LSP and BBN

- Next-to-LSP (**NLSP**) decays to gravitino via gravity \rightsquigarrow long-lived

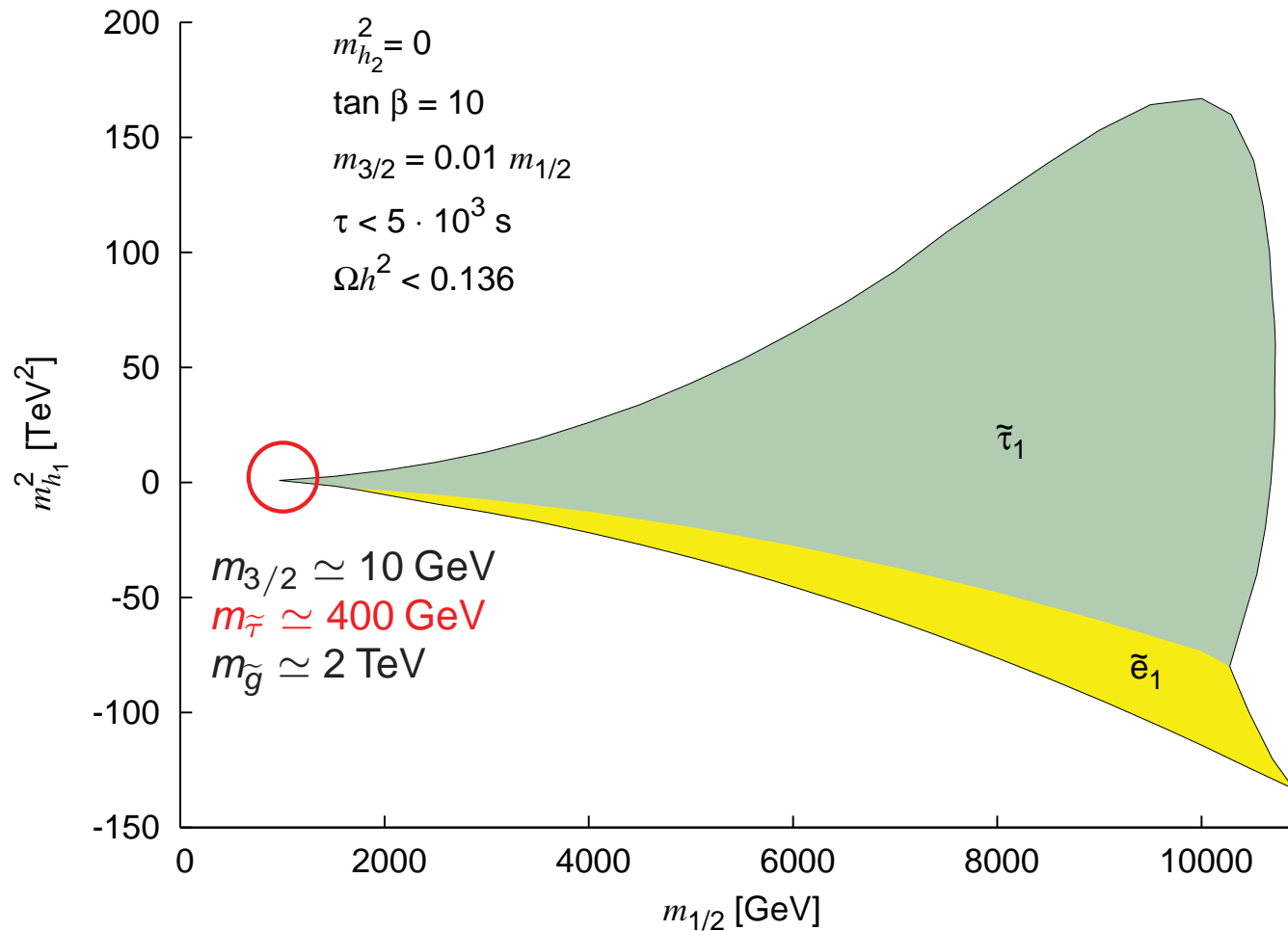
$$\text{E.g., for stau: } \tau_{\tilde{\tau}} \simeq 6 \cdot 10^4 \text{ s} \left(\frac{m_{3/2}}{1 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_{\tilde{\tau}}} \right)^5$$

- **Late NLSP decays** threaten **BBN**
- Stau NLSP **possible**
- **Charged NLSPs** form **bound states** with nuclei
 - \rightsquigarrow **BBN** reaction rates change \rightsquigarrow overproduction of ${}^6\text{Li}$
 - \rightsquigarrow Stau must decay before BBN: $\tau_{\tilde{\tau}} \lesssim 5 \cdot 10^3 \text{ s}$

\rightsquigarrow Relatively **heavy stau**

Gravitino LSP and BBN

Gaugino mediation



→ Maximal stau mass detectable at the LHC?

Scenario

- Gravitino LSP
- Stau NLSP
- R parity \rightsquigarrow pair production
- Stau **metastable**: leaves detector before decay ($\tau_{\tilde{\tau}} \gtrsim 10^{-8}$ s)
- **Charged** \rightsquigarrow direct detection possible, no missing E_T

\rightsquigarrow Signal: pair of slow, charged particles leaving the detector

Production Channels

Cascades: $p\bar{p} \rightarrow \tilde{g}, \tilde{q}, \chi^\pm \rightarrow \tilde{\tau}$

- Large cross section if \tilde{g} etc. not too heavy
- Many parameters relevant

Direct Drell-Yan production: $p\bar{p} \rightarrow Z, \gamma \rightarrow \tilde{\tau}^+ \tilde{\tau}^-$

- 2 parameters relevant: (lighter) stau mass $m_{\tilde{\tau}}$, mixing angle $\theta_{\tilde{\tau}}$
 \rightsquigarrow model-independent analysis possible
- Always present (dominant for large mass gap to heavier sparticles)
 \rightsquigarrow assured discovery potential, strict exclusion limits

\rightsquigarrow Focus on Drell-Yan production here



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Dependence on Mixing Angle

- $$\begin{pmatrix} \tilde{\tau}_1 \\ \tilde{\tau}_2 \end{pmatrix} = \begin{pmatrix} \cos \theta_{\tilde{\tau}} & \sin \theta_{\tilde{\tau}} \\ -\sin \theta_{\tilde{\tau}} & \cos \theta_{\tilde{\tau}} \end{pmatrix} \begin{pmatrix} \tilde{\tau}_R \\ \tilde{\tau}_L \end{pmatrix}$$

- Approximate factorization:

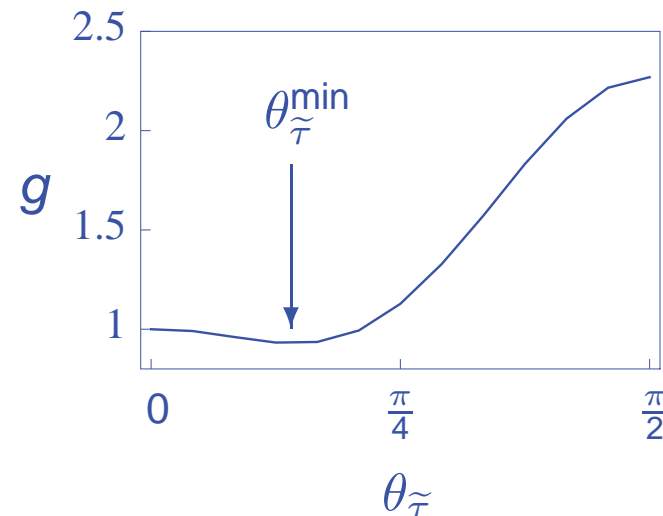
$$\sigma_{\tilde{\tau}}(m_{\tilde{\tau}}, \theta_{\tilde{\tau}}, \text{kinematics}) \simeq f(m_{\tilde{\tau}}, \text{kinematics}) \cdot g(\theta_{\tilde{\tau}})$$

- Smallest cross section
for $\theta_{\tilde{\tau}} = \theta_{\tilde{\tau}}^{\min} \neq 0$

→ **conservative** choice:

$$\theta_{\tilde{\tau}} = \theta_{\tilde{\tau}}^{\min}$$

→ **1** free parameter remaining



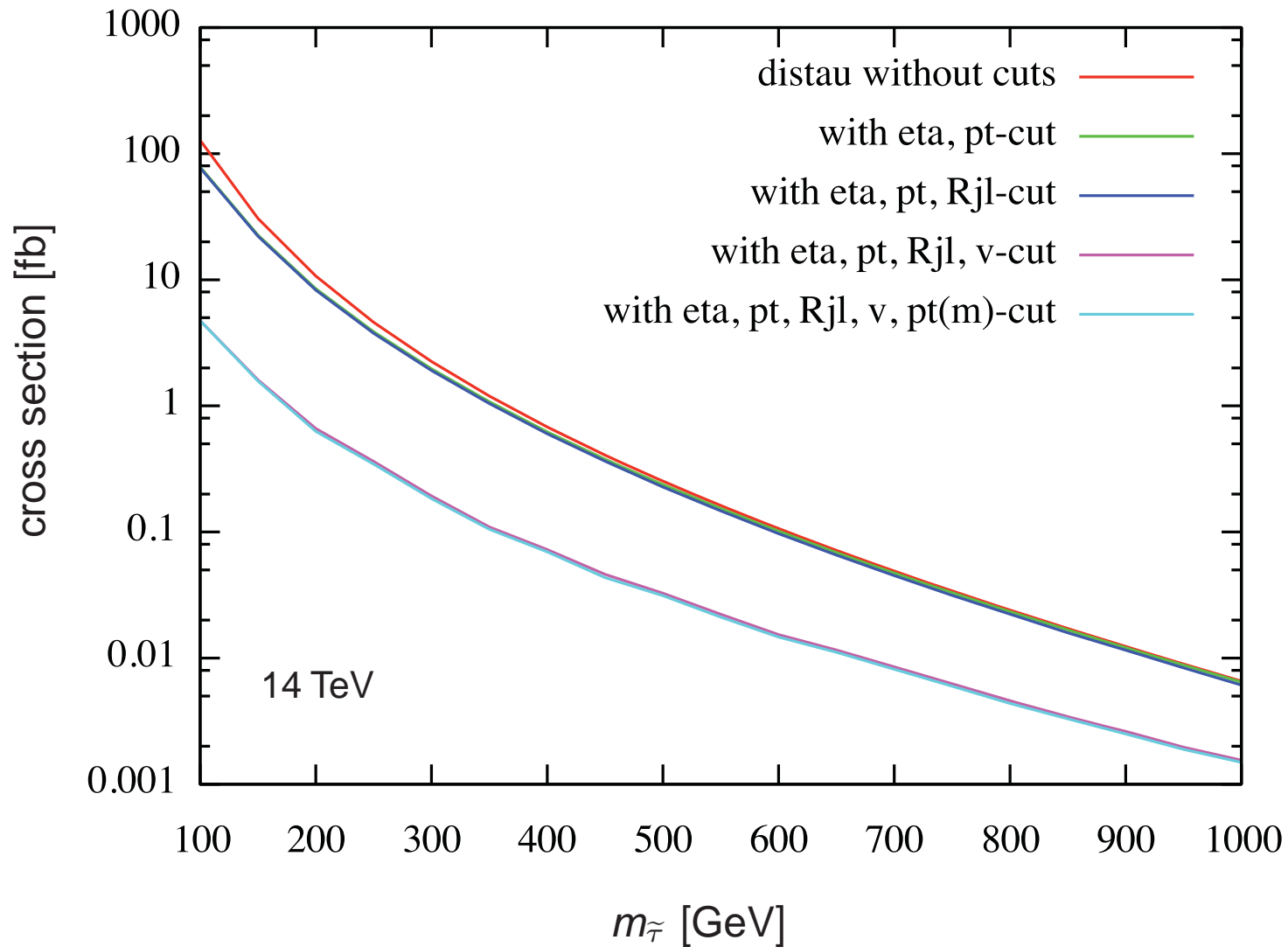
Background

- Background: muons
- Staus much **heavier** \leadsto **slower** \leadsto require **velocity** $\beta < 0.8$
- Velocity measurement:
 - Ionization **energy loss** in tracker
 - **Time of flight** to muon chambers
- Moderate precision for each, but **uncorrelated**
- Reach muon chambers before next bunch crossing $\leadsto \beta > 0.6$
- Combination \leadsto **background rejection factor** $r_\beta \simeq 10^{-7}$ for **single** stau candidate (for one particle) [CMS-PAS-EXO-08-003]
 $\leadsto r_\beta = 10^{-14}$ for 2 particles?
- **Conservative** and **sufficient** choice: $r_\beta = 10^{-10}$
- Also consider worst case $r_\beta = 10^{-7}$

Cuts

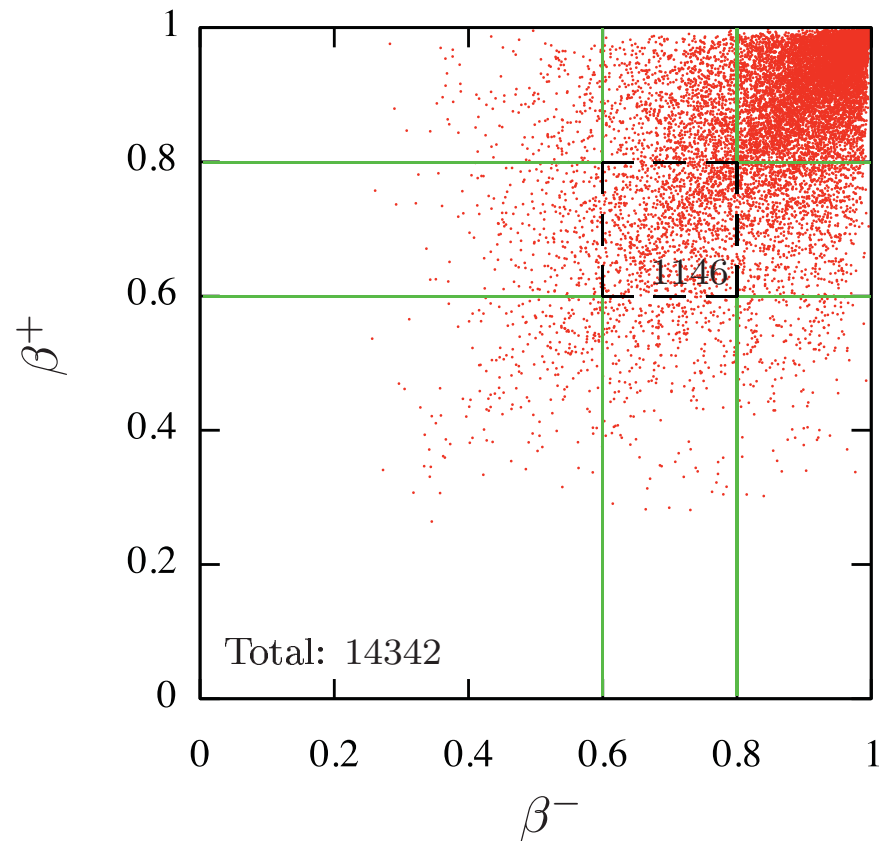
- $0.6 < \beta < 0.8$
- $p_T > 50 \text{ GeV}$
- Pseudorapidity $|\eta| < 2.5$
- Isolation from jets: $\sqrt{\Delta\eta_{\mu,\text{jet}}^2 + \Delta\phi_{\mu,\text{jet}}^2} > 0.5$ for jets with $p_T > 50 \text{ GeV}$

Cross Section for Direct Stau Pair Production

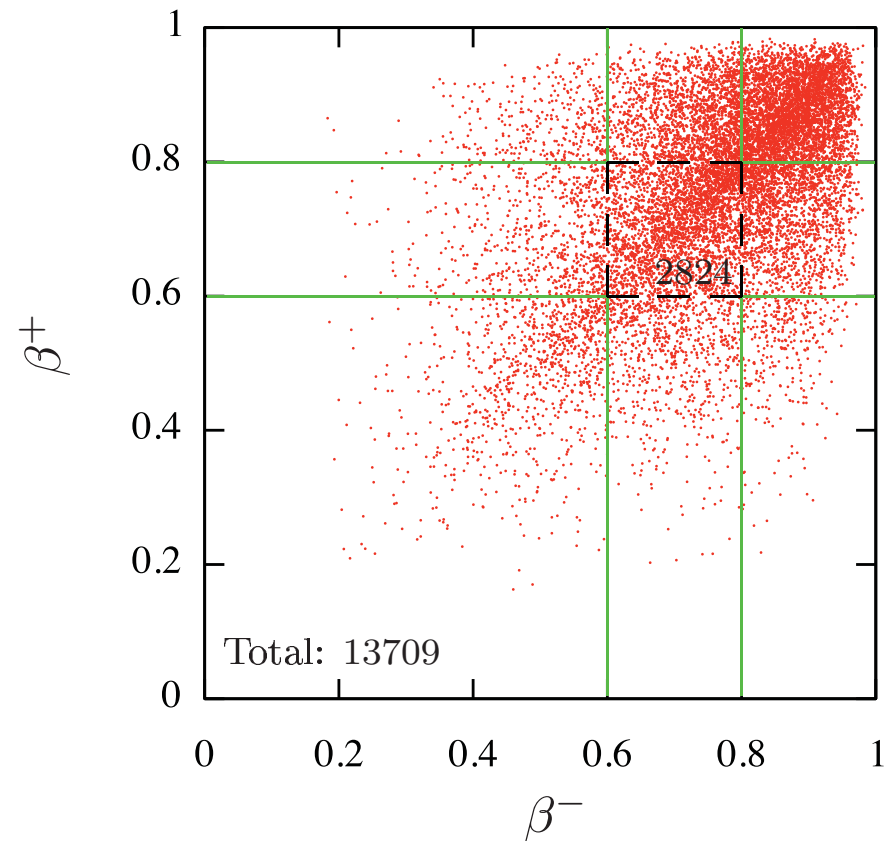


Stau Velocity Distribution

$m_{\tilde{\tau}} = 200 \text{ GeV}$



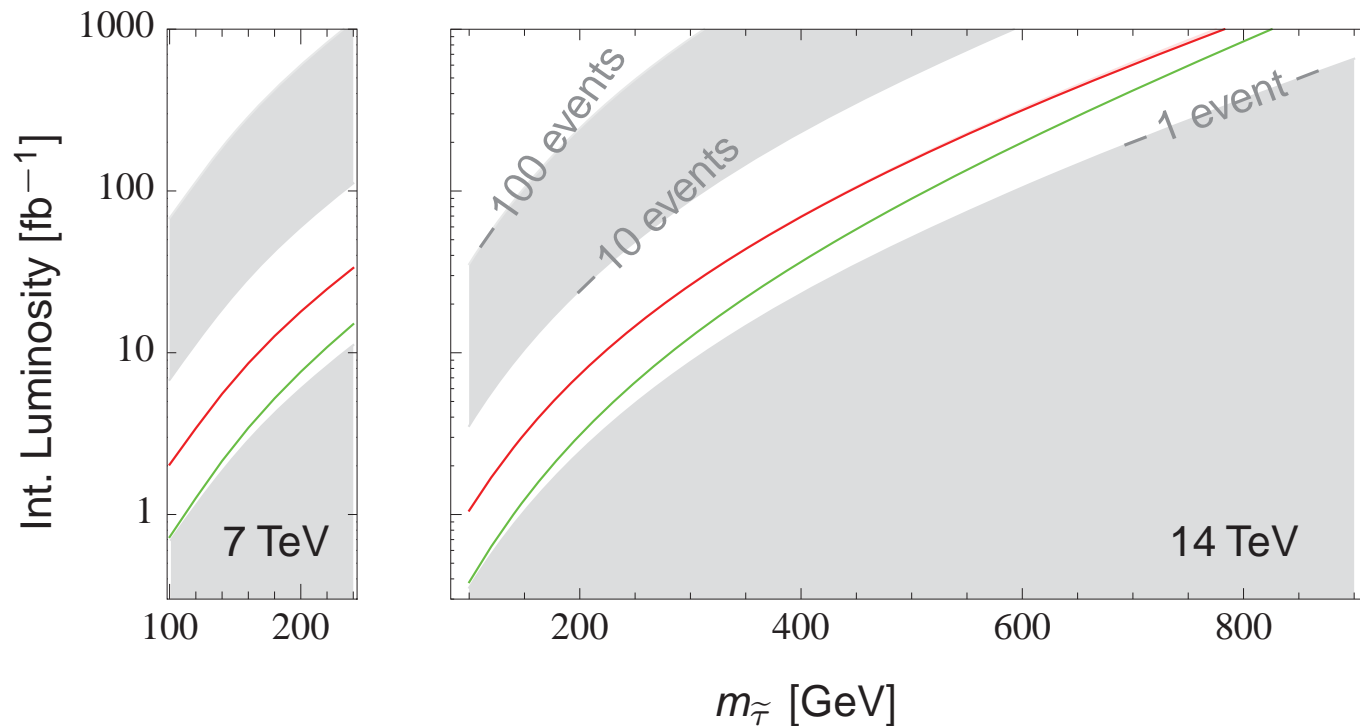
$m_{\tilde{\tau}} = 800 \text{ GeV}$



Discovery and Exclusion Potential

Expected luminosity required for

- 5σ discovery
- 95% CL exclusion

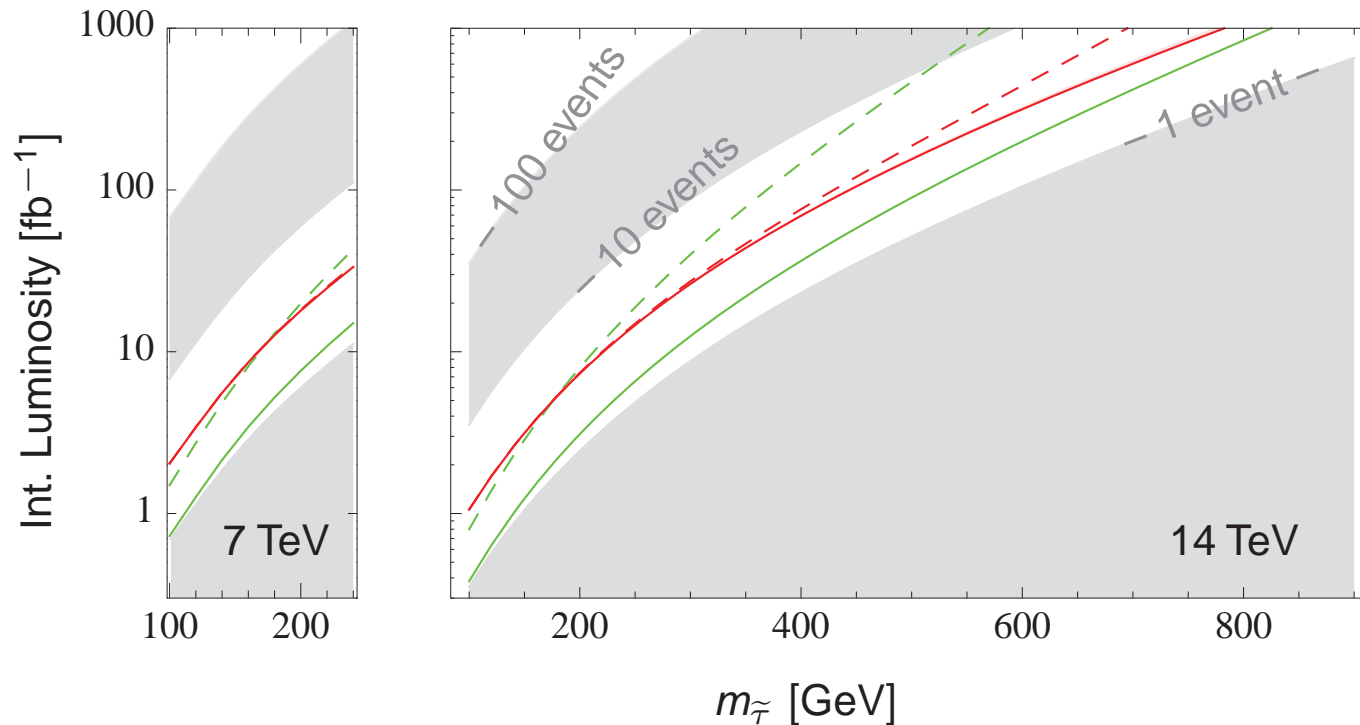


Computed using MadGraph + Pythia

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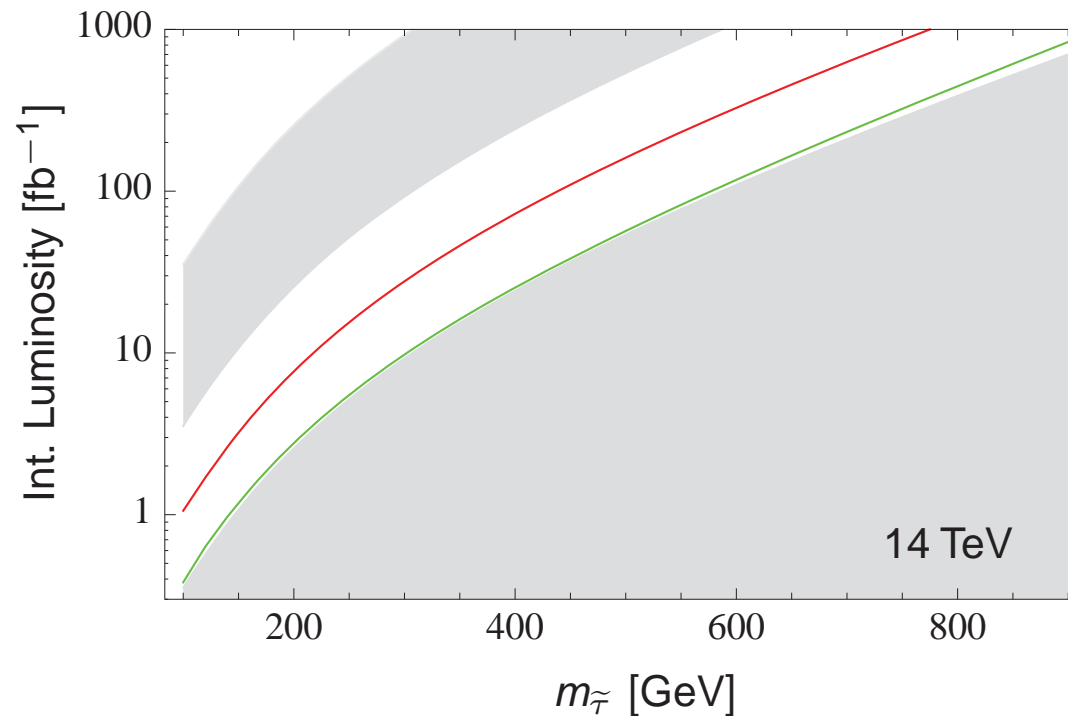


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Optimized p_T Cut

Muon background drops drastically with p_T

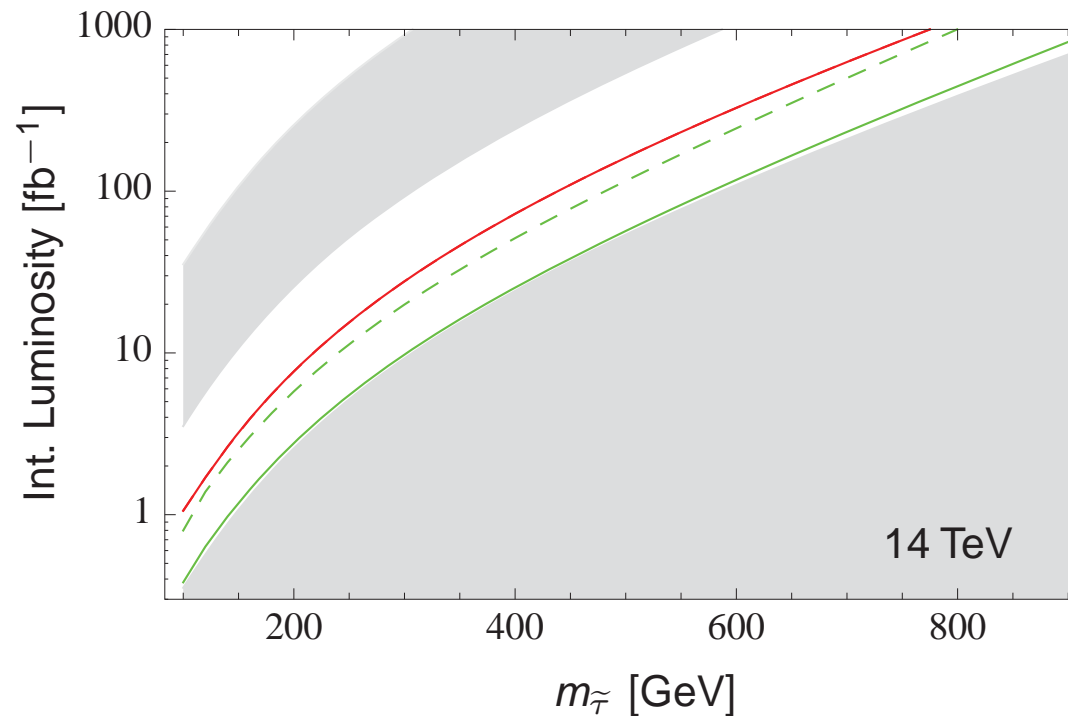
\leadsto Adjust p_T -cut according to $m_{\tilde{\tau}}$: $p_T > 0.4 \cdot m_{\tilde{\tau}}$



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→ Adjust p_T -cut according to $m_{\tilde{\tau}}$: $p_T > 0.4 \cdot m_{\tilde{\tau}}$





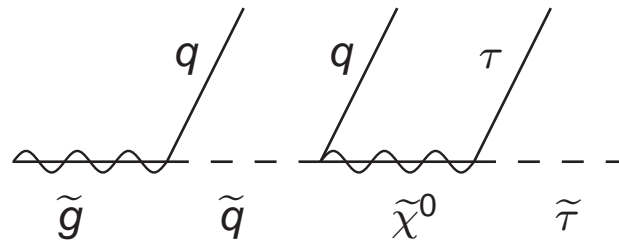
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Production via Cascades

- Production of **colored** sparticles likely to dominate
- More parameters but which are really relevant?
- Consider **simplified scenario**
 - Gluino
 - Squarks $\tilde{u}_L, \tilde{u}_R, \tilde{d}_L, \tilde{d}_R$ with common mass $m_{\tilde{q}}$
 - Lightest neutralino
 - Stau NLSP
- Minimal decay chain



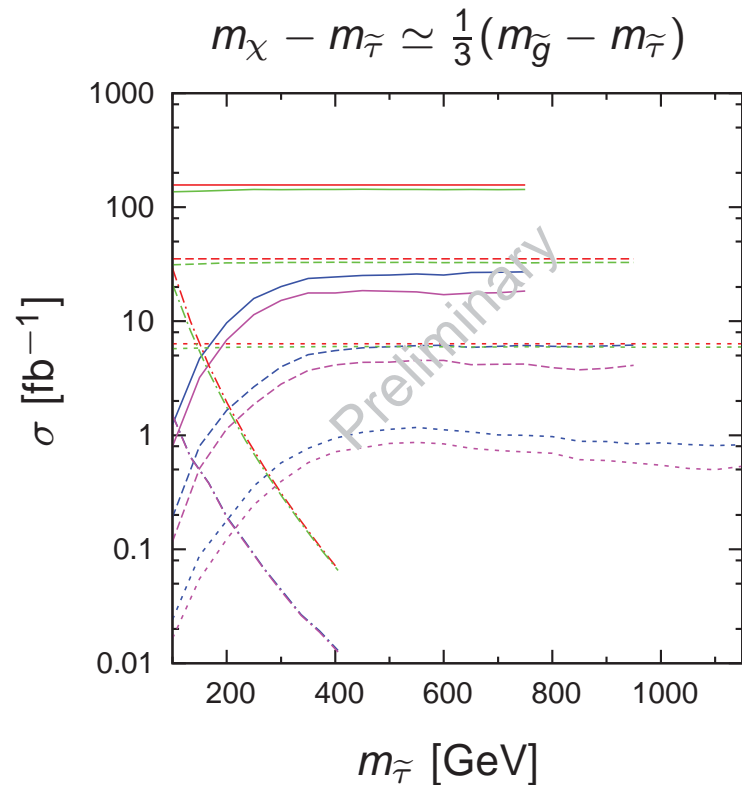
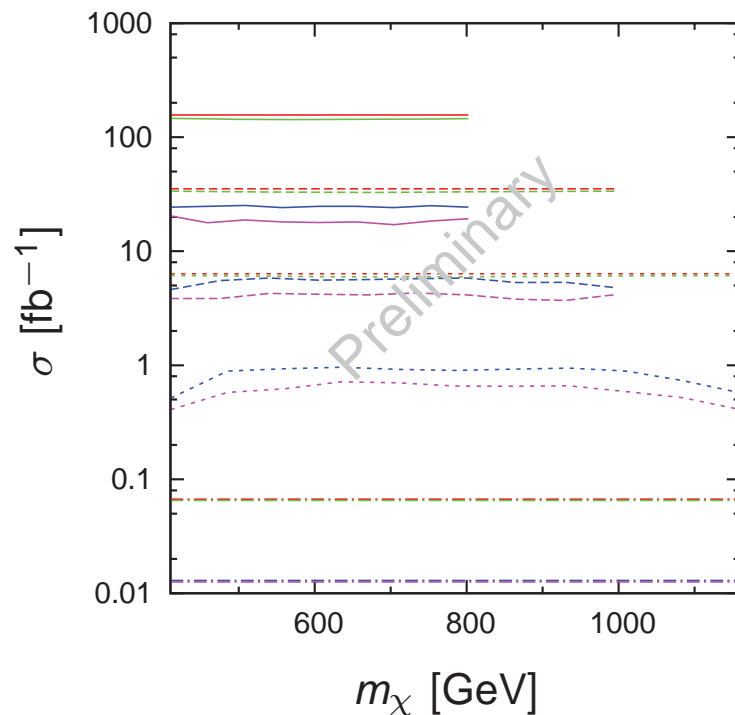
- Same cuts as before
- Consider 7 TeV LHC

Dependence on Stau and Neutralino Masses

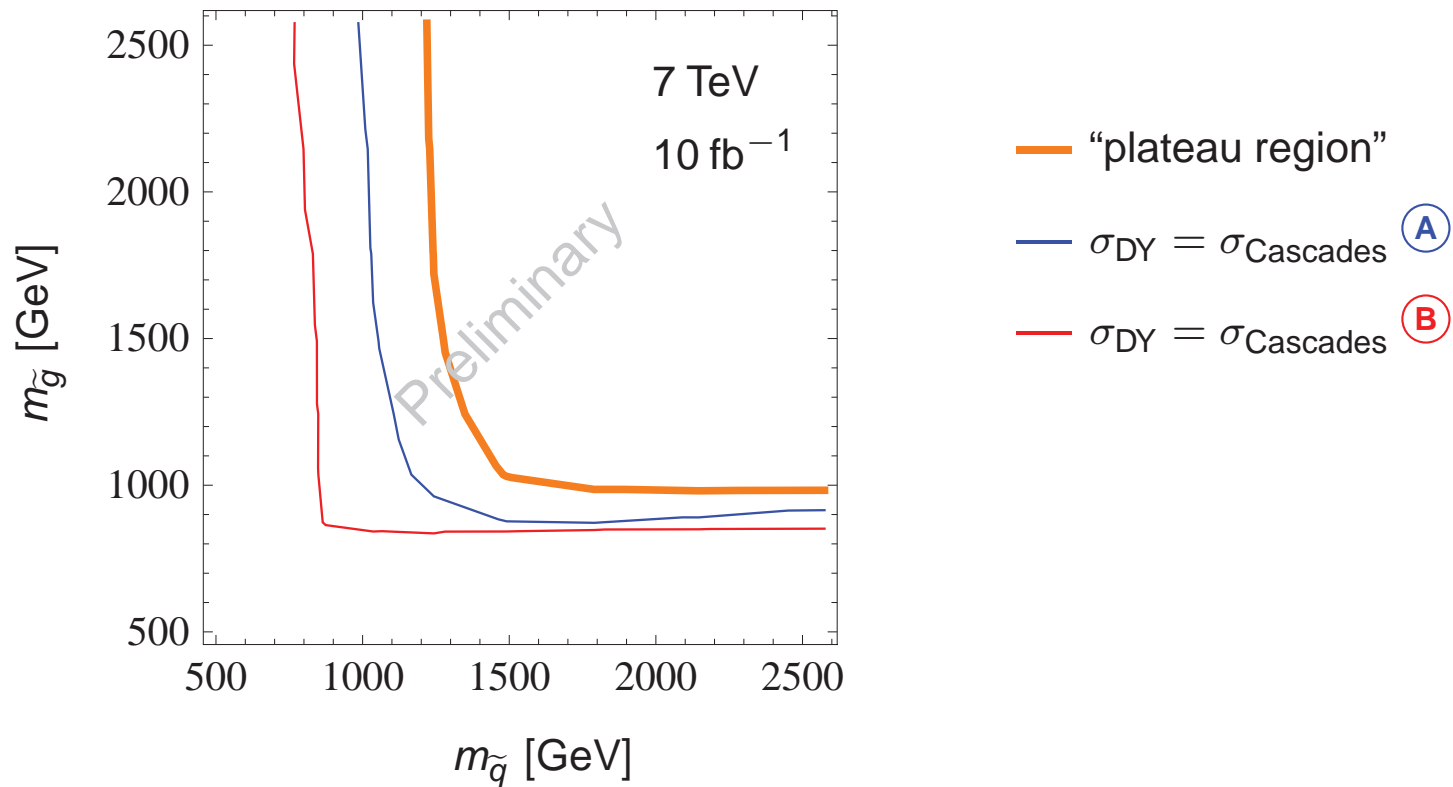
Velocity of final staus **not very sensitive** to $m_{\tilde{\tau}}$ and m_{χ} for sufficiently small mass gap to $m_{\tilde{g}}$ and $m_{\tilde{q}}$

$$m_{\tilde{g}} = m_{\tilde{q}} \in \{900 \text{ GeV}, 1000 \text{ GeV}, 1250 \text{ GeV}\}$$

$$m_{\tilde{\tau}} = 400 \text{ GeV}$$



Exclusion Potential



Conclusions

- Metastable charged SUSY particle gives prominent signature
 \rightsquigarrow **high background rejection possible**
- Upper limit on velocity most crucial for gaining sensitivity
- **Drell-Yan** production
 \rightsquigarrow **model-independent** results (one parameter only)
- With 14 TeV and 300 fb^{-1} : Exclusion up to 600 GeV, discovery up to 750 GeV
- **Cascade production** depends mainly on gluino and squark masses \rightsquigarrow exclusion plot for simplified model seems feasible
- Valid for **all scenarios with a metastable charged slepton**

Muon Background

