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International Centre for Theoretical Physics*



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*II April - 1 May, 2012*

**On the Variability of TEC in the Indian Low Latitude Sector**

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Workshop on Science Applications of GNSS in developing countries  
ICTP, TRIESTE, ITALY  
*11-27 April, 2012*

# On the Variabilities of TEC in the Indian Low Latitude Sector

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Visakhapatnam, INDIA*

## Contents:

- Basic concepts of Electrojet and Equatorial Ionization Anomaly (EIA)
- Morphological features of GPS -TEC measured at four different Indian Latitudes
  - *diurnal, day to day, seasonal and latitudinal variations*
- TEC and solar activity indices
  - *Sunspot Number, Solar Flux (F 10.7), EUV*
- EIA and its association with Equatorial Electrojet (EEJ) using GPS TEC from 9 different stations ( $8.47^{\circ}\text{N}$  to  $31.09^{\circ}\text{N}$ ) along the common meridian of  $77^{\circ}\text{E}$
- Variations in the grid ( $5^{\circ} \times 5^{\circ}$ ) based measurements of GPS TEC in the Indian low latitude sector



# EQUATORIAL ELECTROJET

## *Definition*

- The **equatorial electrojet (EEJ)** is a narrow band of current at the E-region altitudes of the ionosphere flowing eastward along the dayside dip equator.

## *Generation Mechanism*

- The worldwide solar-driven wind flowing across field lines results in the E-region Sq-current system that produces an intensified eastward electric field at the equatorial day side of the ionosphere.

## *Measurement*

- This manifests itself in the enhancement of the horizontal component (H) of the earth's magnetic field at the magnetic equator with the maximum occurring at  $\pm 3^{\circ}$  on either side of the equator.

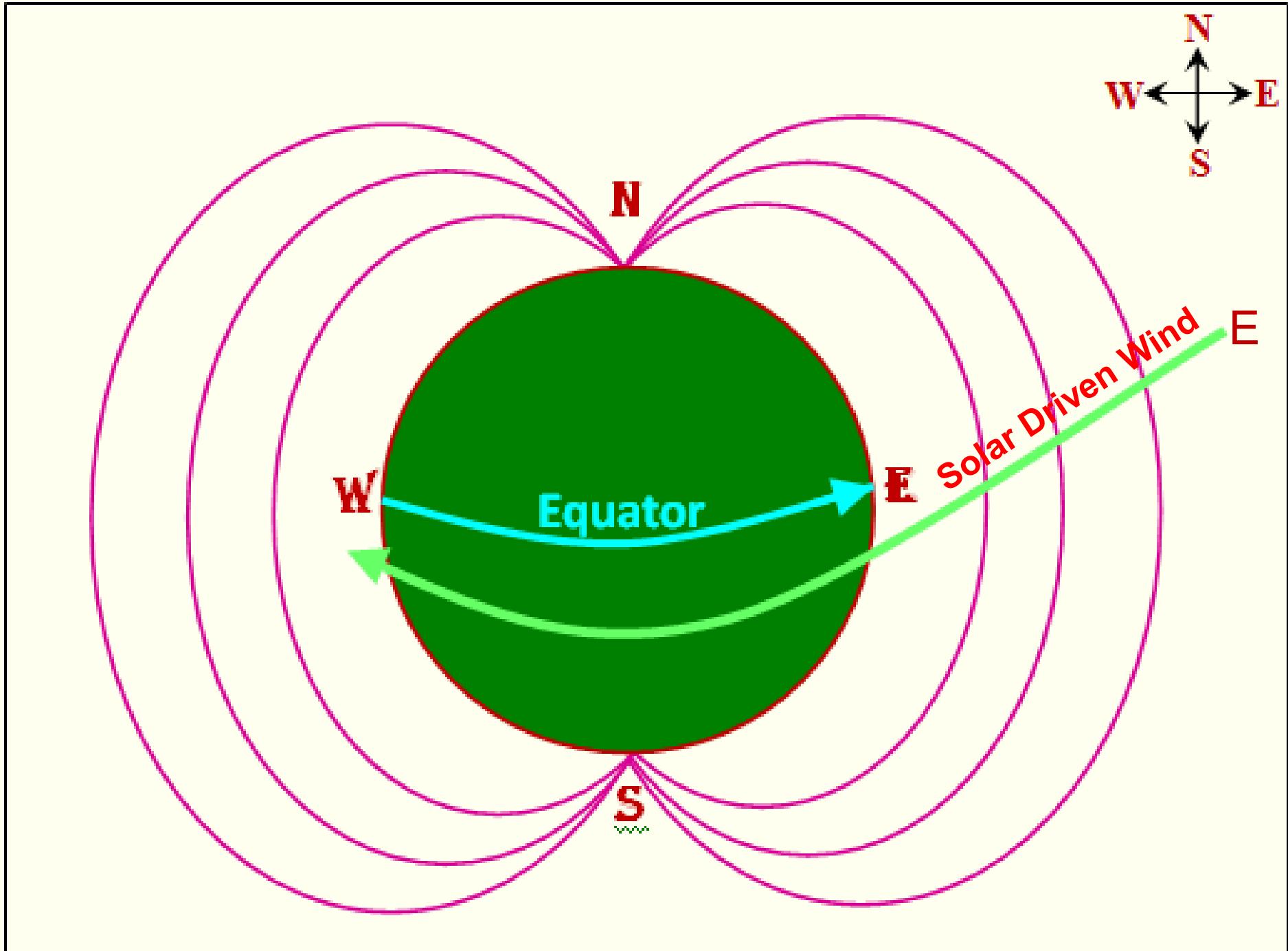


- Why the EEJ is confined to  $\pm 3^{\circ}$  geomagnetic latitude

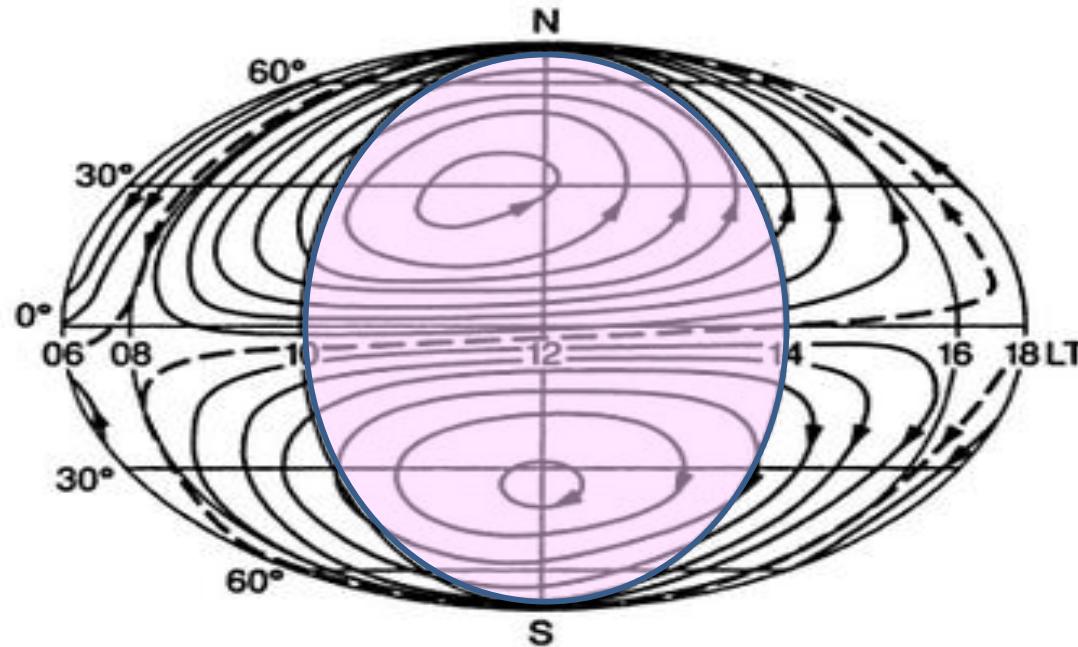
*It is dictated by the magnetic field line geometry which is horizontal and perpendicular to the magnetic equator.*

- Why only at the E-region altitudes
- Because the conductivity is maximum at the E-region altitudes of 100 to 110 km where the gyro frequency ( $\omega$ ) and collision frequencies ( $v$ ) are nearly equal.*





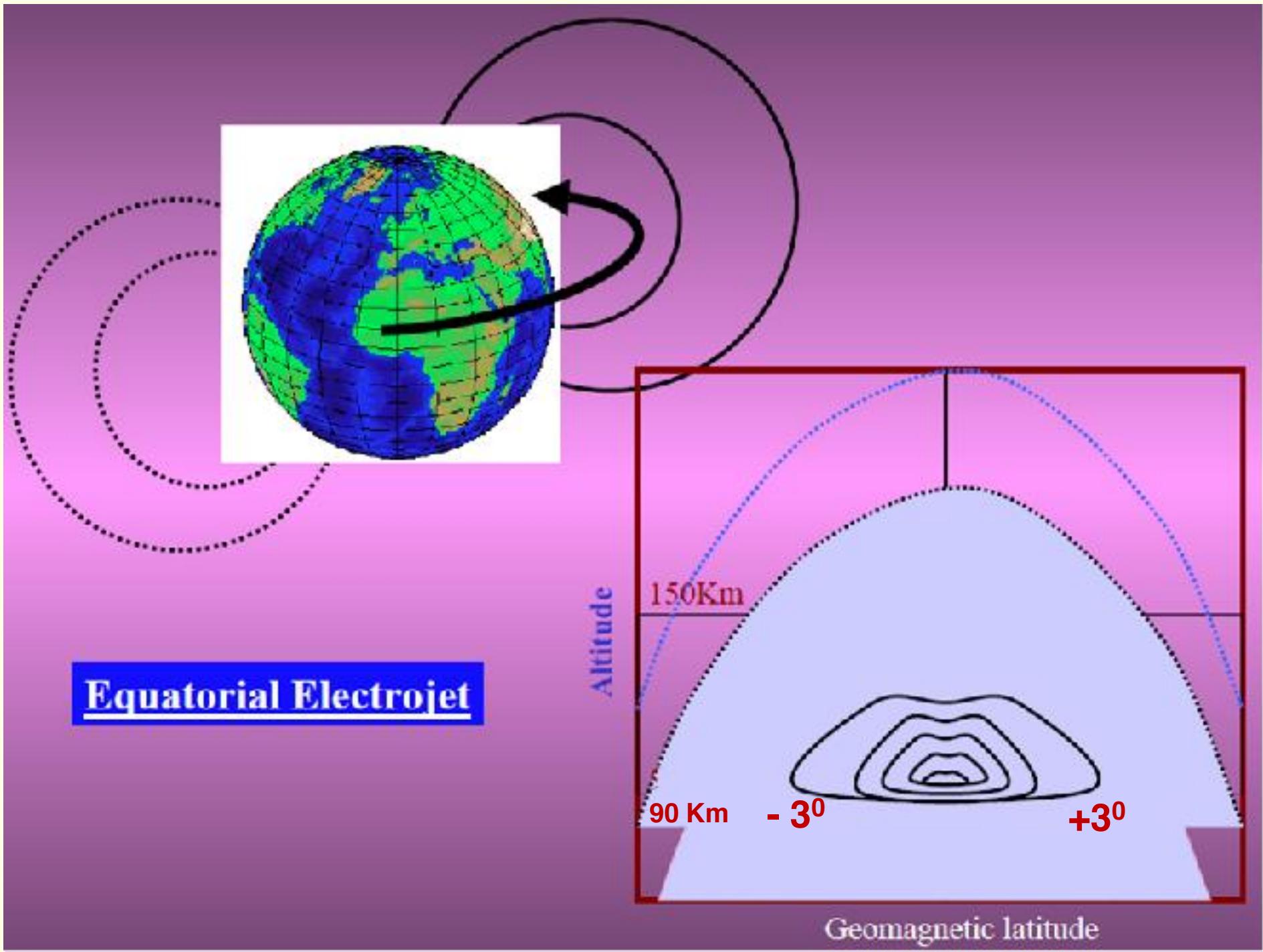
## Sq Current System



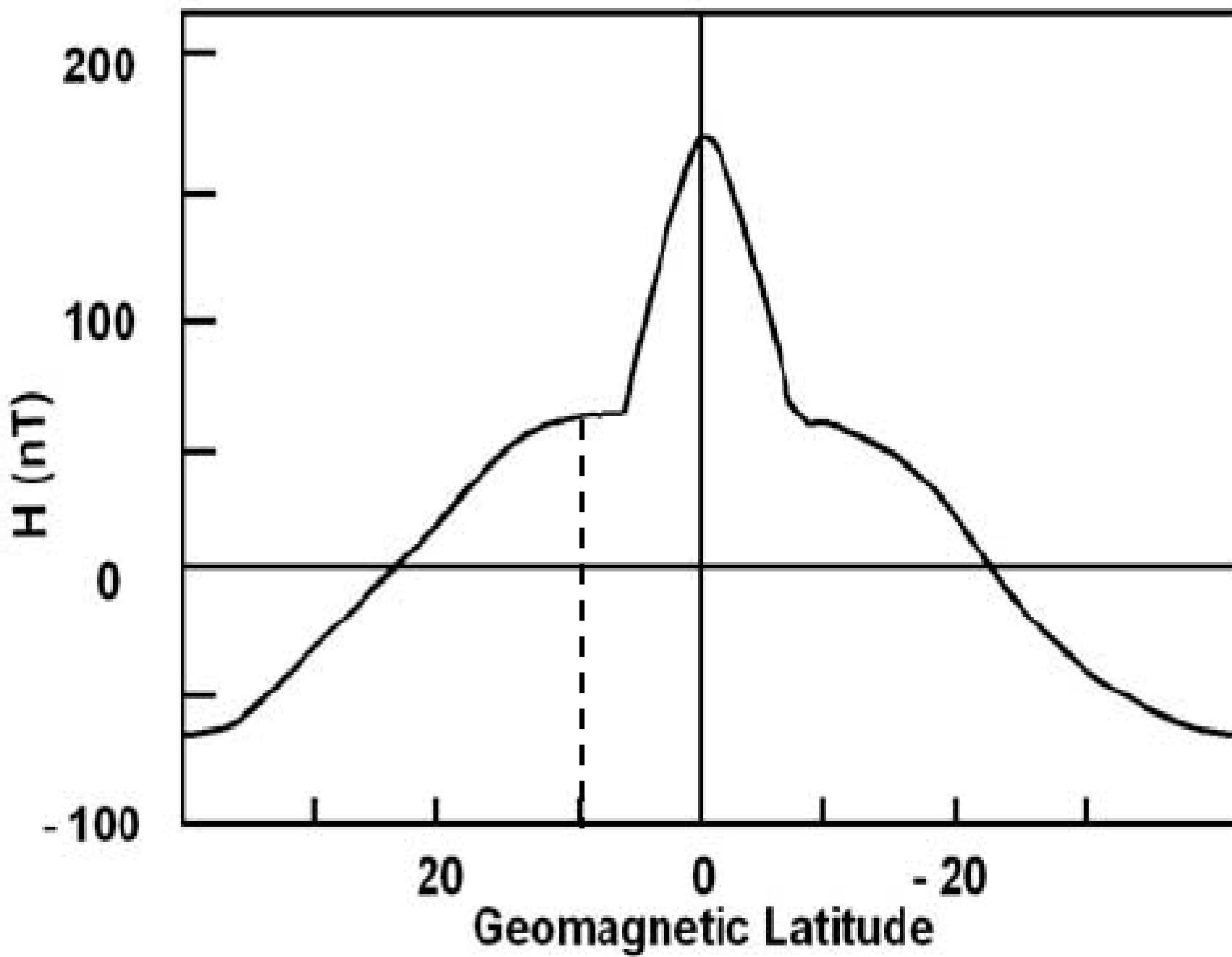
Global view of the average Sq current system

- At the geomagnetic equator, the Sq current systems of the southern and northern hemisphere merge with each other and intensify to form a ***jet-like current*** in the E-region of the ionosphere, which is called as ***Equatorial Electrojet***.
- The special geometry of the magnetic field over the equator (nearly  $\perp$  to the wind) together with the nearly perpendicular incidence of solar radiation causes an enhancement in the effective conductivity which leads to an amplification of the jet current over the equator.





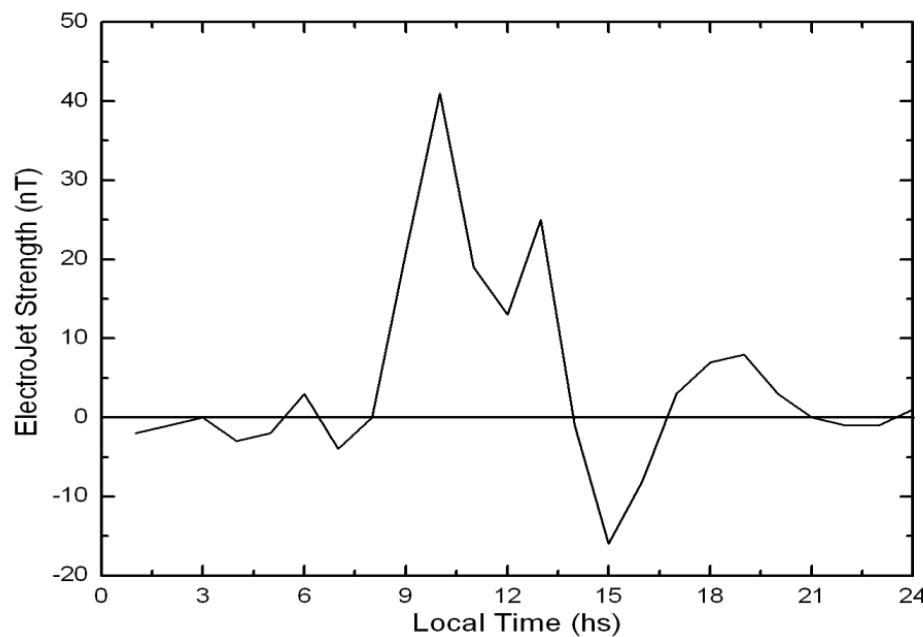
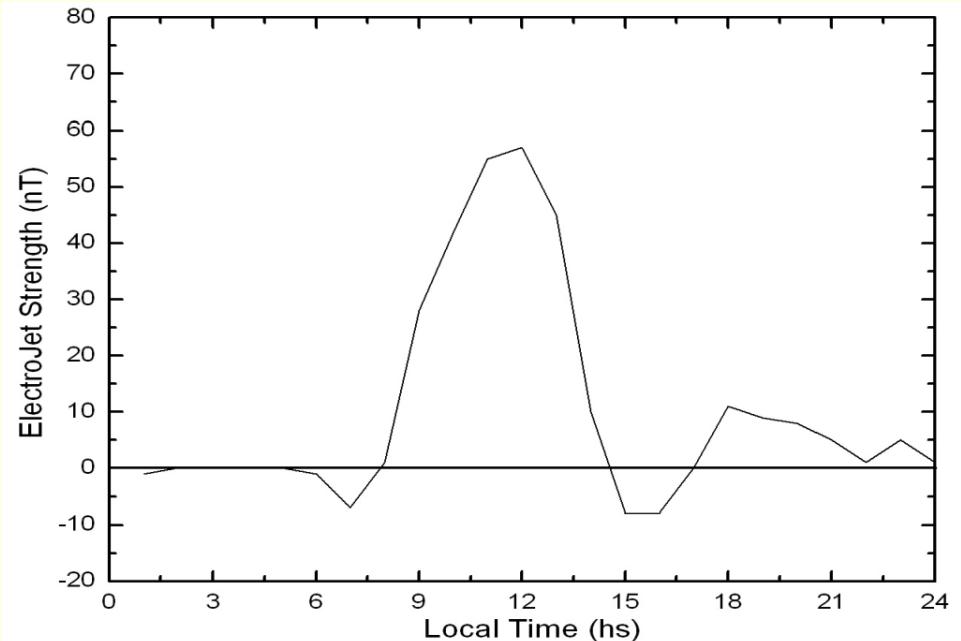
## HORIZONTAL COMPONENT OF EARTH'S MAGNETIC FIELD (H)



- The Electrojet is measured in nano Tesla (nT) by taking the difference in the horizontal components of the earth's magnetic field at the magnetic equator and at a location about  $10^{\circ}$  away from the equator so as to subtract the components of the magnetic fields from the inter planetary magnetic field (IMF), etc.,
- In the case of Indian sub continent Trivndrum ( $0.7^{\circ}$ ) – Alibag ( $10^{\circ}$ ) are generally used to derive the electrojet strength



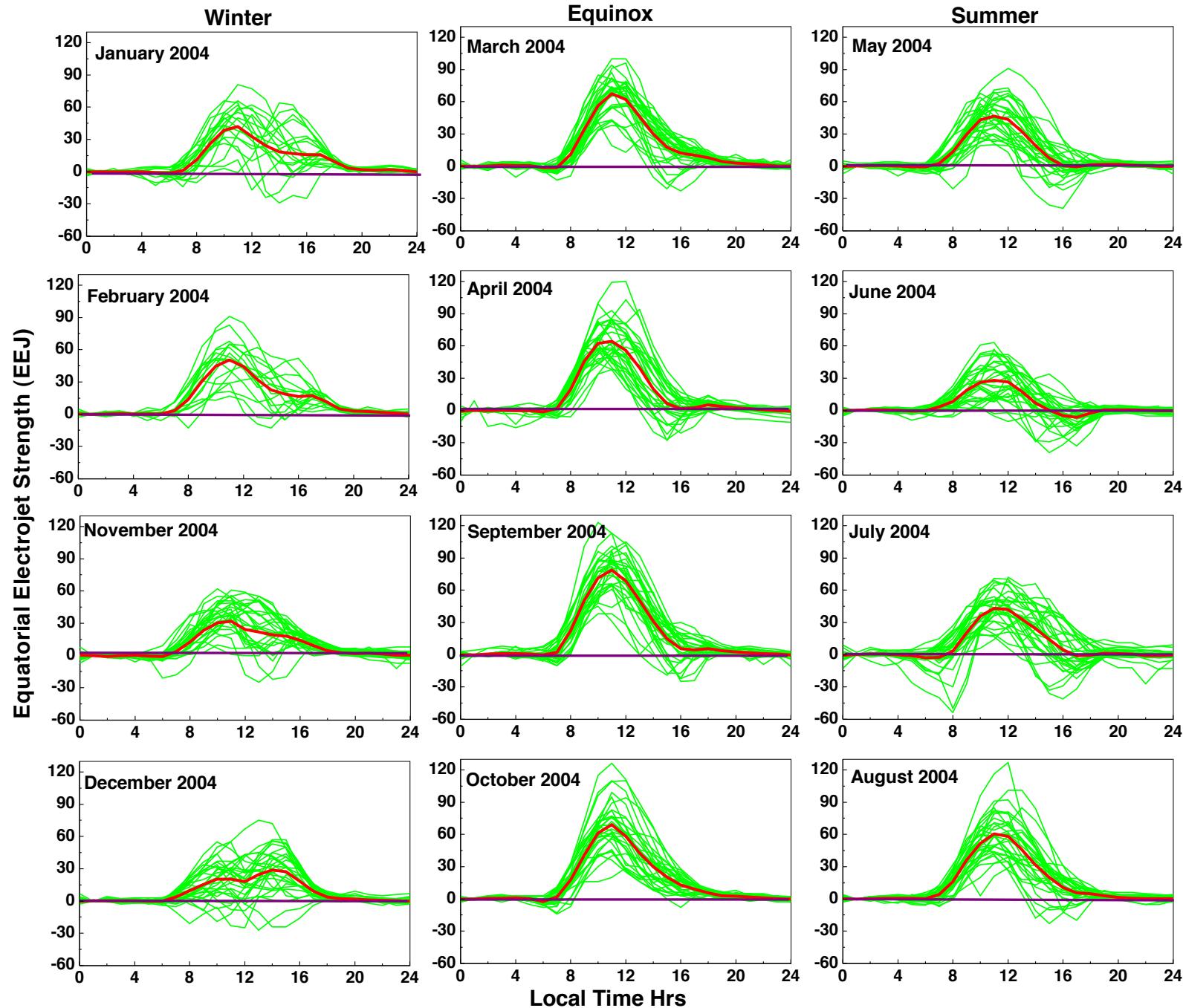
# Equatorial ElectroJet (EEJ)



# Counter ElectroJet (CEJ)

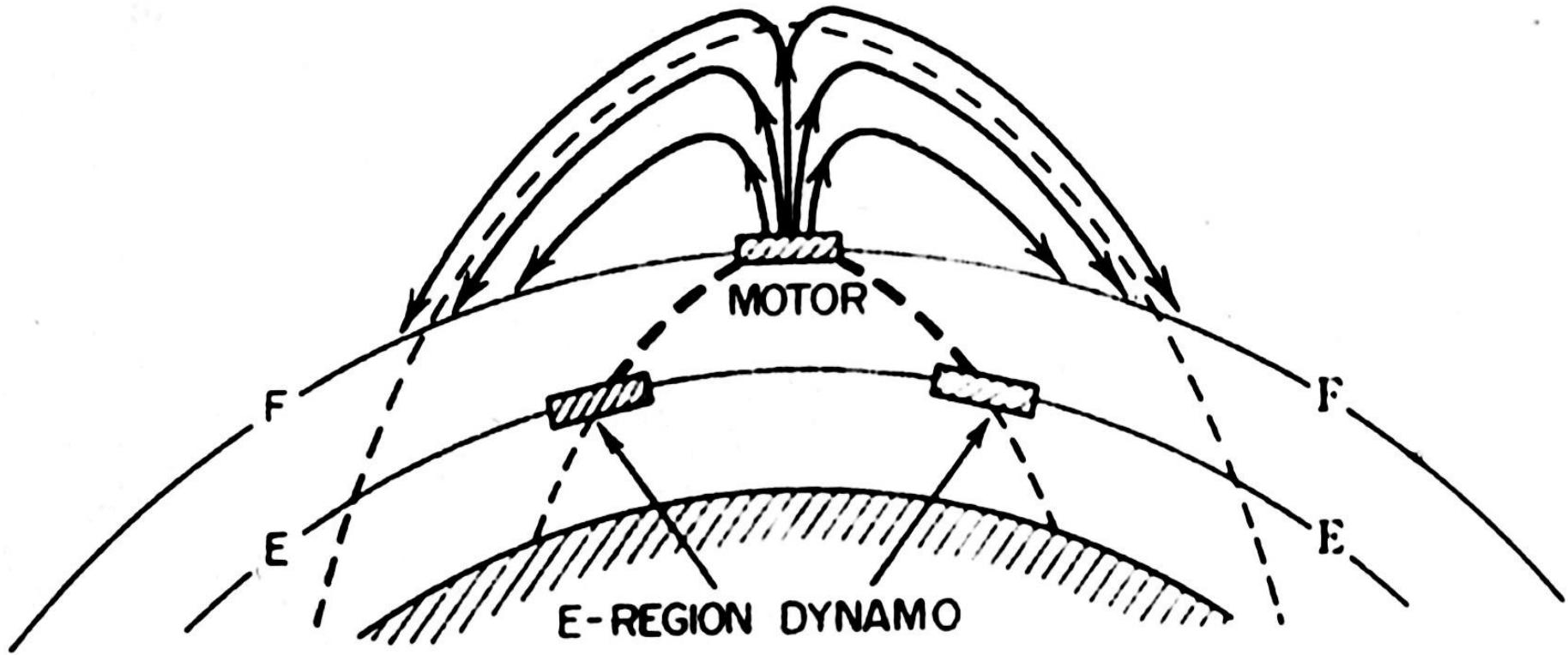


# Electrojet Strength at an equatorial station, Trivandrum



# Equatorial Ionization Anomaly





- The F-region geomagnetic anomaly. Near the equator the electric fields of the atmospheric dynamo in the E layer are conveyed upwards along geomagnetic lines of force to the motor in the F layer where they produce an upwards movement of the plasma during the day. The raised plasma then diffuses down lines of force to produce enhanced concentration at places on each side of the equator and decreased concentration at the equator itself. [Ratcliffe. 1972].



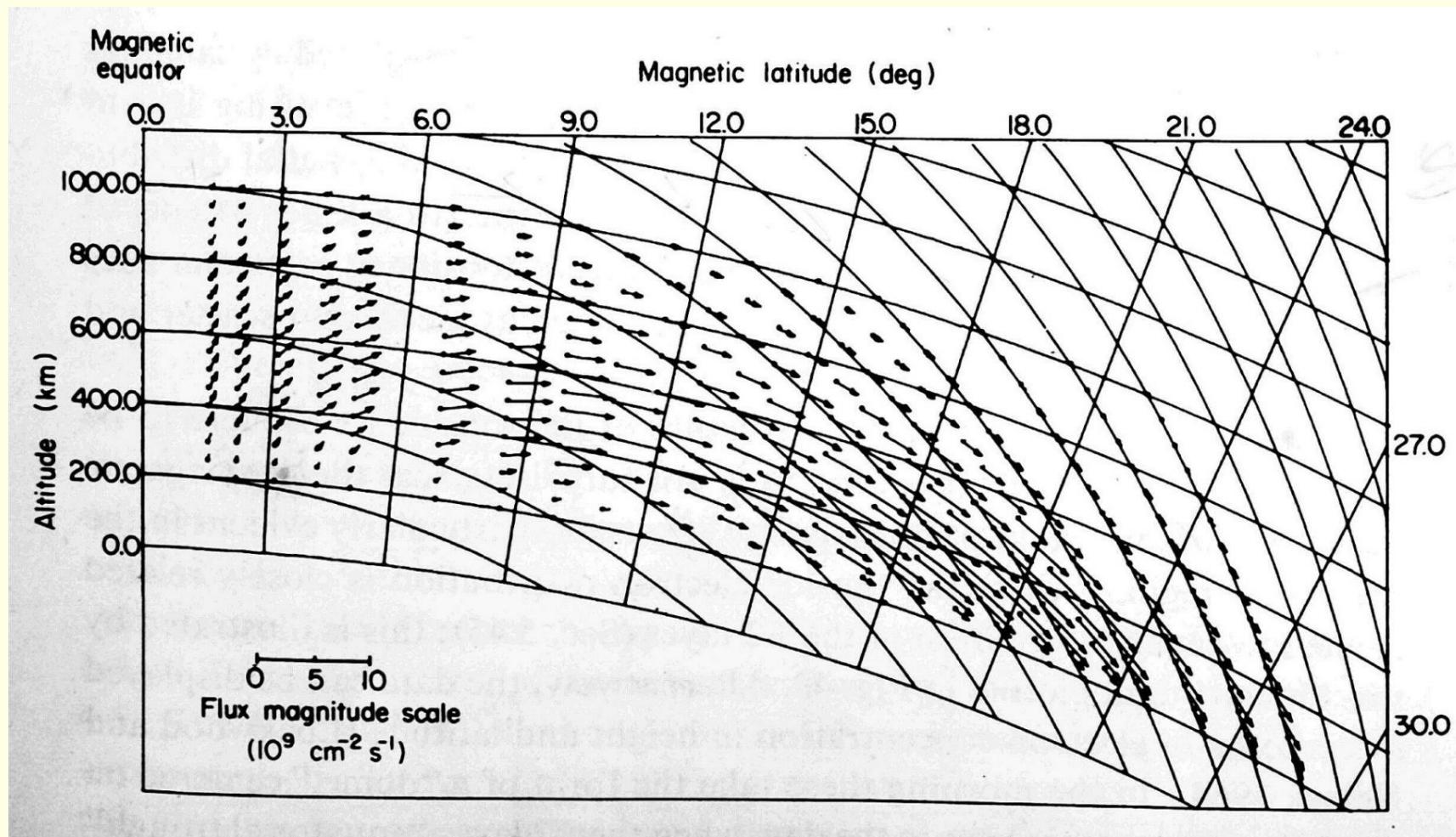
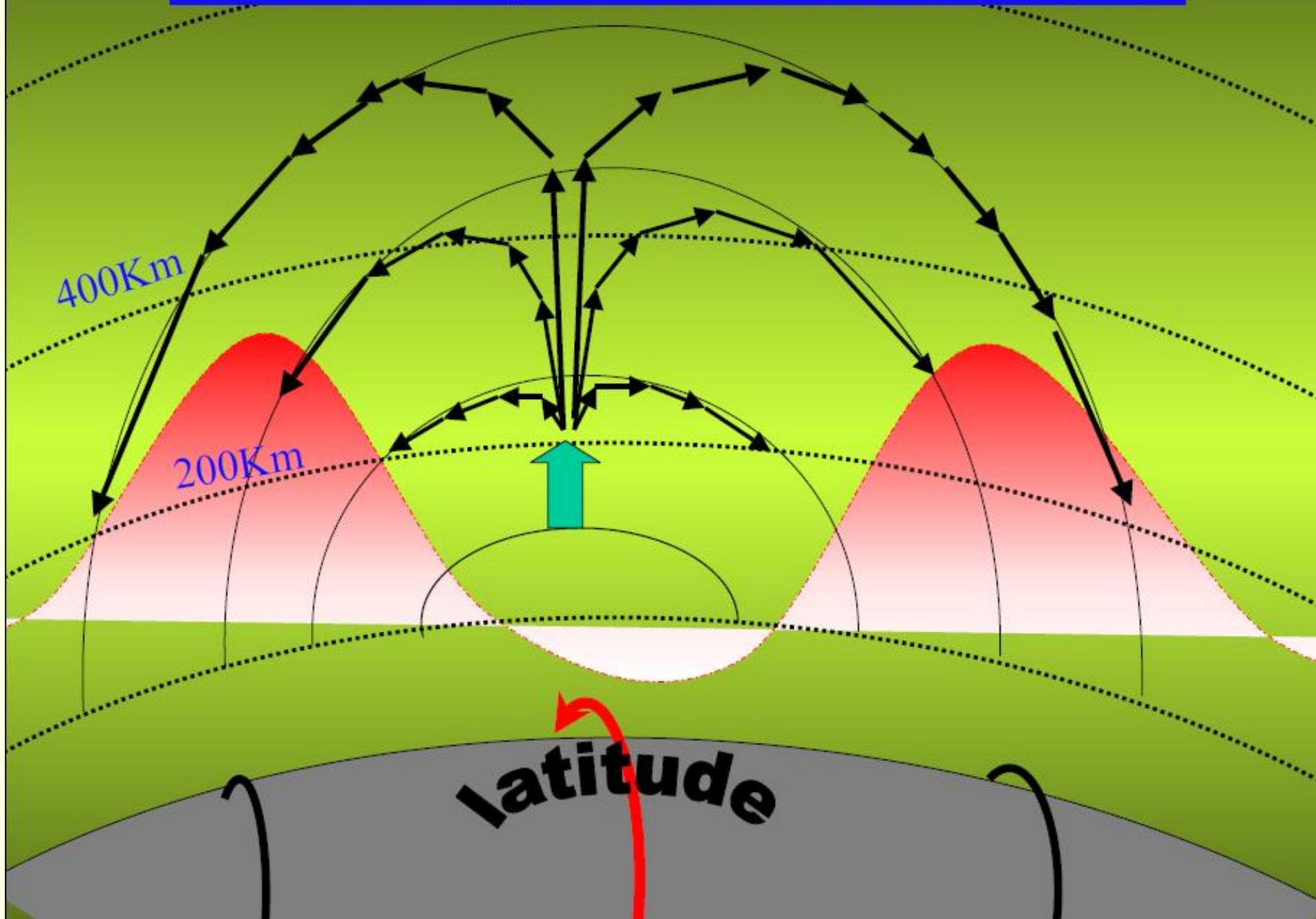
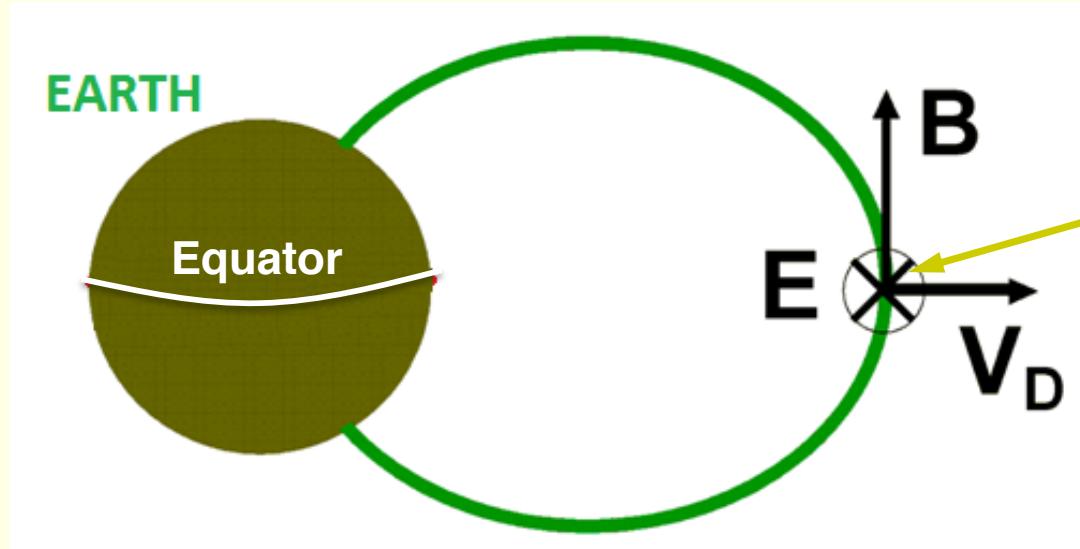


Fig. The “equatorial fountain.” Vector plot of electron flux in the meridian plane, in a theoretical steady-state model of the equatorial F region. The motions are due to the combined action of plasma diffusion along magnetic field lines and electromagnetic drift across field lines, produced by an assumed distribution of eastward electric field. The magnetic field lines are shown every 200 km above the equator [Hanson and Moffett (1966)].



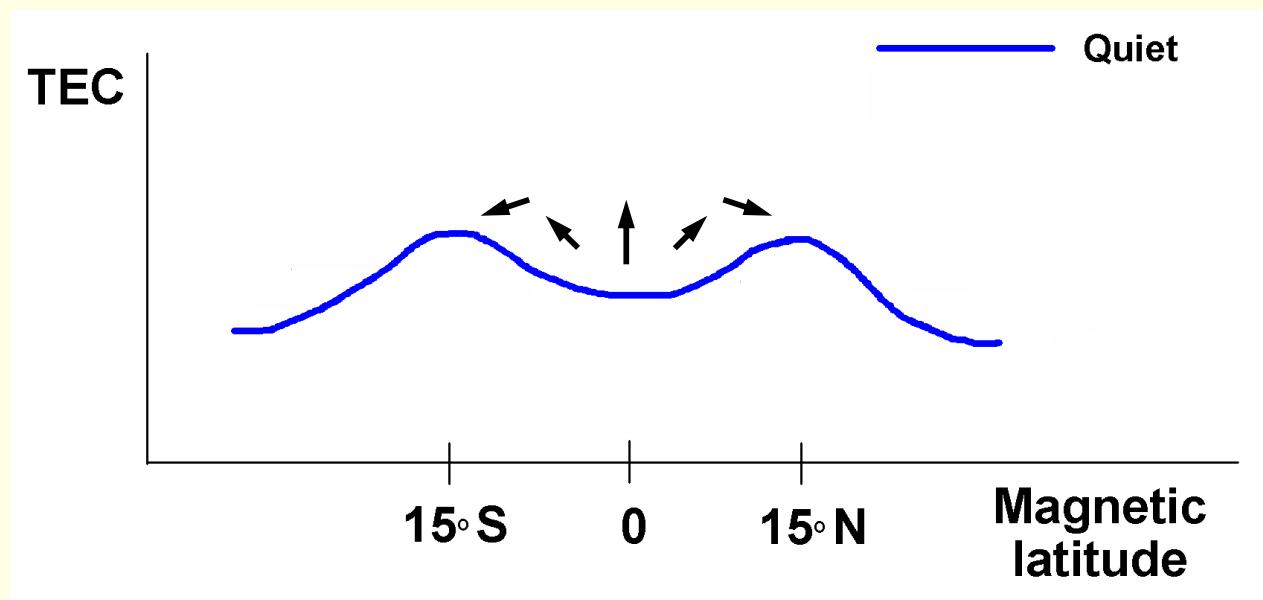
# Generation of Equatorial Ionisation Anomaly (EIA)

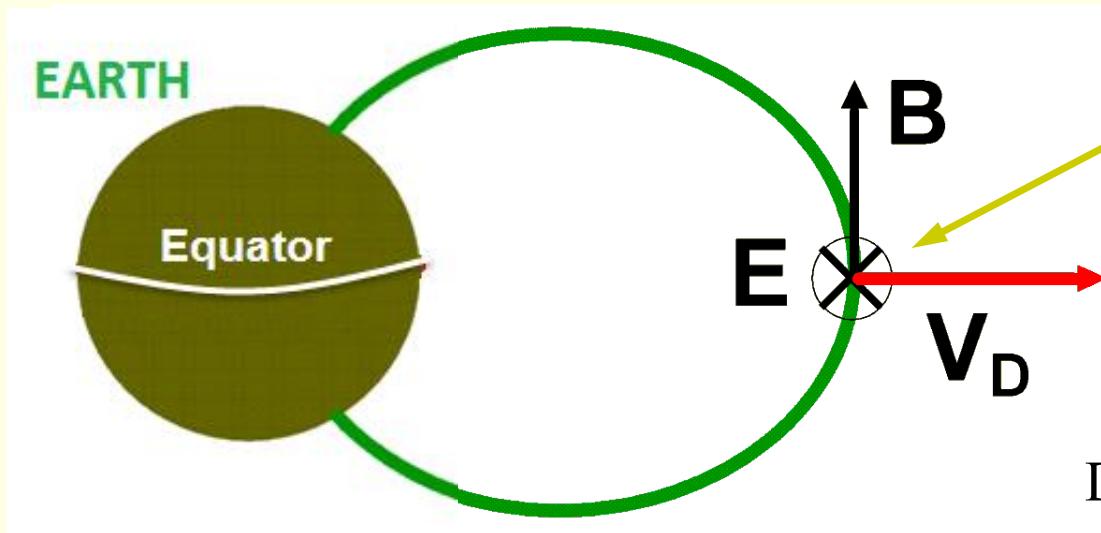




Eastward  
Electric Field

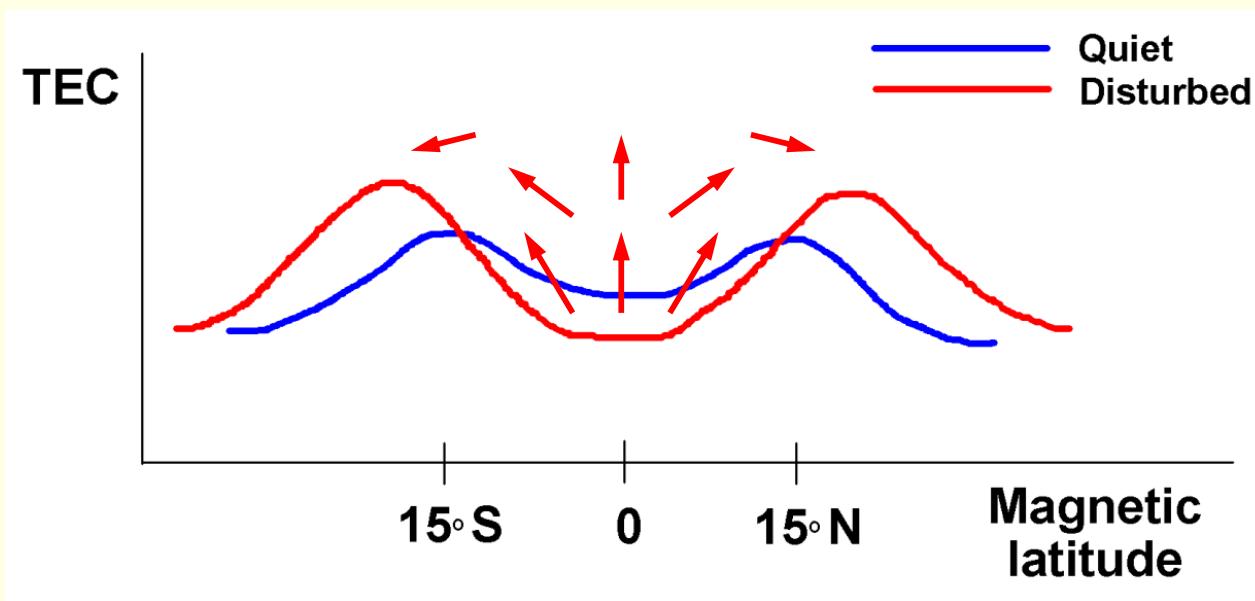
Daytime ~ 9 LT to 21 LT

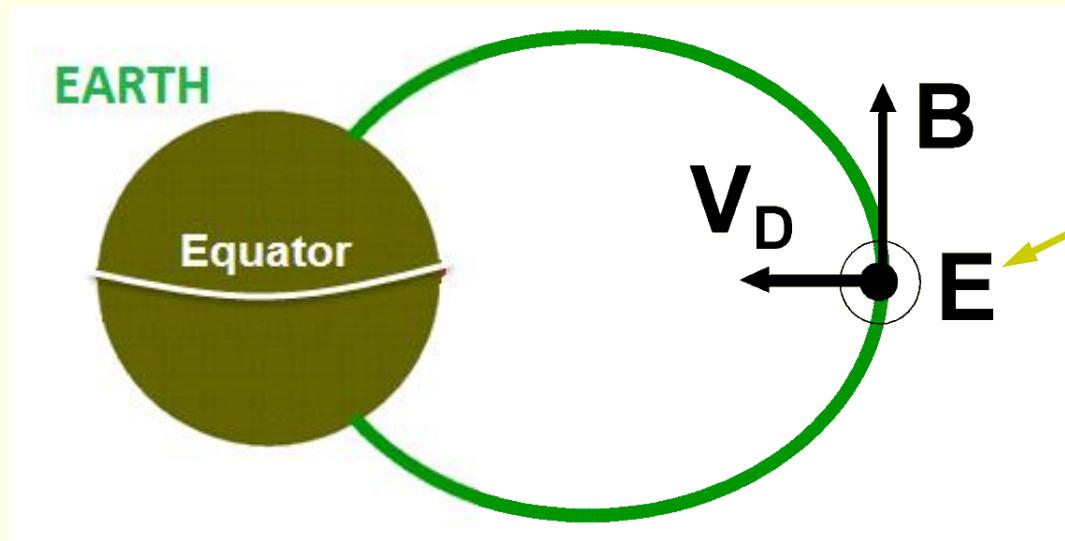




Intensified  
Eastward  
Electric Field

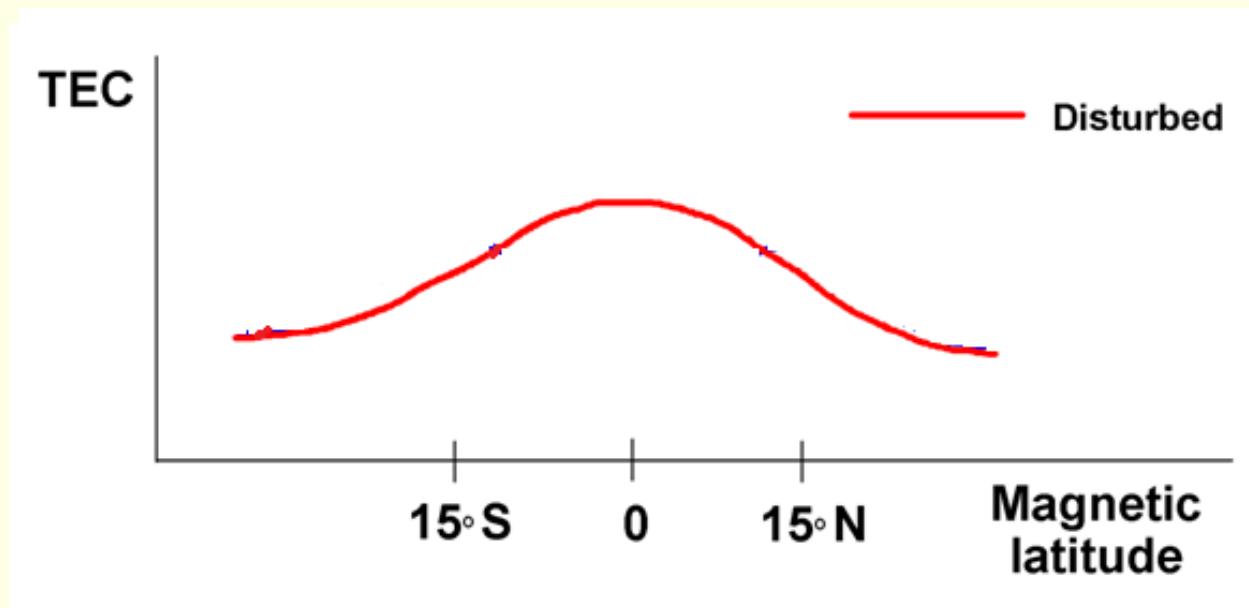
Daytime  $\sim 9$  LT to  $21$  LT





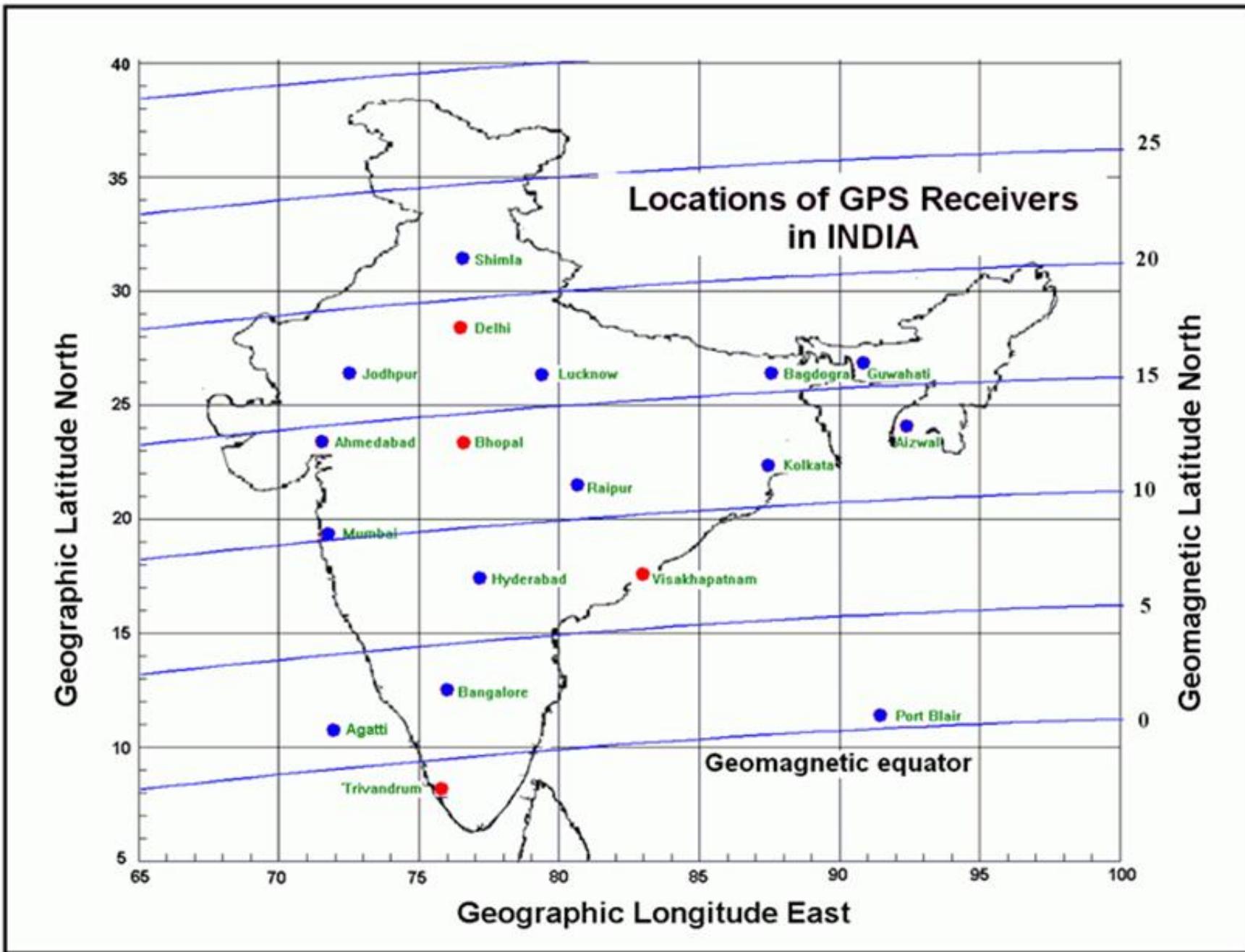
Westward  
Electric Field

Daytime ~ 9 LT to 21 LT  
Daytime ~ 9 LT to 21 LT



# **Morphological features of GPS-TEC in the Indian equatorial and low latitudes**

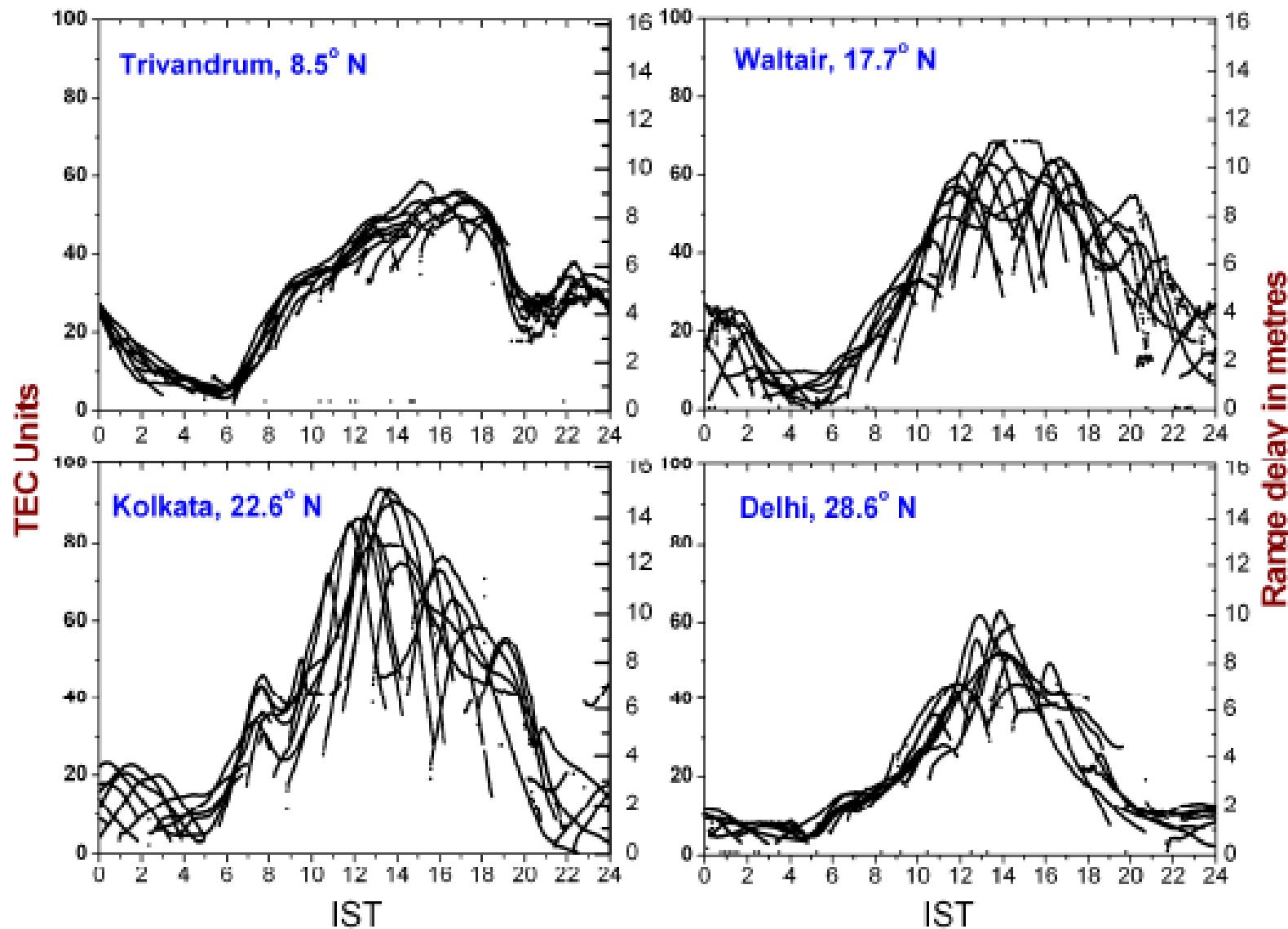




Locations of GPS Receivers in INDIA



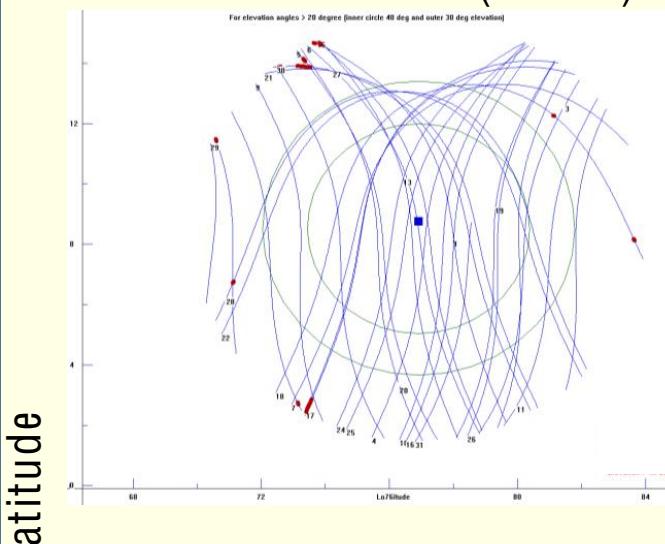
2<sup>nd</sup> APRIL 2004



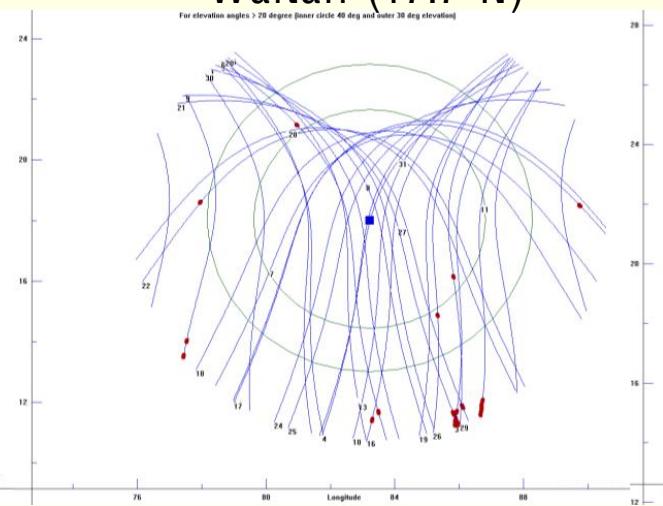
*The diurnal variation of vertical TEC/range delays measured using GPS receivers from four different stations near the equator (Trivandrum), the low latitude (Waltair), the anomaly crest region (Kolkata) and beyond the crest region (Delhi)*



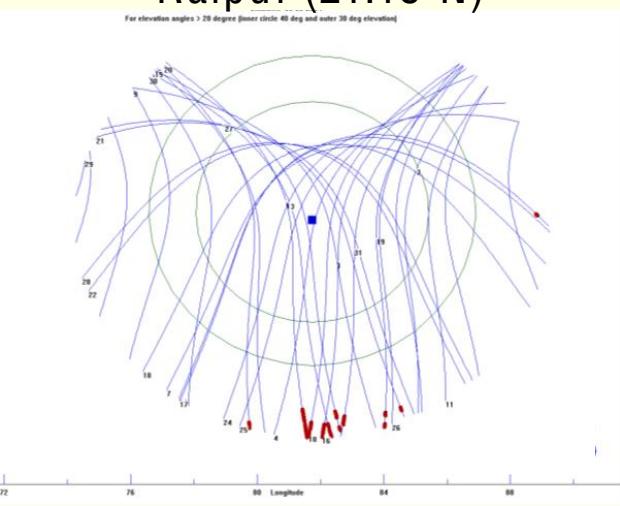
### Trivandrum ( $8.47^{\circ}\text{N}$ )



### Waltair ( $17.7^{\circ}\text{N}$ )

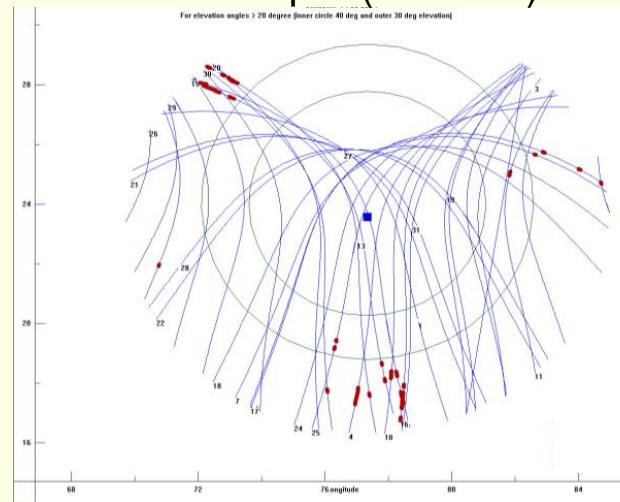


### Raipur ( $21.18^{\circ}\text{N}$ )

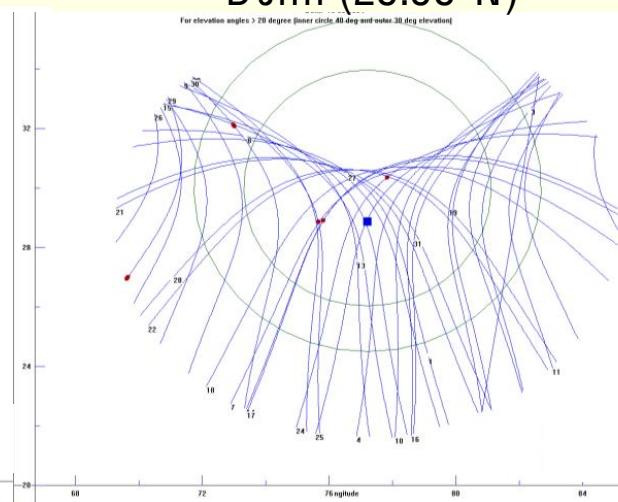


Latitude

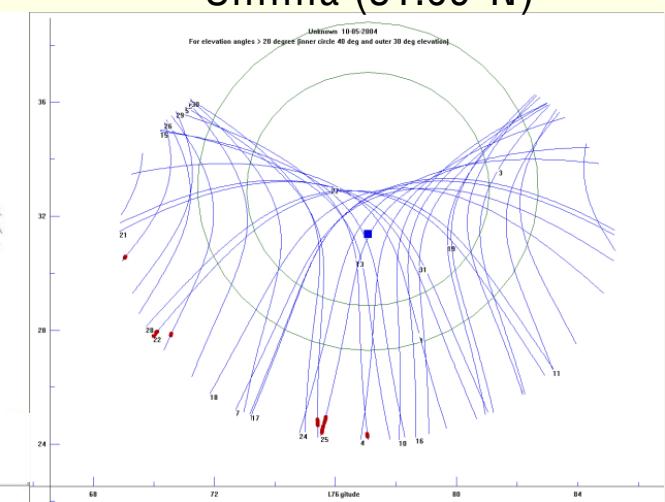
### Bhopal ( $23.28^{\circ}\text{N}$ )



### Delhi ( $28.58^{\circ}\text{N}$ )

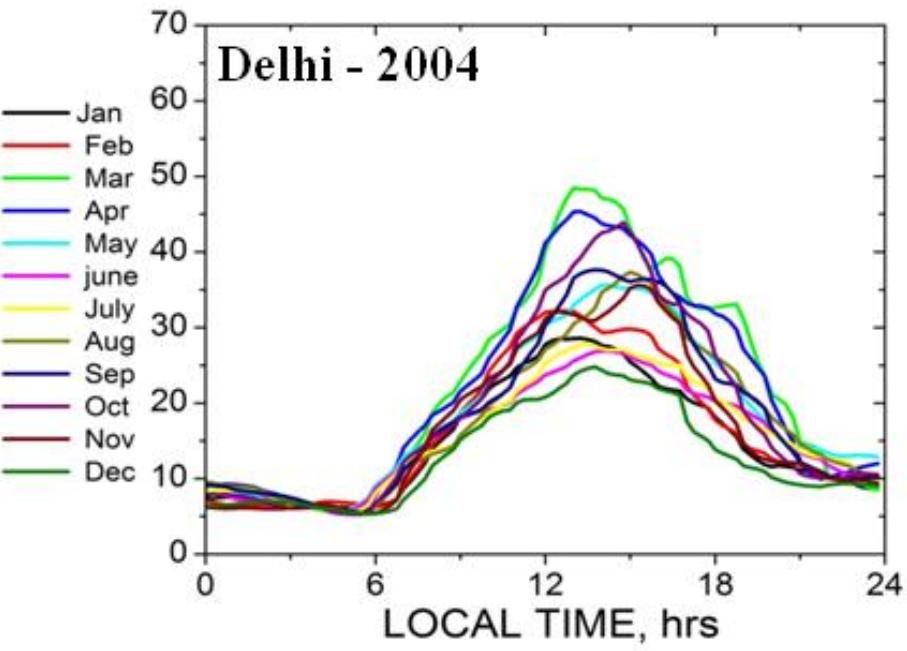
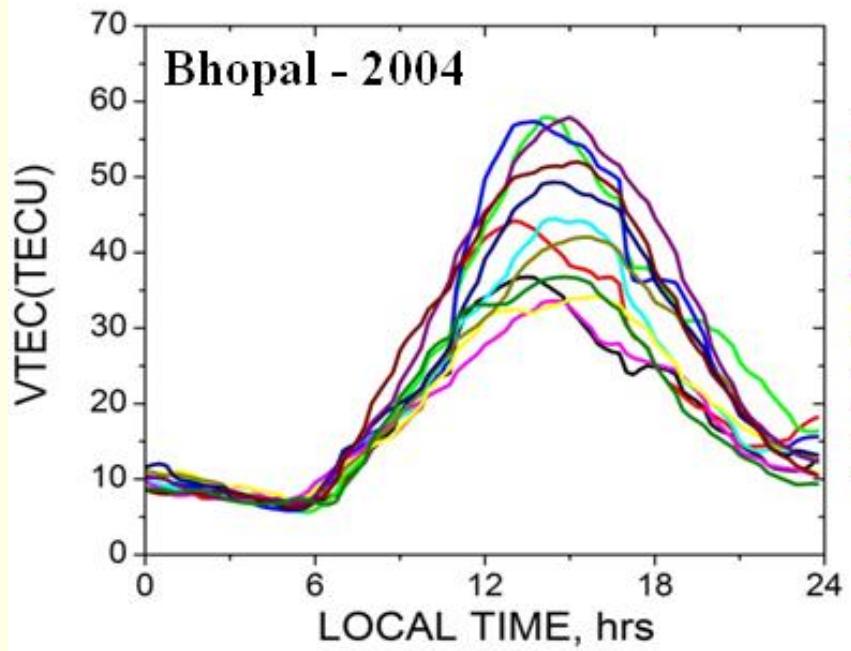
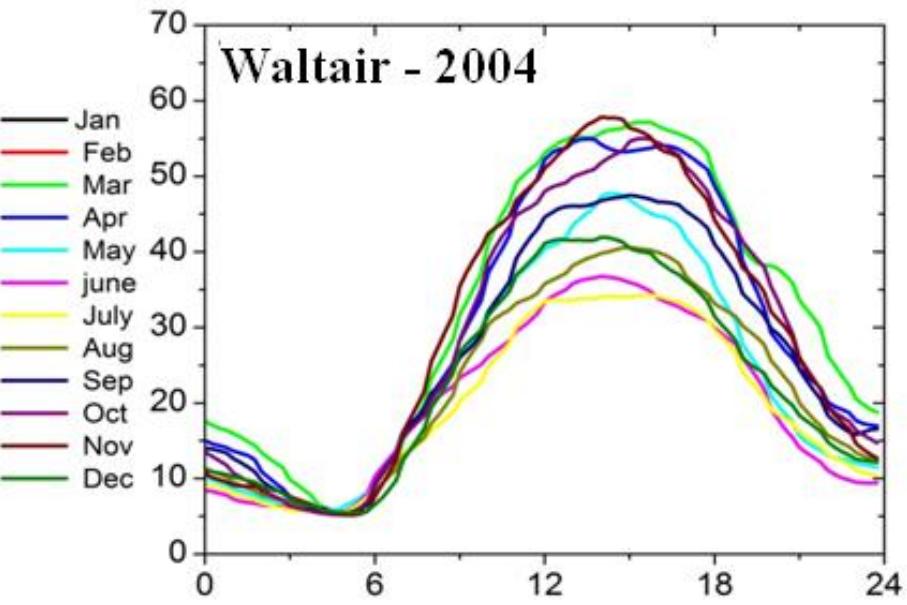
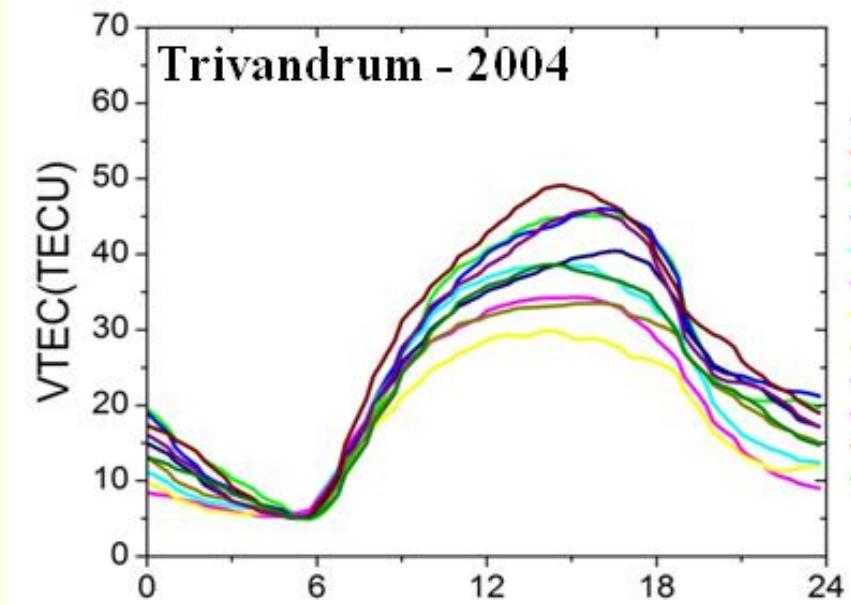


### Shimla ( $31.09^{\circ}\text{N}$ )



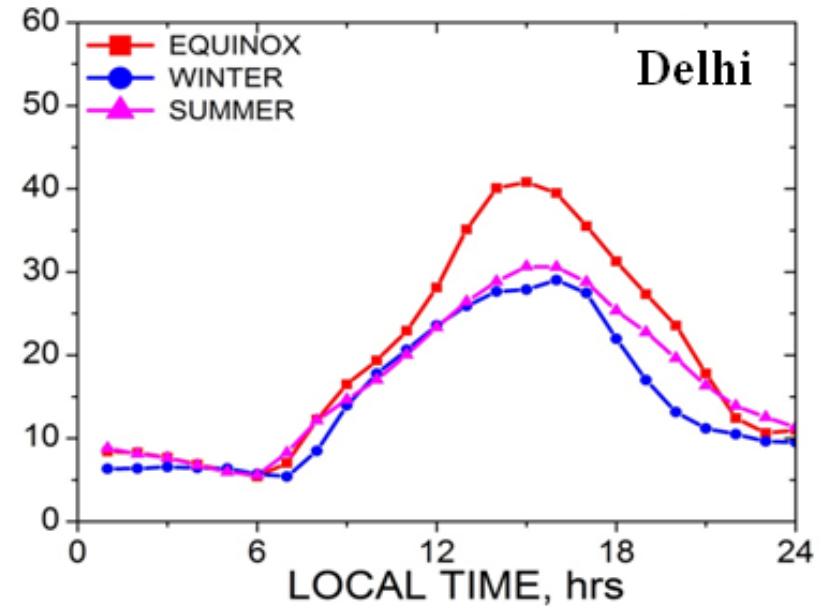
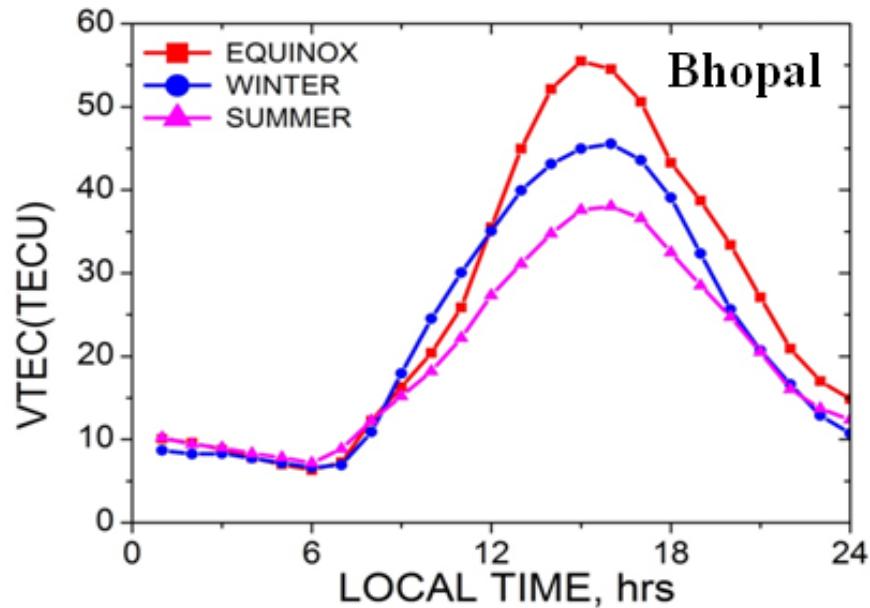
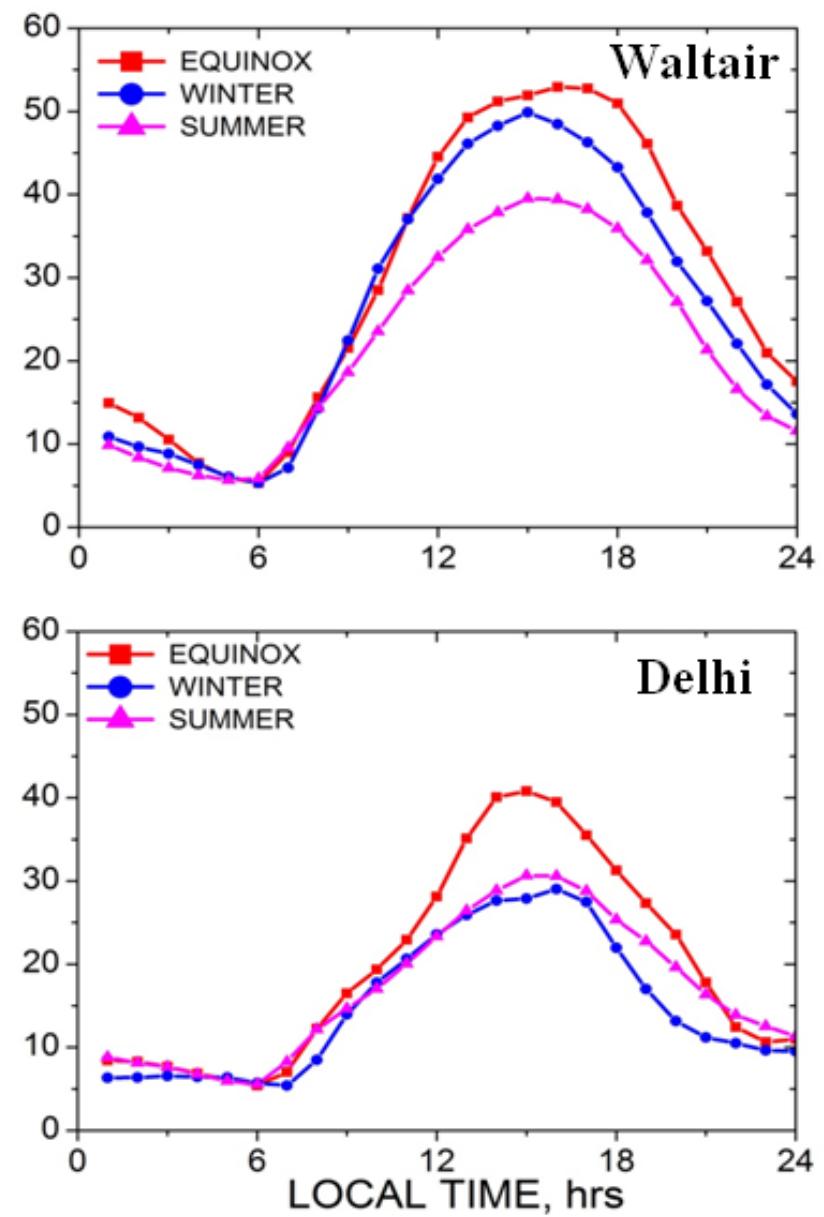
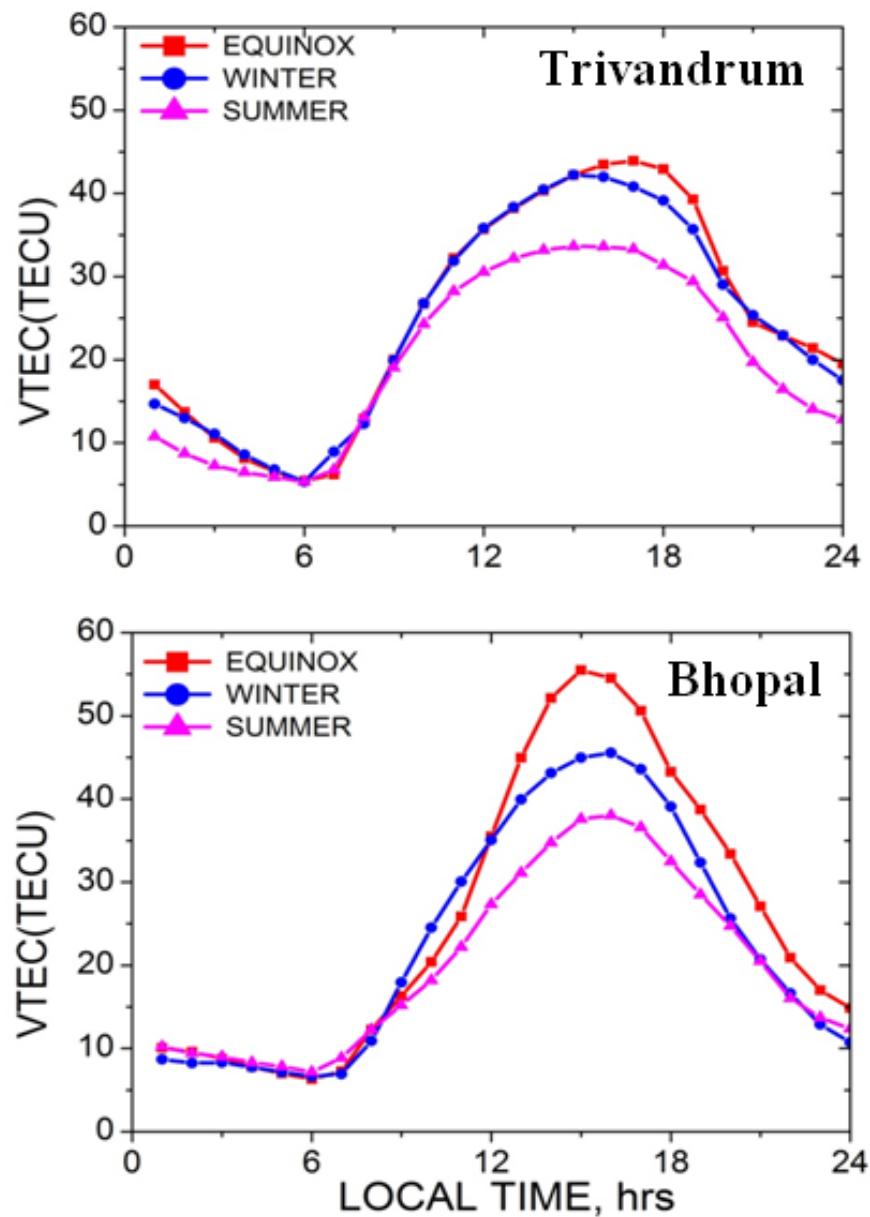
Longitude





*Monthly mean diurnal variation of the TEC over four different stations, Trivandrum, Waltair, Bhopal and Delhi during the period from January 2004 to December 2004*

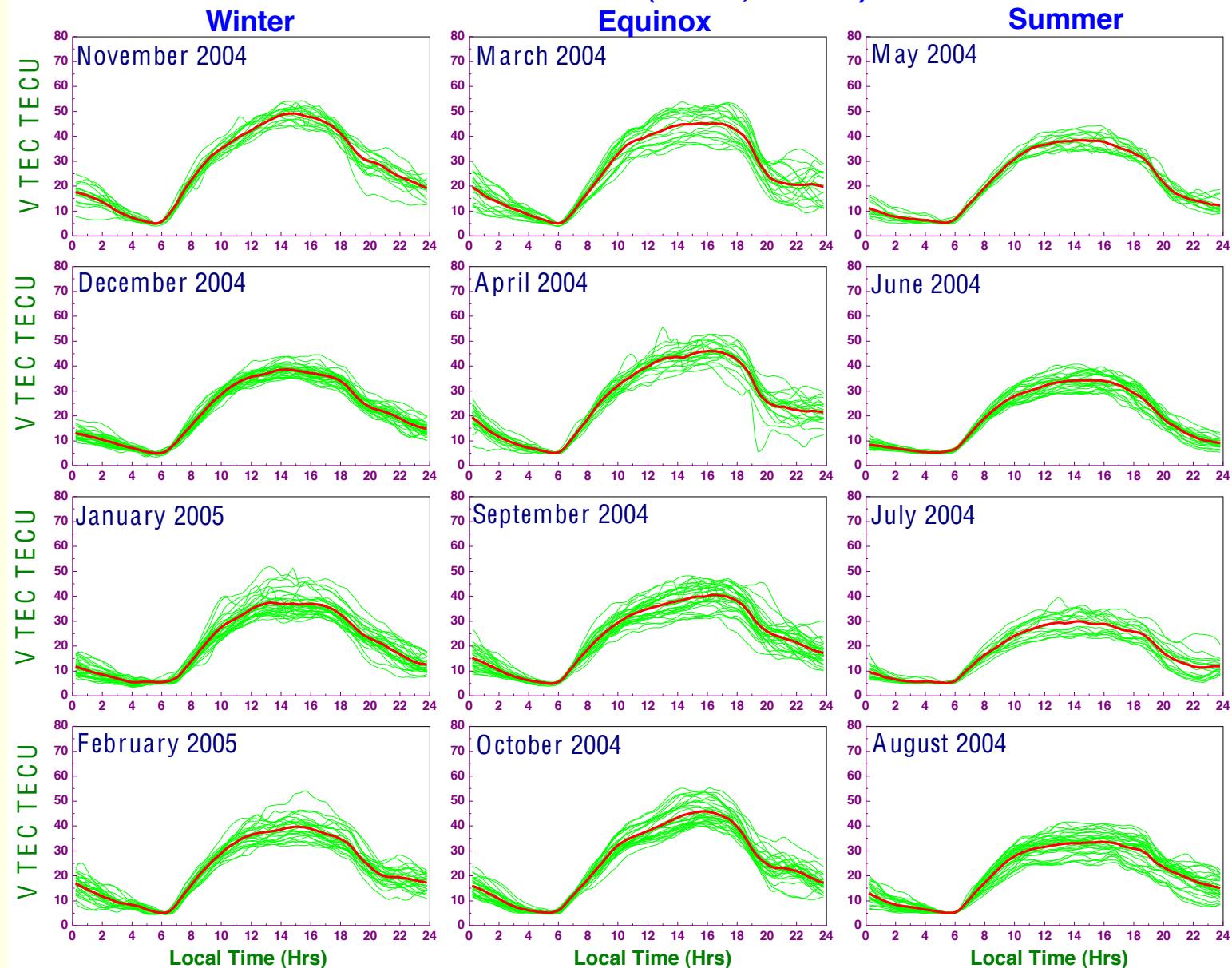




*The mean diurnal variation of TEC over four different stations, Trivandrum, Waltair, Bhopal, Delhi for three different seasons equinox, winter and summer during the period from January 2004 to December 2004.*

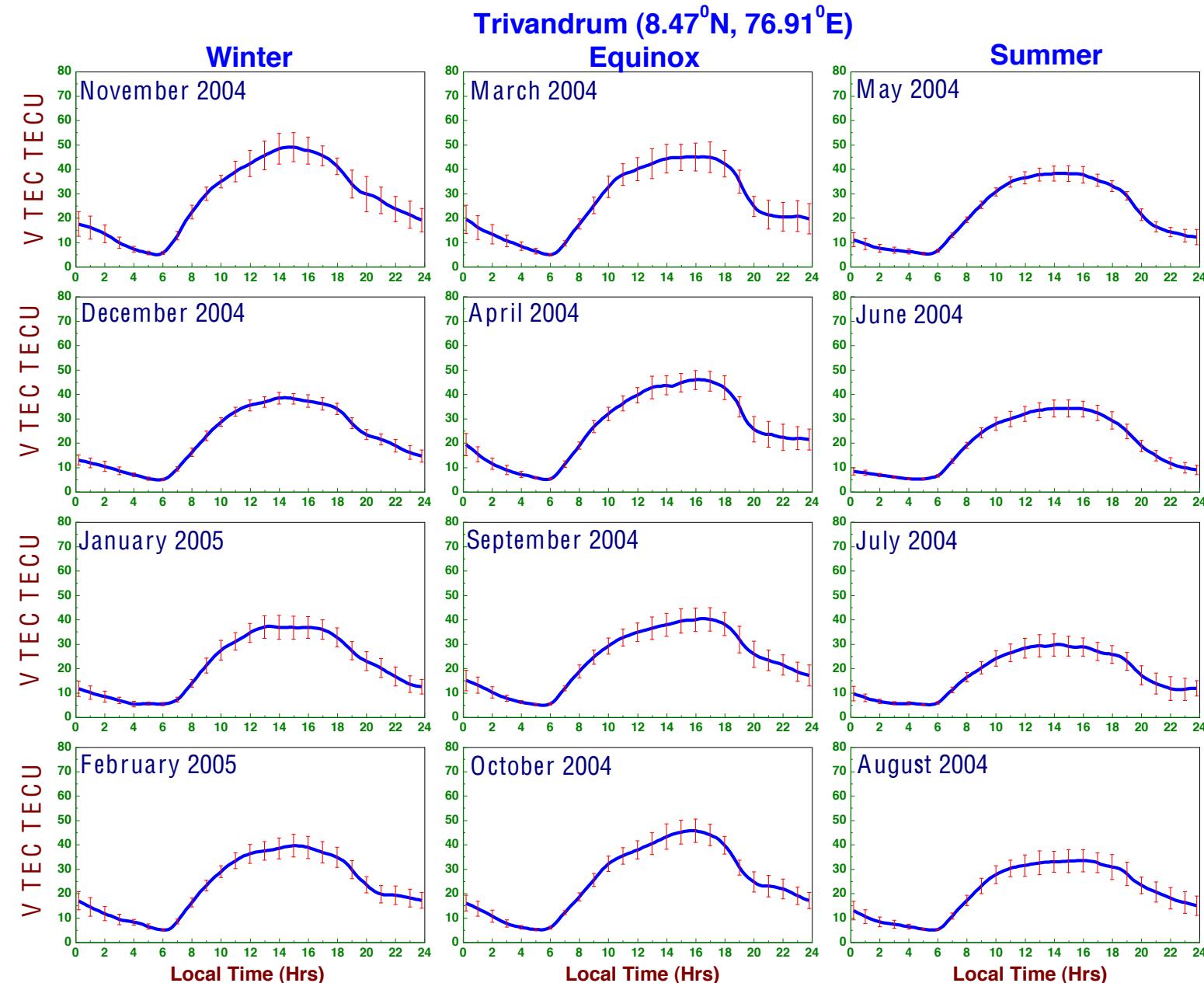


**Trivandrum ( $8.47^{\circ}\text{N}$ ,  $76.91^{\circ}\text{E}$ )**



*Mass plots of TEC for each month (showing the typical day to day variability) measured from the GPS signals at the equatorial station Trivandrum.*

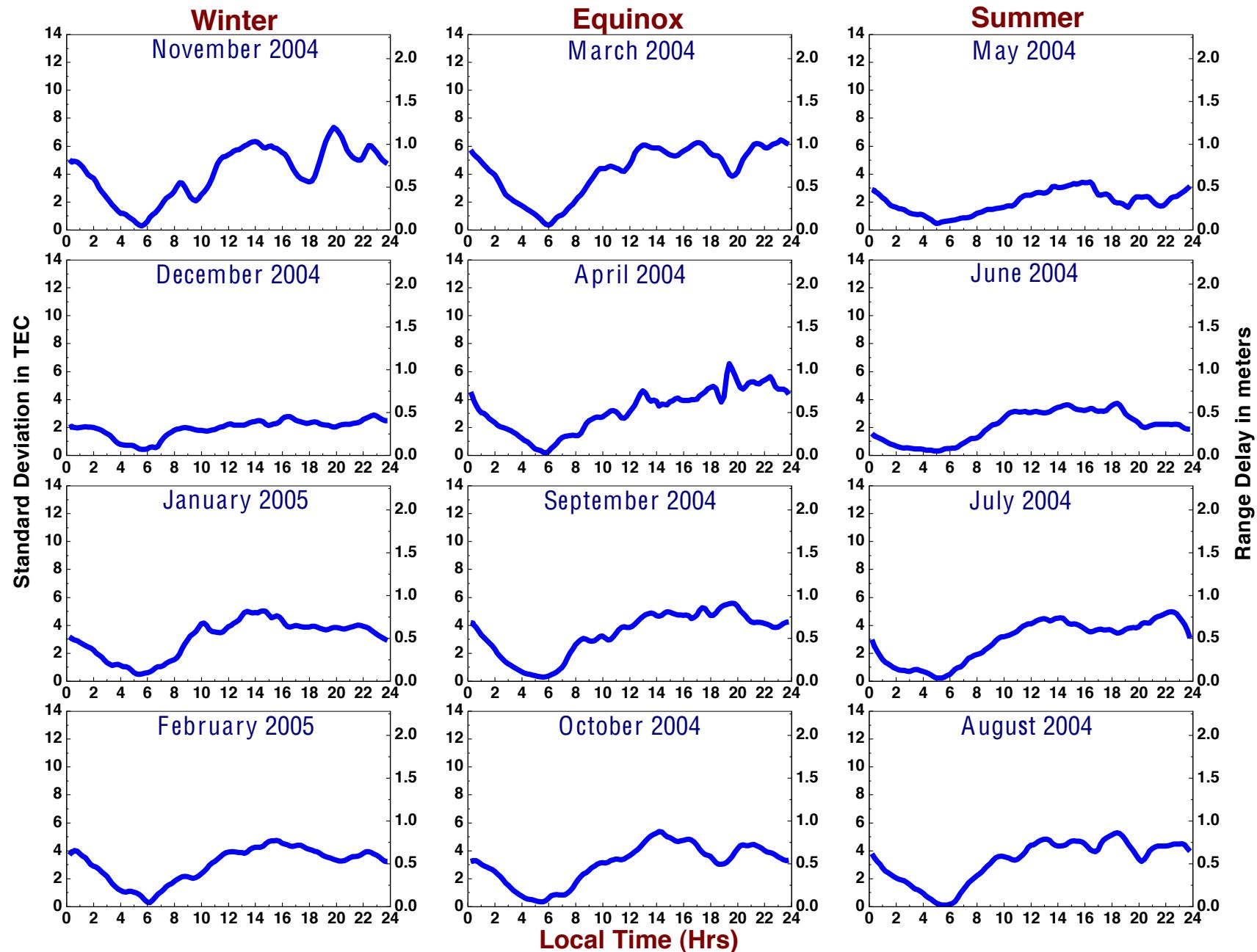




*Average diurnal variation plots showing the standard deviations of the day-to-day variability in TEC over Trivandrum*

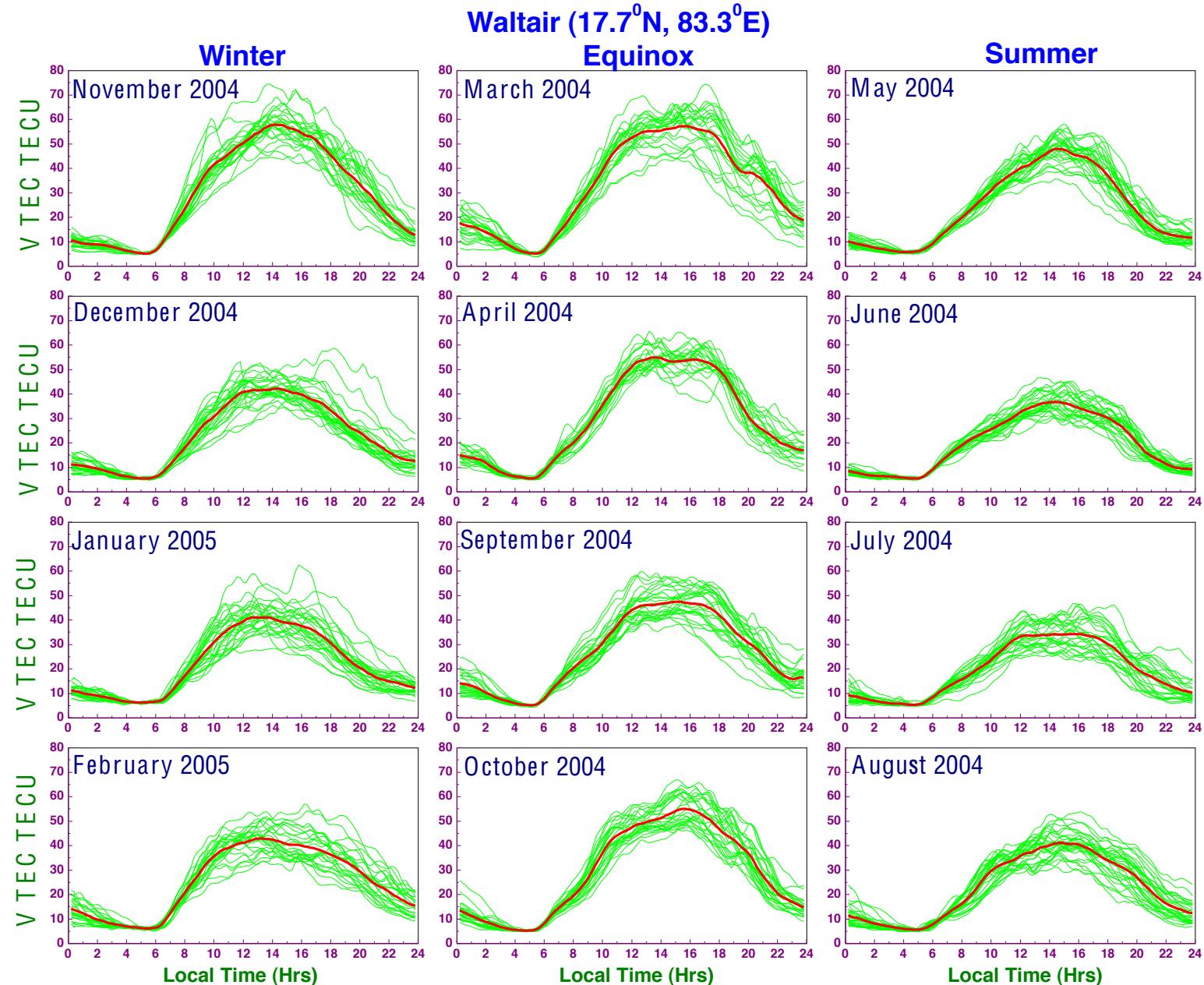


Trivandrum ( $8.47^{\circ}\text{N}$ ,  $76.91^{\circ}\text{E}$ )



*Monthly average diurnal variation of standard deviations in TEC over Trivandrum*

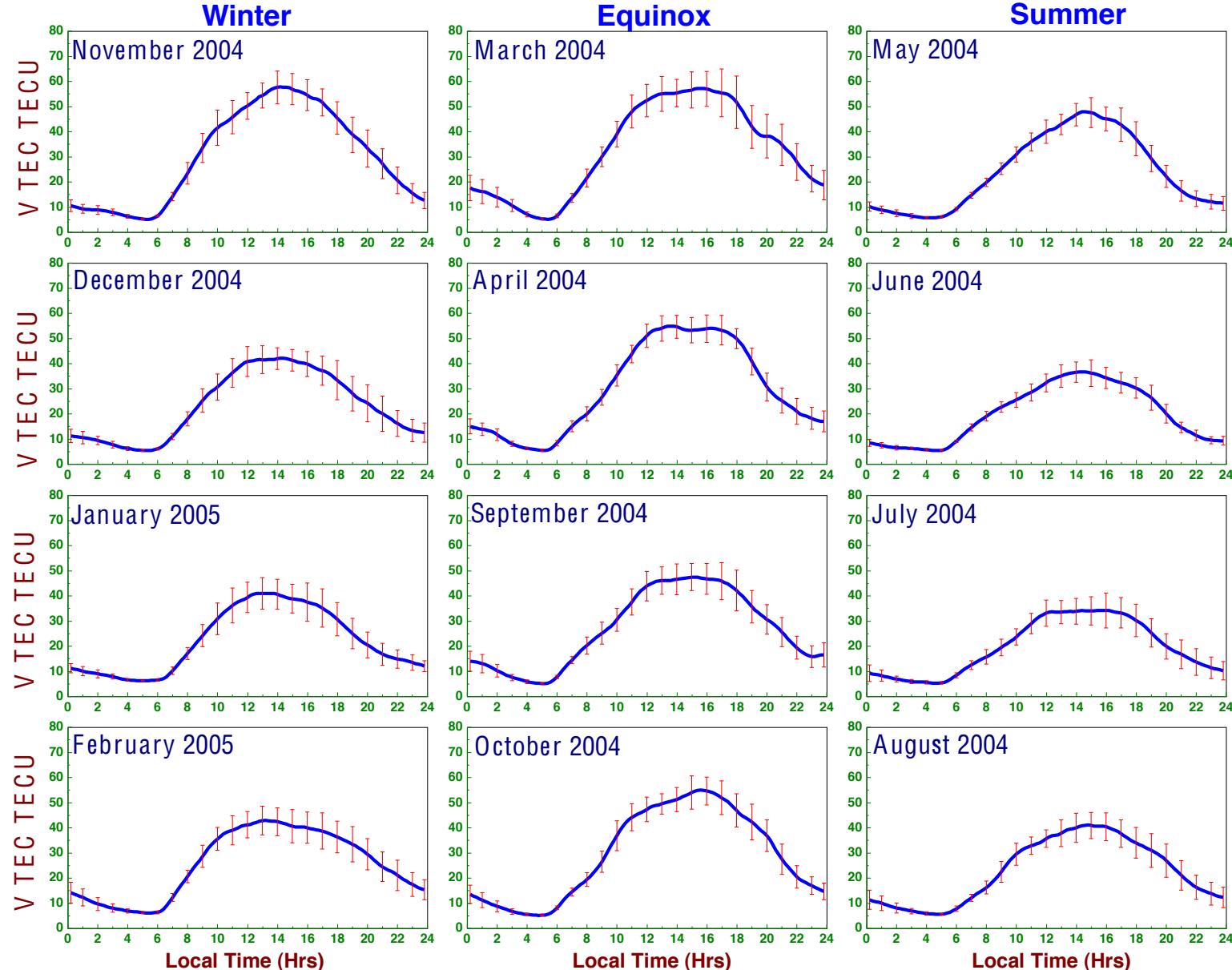




*Mass plots of TEC for each month (showing the typical day to day variability) measured from the GPS signals at Waltair, located at the inner edge of the EIA*



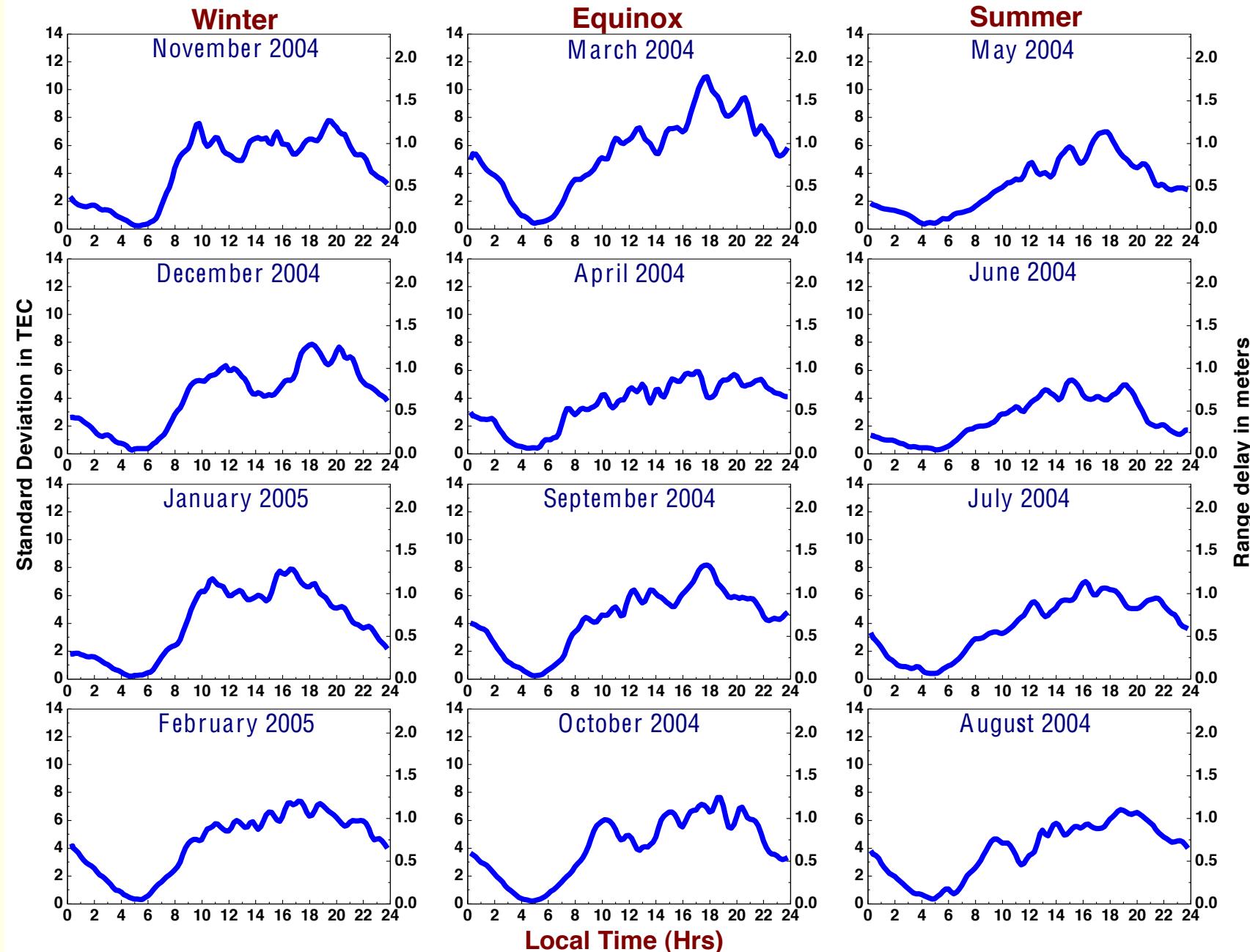
Waltair ( $17.7^{\circ}\text{N}$ ,  $83.3^{\circ}\text{E}$ )



*Average diurnal variation plots showing the standard deviations of the day-to-day variability in TEC over Waltair*

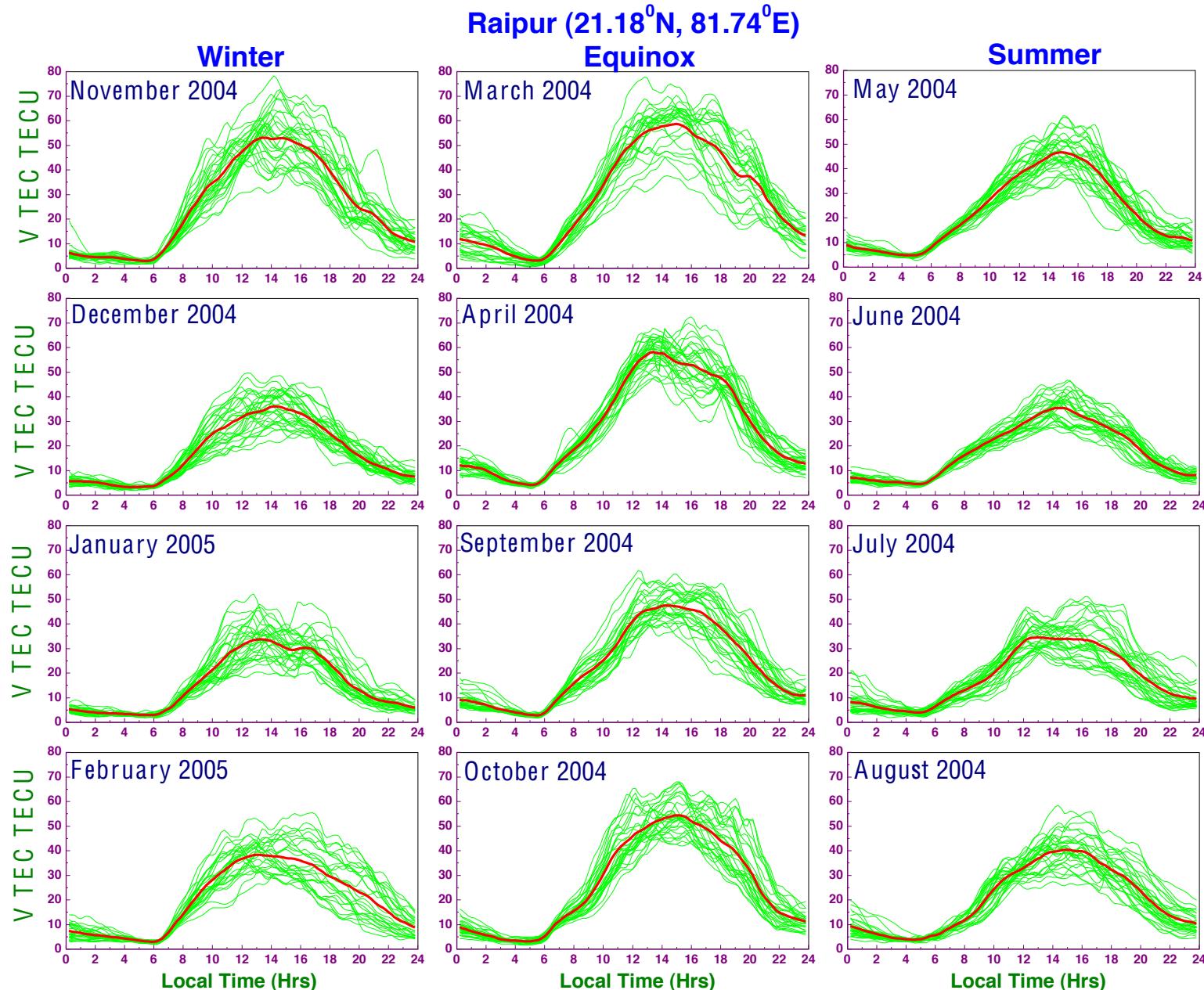


Waltair ( $17.7^{\circ}\text{N}$ ,  $83.3^{\circ}\text{E}$ )



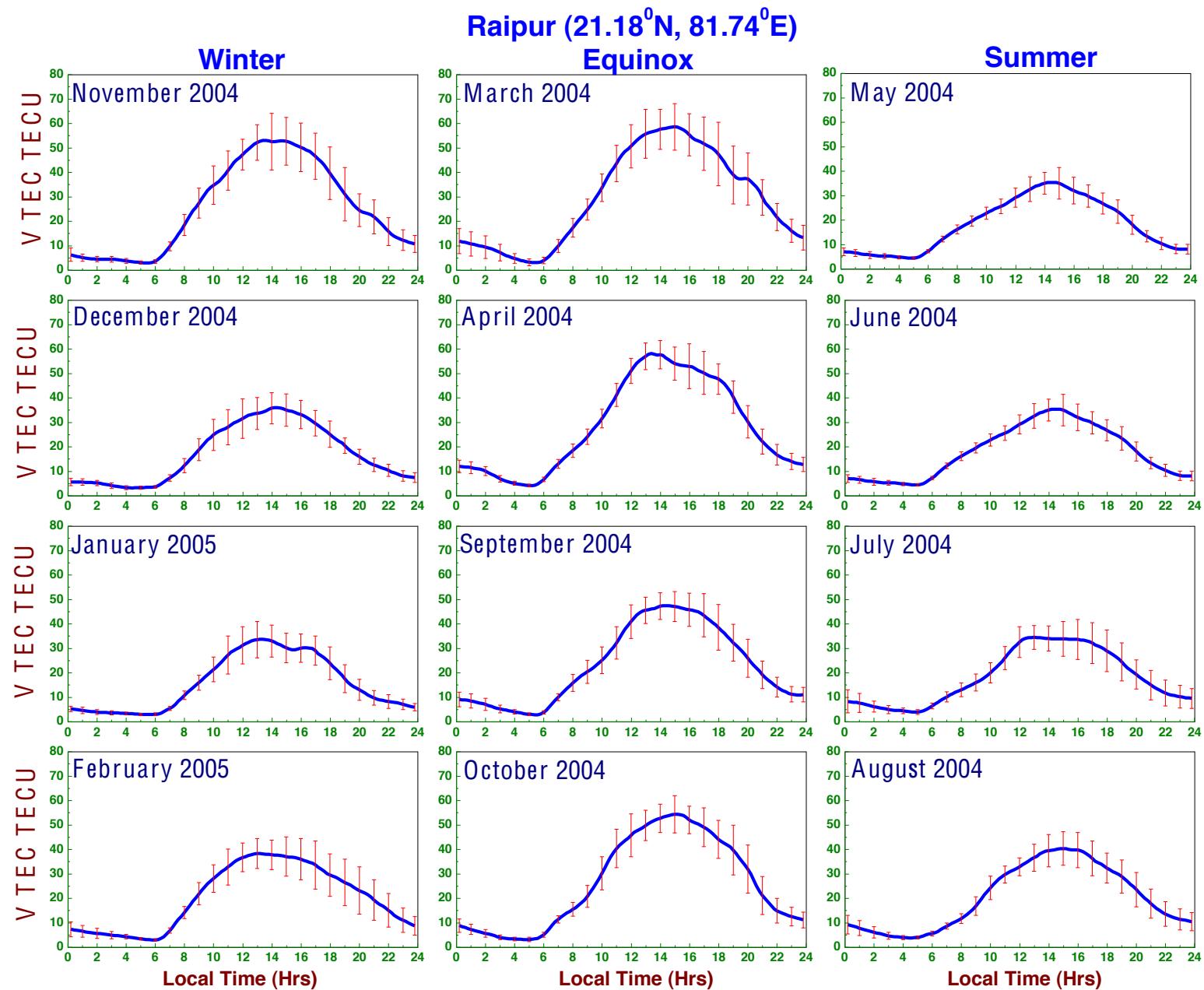
Monthly average diurnal variation of standard deviations in TEC over Waltair





*Mass plots of TEC for each month (showing the typical day to day variability)  
measured from the GPS signals at Raipur, located at the crest of the EIA*

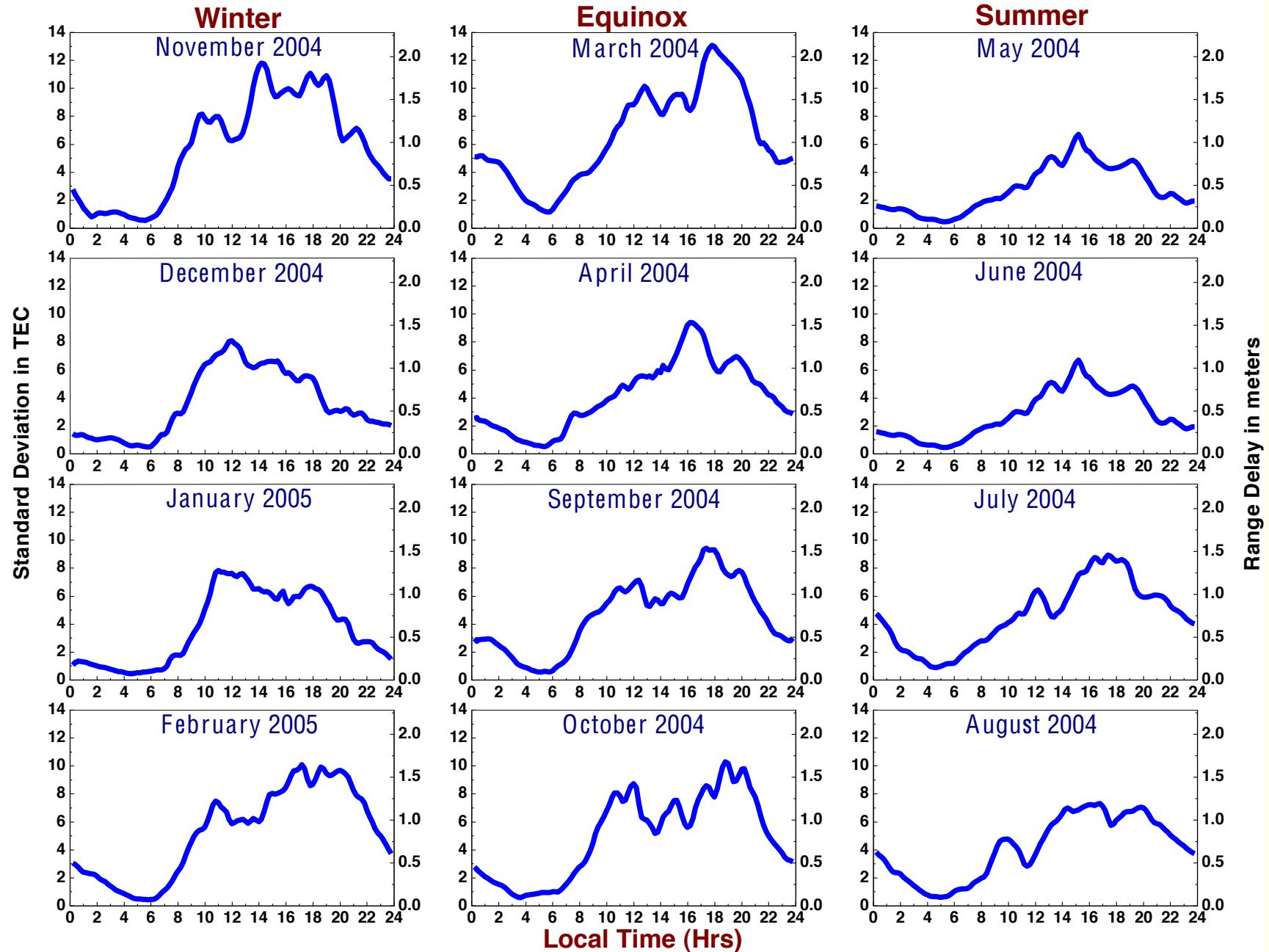




*Average diurnal variation plots showing the standard deviations of the day-to-day variability in TEC over Raipur*



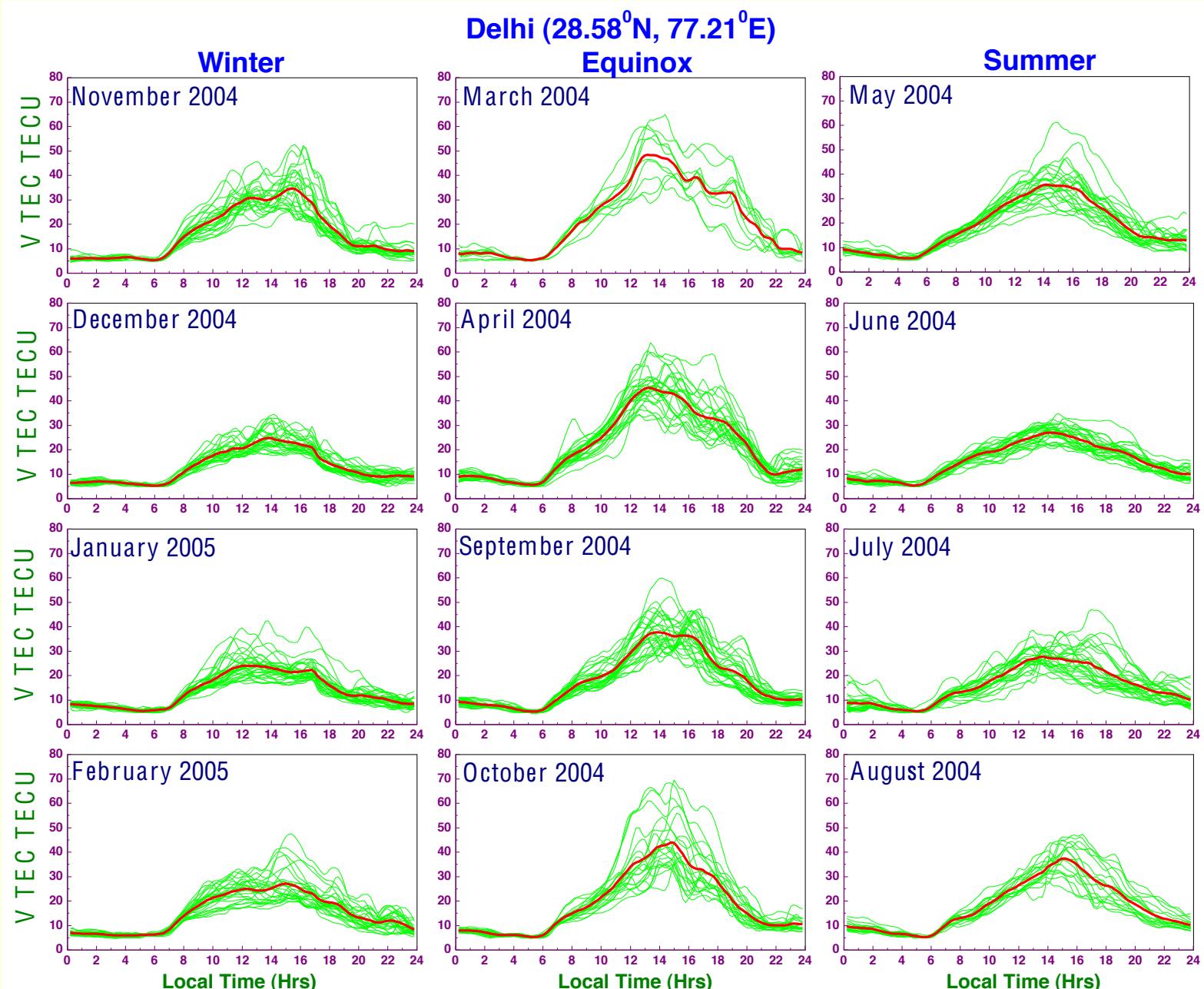
Raipur ( $21.18^{\circ}\text{N}$ ,  $81.74^{\circ}\text{E}$ )



*Monthly average diurnal variation of standard deviations in TEC over Raipur*

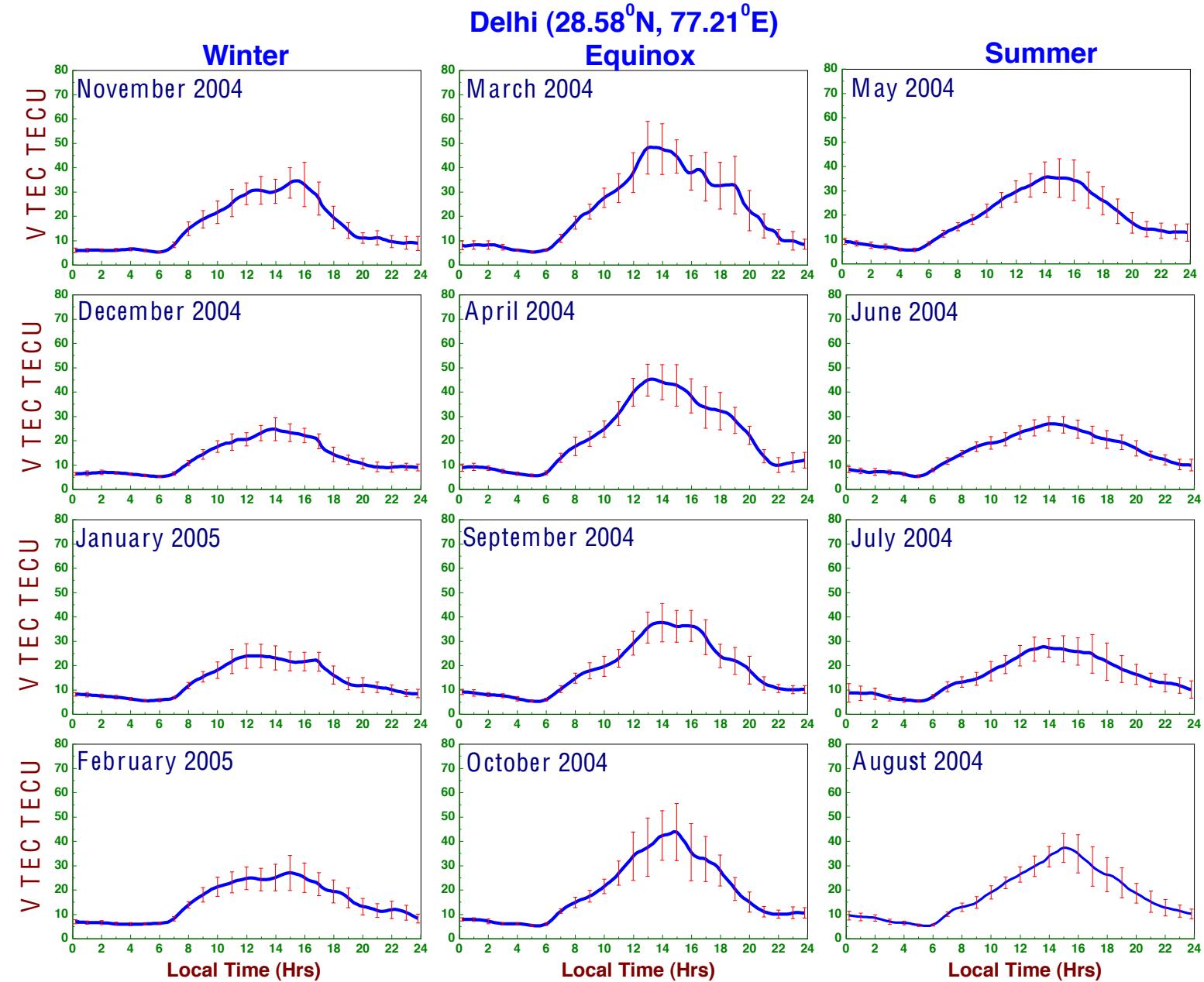
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*Mass plots of TEC for each month (showing the typical day to day variability) measured from the GPS signals at Delhi, located at the outer edge of the EIA*

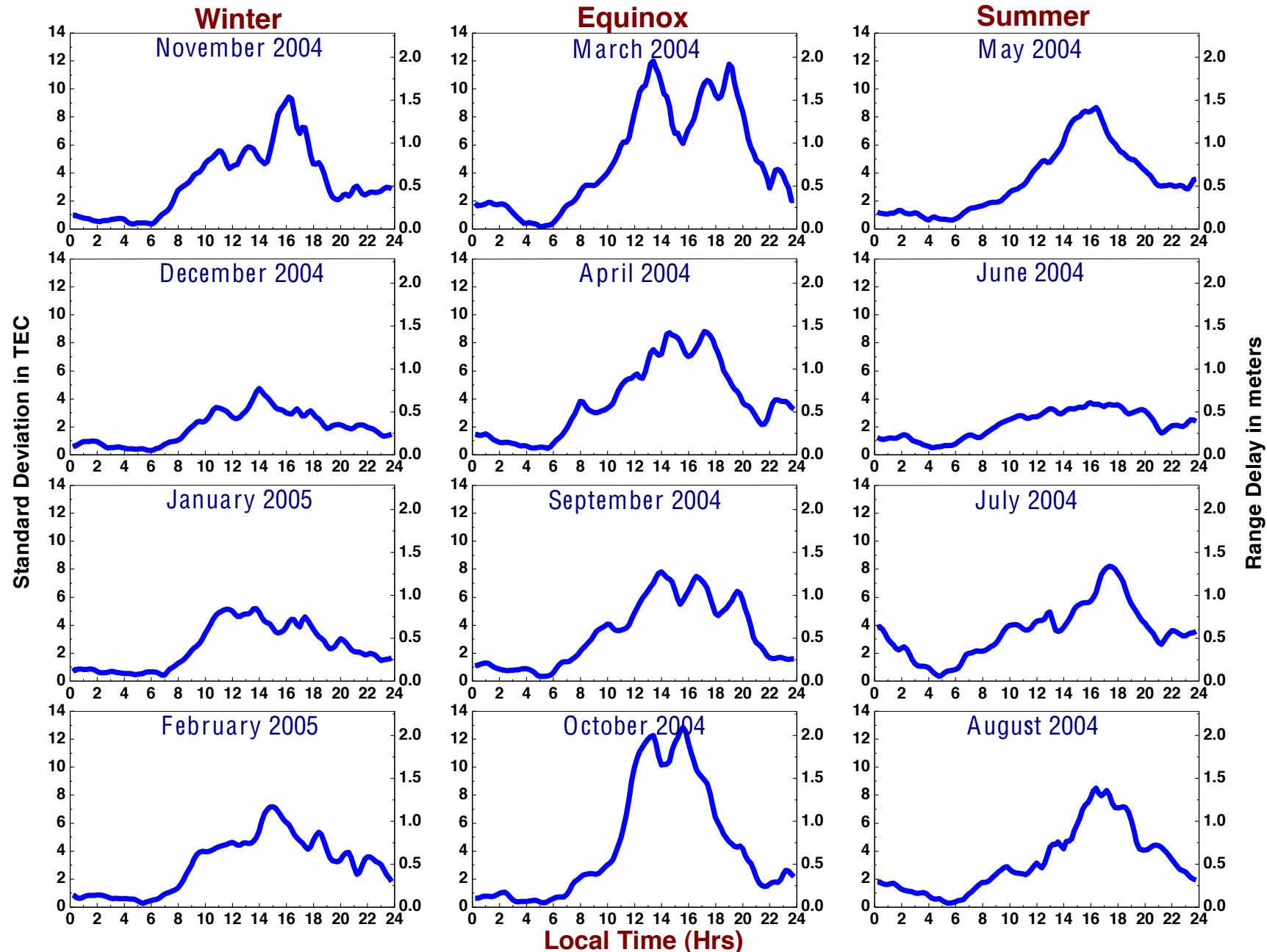




*Average diurnal variation plots showing the standard deviations of the day-to-day variability in TEC over Delhi*



Delhi ( $28.58^{\circ}\text{N}$ ,  $77.21^{\circ}\text{E}$ )

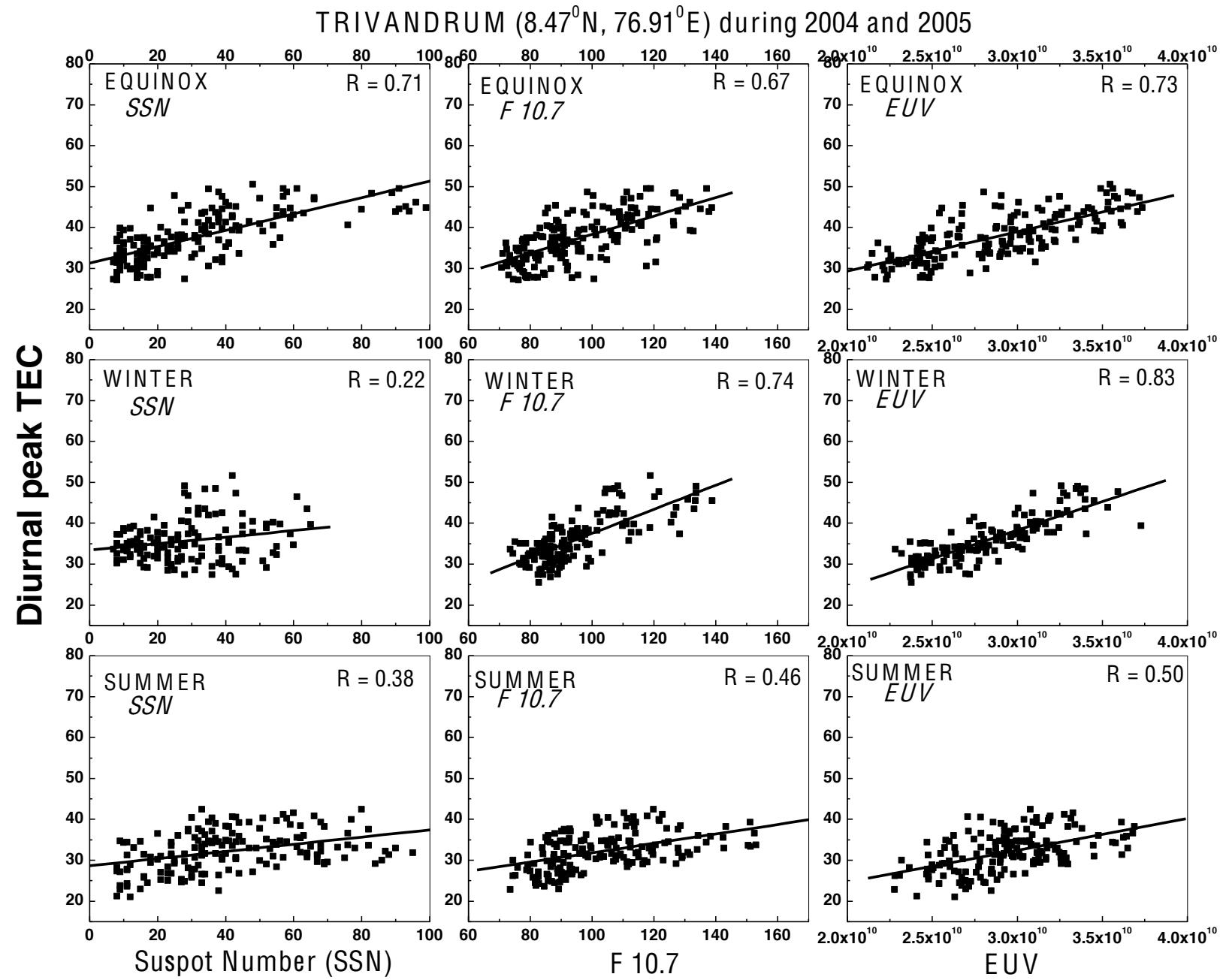


Monthly average diurnal variation of standard deviations in TEC over Delhi



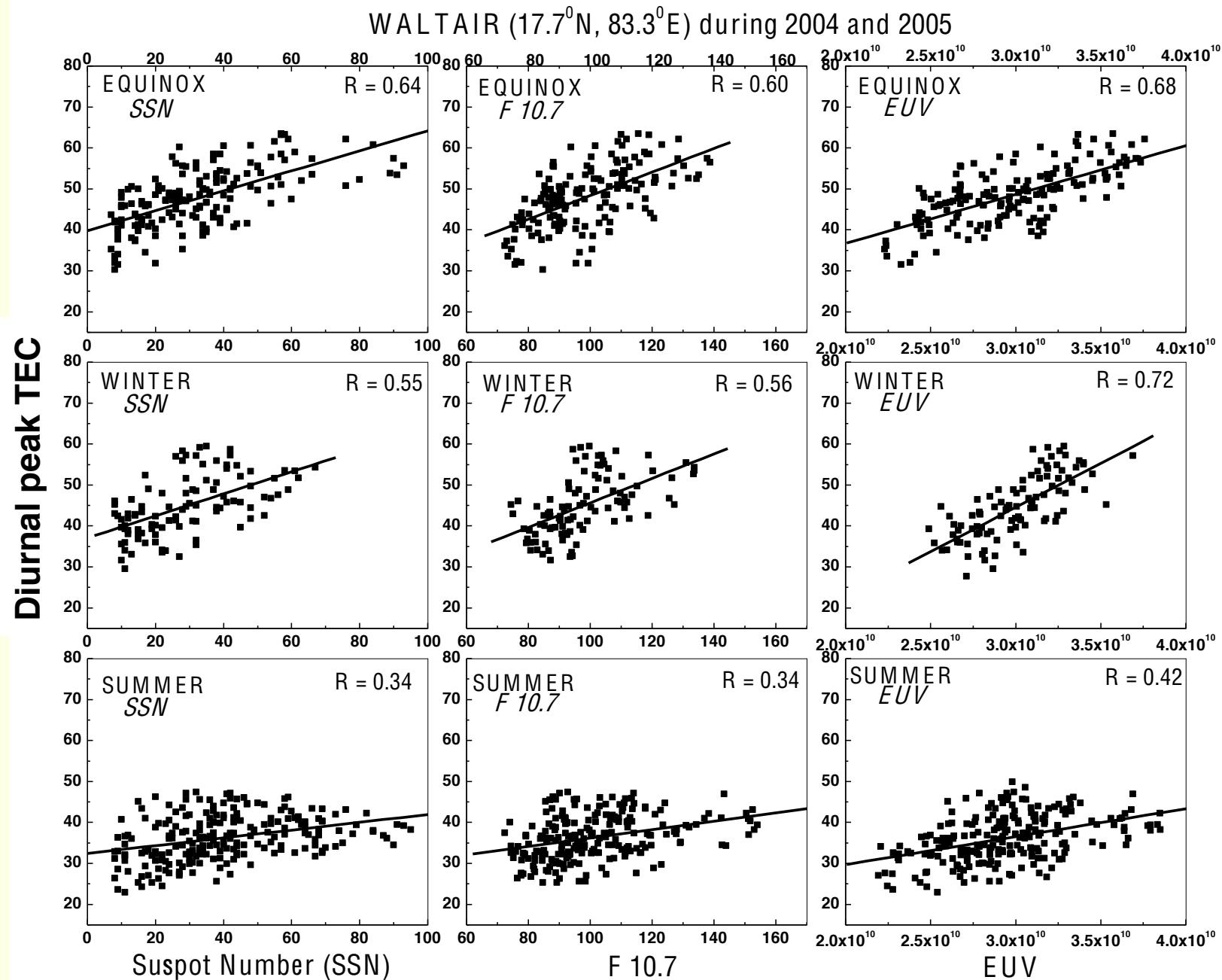
# **Variation of durnal peak Total Electron Content with solar activity indices**





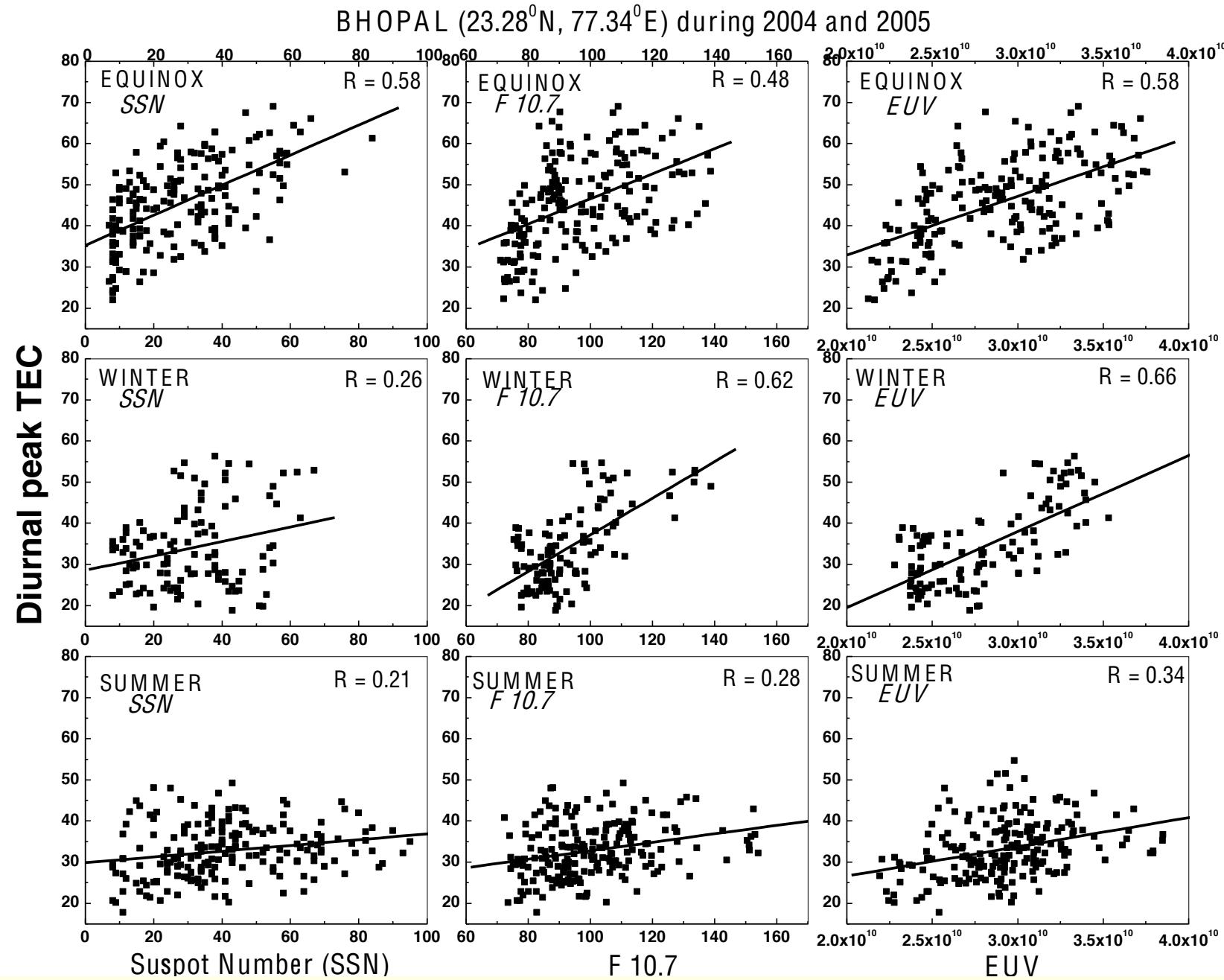
*Plots showing the correlation between TEC and SSN, F10.7, and EUV for three different seasons over Trivandrum during the period from 2004 to 2005.*





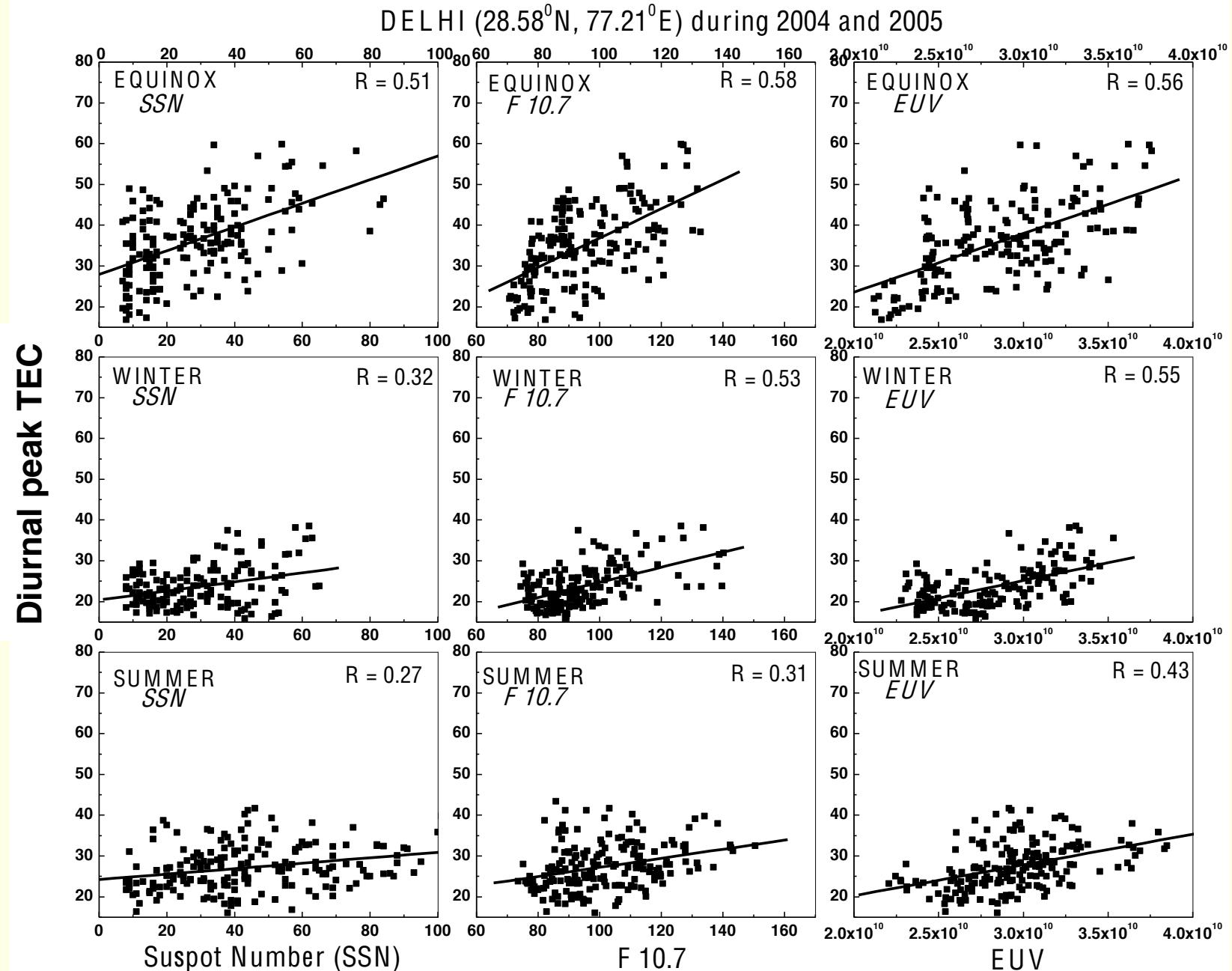
*Plots showing the correlation between TEC and SSN, F10.7, and EUV for three different seasons over Waltair during the period from 2004 to 2005.*





*Plots showing the correlation between TEC and SSN, F10.7, and EUV for three different seasons over Bhopal during the period from 2004 to 2005.*



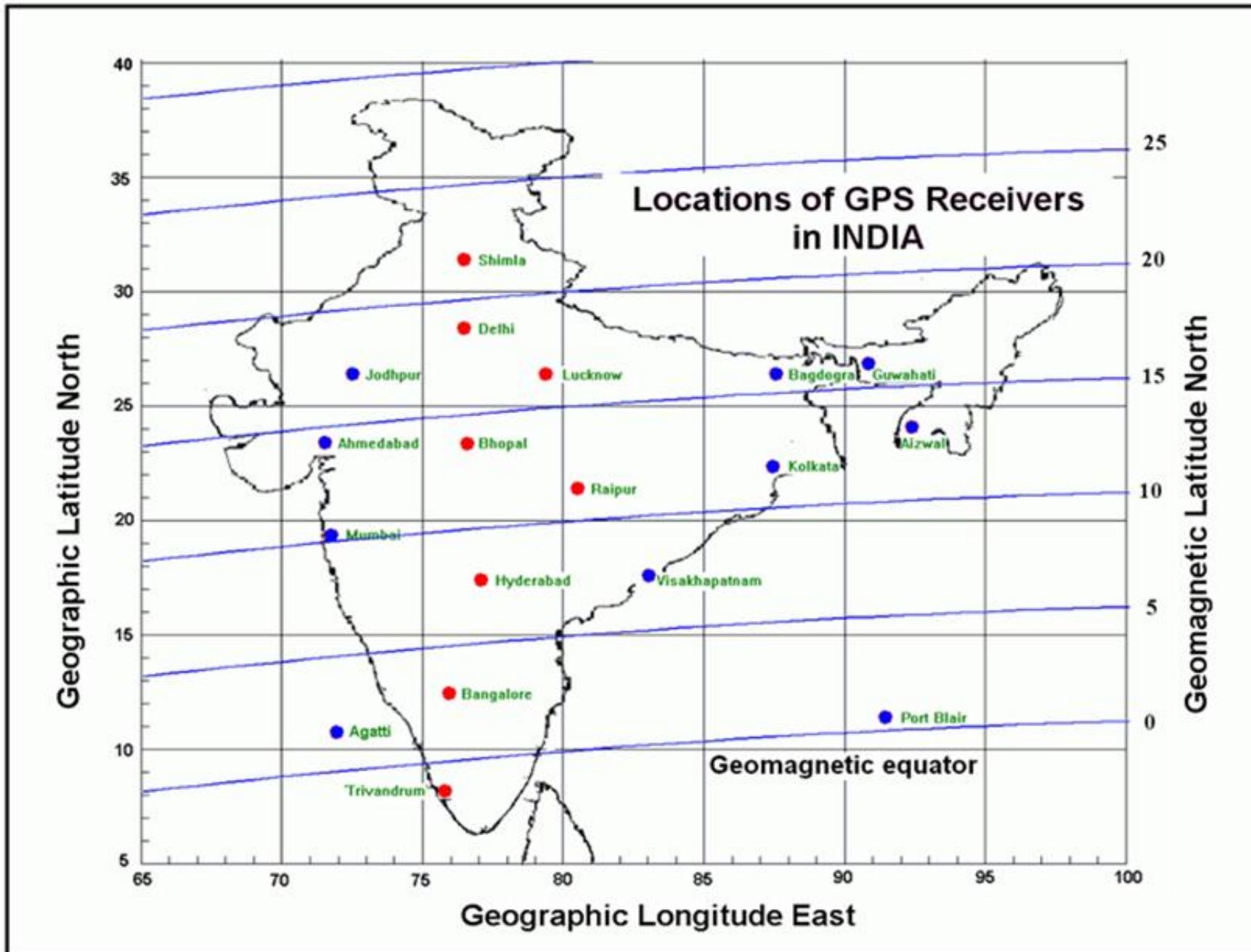


*Plots showing the correlation between TEC and SSN, F10.7, and EUV for three different seasons over Delhi during the period from 2004 to 2005.*



# **Equatorial Ionization Anomaly and its association with Integrated Equatorial Electrojet Strength**

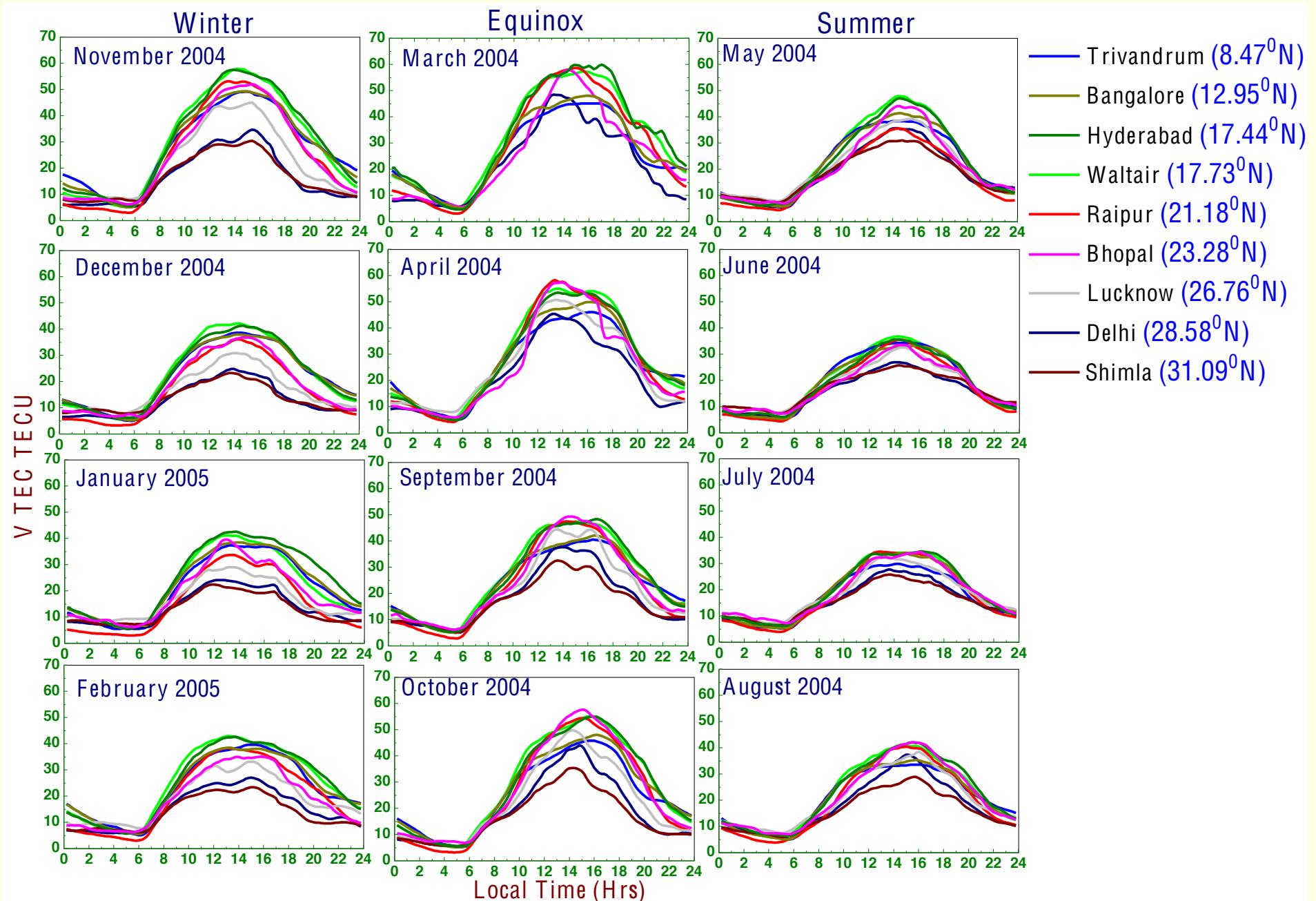




*Locations of GPS Receivers in India*

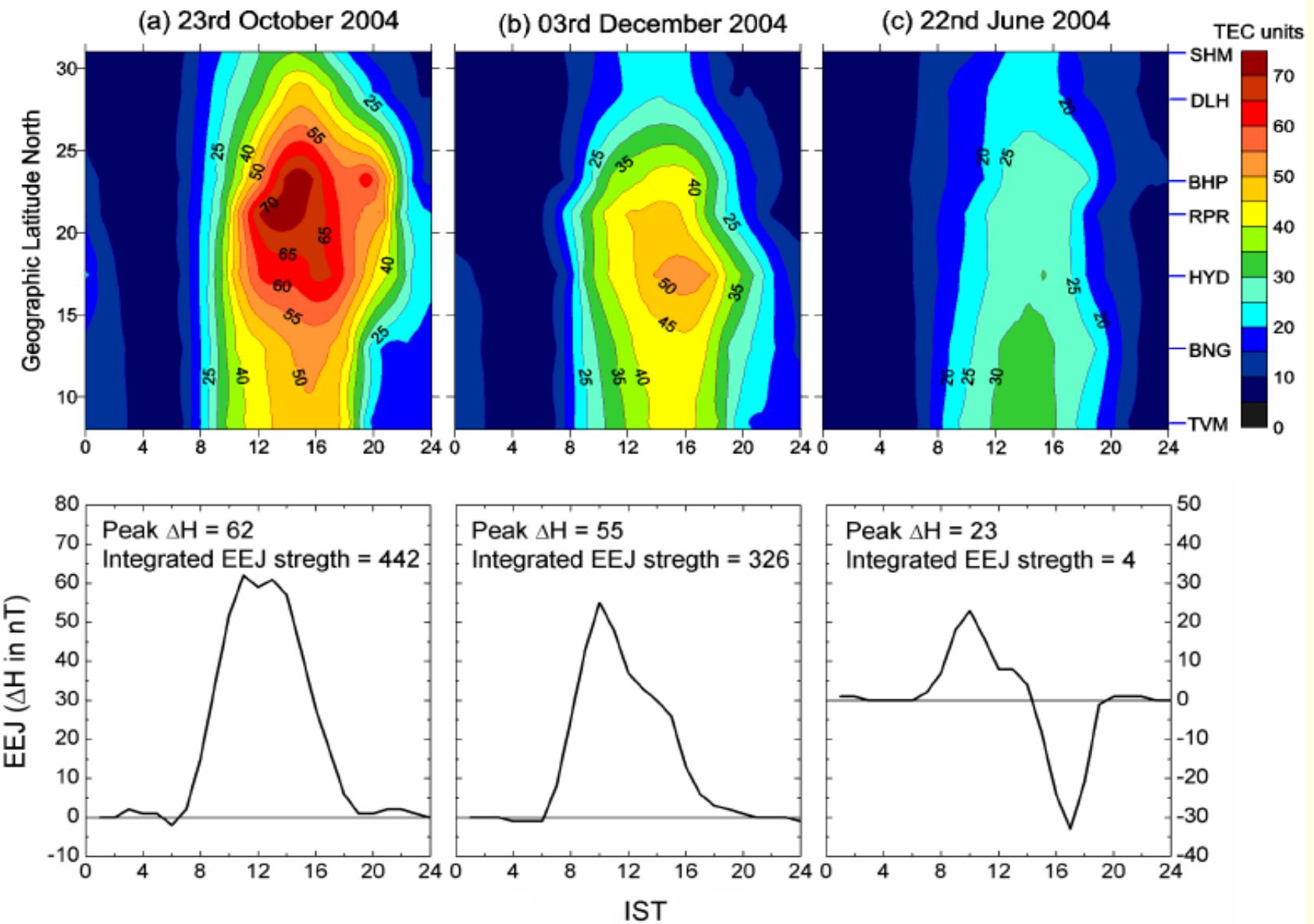
Space Physics Laboratory, Andhra University, Visakhapatnam, INDIA





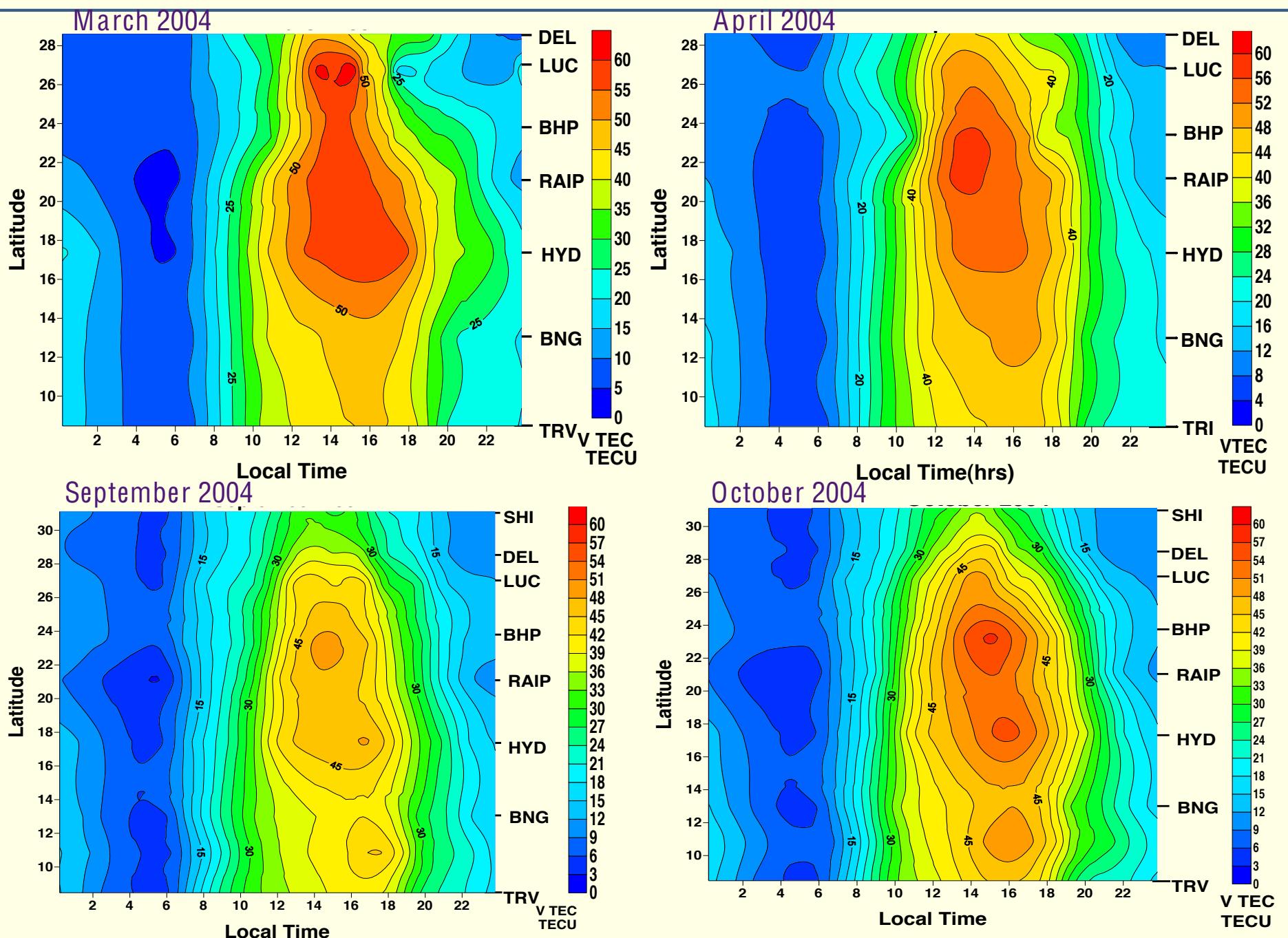
*Monthly average diurnal variations at 9 different stations along  
a common longitude of  $77^{\circ}\text{E}$*





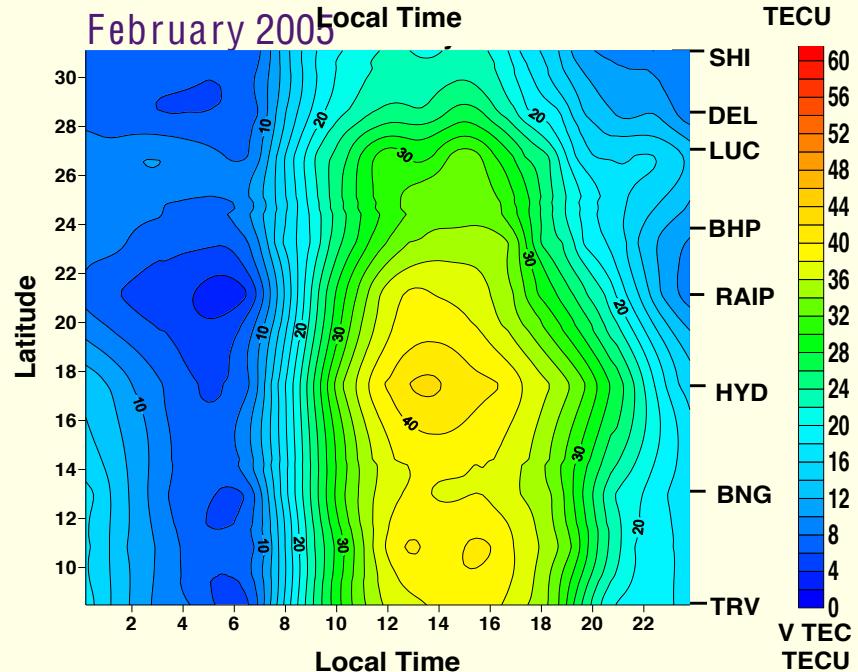
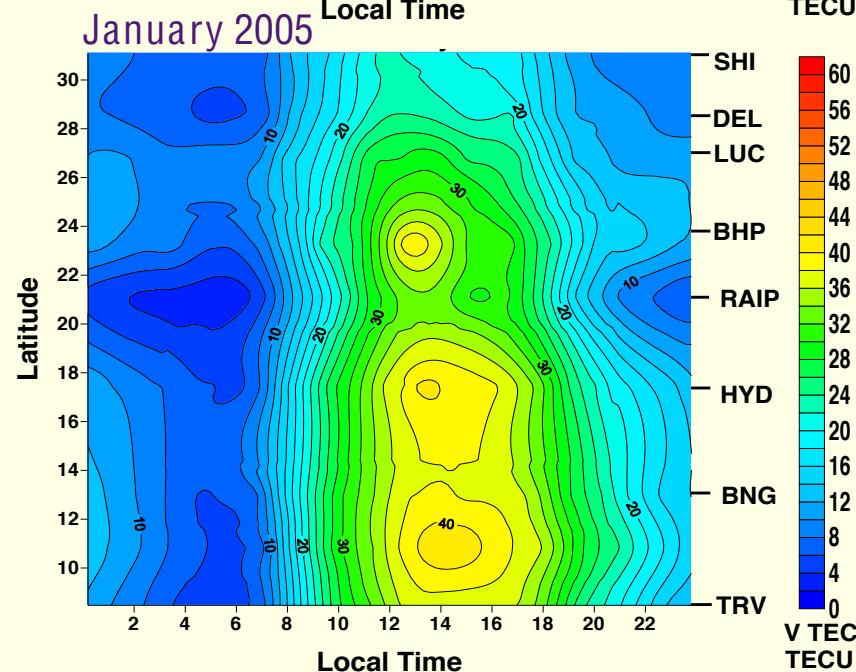
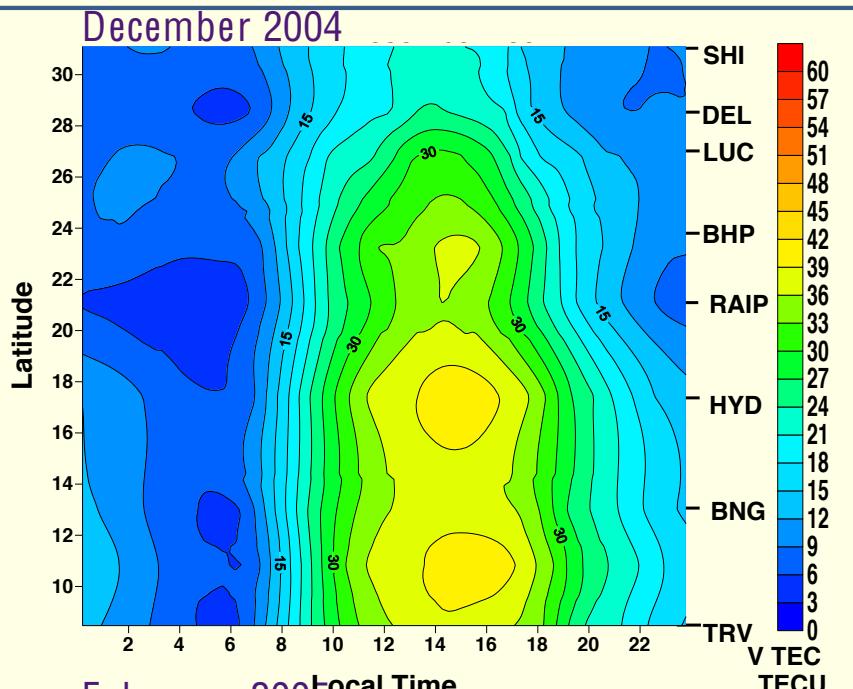
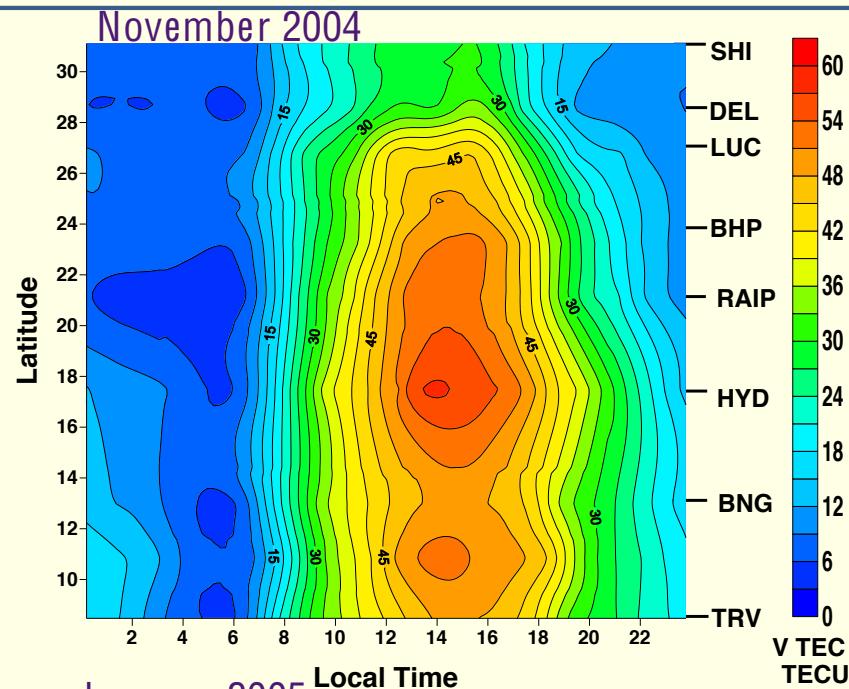
*Typical contour plots of the diurnal variations of TEC, drawn as a function of latitude for (a) an equinoctial day, (b) a winter day and (c) a summer day along with the corresponding diurnal variations of the equatorial electrojet strengths.*





Contour plots of the diurnal variation of TEC for 9 stations ( $8.47^{\circ}\text{N}$  to  $31^{\circ}\text{N}$ ) along the common longitude of  $77^{\circ}\text{E}$  during Equinoctial months

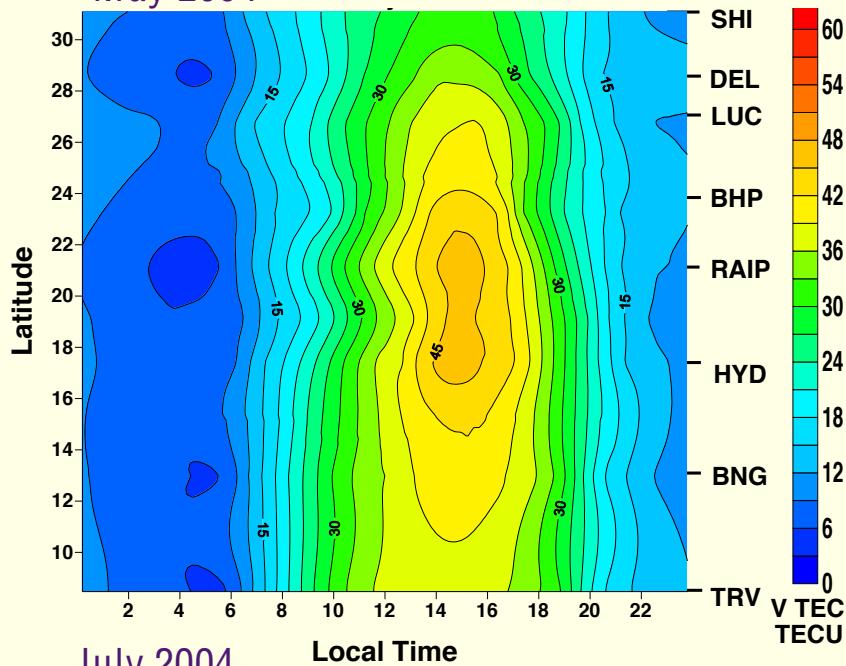




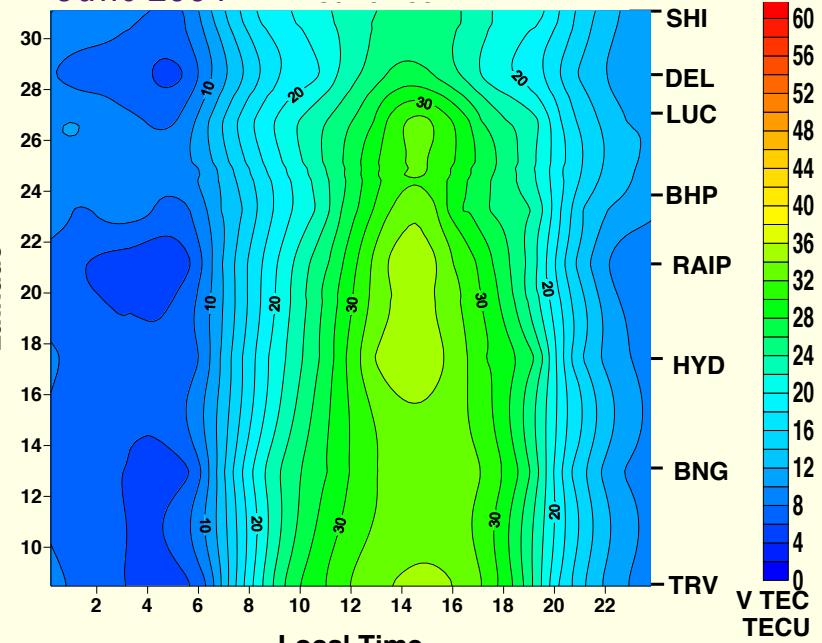
Contour plots of the diurnal variation of TEC for 9 stations ( $8.47^{\circ}\text{N}$  to  $31^{\circ}\text{N}$ ) along the common longitude of  $77^{\circ}\text{E}$  during Winter months



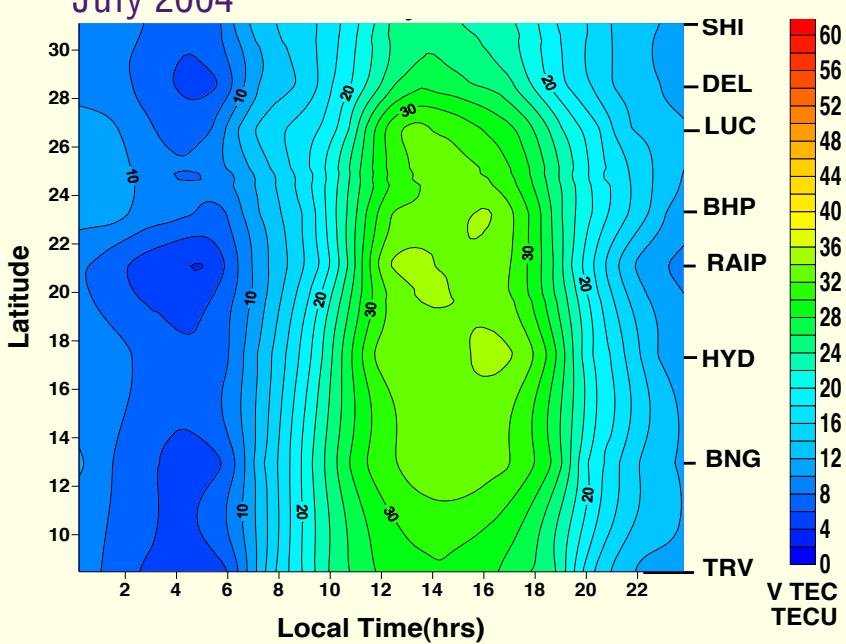
May 2004



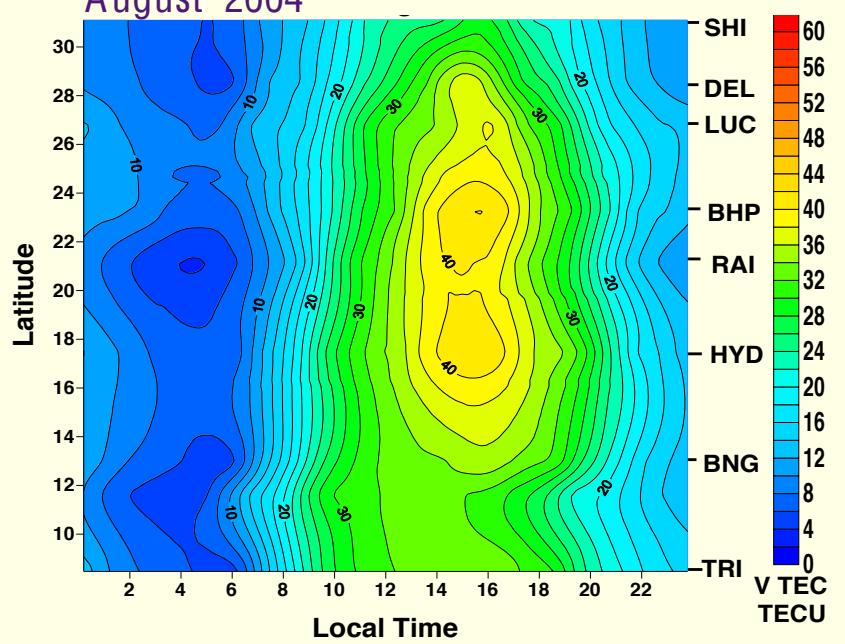
June 2004



July 2004



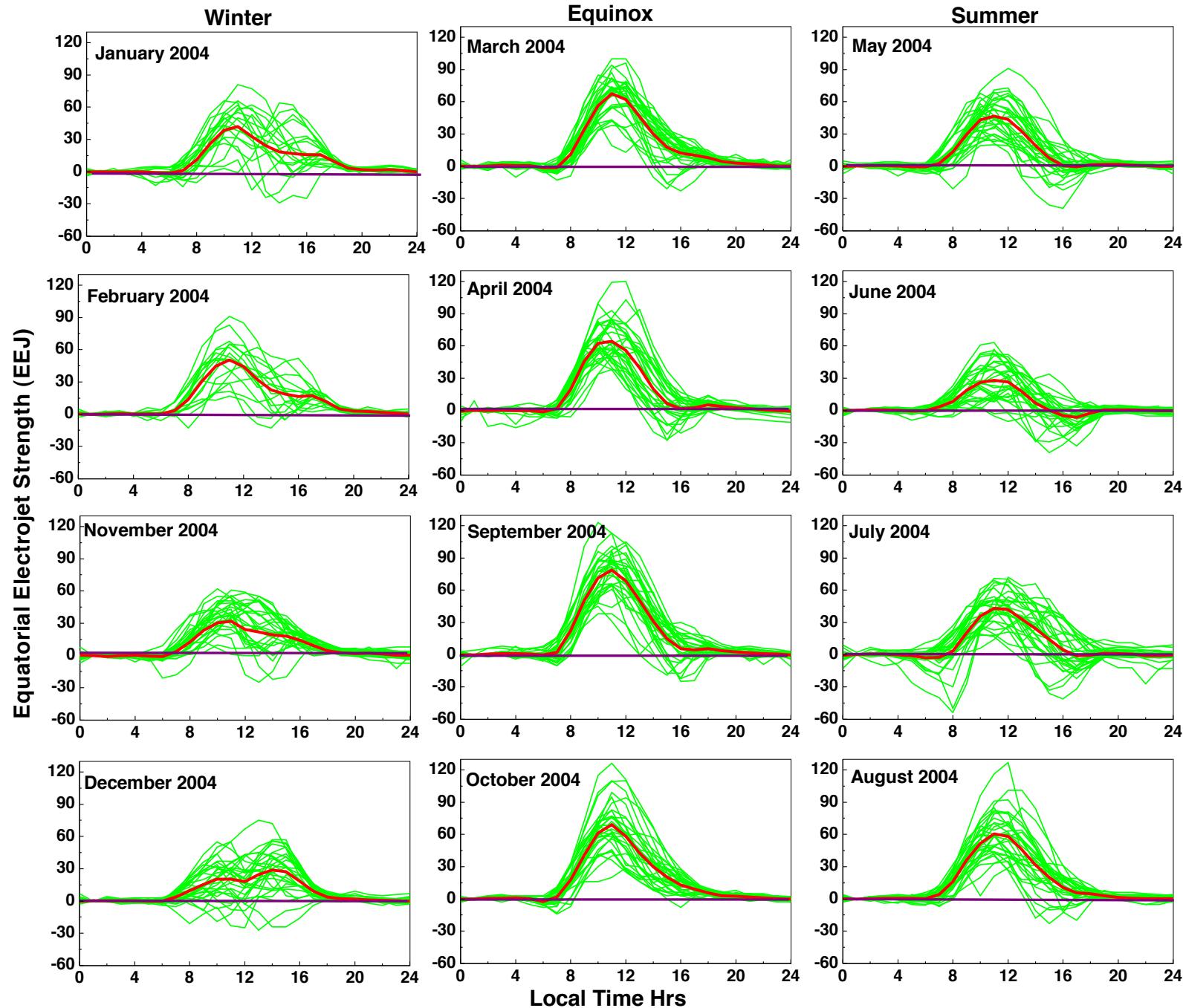
August 2004

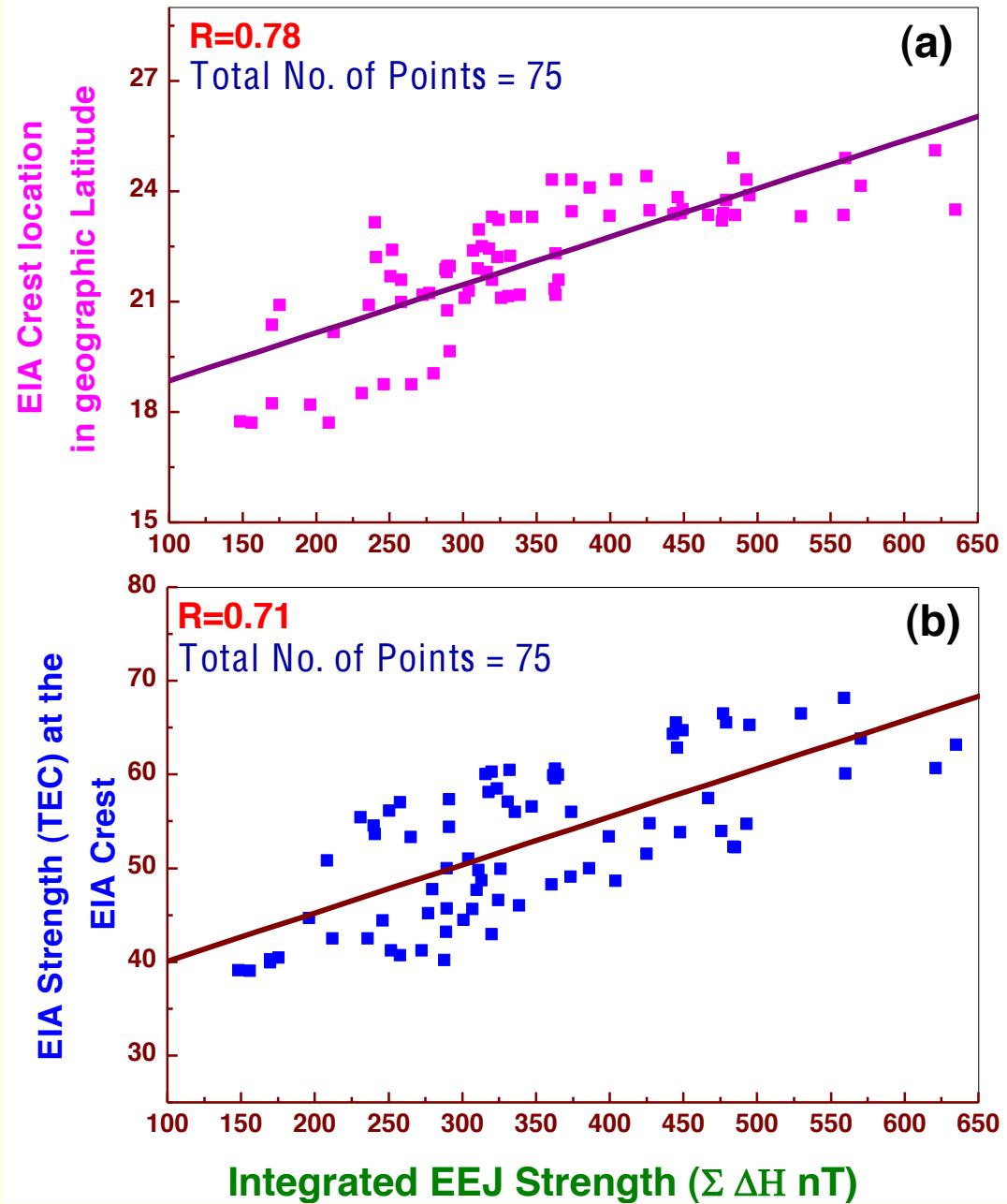


Contour plots of the diurnal variation of TEC for 9 stations ( $8.47^{\circ}\text{N}$  to  $31^{\circ}\text{N}$ ) along the common longitude of  $77^{\circ}\text{E}$  during Summer months



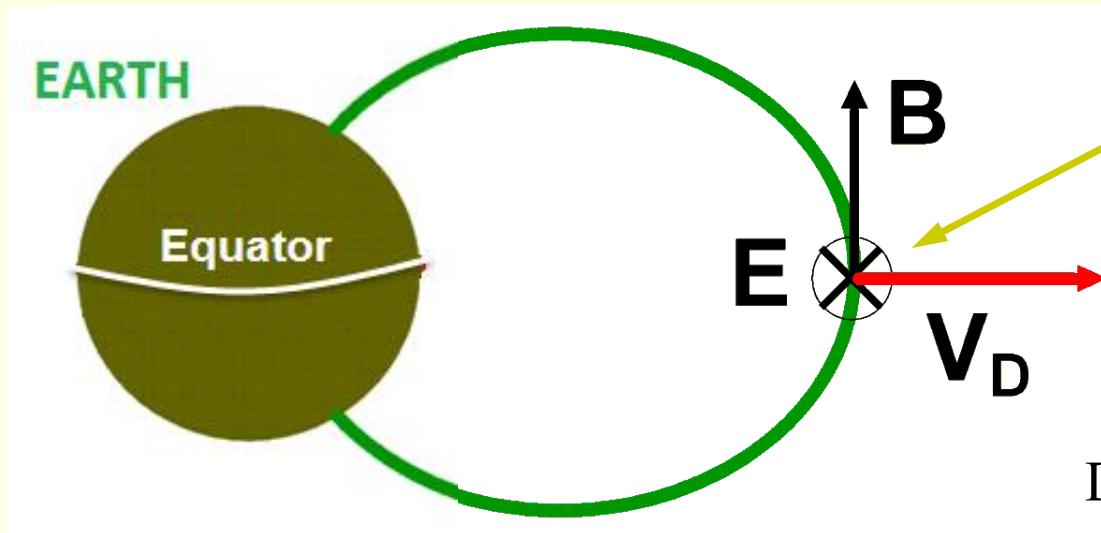
# Electrojet Strength at an equatorial station, Trivandrum





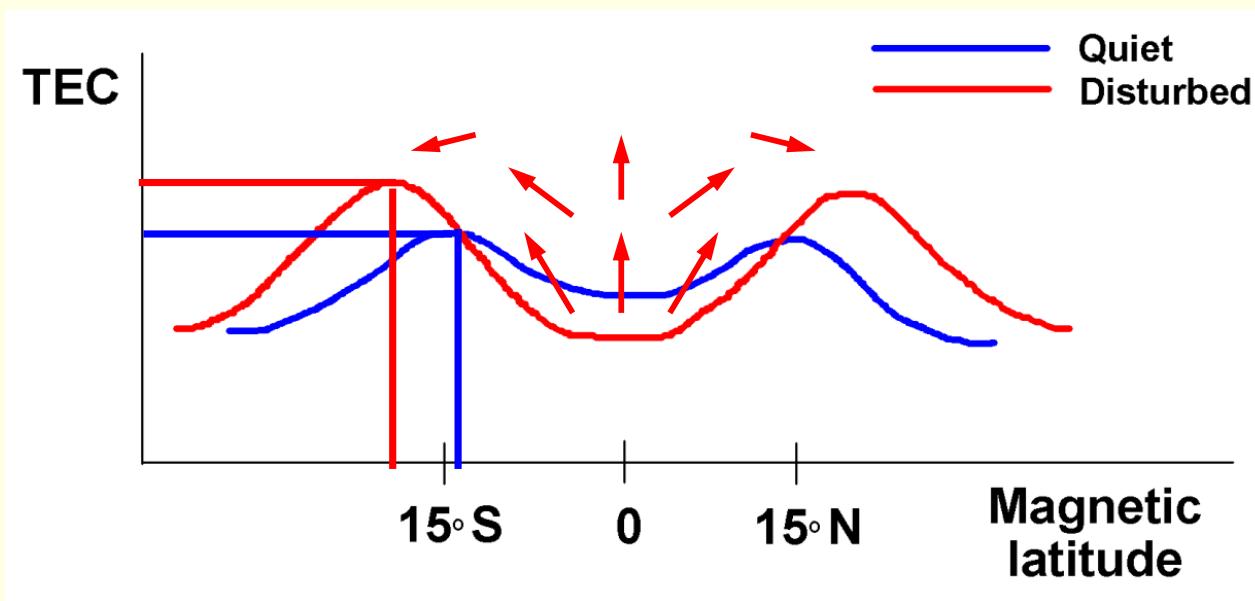
*Variation of (a) Location and (b) peak electron density of the crest of EIA as a function of Integrated Equatorial Electrojet Strength during geomagnetically quiet days*





## Intensified Eastward Electric Field

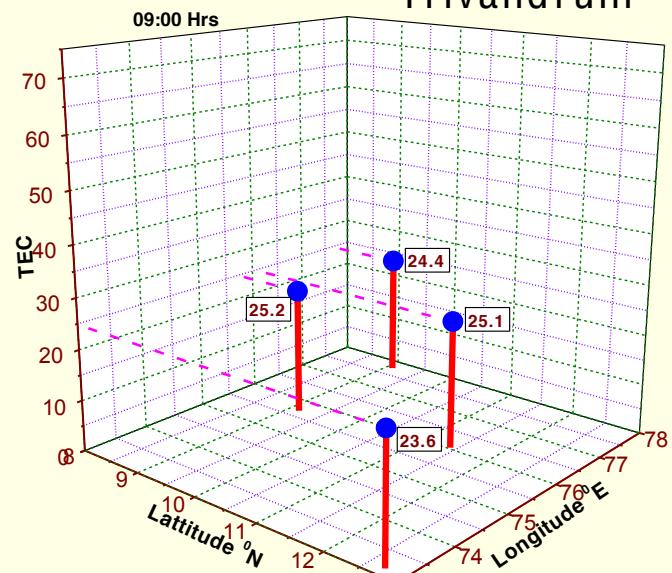
Daytime  $\sim 9$  LT to  $21$  LT



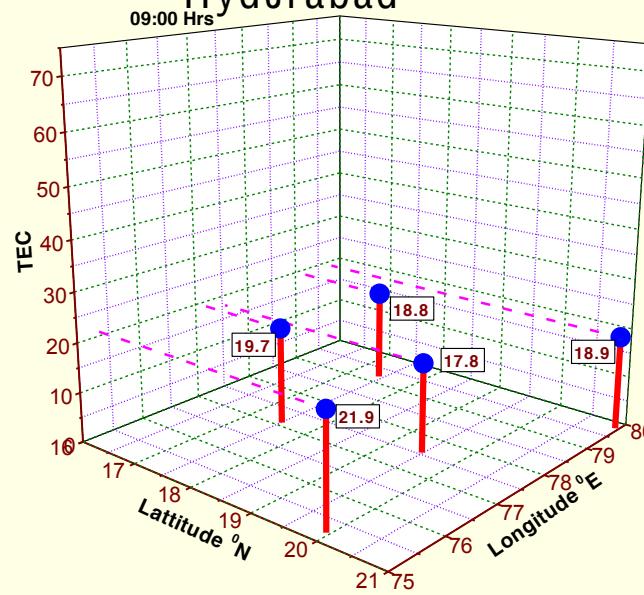
# **Variations of the GPS measured Total Electron Content in the $5^{\circ} \times 5^{\circ}$ grid**

5<sup>th</sup> May 2004 (09:00 Hrs LT)

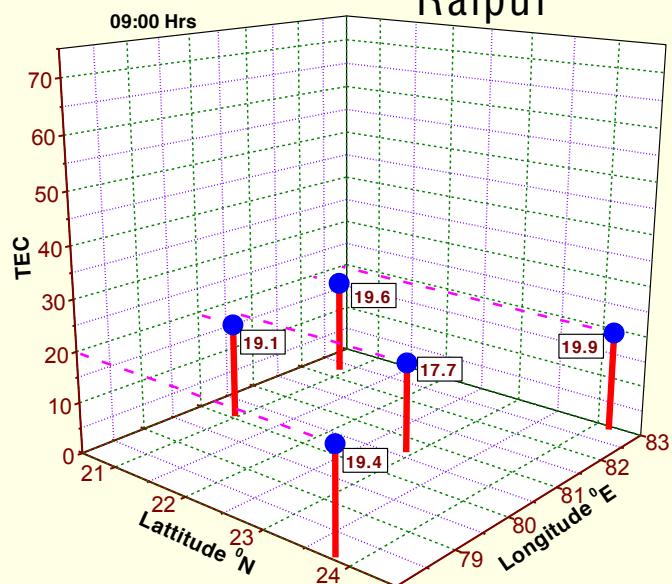
Trivandrum



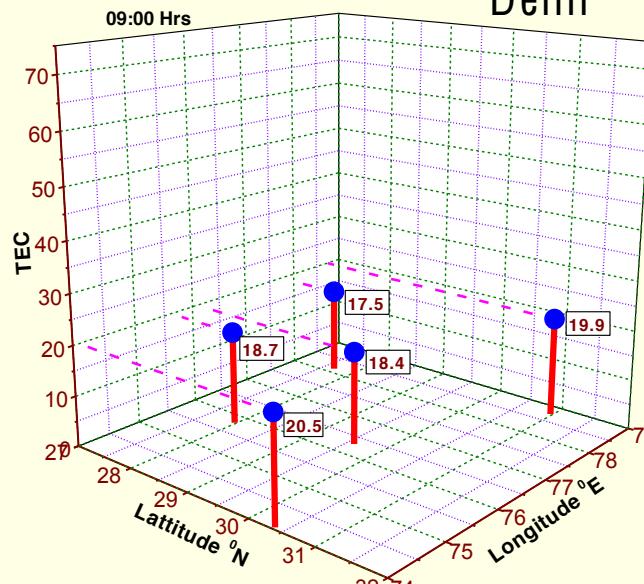
Hyderabad



Raipur



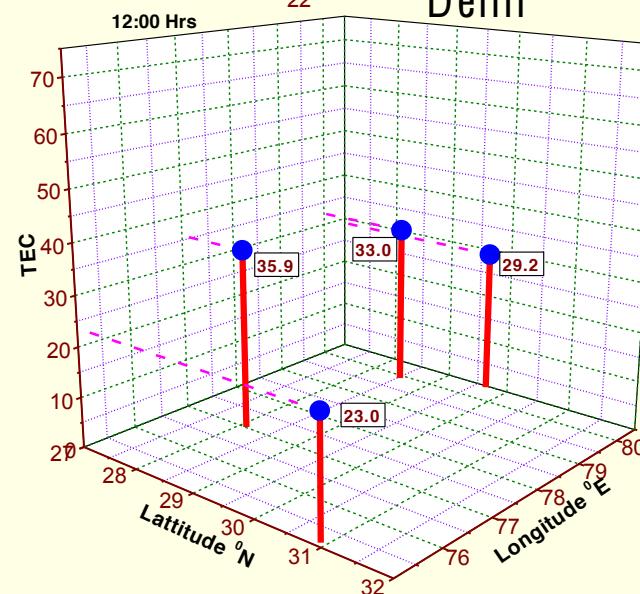
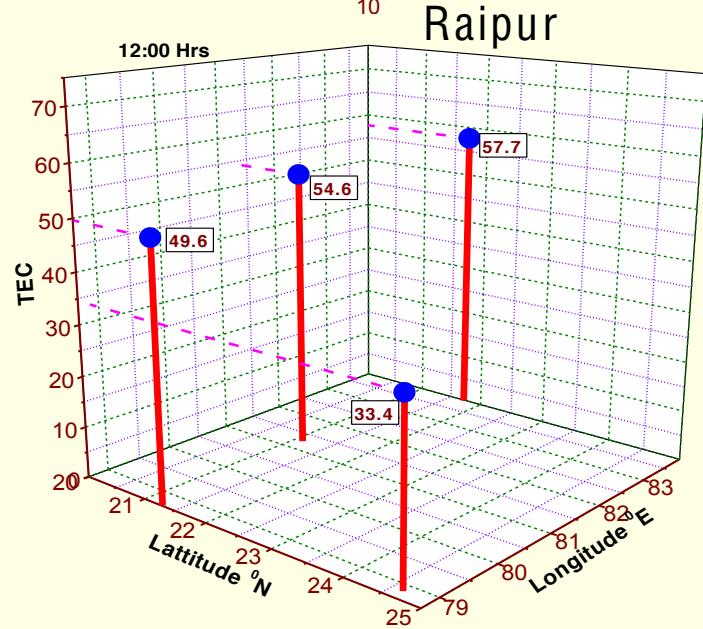
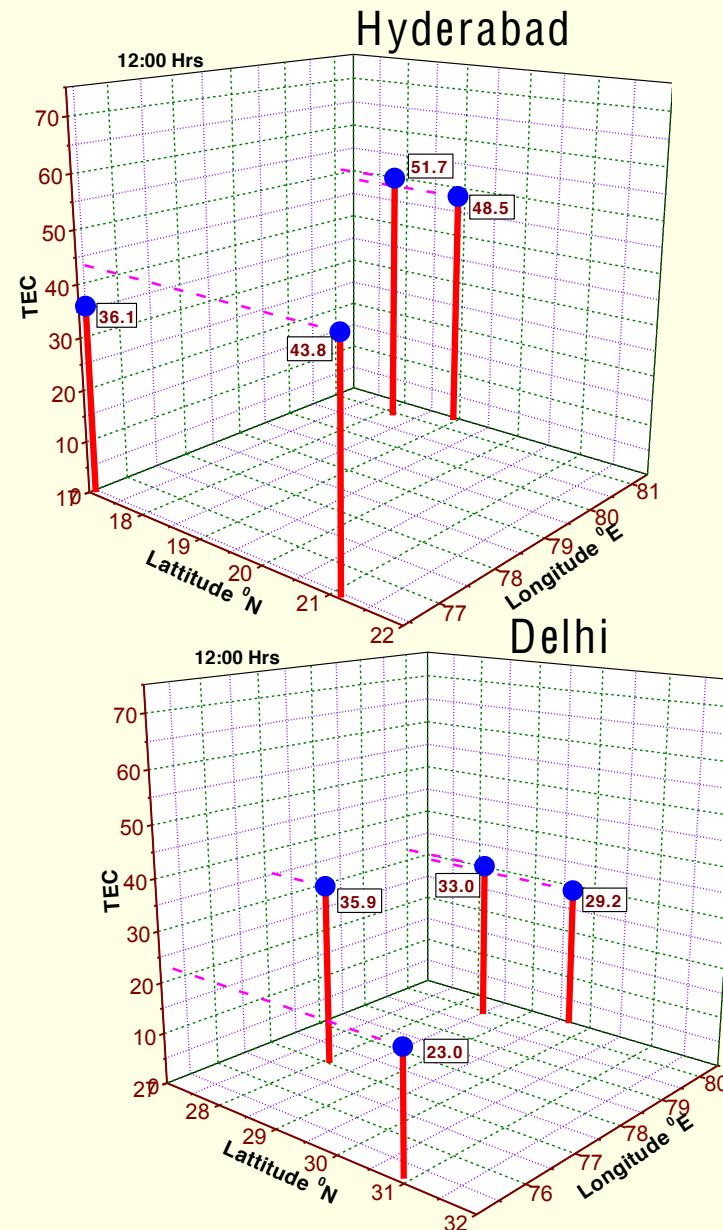
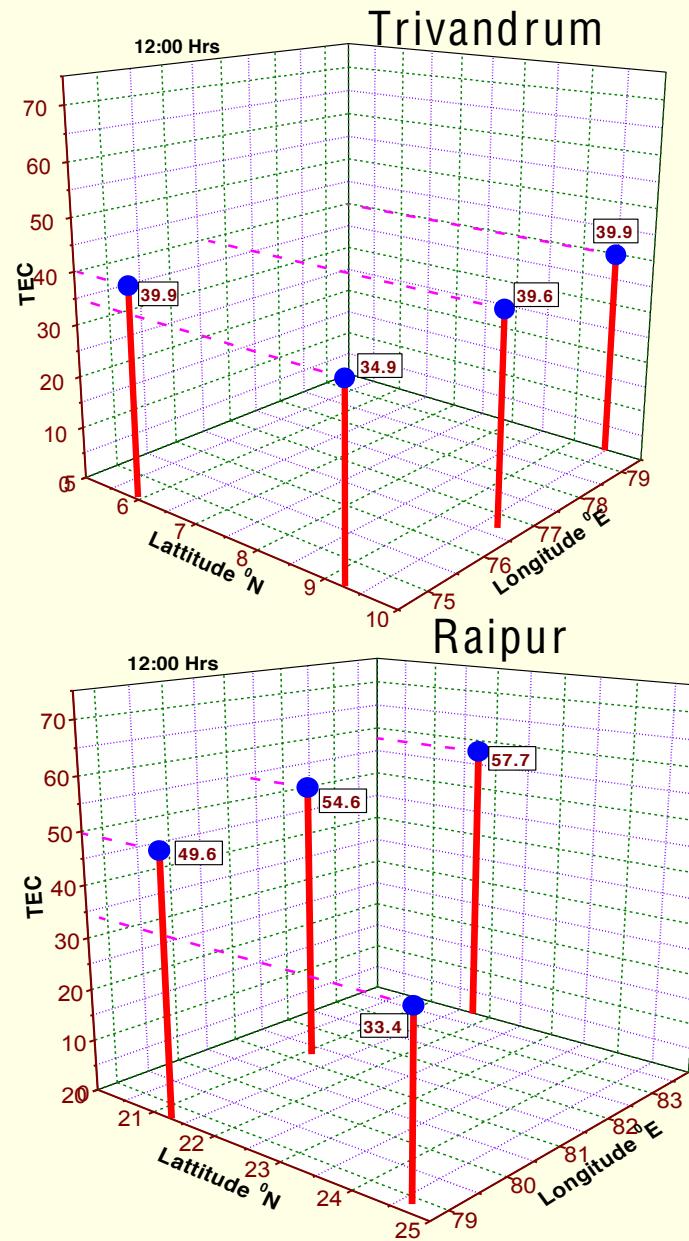
Delhi



Typical TEC values at the respective subionospheric points in the region of 5 X 5 degree Latitude and Longitude at four Indian typical stations Trivandrum, Hyderabad, Raipur and Delhi at a specific local time of 09:00 Hrs on 5th May 2004



5<sup>th</sup> May 2004 (12:00 Hrs LT)

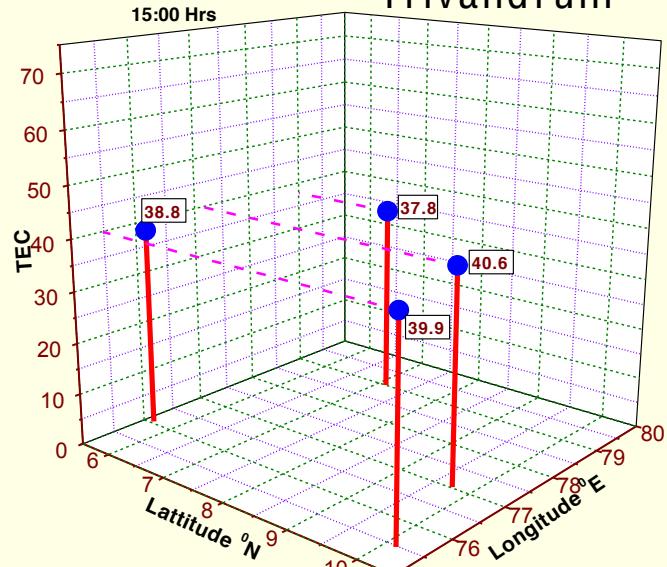


Typical TEC values at the respective subionospheric points in the region of  $5 \times 5$  degree Latitude and Longitude at four Indian typical stations Trivandrum, Hyderabad, Raipur and Delhi at a specific local time at 12:00 Hrs noon on 5<sup>th</sup> May 2004

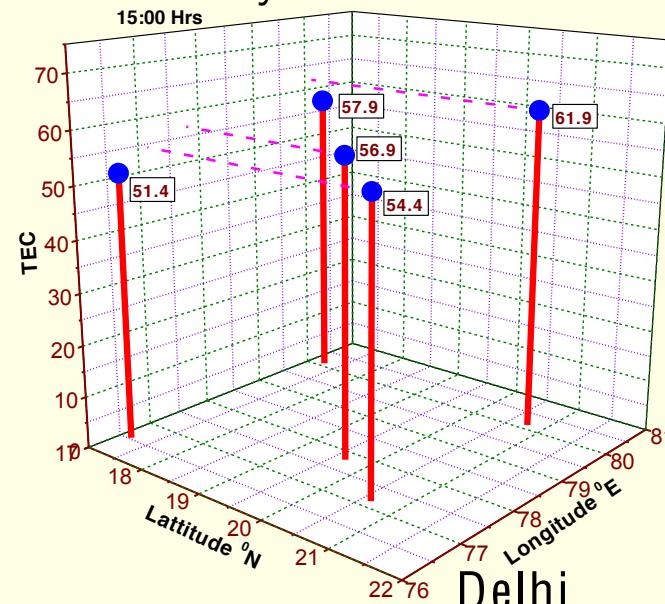


5<sup>th</sup> May 2004 (15:00 Hrs LT)

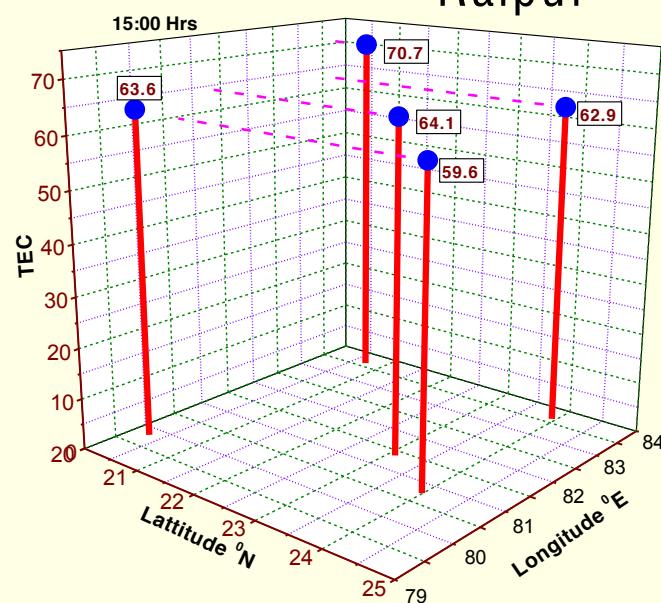
Trivandrum



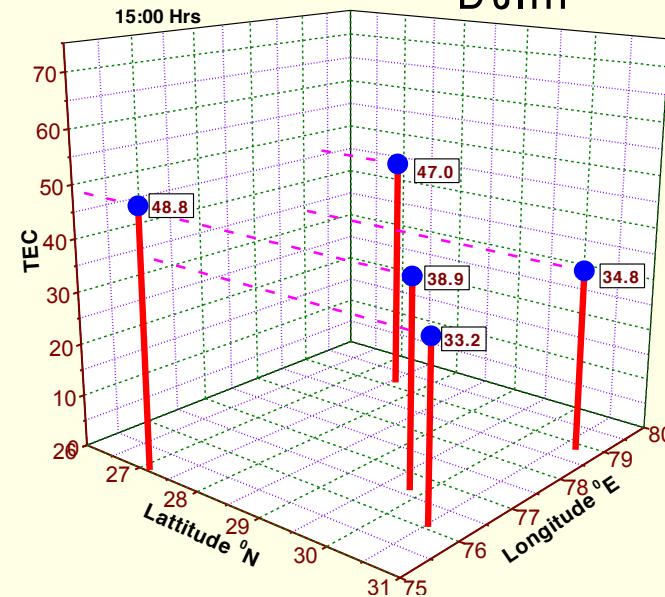
Hyderabad



Raipur



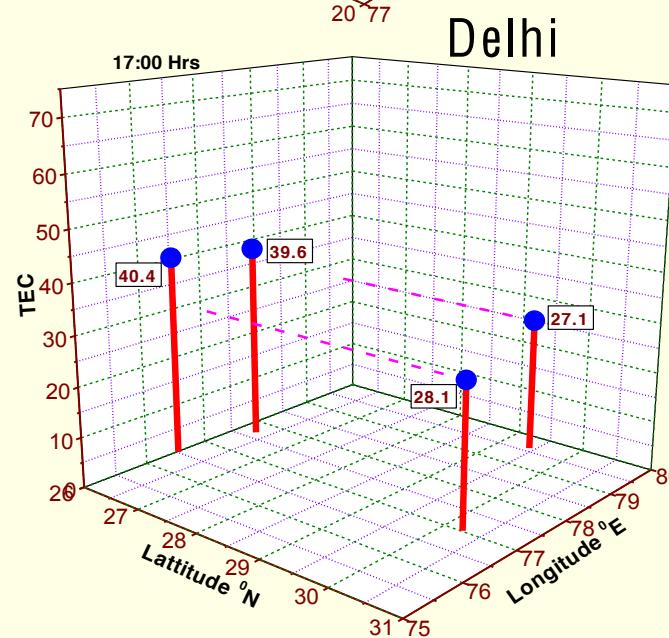
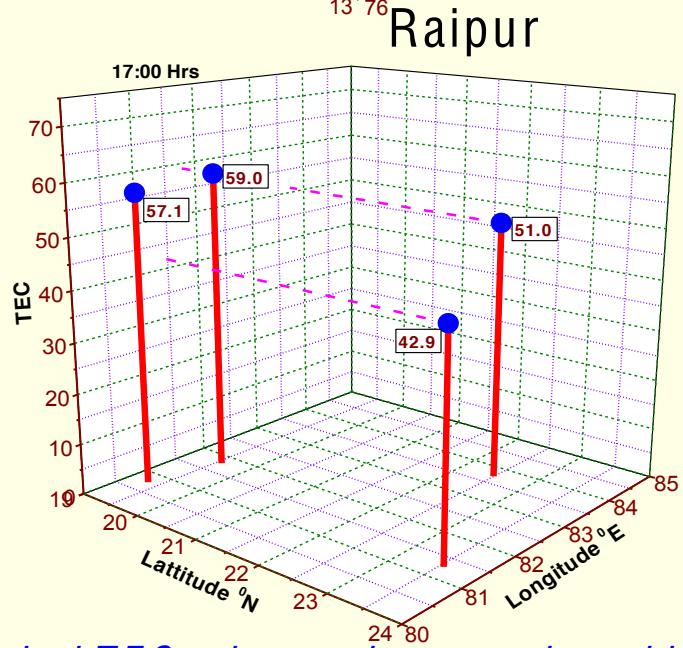
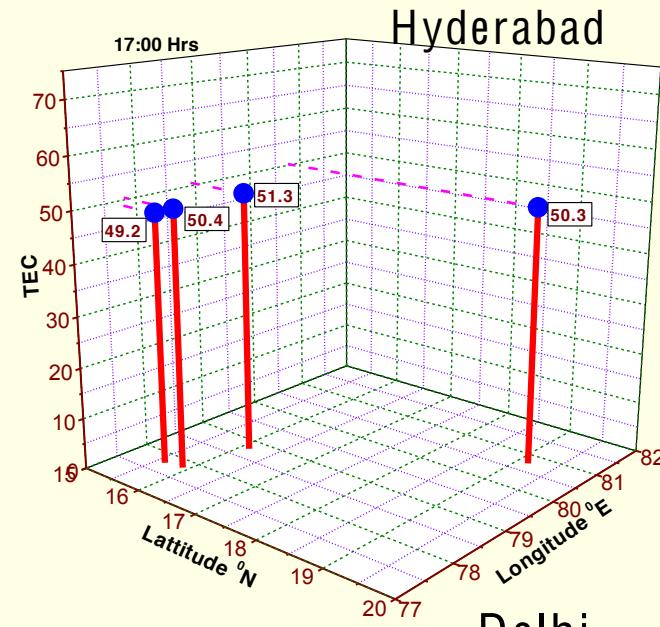
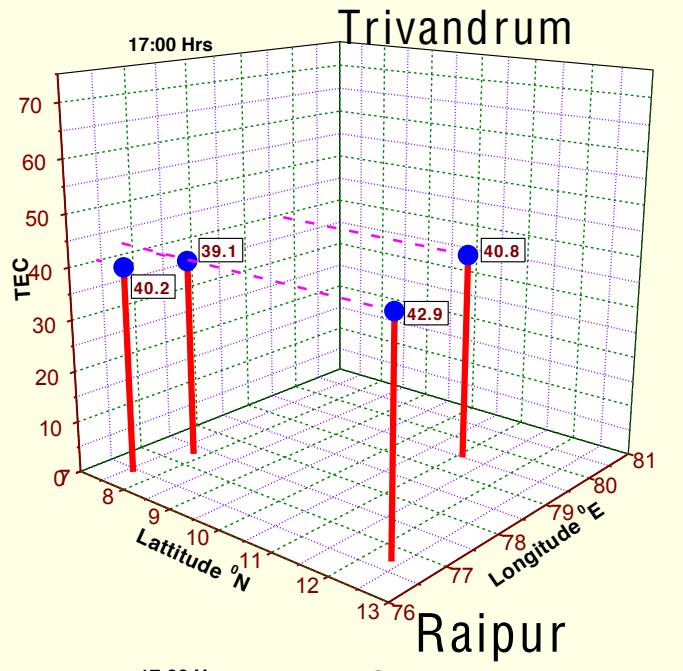
Delhi



Typical TEC values at the respective subionospheric points in the region of 5 X 5 degree Latitude and Longitude at four Indian typical stations Trivandrum, Hyderabad, Raipur and Delhi at a specific local time at 15:00 Hrs on 5th May 2004

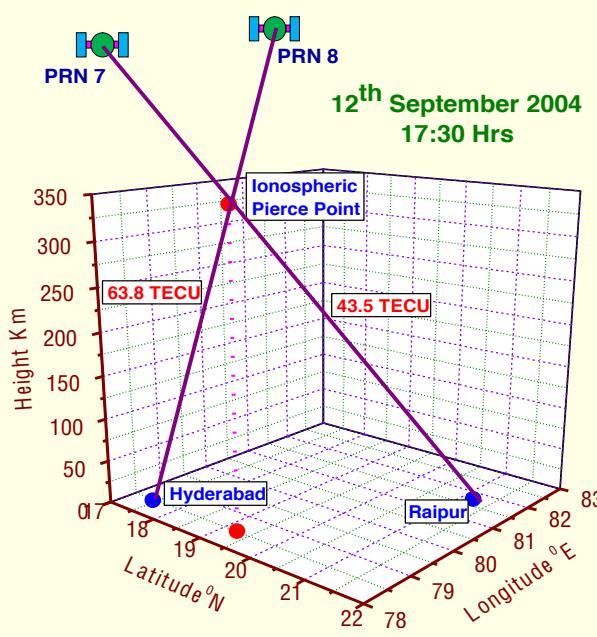
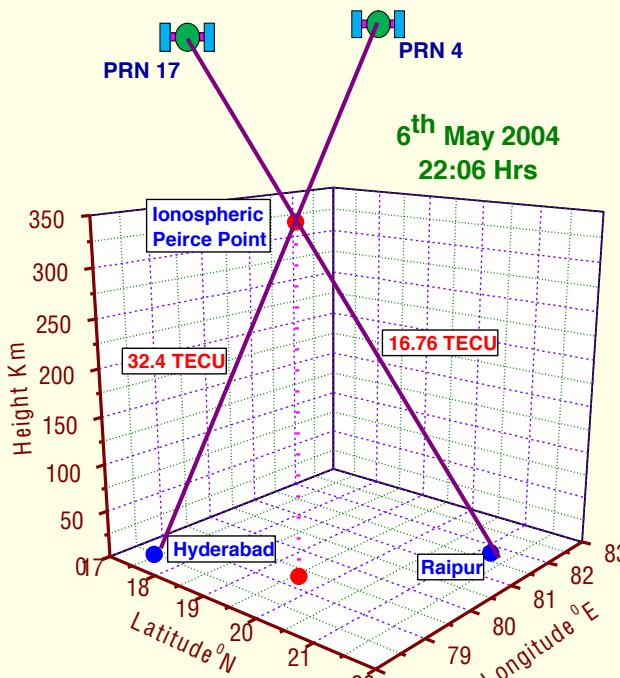
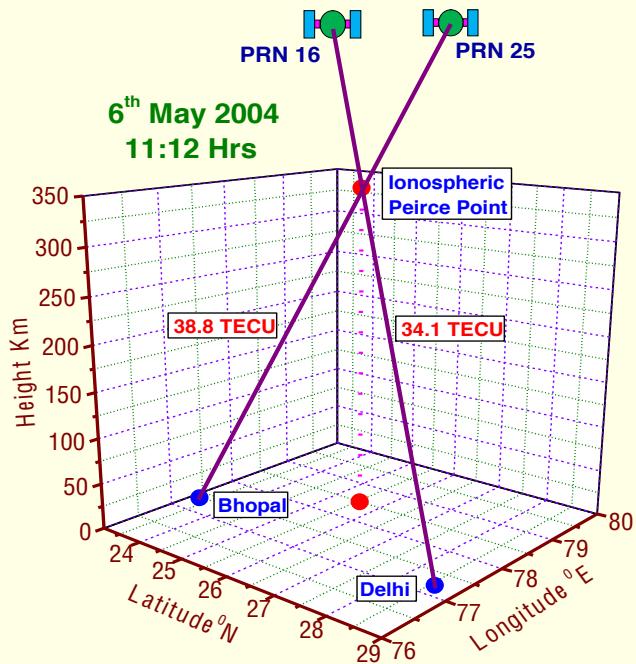
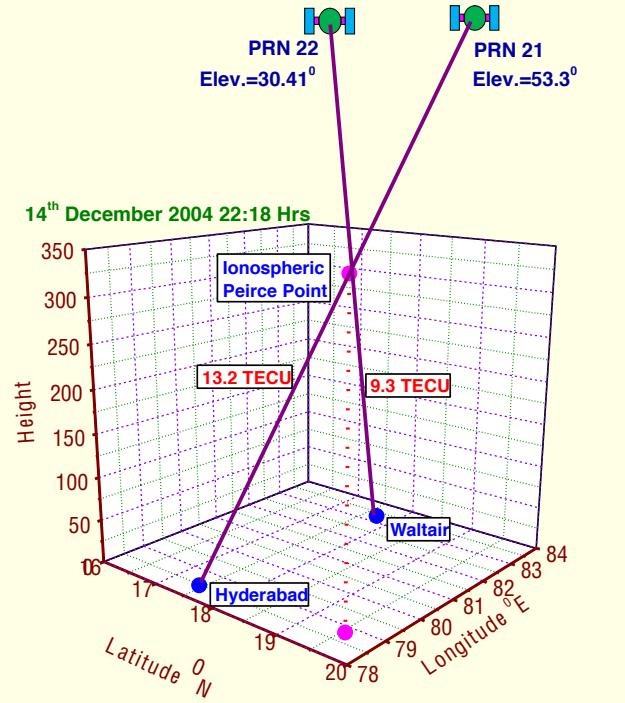


5<sup>th</sup> May 2004 (17:00 Hrs LT)



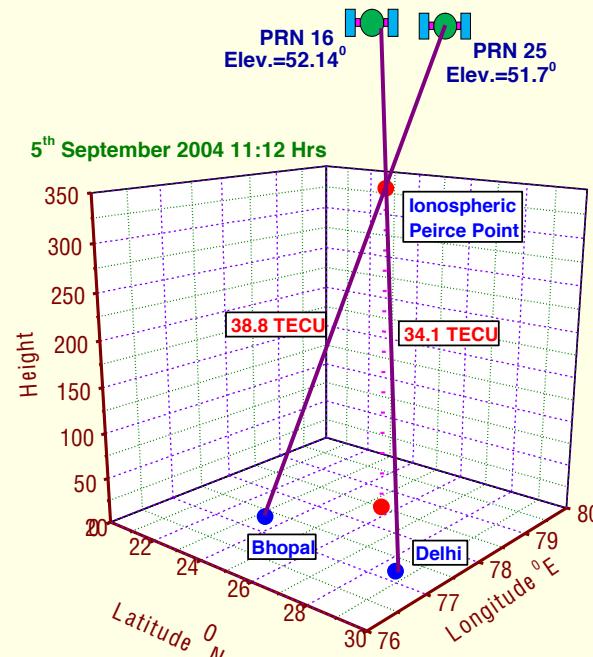
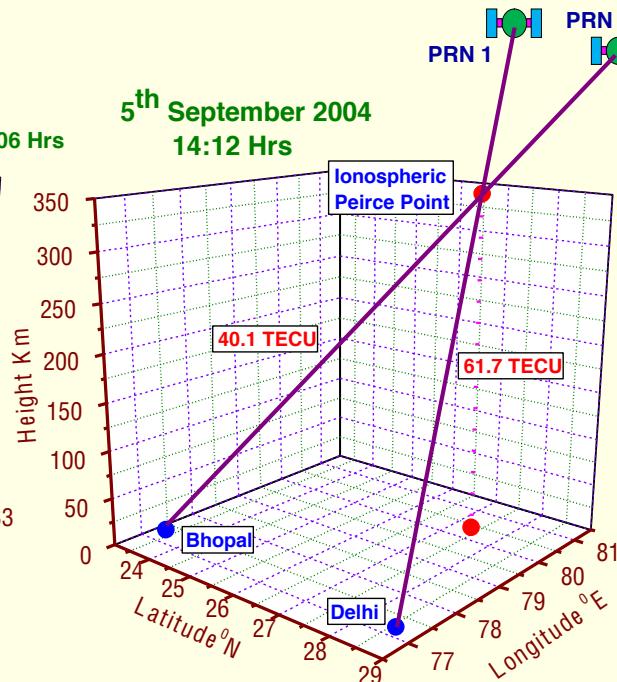
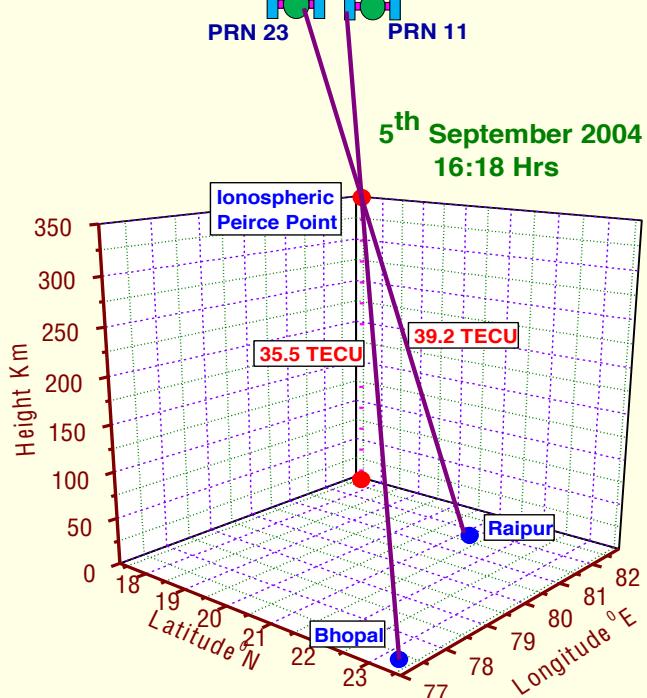
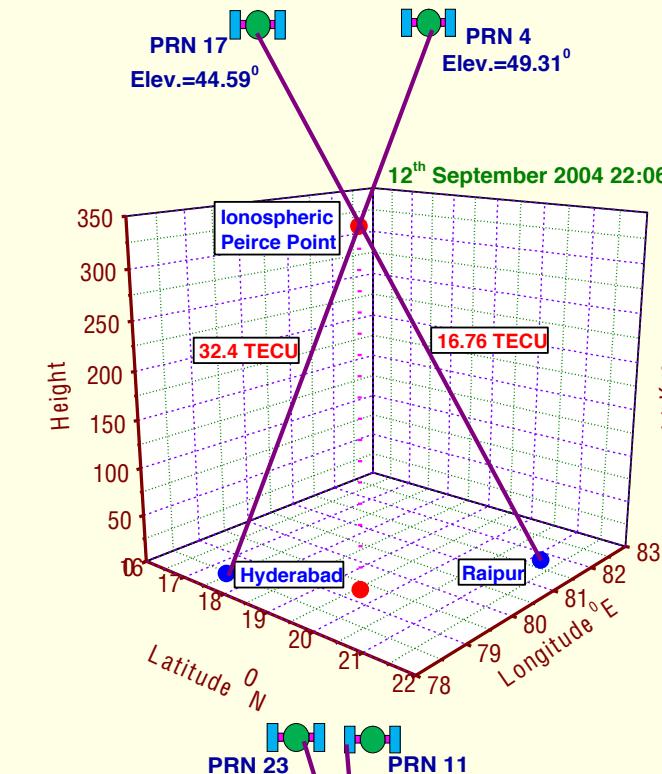
Typical TEC values at the respective subionospheric points in the region of 5 X 5 degree Latitude and Longitude at four Indian typical stations Trivandrum, Hyderabad, Raipur and Delhi at a specific local time of 17:00 Hrs PM on 5th May 2004





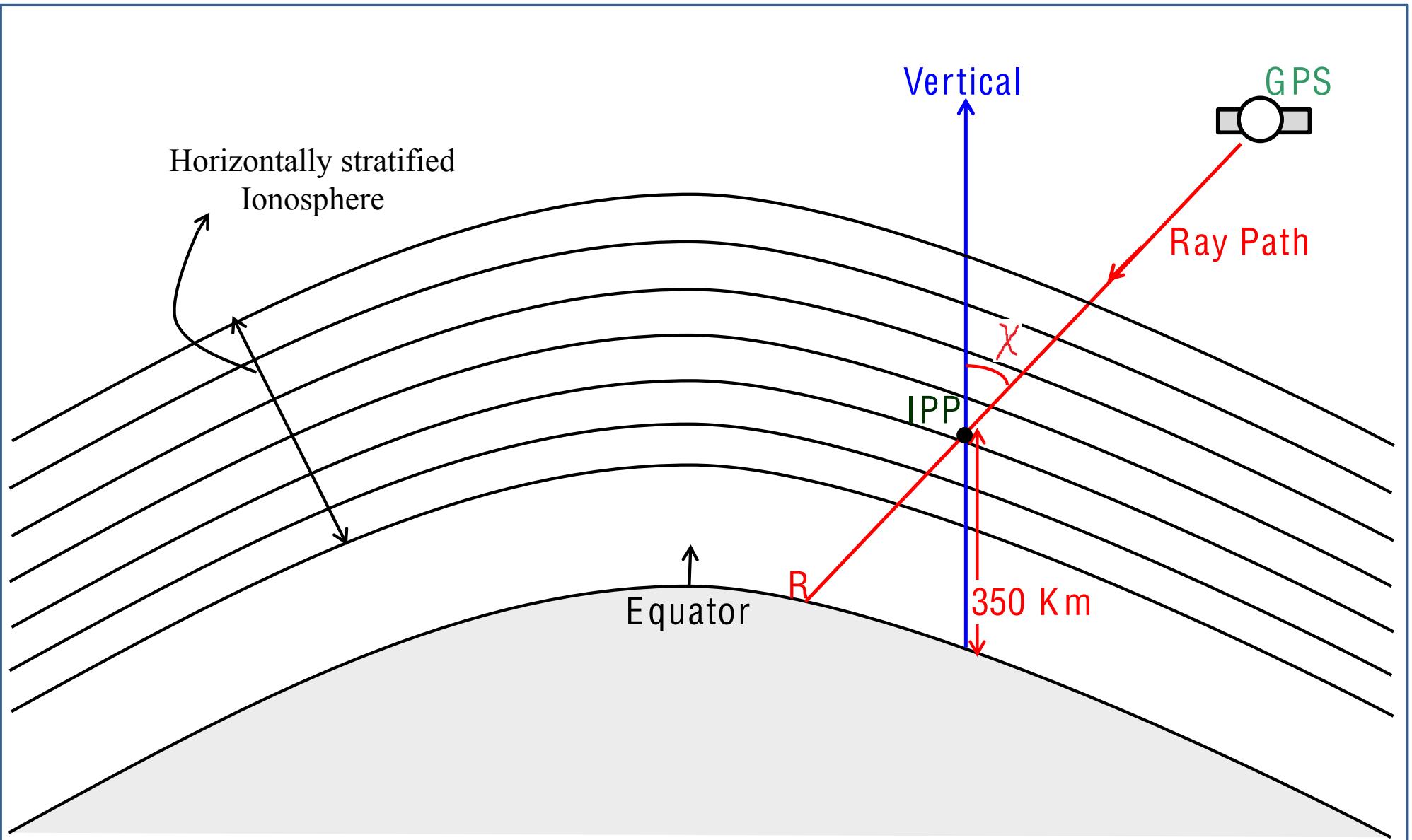
*TEC derived from two different and simultaneous ray paths at the same location of the sub ionospheric point showing significant variability in their values*





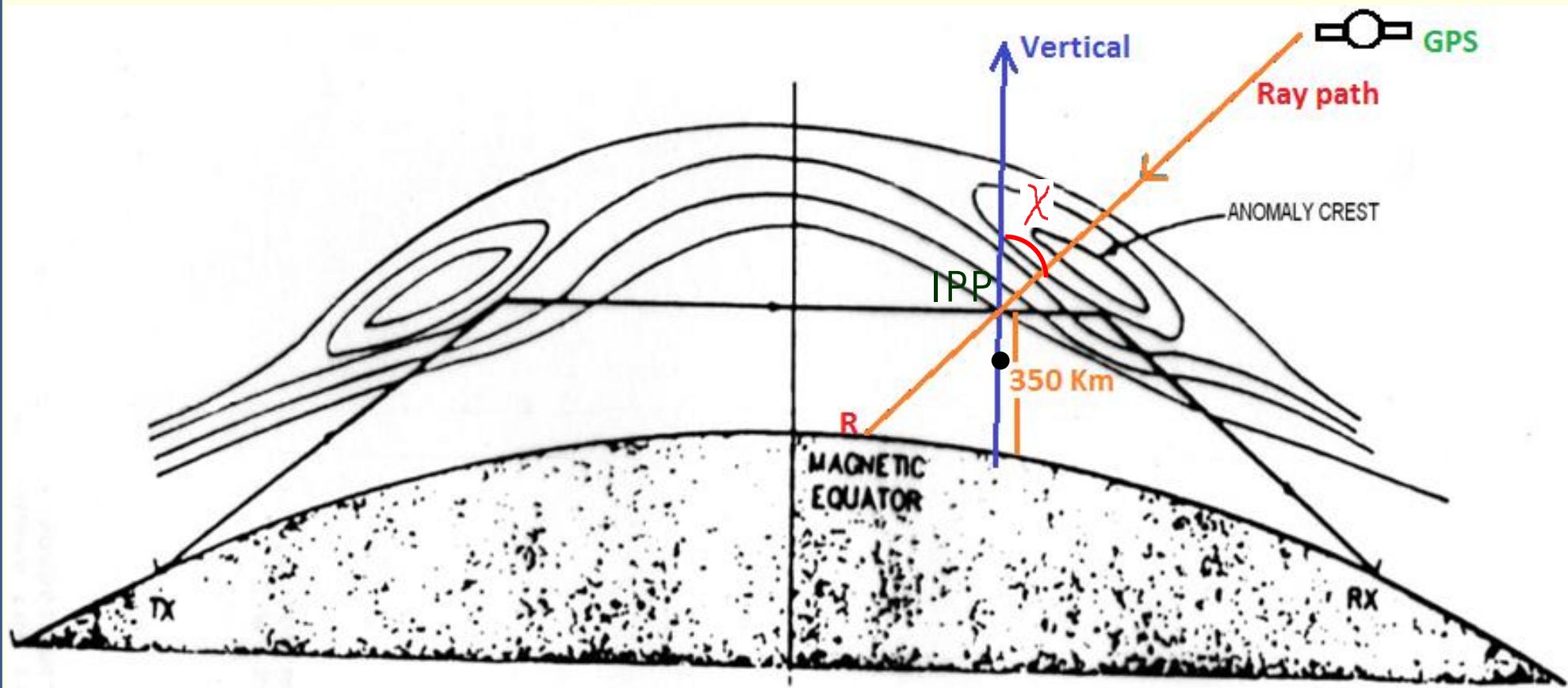
*TEC derived from two different and simultaneous ray paths at the same location of the sub ionospheric point showing significant variability in their values*



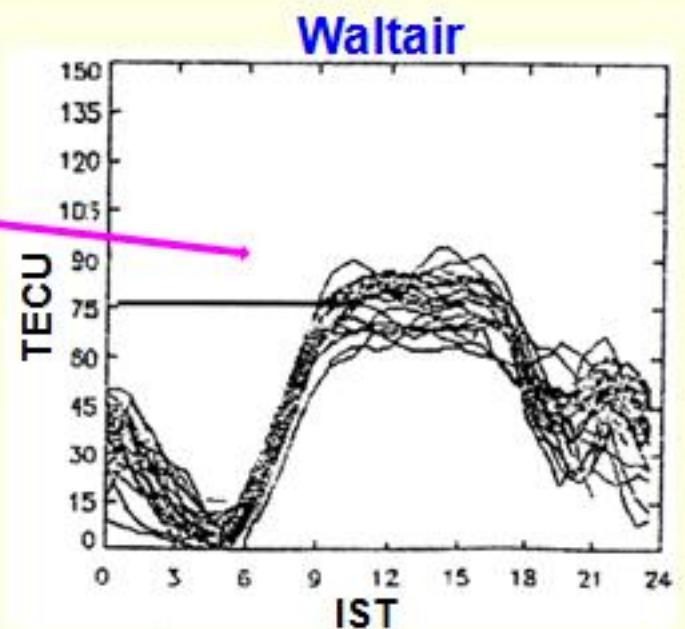
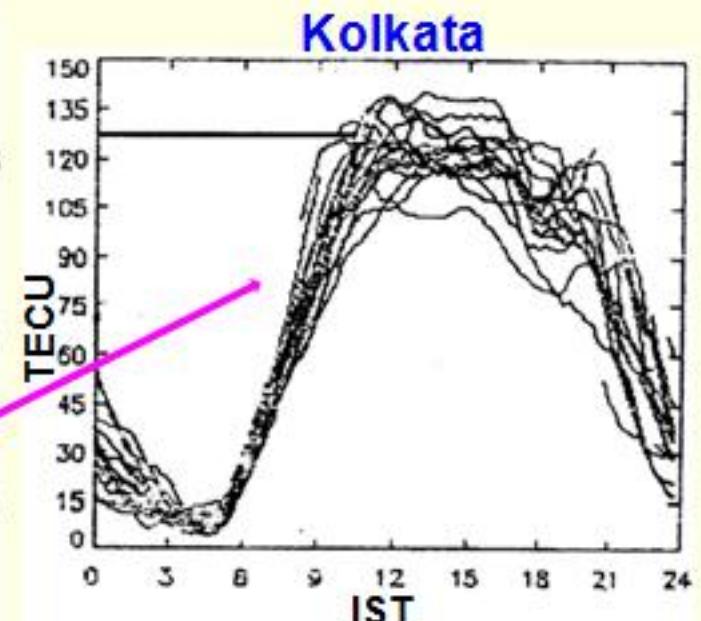
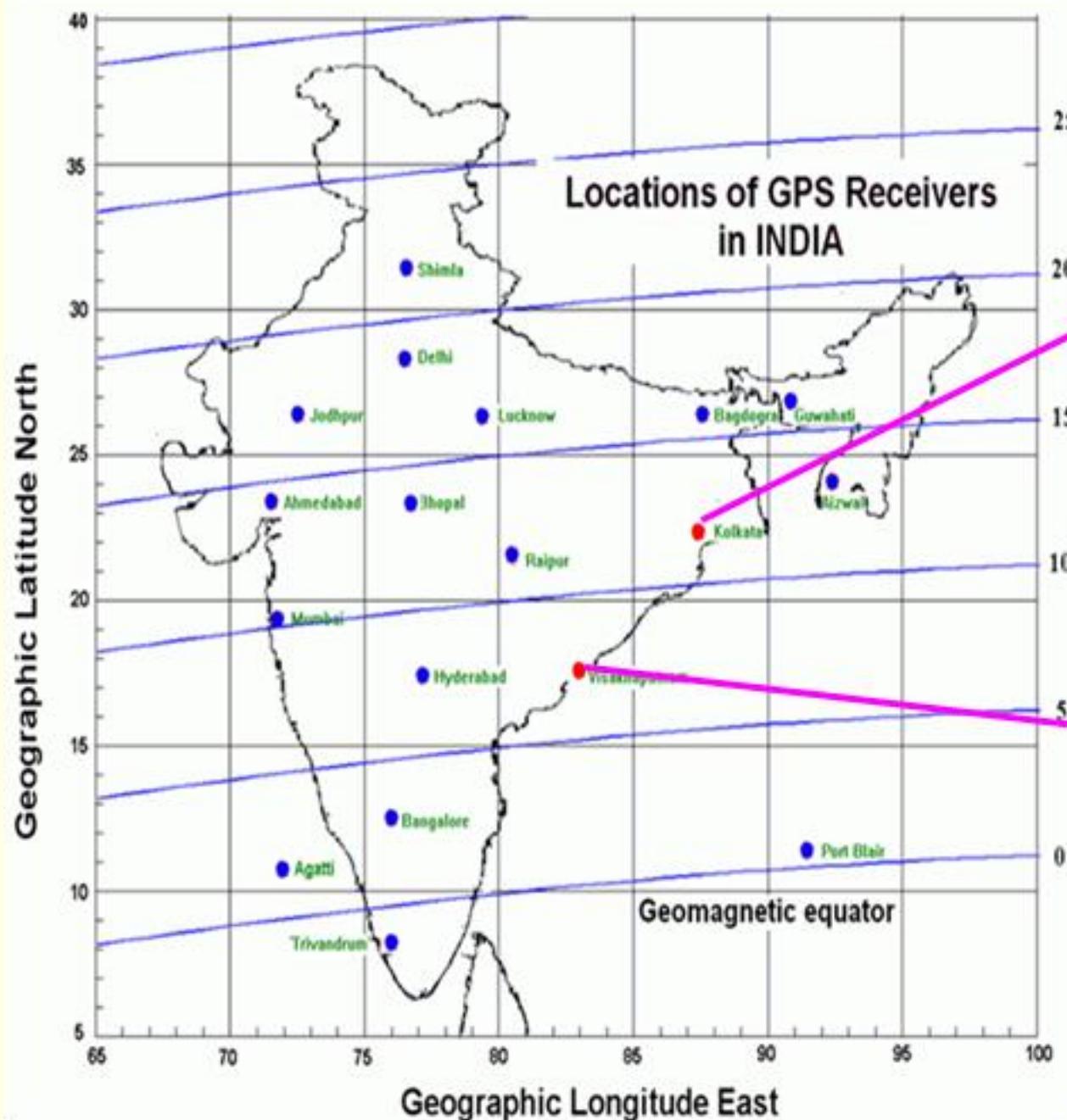


*View Plot showing the conversion of STEC to VTEC  
on a horizontally stratified ionosphere*





# Diurnal variation of TEC over Kolkata and Waltair during high solar activity years



## **Current Challenges in Beacon Research with respect to Satellite Based Navigation systems in the low latitude sector**

- 1) Measurement of absolute TEC for accurate evaluation of the range delays in the trans-ionospheric signals
- 2) Prediction of TEC ahead of time with reasonable accuracy demanded by satellite based navigation
- 3) SBAS applications –  $5^{\circ} \times 5^{\circ}$  grid based interpolation technique – inadequacy in the equatorial and low latitudes because of large spatial gradients due to EIA (8 to 10 TEC units per each degree latitude) during high solar activity period.
- 4) Mapping function errors in the conversion of Vertical TEC to Slant TEC and vice versa due to large gradients around the EIA region.
- 5) Day-to-day variability in TEC
- 6) Storm time variations in TEC



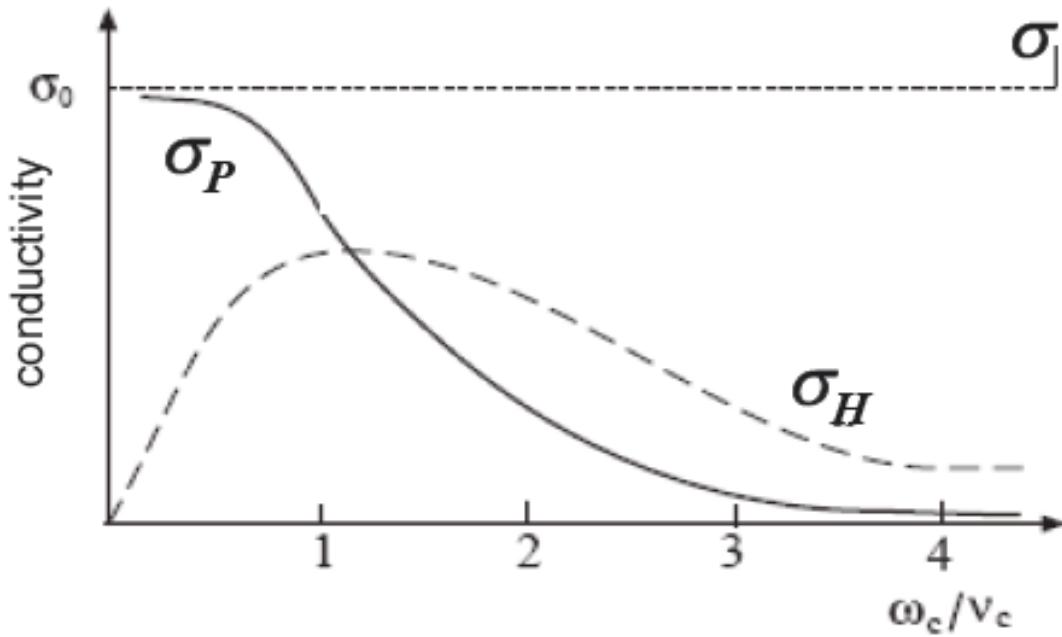
- 7) To the existing constellation of GPS, if GLONASS and GALELIO satellites also become operational, more vertical ray paths can be used with minimum mapping function errors for TEC measurements at any location
- 8) Increased L-band scintillation activity is expected at the EIA crest region with the increase in the solar activity
- 9) Loss of lock of receivers – To identify threshold amplitudes of scintillations
- 10) In the equatorial and low latitude sectors over different parts of the globe, measurements of TEC and scintillations to be intensified to buildup strong database that will help in modeling and in increasing the prediction capability of TEC and Scintillations.





Space Physics Laboratory, Andhra University, Visakhapatnam, INDIA





The significance of different elements of the conductivity tensor depends on the relation between the gyro frequency ( $\omega$ ) and collision frequency( $\nu$ ).

$$\omega \ll v_c$$

Collisions prevent the gyro motion  
 $\rightarrow$  particles move  $\parallel E$   
 $\rightarrow \sigma_p$  : **Pedersen** Current ( $\perp B \parallel E$ )

$$\omega \gg v_c$$

Particles able to make many gyrations before scattering by a collision  
 $\rightarrow E \times B$  drift dominates  
 $\rightarrow \sigma_H$  : **Hall** Current ( $\perp B \perp E$ )



# IONO-SBAS MEETING: INDIAN GAGAN Programme

ICTP – TRIESTE, ITALY (20-21 April, 2012)



*Prof. P.V.S. Rama Rao*  
Andhra University, INDIA



# GAGAN

Indian Space Research Organisation

## OBJECTIVE

To provide Satellite-based Navigation services with accuracy and integrity required for civilian and aviation applications over Indian Air Space.

Better Air Traffic Management over Indian Airspace.

# GAGAN CONFIGURATION



Indian Space Research Organisation

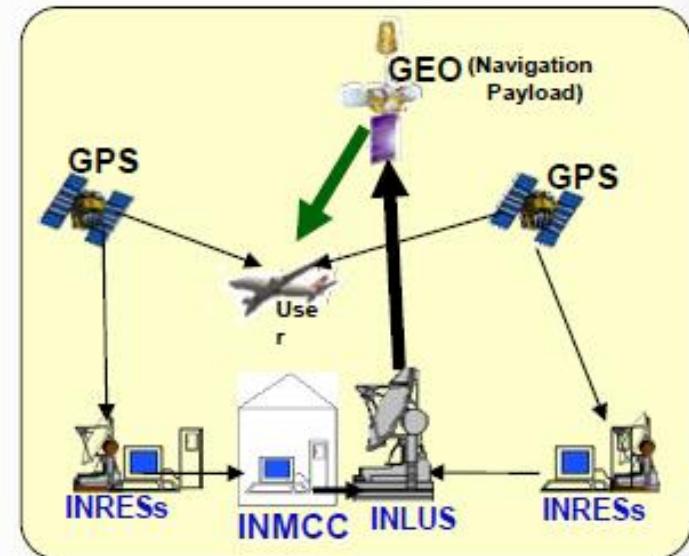
## SPACE SEGMENT

### Three GEO Satellites

- Two operational Navigation Payloads
- One in-orbit spare Navigation Payload

## GROUND SEGMENT

- Indian Master Control Centre (INMCC)
- Indian Navigation Land Uplink Station (INLUS)
- Indian Reference Stations (INRES)
- Communication links



# GAGAN – INMCC Facility

## Bangalore



Source: ISRO and Airport Authorities of India

## Indian Master Control Centre



Source: ISRO and Airport Authorities of India  
*Space Physics Laboratory, Andhra University, Visakhapatnam, INDIA*



## 11 Meter Up-Link Antenna



Source: ISRO and Airport Authorities of India  
*Space Physics Laboratory, Andhra University, Visakhapatnam, INDIA*



# Application of SBAS Receiver



## Applications

- For Aviation

- High Accuracy, Global Coverage, Direct En-route, Reduced and Simplified equipment on-board & ground

- Other Potential Users

- Survey, Land Management Scientific Research, Business solutions, Geodynamics etc.,

## Tasks Ahead

- Signal In Space tests including air-borne receivers
- Production of SBAS Receivers
- User Meet

**The implementation of the GAGAN Programme is in two phases**

- **GAGAN-TDS (Technology Demonstration System)**
- **GAGAN – FOP (Full Operation Phase)**



## **GAGAN-TDS –Performance**

- Only accuracy has been verified (7.6m Vertical and Horizontal 95% of the time within the perimeter of INRESSs). Verification results were well within the performance requirements
- SBAS Receiver is able to lock GAGAN PRN 127 (GSAT-8) and process the GAGAN messages for providing SBAS position. Results were satisfactory
- Ranging Capabilities were Demonstrated
- Perimeter sites exhibited good performance also
- In TDS phase there were no Integrity and Level of service requirements

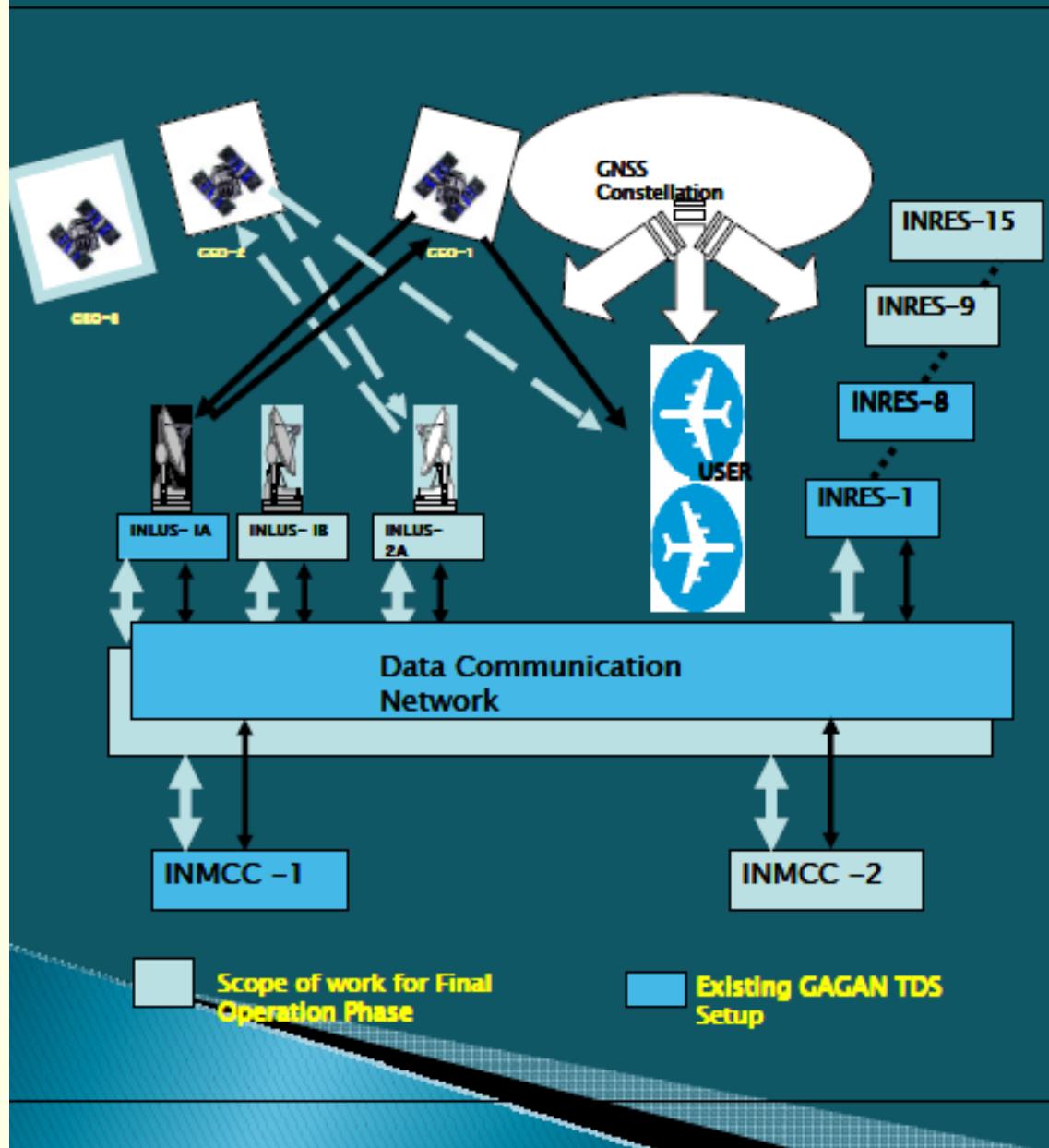


# **GAGAN - FOP Objective**

- 1) To establish a fully operational SBAS with redundancies and to
  - 2) Establish suitable region specific IONO model and to get
  - 3) safety Certification for the system for the Civil Aviation use.
- 
- The GAGAN Full Operational Phase (FOP) objective is to provide a certified navigation system for all phases of flight by augmenting the TDS system suitably



# GAGAN-FOP Configuration



## System/Activities Completed

System	FOP
INRES	15 (3 chains)
INMCC	2
INLUS-SG	2
O&M	2
Iono Model (Equatorial)	IGM-MLDF
GEO	1 (GSAT-8)
INLUS-RF	1
COM Links	2

## System/Activities Planned

System	FOP
INLUS-RF	2
GEO	GSAT-10, 15
Comm. Links	1



# GAGAN: Performance Objective

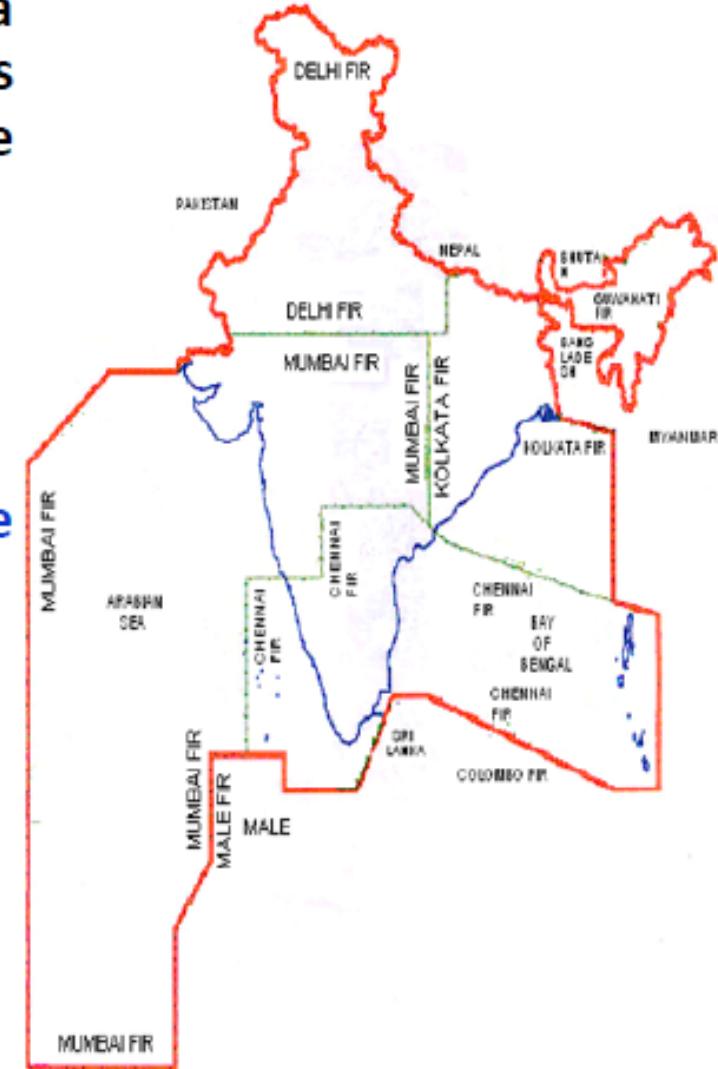
The objective of GAGAN-FOP is to realize a certified and operational SBAS for all phases of flight path **over the Indian FIR** to provide the air navigation services of

- ❖ RNP 0.1 en route navigation within Indian FIR
- ❖ APV-1/1.5 precision approach over the landmass of Indian FIR

Performance Requirements

Flight Phase	Accuracy	HAL	VAL
RNP 0.1	72 m (H)	185 m	N/A
APV 1.5	33.5 m (H) 9.8 m (V)	40 m	50 m

lon.  
lat.  
model

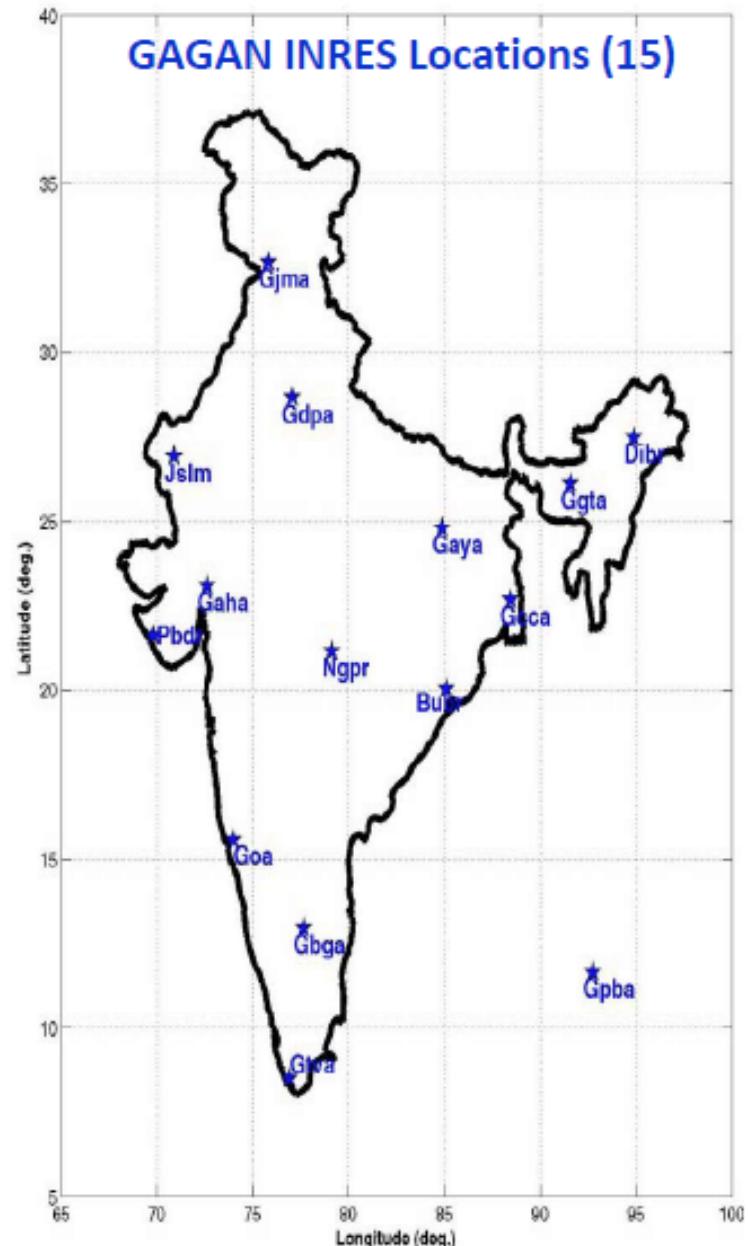




# Grid Based Ionosphere Model For GAGAN

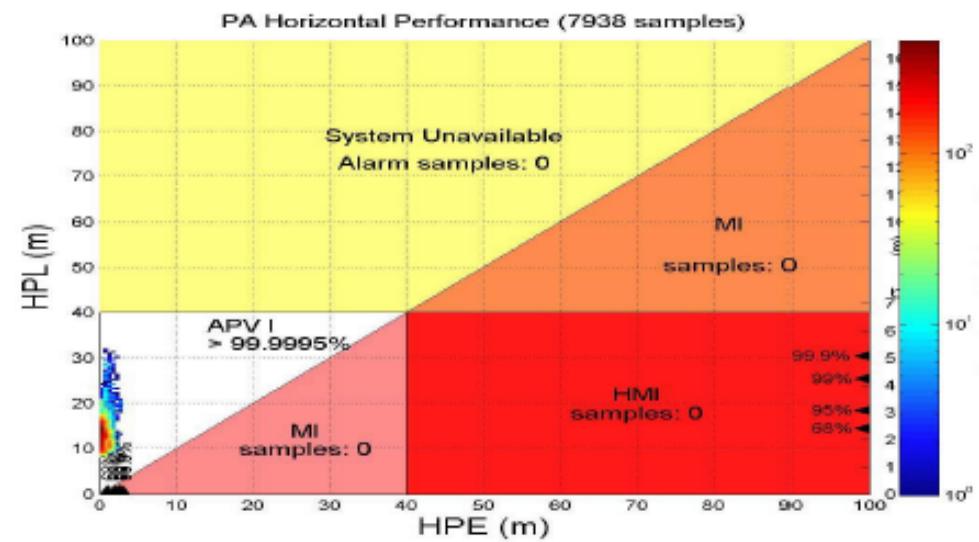
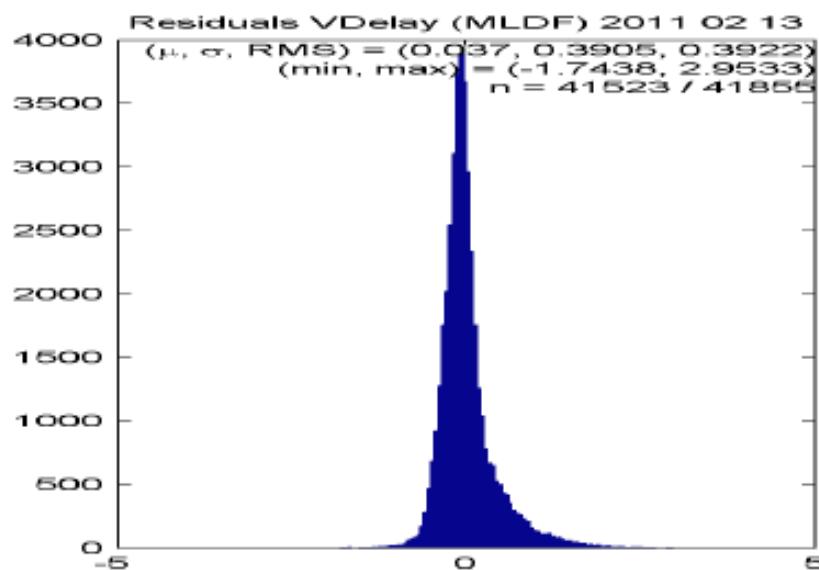
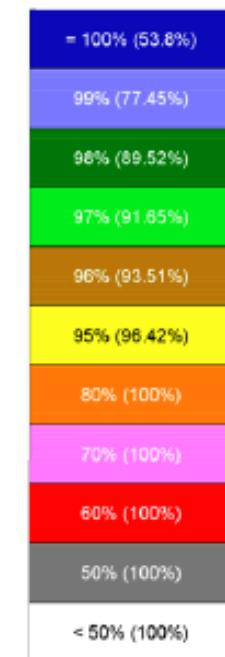
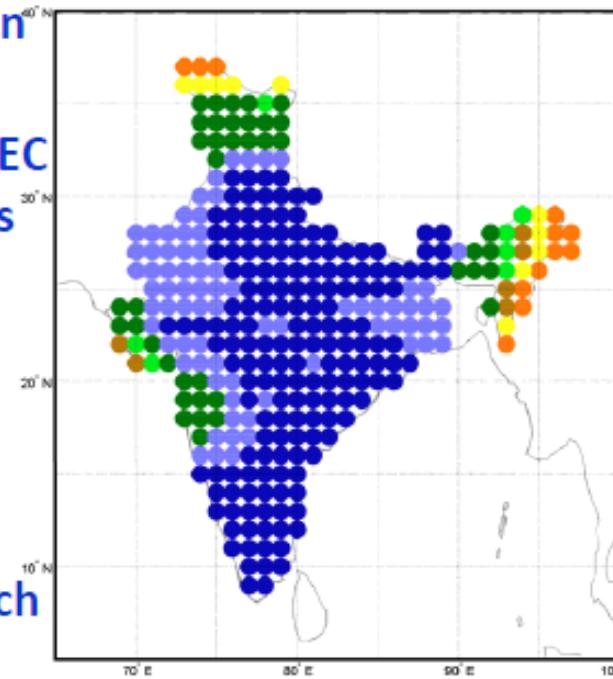
## IGM-MLDF (ISRO GIVE Model-Multi Layer Data Fusion):

- New Multi layer Grid Based Model
- Uses Data Fusion Technique
- Provides GIVD, GIVE at 350 km (as per MOPS)
  - This model does not call for any change to the existing SBAS message structure
  - No change for the legacy users



# IGM-MLDF: Preliminary Results

- Joint activity between ISRO and Raytheon
- Algorithm evaluated using the measurement based model data (ISRO-TEC data) for selected nominal & stormy days over 2004-07
- Algorithm is under evaluation with high quality live data collected over 15 INRES locations
- The live data (supertruth) contains measurements from 2/3 INREEs from each of the INRES





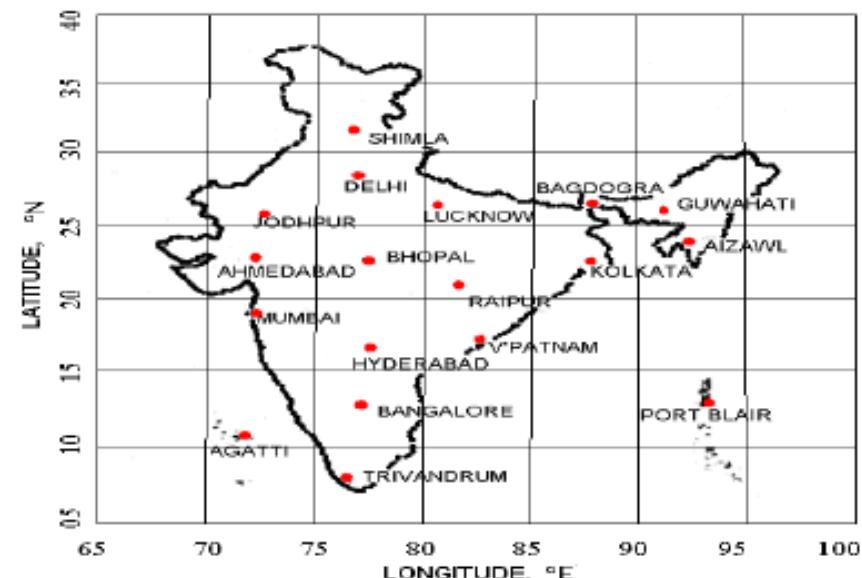
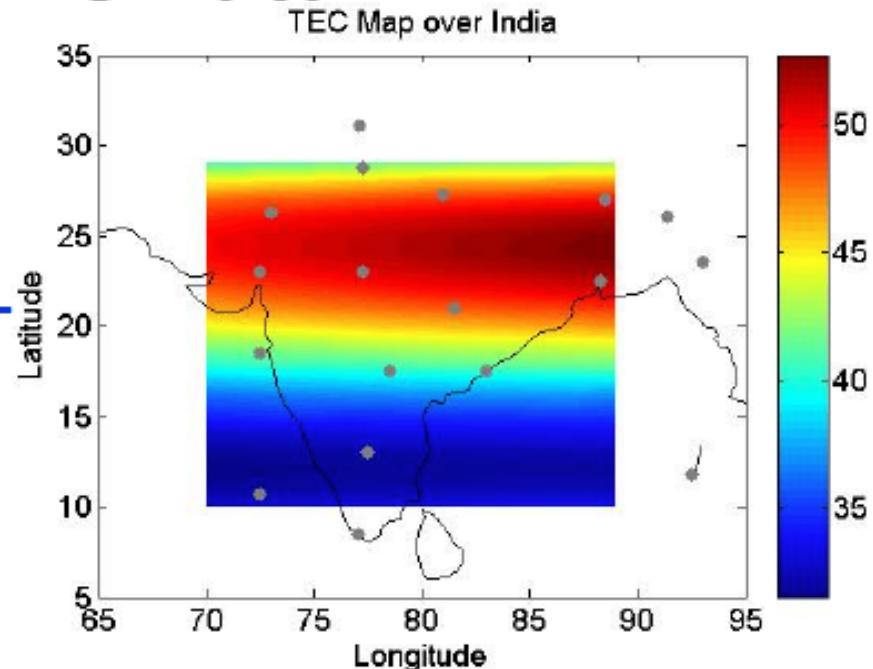
# GAGAN TEC Data

## Measurement Based Ionosphere Model - ISRO TEC Model

- Based on real data from GAGAN TEC stations combined with Physics-based Semi-empirical Model
- Provides TEC Map over Indian region

### Activities include:

- Data Collection and Archival
- Model Comparison and Validation
- Testing New Algorithms



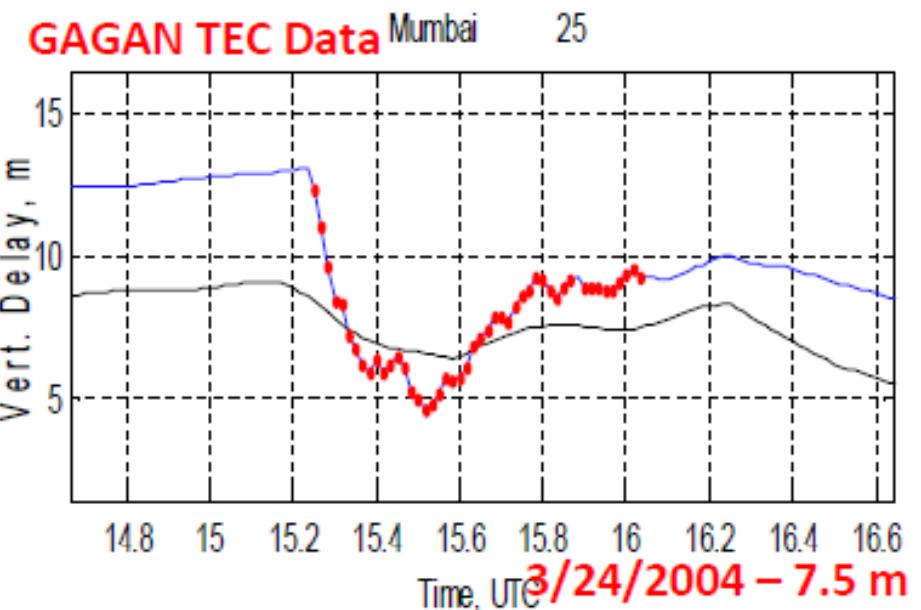
# Depletion Studies

- **Depletion:**

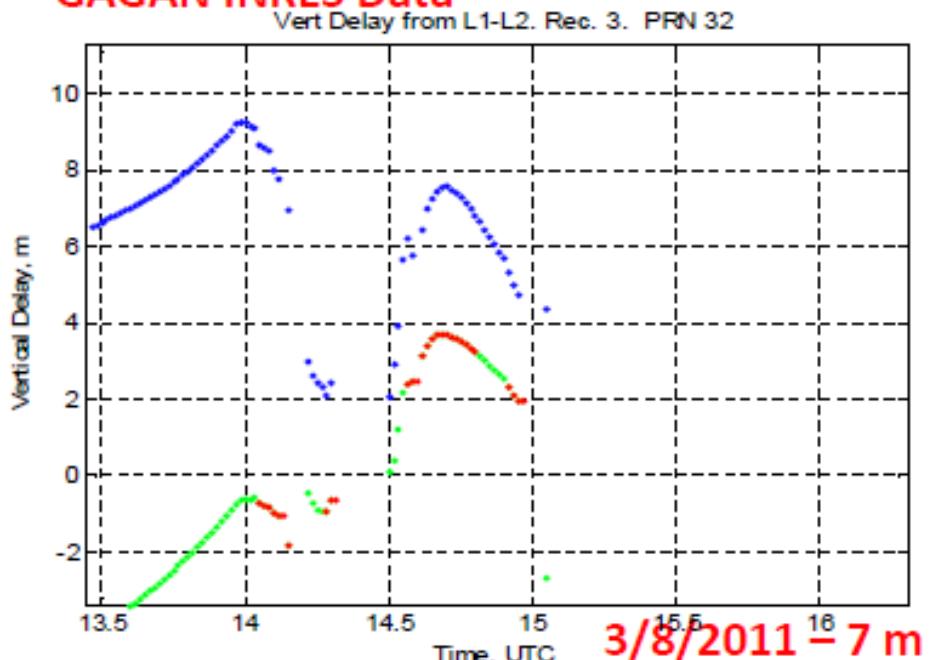
- Post sunset events
- Local phenomena (i.e., Depletion observed by an user may/may not be observed by any of the INRES )
- User protection against depletion: Mitigated through the ground model

- **Observations:**

- Indicates that the depletion with >0.5 m magnitude occur during 5:30 p.m. – 3:30 a.m. local time.
- Maximum observed depth: 7 - 7.5 m



**GAGAN INRES Data**



# GAGAN payload on GSAT-8: Completion of ground and in orbit test



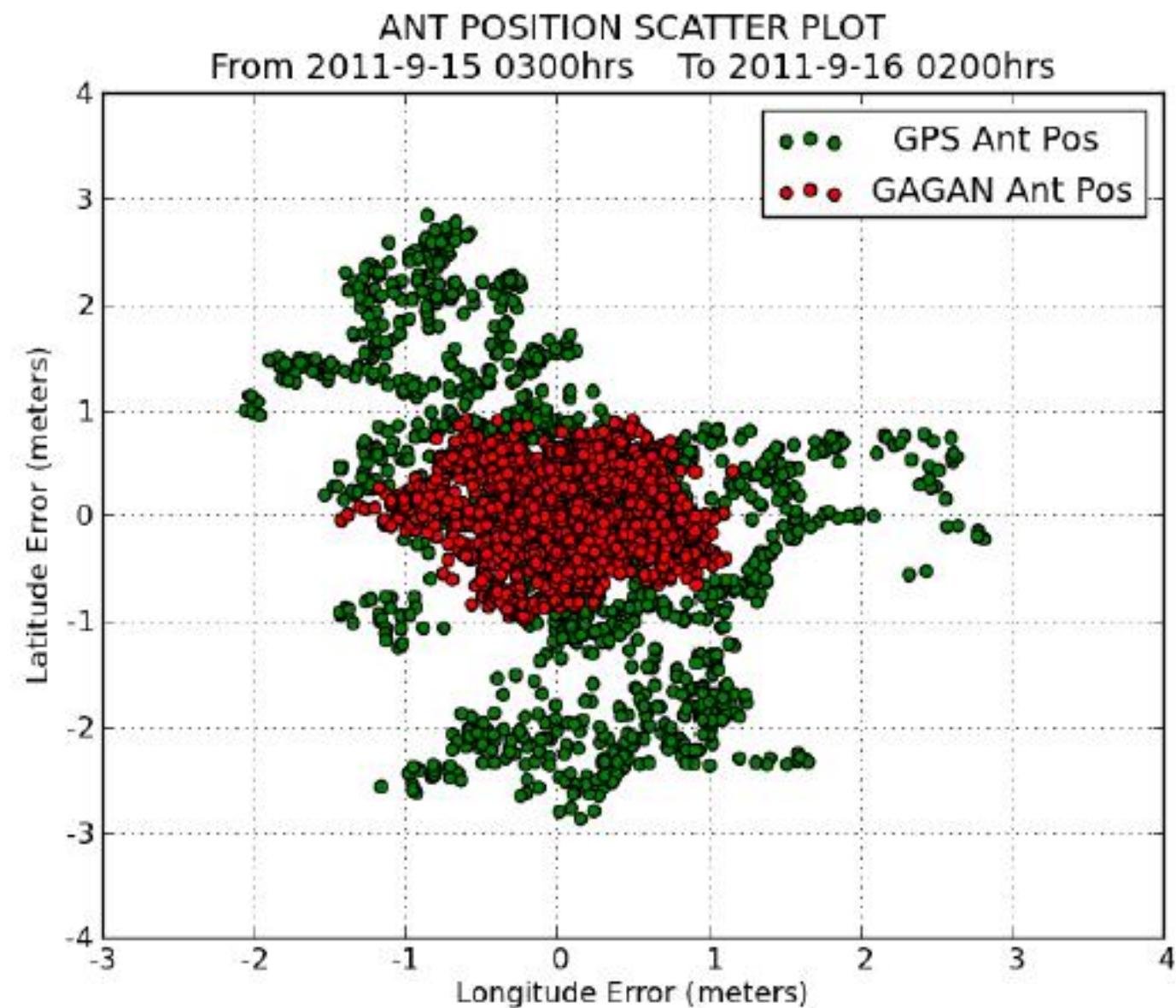
- GSAT-8 carrying GAGAN Payload launched on 21 May 2011
- INLUS RFU Subsystem Tested and Evaluated
- GSAT-8 GAGAN Payload IOT Completed



1<sup>st</sup> Signal In Space from GAGAN payload received on 13-Sep-2011

# Preliminary Performance using SBAS Receiver

GAGAN versus GPS scatter plot and position uncertainties (15-Sep-2011)

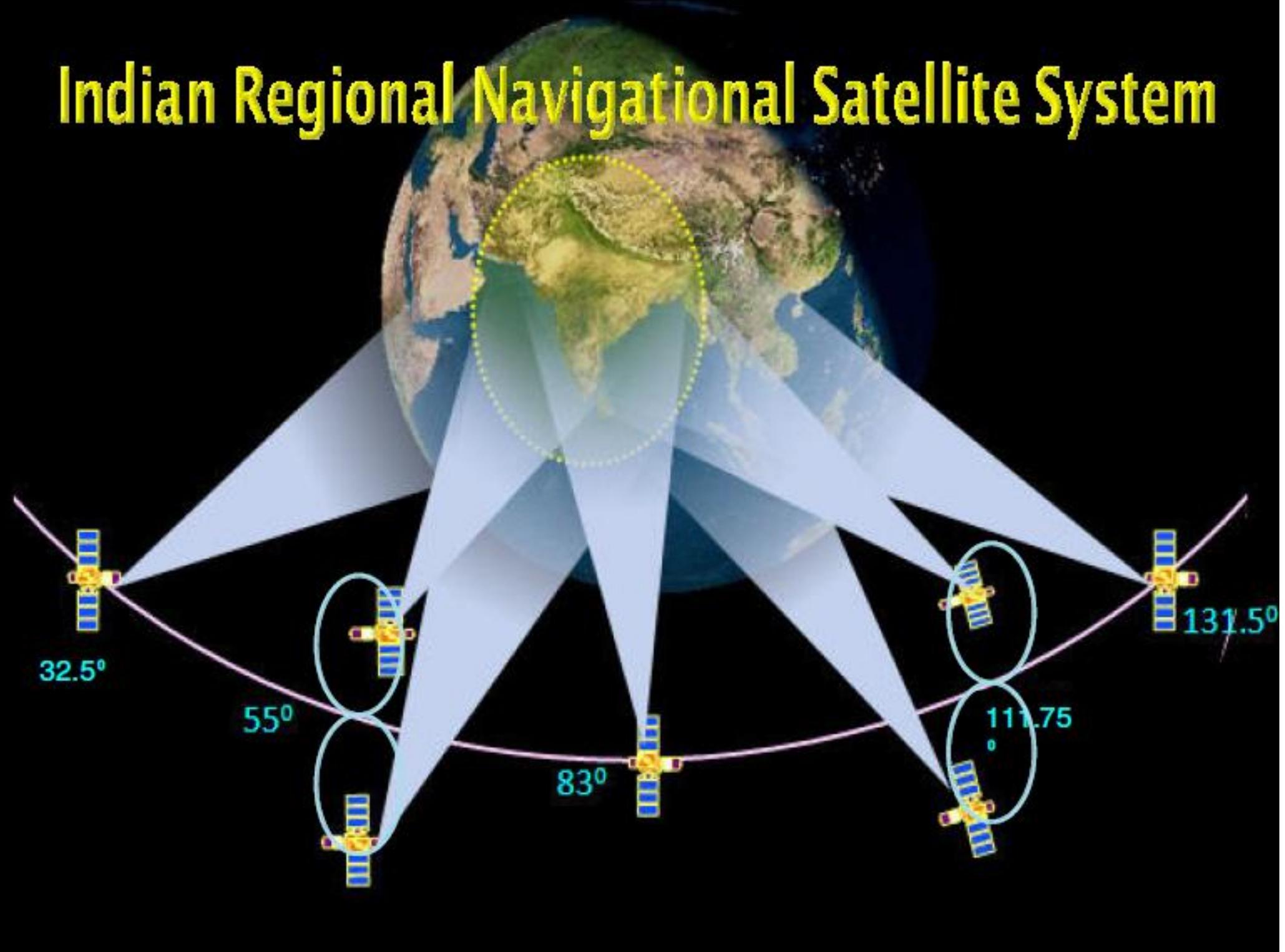


## GAGAN - CURRENT STATUS

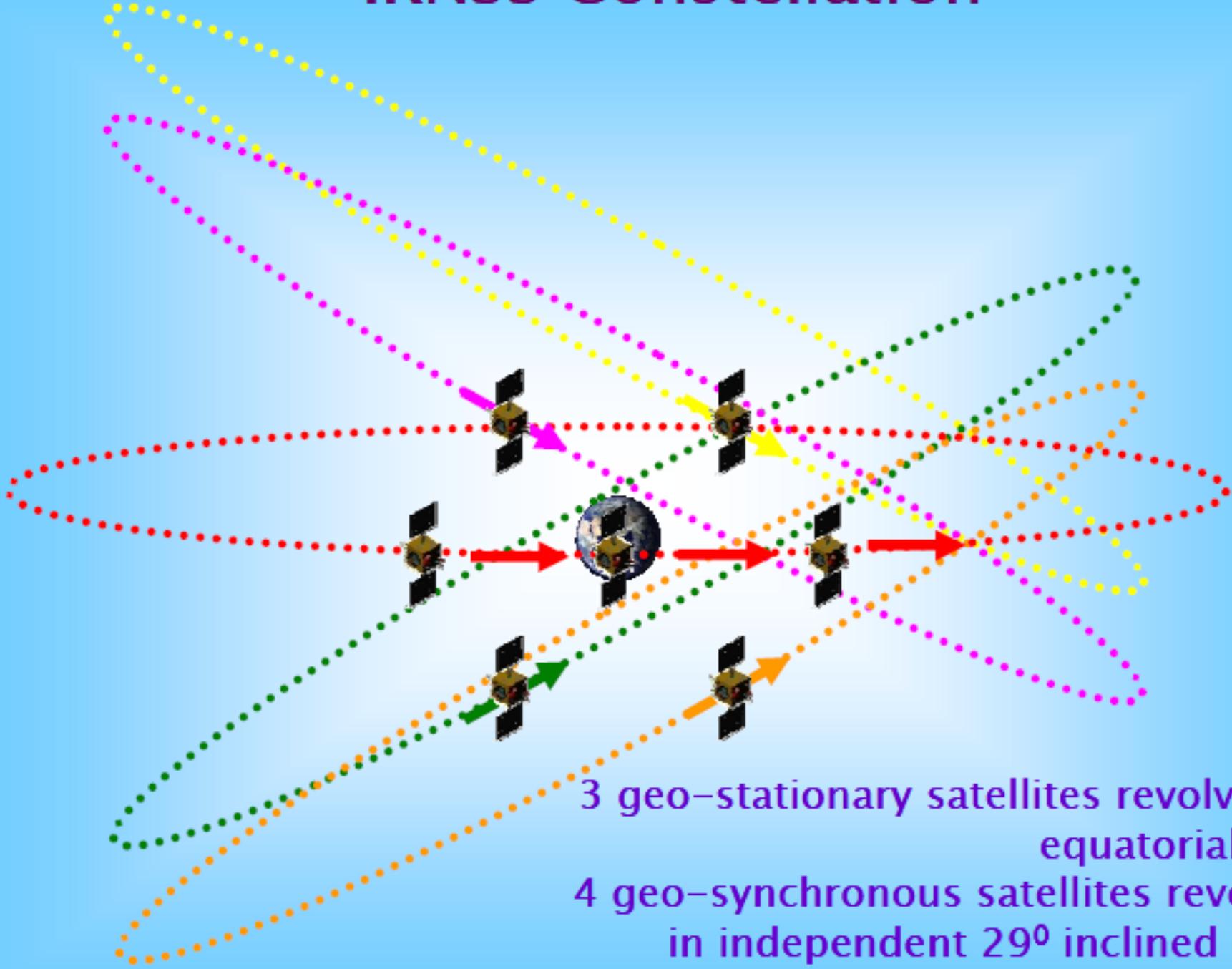
- After the GSAR-8 is launched, flight test with small aircraft has been carried out for all the four parameters Accuracy, availability, integrity and continuity.
- Satisfactory results are achieved for CAT- I & II approaches.
- Enroute – 250 m Decision Height
- Integrity (MCS – Satellite – User) – 6 Sec has been achieved.
- GSAT -10 (PRN 128) will be launched between July-December 2012.
- The three Ionospheric models developed by ISRO are successfully tested.
  1. MLDF (Multi Layer Data Fusion) Model
  2. TEC model
  3. Depletion model
- The installation of ground based equipment are in the final stage
- The system is expected to be certified and operational in the year 2013



# Indian Regional Navigational Satellite System



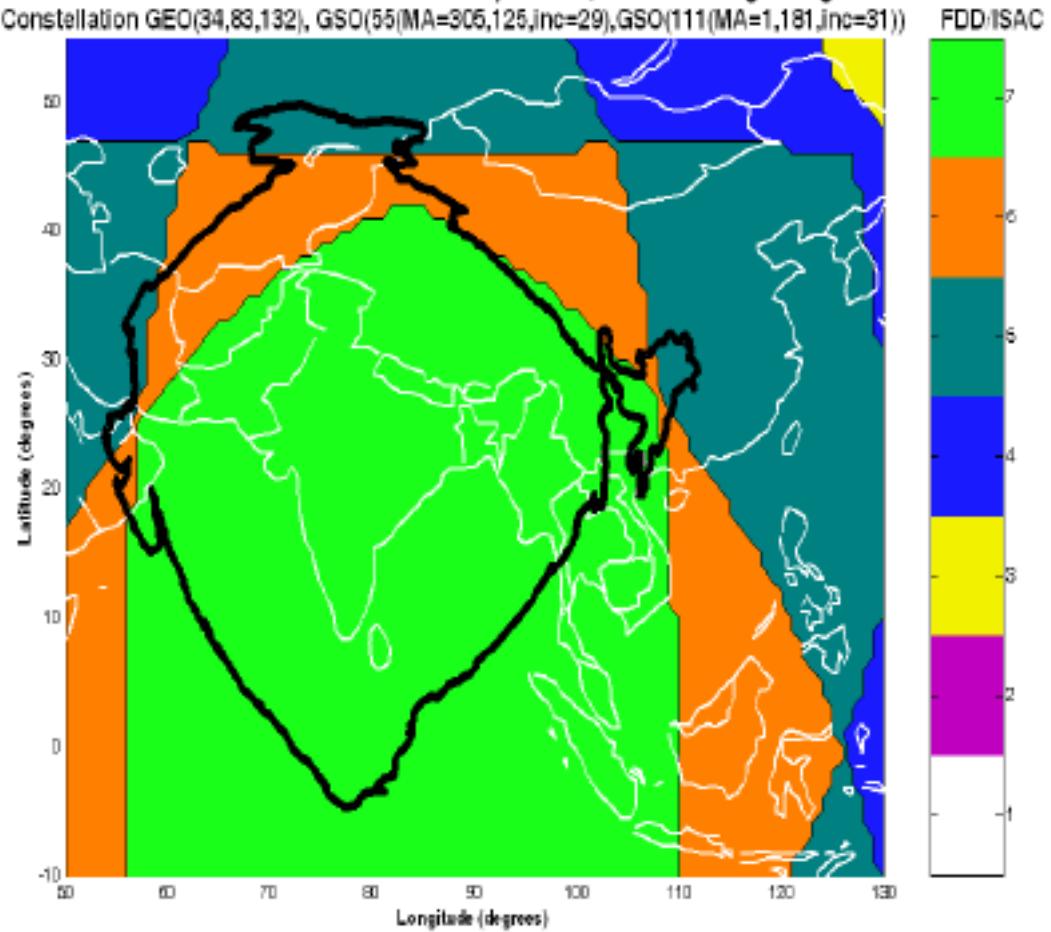
# IRNSS Constellation



# IRNSS SERVICE AREA DEFINITION

## IRNSS Visibility

IRNSS Constellation Visibility for 95%, User Mask Angle 5deg  
Constellation GEO(34,83,132), GSO(55(MA=305,125,inc=29),GSO(111(MA=1,181,inc=31))



S. No	Service Area	Description of the service area
1	Primary Service Area	Indian land mass and 1500 km from Indian geopolitical boundary

- Targeted position accuracy better than 20 metres



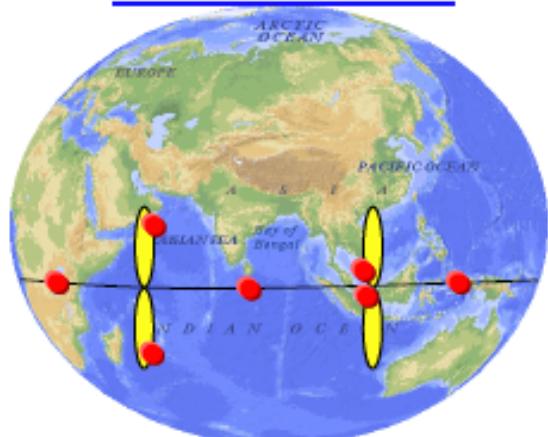
# IRNSS: INDIAN REGIONAL NAVIGATION SATELLITE SYSTEM

- Refers to an independent Indian Regional Navigation Satellite System providing navigation services over Indian Region.

## IRNSS ELEMENTS

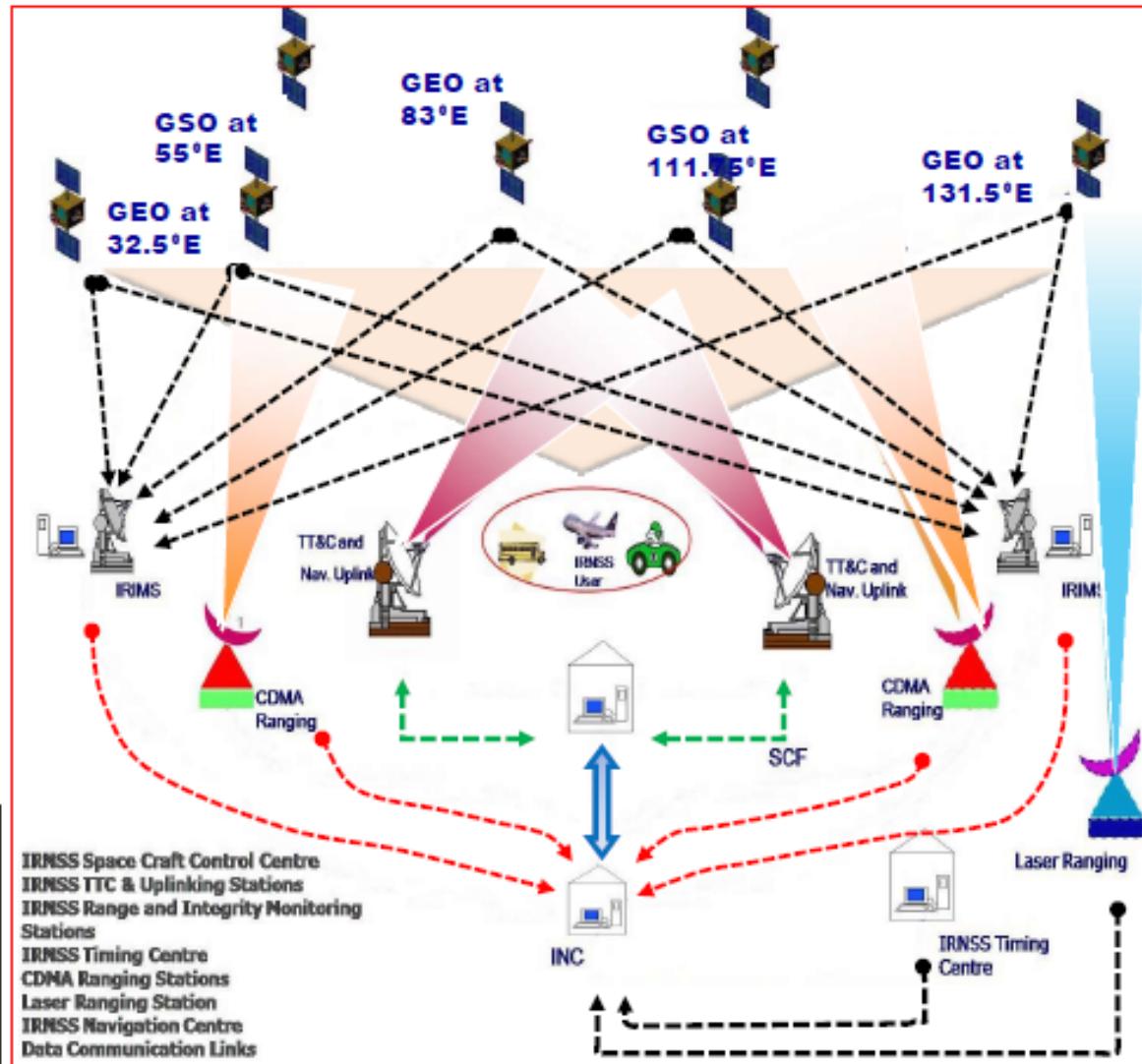
- Space Segment
- Ground Segment
- User Segment

## SPACE SEGMENT



## IRNSS SERVICES

Service Type	Freq. Band
Standard Positioning Service	L5(1176.45Mhz) S(2492.028hz)
Restricted Services for Special Users	L5 S



**IRNSS Configuration**

# Space Segment

- ▶ Space Segment consists of Seven satellites
- ▶ 3 Satellites in Geo-Stationary orbit at  $32.5^\circ$ ,  $83^\circ$  and  $131.5^\circ$  East.
- ▶ 4 Satellites in GEO Synchronous orbit placed at inclination of  $29^\circ$  with Longitude crossing at  $55^\circ$  and  $111.75^\circ$  East.
- ▶ Two spare satellite are also planned.
- ▶ The Satellites are specially configured for the Navigation. Same configuration for GEO and GSO which is desirable for the production of the satellites. Production plan & schedule are worked out.
- ▶ IRNSS Satellites are to be launched by the Indian launcher PSLV.
- ▶ The first Satellite will be launched in 2012. The subsequent launches are planned once in Six months. The full constellation will be operational by 2015.

# IRNSS SPACE SEGMENT

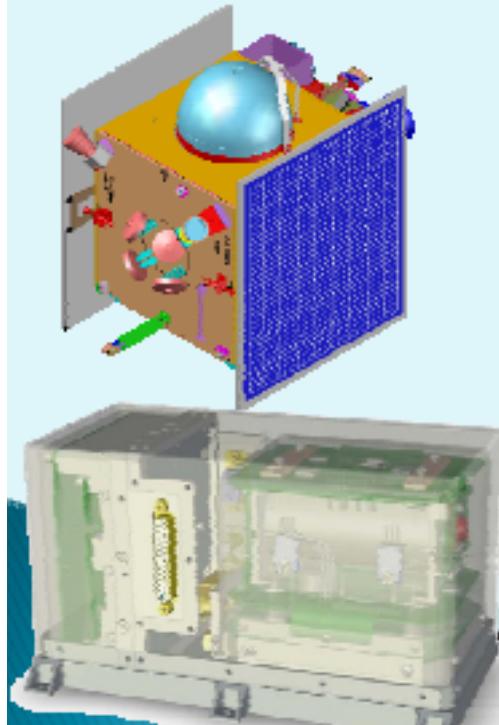
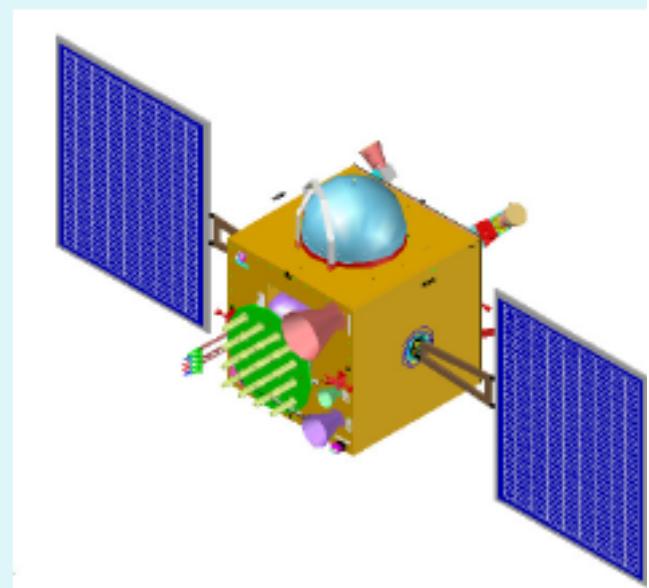


Indian Space Research Organisation



# IRNSS Satellite

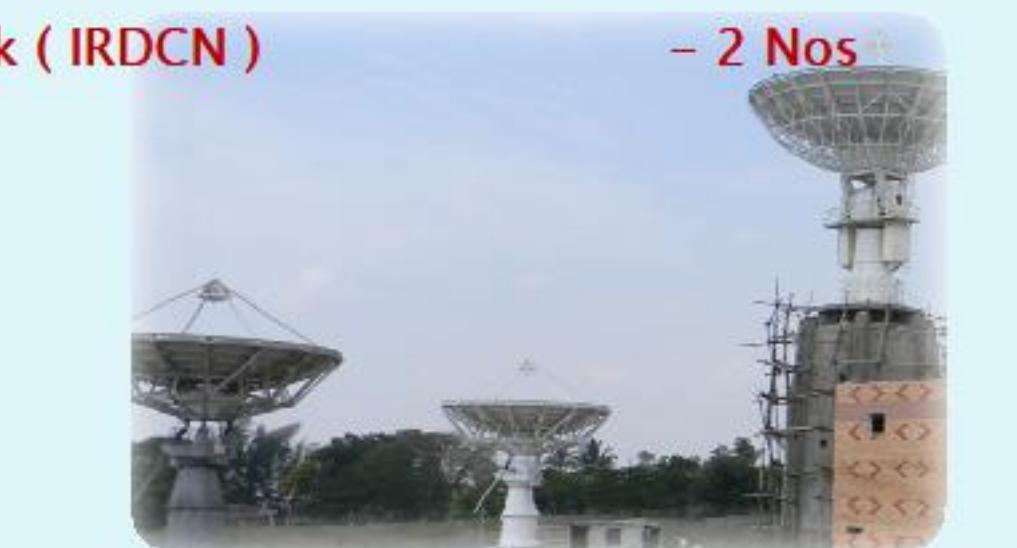
- ▶ IRNSS Satellites are designed around I-1K bus.
- ▶ Dry mass of around 600 kgs and lift off mass of 1425 kgs
- ▶ Power generation capability of 1600 W and payload power requirement of 901 W
- ▶ Navigation Payload Transmits SPS and RS signals in L5 and S Bands.



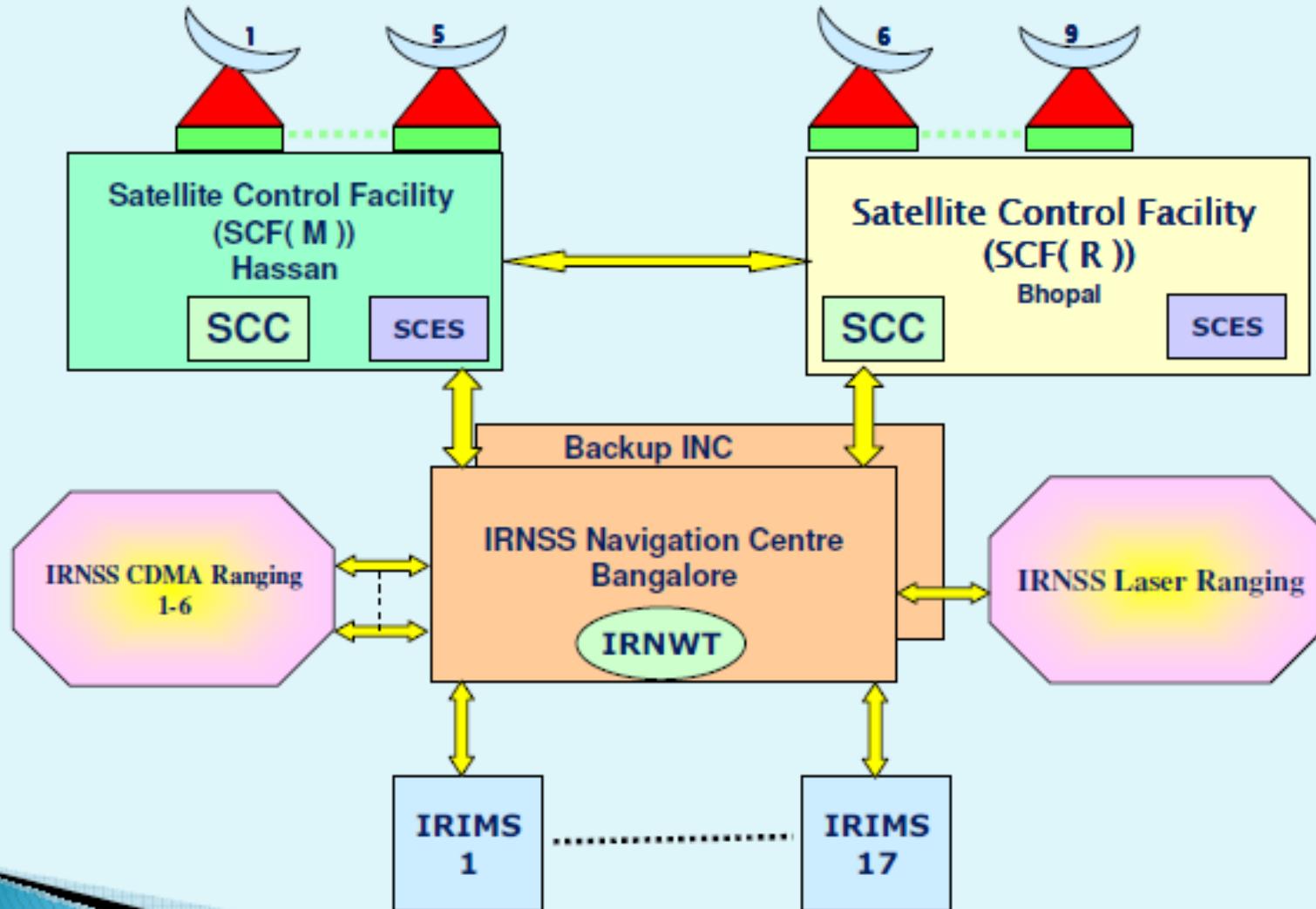
- Payload utilizes highly stable Atomic Frequency Standards for generation of Navigation Signals.
- 3 Axis control of the satellite with Yaw steering capability to optimize the use of Solar Panels and to support the thermal control of the satellite

# Ground Segment Subsystems

- IRNSS Satellite Control Earth Stations – 9 Nos
- IRNSS Satellite Control Centre ( IRSCL ) – 2 Nos
- IRNSS Range and Integrity Monitoring Stations ( IRIMS ) – 17 Nos
- IRNSS Navigation Centre ( INC ) – 2 Nos
- IRNSS Network Time ( IRNWT ) – 2 Nos
- IRNSS CDMA Ranging Stations ( IRCDR ) – 4 Nos
- IRNSS Data Communication Network ( IRDCN ) – 2 Nos



# Ground Segment Architecture

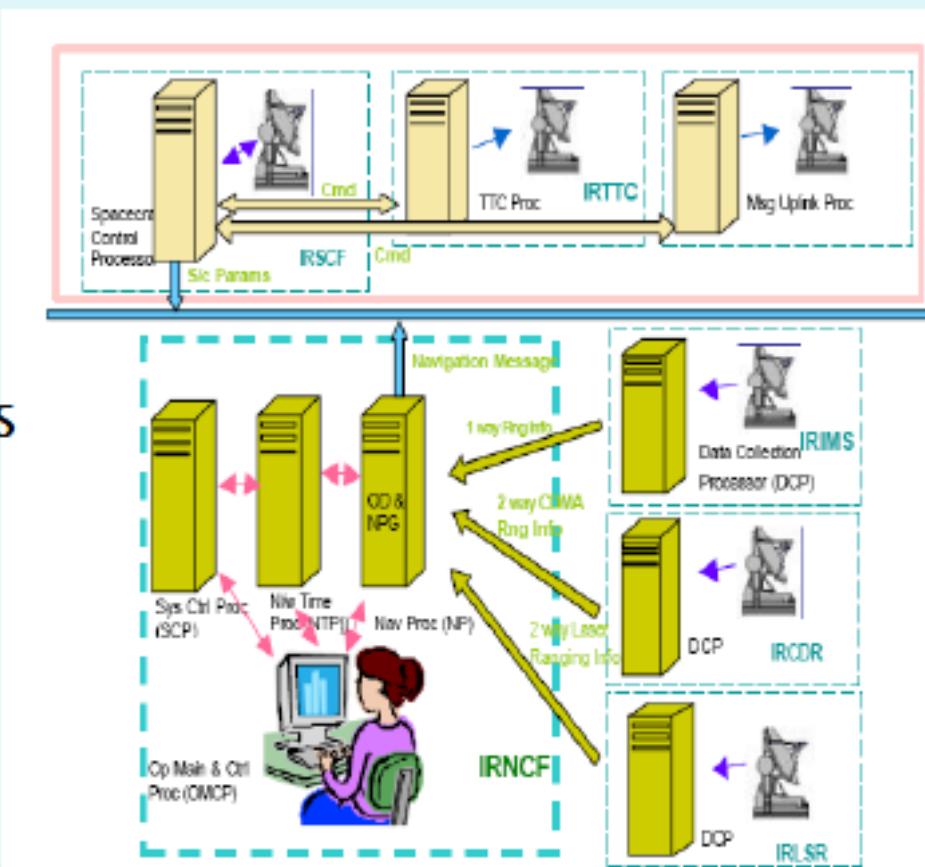


# Major Elements Of Navigation Software

Navigation software will be deployed in the IRNSS Navigation Centre. The software modules interface with various subsystems of the ground segment and generate navigation parameters, required for broadcast from spacecraft.

Navigation software generates the following in 4 sub-frames:

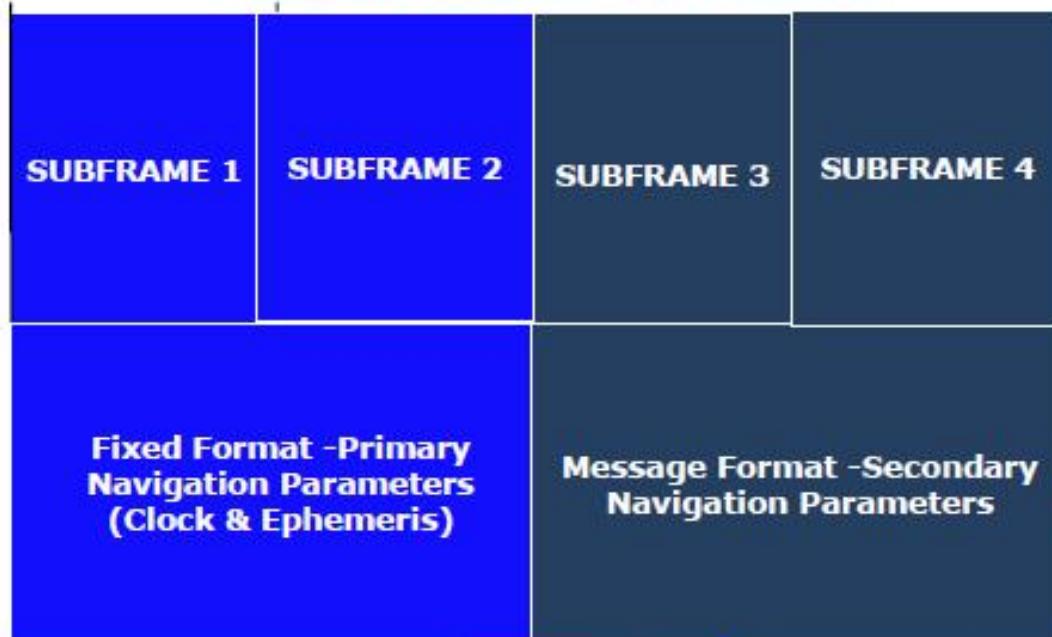
- Primary Parameters
  - Satellite ephemeris , clock
  - Satellite health status & accuracy
- Secondary Parameters
  - Satellite almanac
  - Ionospheric corrections-coefficients
  - IRNSS time difference w.r.t UTC /GNSS
  - Ionospheric gird delay parameters
  - Encryption keys
  - Text messages
  - Differential corrections
  - Earth orientation parameters
  - Auto Nav. Parameters





# Ionosphere Corrections In IRNSS

## IRNSS Data Structure



1200 bits



1. Co-efficient based ( $4\alpha$  &  $4\beta$ ) Iono. Corrections
2. Grid Based Iono. Corrections (on L5)



# User Segment

- ▶ The user segment consists of IRNSS receivers operating in
  - Single Frequency ( L5 or S band)
  - Dual Frequency (L5 and S band)
- ▶ Single frequency and dual frequency receivers shall receive both SPS and RS signals.



## IRNSS SERVICES

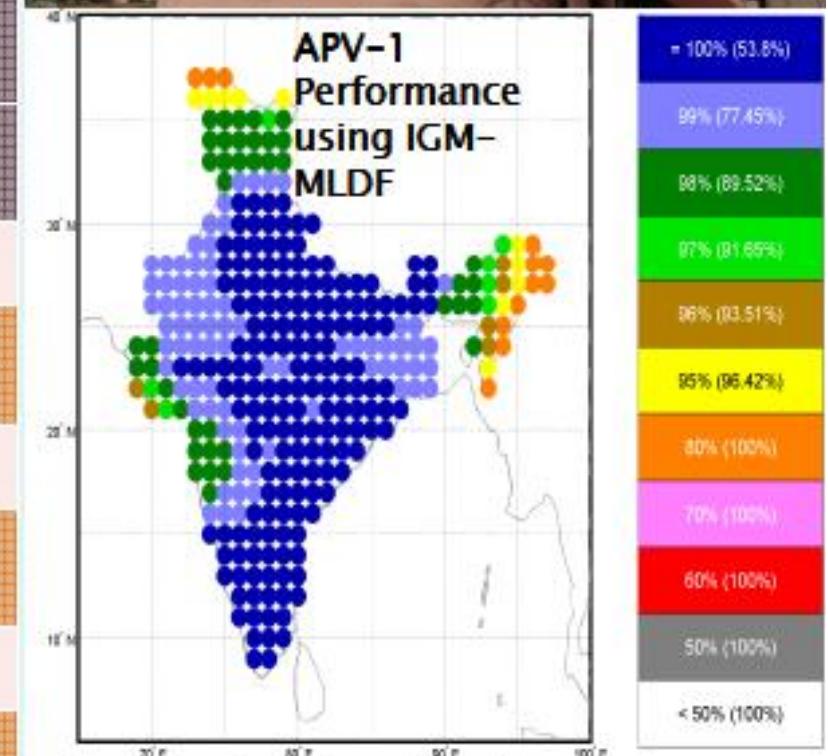
Service Type	Signals	Frequency Band
Standard Positioning Service	1 MHz BPSK	L5 (1176.45 MHz) S (2492.08 MHz)
Restricted Services for Special Users	BOC(5,2)	L5 (1176.45 MHz) S (2492.08 MHz)

# ISRO Participation to International GNSS Forum

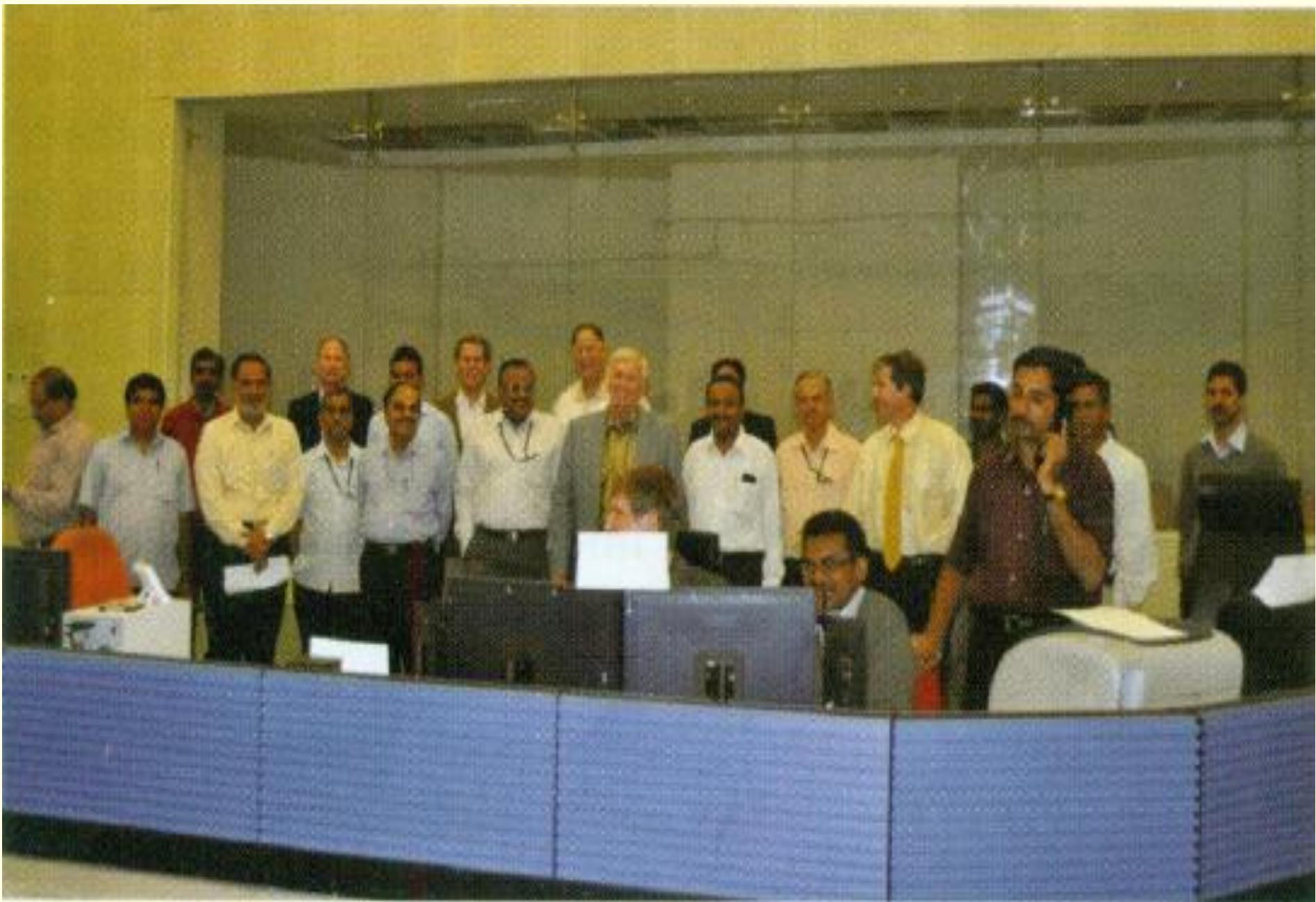
- ▶ ISRO is an active member of International Committee of GNSS (ICG).
- ▶ ISRO participated in many bilateral discussions with GPS, GALILEO etc. on Interoperable and compatible signal structure.
- ▶ ISRO is hosting a QZSS Monitoring Station site in ISTRAC.

# Major Activities (GAGAN)

Completed	Planned Activities
<b>Ground Elements</b>	
INLUS-2 at Bangalore: Completed	Test & Evaluation
Proto M&C system for INLUS RFU	Development in DO-178B
DELHI (Nangloi) INLUS-3 FMA: Civil works initiated	Commissioning by Apr'12
<b>Ionosphere Algorithm</b>	
ISRO developed IGM-MLDF model  Agreed by AAI, DGCA, RTN	Incorporation into operational system  Review by TRT for Certification
<b>Space Segment</b>	
GSAT-8 launch, IOT & integration	GSAT -10,15 launch & integration
<b>System Acceptance Test</b>	
Preliminary Completed in Dec'10	Final during Jun'12
<b>Certification by DGCA</b>	
Work Initiated	Completion by Jun'13



# GAGAN Task team with Raytheon team



Space Physics Laboratory, Andhra University, Visakhapatnam, INDIA



