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TIDAL COSMIC RAYS IN GALACTIC FLY BYS Tijana Prodanović, University of Novi Sad

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SEARCHING FOR DARK MATTER

- × (In)Direct
- Experiments DAMA, LHC, SuperCDMS, IceCube etc.
- Gamma-ray signals

 lines or excess in background



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- × (In)Direct
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x See anything weird?





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I SEE WEIRD LITHIUM

- One of the primordial elements (D,He,Li)
- Primordial nuclear
 reactor
- × Initial conditions+physics
 →abudances↔observations
- Big Bang Nucleosynthesis testing physics and cosmology





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PRIMORDIAL ABUNDANCES

- × Schramm plot
- Abundance vs baryon density
- × Curves: BBN predictic_z
- × 1σ uncertainty nuclear reaction rates
- × CMB (*WMAP*: Dunkley 2008)
 - fixes baryon density
- Read off abundance
- x Simple!(?)



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THEORY VS OBSERVATIONS?

- × BBN theory curves
- CMB baryons
- Observations boxes
- ⁴He OK ⁽ⁱ⁾
 D right on! ⁽ⁱ⁾
 ⁷Li problem! ⁽ⁱ⁾
 Factor of 3-4 discrepancy!

LITHIUM PROBLEM! Note: Planck gives lower baryon density – slightly lower Li discrepancy!







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SOURCES OF LITHIUM

× ⁷Li sources

+ BBN

+ cosmic-ray interactions (ingredients: shock waves, magnetic field, charged particles) (Reeves 1970)

- × fusion $\alpha + \alpha \rightarrow^{6,7} Li$
- × spallation $p, \alpha + \text{CNO} \rightarrow \text{LiBeB}$

 + type II supernovae (neutrino process – neutrinoinduced spallation in C shell – ¹¹B, ¹⁰B i ⁷Li)

× ⁶Li sources

- + cosmic-ray interactions mainly (Reeves 1970)
- + in accretion disks of micro quasars (locco & Pato 2012)



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MEASURING PRIMORDIAL LITHIUM

- × Spite & Spite 1982
- × Old, warm, metal-poor halo dwarfs
- (Almost) uniform Li abundances over a range of low metallicities
- × Very low scatter!× Primordial plateau





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plateau

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LITHIUM PROBLEM



Li is destroyed instead of expected post-BBN production!?



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WHO IS TO BLAME?

- × Observations?
- Measured abundances are wrong?
- Stellar modeling problems?

- **×** Theory?
- Measured aboundances are ok?



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WHO IS TO BLAME?

- × Observations?
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Blame it On Dark Matter Everyone Else Does

- × Theory?
- Measured aboundances are ok?

 Find a way to destroy Li pregalactically
 New physics?



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MAIN ISSUES

- × How big of a problem?
 - + Contamination due to other sources of 6,7Li?
 - + Lithium problem even worse?
- Sure we should not blame stellar observations?
 + Need idependent confirmation
- x Stellar or early, new physics destruction?
 - + Measured abundances are below predicted values over a large range of low metallicities
 - + Uniform destruction of Li in different stars at different metallicities?



CONTAMINATION?

- Cosmological cosmic rays
 - + Shock waves during structure formation (eg. Miniati 2000)
 - Hope to observe on clusters (eg. Colafrancesco & Blasi 1998, Brunetti et al 2012)
 - + Li (6,7) production
 - Such Li is a contaminat to primordial Li content in halo stars (Suzuki & Inoue 2002)
 - + Must correct for larger discrepancy!
 - Note: Also adding to extragalactic gamma-ray background – foreground for DM signals! (Loeb & Waxman 2000)









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CONSTRAIN CONTAMINATION!

- x Can use gamma-rays as a probe!
- Same source from hadronic CRs
- x Lithium-gamma-ray connection

$$\frac{\gamma_{\pi}}{\text{Li}} \propto \frac{1}{y_{\alpha,ism}y_{\alpha,cr}} \frac{\langle \sigma_{\gamma} \rangle}{\langle \sigma_{\alpha\alpha} \rangle}$$

(Fields & Prodanović 2005)



- Use extragallactic gamma-ray background (EGRB) to constrain pre-Galactic Li production
- But must have a model of structure formation cosmic ray contribution to the EGRB (Dobardžić & Prodanović 2012, Kakabaze et al. in preparation)



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FIND NEW SITES!

- × New site where Li was measured !
- × Small Magellanic Cloud (Howk et al 2012)
 - + Metallicity ~ 0.20 solar
 - Measured gas phase Li abundance in its ISM (and isotopic ratio!)
 - + No worry about stellar modeling
 - + Independent test of the state of the Lithium problem!





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SMALL MAGELLANIC CLOUD

- × Howk et al. 2012, Nature
- Measured abundance consistent with teoretical primordial value (above observed plateau)
- x Li problem cause stellar modeling?





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BUT...CLOSE FLY-BYS

- Interractions, collisions and close fly-bys impact galaxy morphology
- Especially pronounced on small satellite galaxies disrupted by large ones (eg. SMC & MW)





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TIDAL SHOCK WAVES

- Solve the second stress of the second stress of
- Close fly-bys tidal shock waves
- * "Magellanic stream" due to tidal interaction between Magellanic clouds ~ 2Gyr ago? (Diaz & Bekki 2012)





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TIDAL COSMIC RAYS

Tidal shock waves + magnetic fields = tidal cosmic rays (TCR)

- × Nucleosynthesis LiBeB production
- In low metallicity systems
 - + no CNO
 - + only ^{6,7}Li production contamination! Must correct for!

 Gravitational interactions energetically sufficient

 $E_{kin,SMC-MW} \sim 10^{58} \mathrm{erg}$

$$E_{^{6}Li,SMC} \sim 2 \times 10^{56} \mathrm{erg}$$





- How much Li can be produced in these interactions?
- Toy model compare to/scale with supernova shock waves









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VS



- Short but intensive cosmic-ray injection
- Large scale shock waves
 Not dirrectly/imediately accompanied with metallicity increase



- Continuous injection of cosmic rays
- Small scale shock waves
- Directly accompanied with
- metallicity increase



TCR & LITHIUM

× To have Li_{TCR} = Li_{GCR}



 $N_{\rm Li} \propto n_{\rm ISM} N_{\rm CR} \tau_{\rm CR}$

- + how large TCR flux would be needed?
- + how large should a scale of tidal shock be?
- Fraction of total gas mass of the system that tidal shocks sweep over





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TCR, LITHIUM & SMC

- * At solar metallicity, entire system needs to be tidally shocked 8 times in order for TCRs to produce same amount of Li as GCRs!
- **SMC** (metallicity ~ 1/5 solar)
 - + Only 2 interactions are sufficient for TCRs to make as much Li as GCRs!

 $M_{TCR,gas}/M_{gas,SMC} \approx 2$

+ SMC has suffered at least 2 interactions with LMC and 1 with MW!

 But main caveat: TCR duty-cycle could be 10x lower and TCRs 10x less efficient (need numerical modeling)



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IMPLICATIONS: ISOTOPIC RATIO

- × If no TCRs were accelerated in SMC
 - + Expected isotopic ratio $({}^{6}Li/{}^{7}Li)_{SMC,GCR} \approx 0.06$
- If significant fraction (eg. 50%) of total Li abundance in the SMC gas was due to
 - + Isotopic ratio $({}^{6}\text{Li}/{}^{7}\text{Li})_{SMC,GCR+TCR} \approx 0.1$
- × Observed isotopic ratio (Howk et al. 2012)

 $\left({}^{6}\mathrm{Li}/{}^{7}\mathrm{Li}\right)_{SMC,obs} \approx 0.13 \pm 0.05$

Consistent with having another cosmic ray population in SMC!



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IMPLICATIONS: CONTAMINATION

- If significant fraction of total Li abundance in the SMC is due to TCRs
- × Li problem remains!
- × Solution in non-standard physics?





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CONCLUSION

× Primordial lithium problem

+ predict 4x more Li than observed

+ stellar destruction or new physics?

- Li observed in SMC gas ok if post-BBN production low
- × But SMC "survived" few close encounters
- x Tidal shock waves Tidal cosmic-rays!
- * 2 close fly-bys sufficient for TCRs to make same quantity of Li as GCRs made over the system history!
- x Truly primordial SMC Li even lower!
- Implies solution in the form of new physics!?







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QUESTIONS?

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BACKUPS

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FIND NEW SITES!

× Find a better site for measuring Li!

- × High Velocity Clouds (Prodanovic & Fields 204)
 - + (In some cases) Low metallicity (~ 10% solarne)
 - + Little dust
 - + Not complicated by stellar modeling
 - + Idependent test of pre-galactic
 - Li production

 But High ionization, low density, difficult to observe





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NEW (ASTRO) PHYSICS

- × Lithium problem not solved
- No conventional solution...so far
- * ⁷Li problem must be solved without endangering concordance with other elements!
- Sonus: Find sink of ⁷Li which is a source of ⁶Li?
- Dark matter affecting BBN epoch?



- × ⁶Li "plateau"? Not made during BBN!
- × Post-BBN production in CR interactions
 - + Galactic CRs ⁶Li increases with metallicity
 - + Cosmological CR? ⁶Li independent of metallicity







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LITHIUM OBSERVATIONS

Easy in insterstellar medium gas! ⁽ⁱ⁾ + Both isotopes nicely deparated





Complicated in stars 🛞

Asplund et al. 2006

 ⁶Li shows as small asymmetry on ⁷Li line – isotopic ratio infered from this



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SOLUTION?: LITHIUM OBSERVATIONS

- × Absorption lines in stellar atmospheres
- × Modeling! Non-LTE, 1D vs. 3D
 - + Li mostly ionized in atmospheres
 - + Must determine Li II/Li I ratio
 - + Strong temperature dependance



- Much larger temperature needed to solve Li problem? ΔT ~ 500-600K?
 - + Affects other element abundances (Be, B, O)
 - + Casagrande et al. 2010 new, detailed analysis of T scale gives at most $\Delta T \sim 200 K$



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SOLUTION?: MIXING LITHIUM

× Lithium easily burned in stars $T > 2.5 \times 10^6 \Rightarrow^7 Li + p \rightarrow 2^4 He$ $T > 2 \times 10^6 \Rightarrow^6 Li + D \rightarrow 2^4 He$



Destroyed by convection!
 Deeper mixing of surface material – Li burning
 If ⁷Li is destroyed – ⁶Li destroyed even more!
 But not enough and not uniformily!
 Different stars – different convective zones
 Should see larger scatter around "plateau"!



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IMPLICATIONS: RADIO EMISSION

× Enhanced radio signal due to TCR presence

- + expect strong synchrotron emission
- Note: tidal shock waves would trigger star-formation, followed by enhanced GCR flux, but at the later epoch
- Enhanced radio signal detected from M51 interacting system!
 + especially pronounced in smaller of interacting galaxies!
 + due to TCRs?



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IMPLICATIONS

If significant fraction of total Li abundance in the SMC is due to TCRs

 Li measured in SMC gas is not representatiove of close-to-primordial abundance and should not be compared to MW as it is – must make

downward correction for this contamination!

+ Li problem remains!
+ Observations and stellar destruction not a problem!





CONCLUSION

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SMC "survived" few close encounters with LMC and MW



- × Tidal shock waves Tidal cosmic-rays!
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- × SMC Li contaminated by TCRs
- x Truly primordial SMC Li even lower!
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