

Topics in theoretical physics I

self-organization on scale hierarchy

Z. Yoshida (The University of Tokyo)

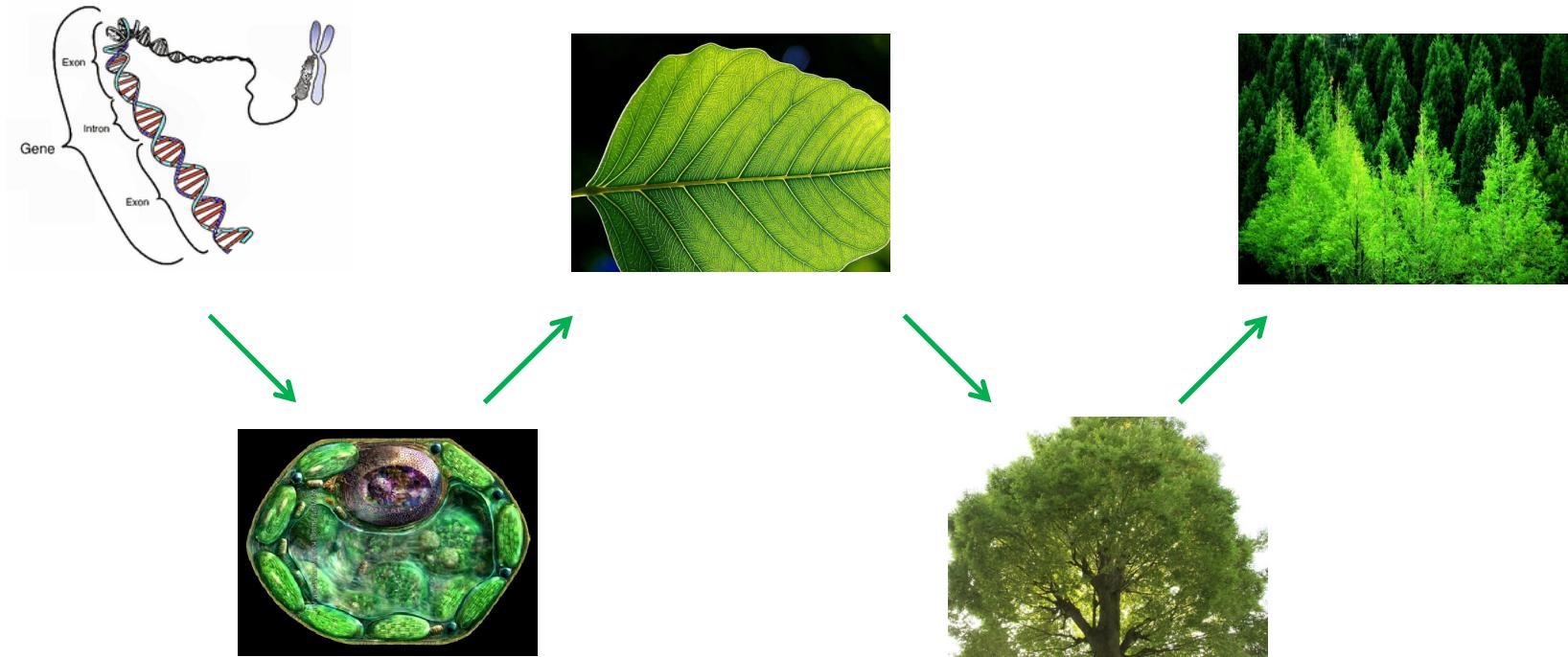
Collaborators:

S.M. Mahajan (U. Texas), P.J. Morrison (U. Texas),

R.L. Dewar (ANU), F. Dobarro (Trieste)

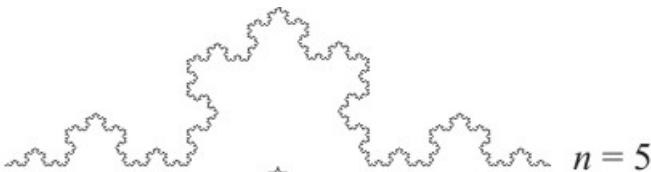
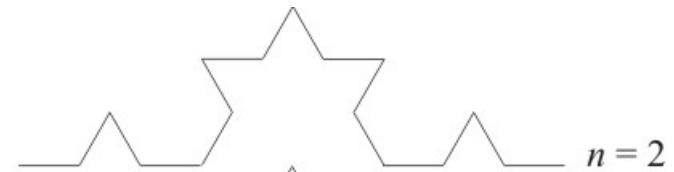
N. Shatashvili (TSU)

Scale hierarchy



- Surprisingly diverse structures and mechanisms on different scales
- How “macro” can be different from “micro” ?
- Physics: What is the mechanism that works without **blueprints** ?

subject defines a scale



$n = 1$

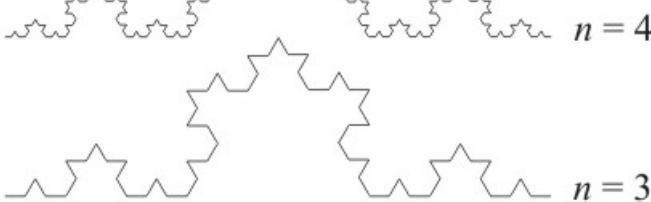
$n = 5$



$n = 4$



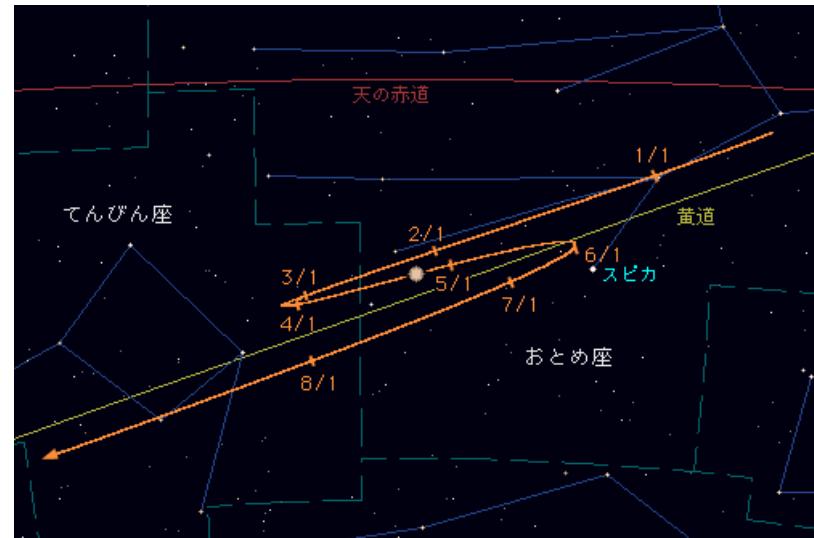
$n = 3$



The definition (representation) of matter depends on the subject=scale.

Space-time: *a priori* of theory (?)

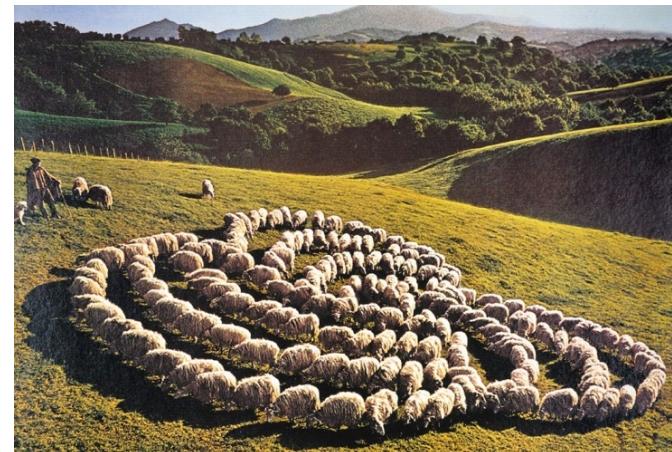
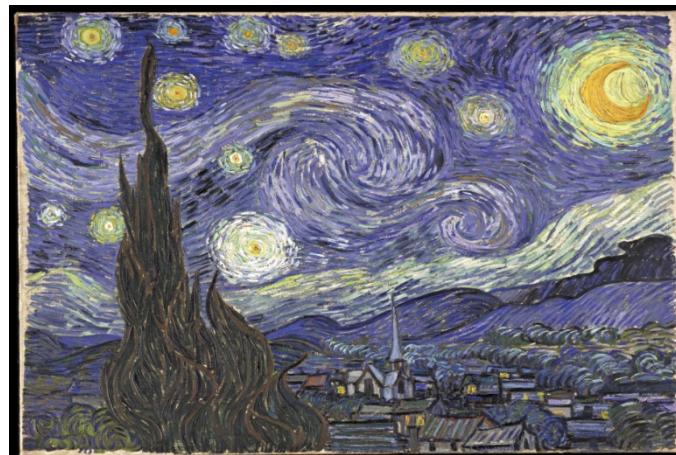
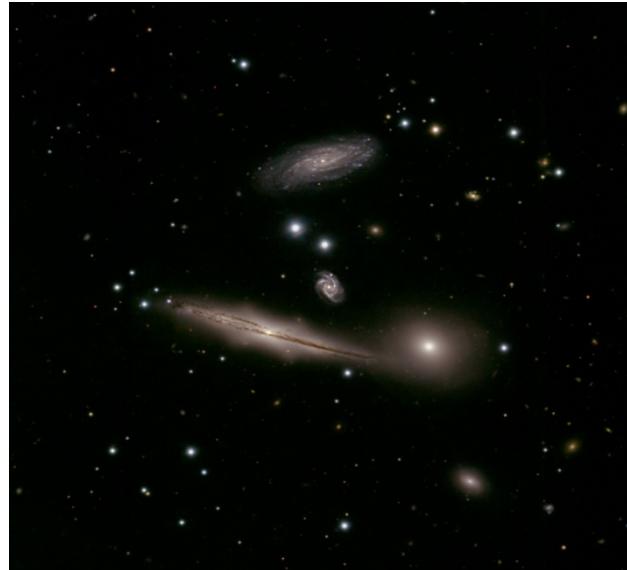
- *Kopernikanische Wendung*
a priori of recognition



Immanuel Kant
(1724-1804)

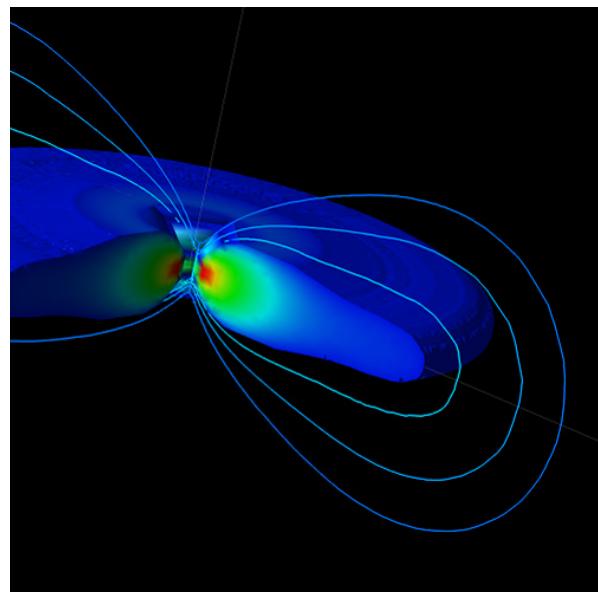
phenomena (event) → framework of recognition (space-time)

Vortex = space-time distortion

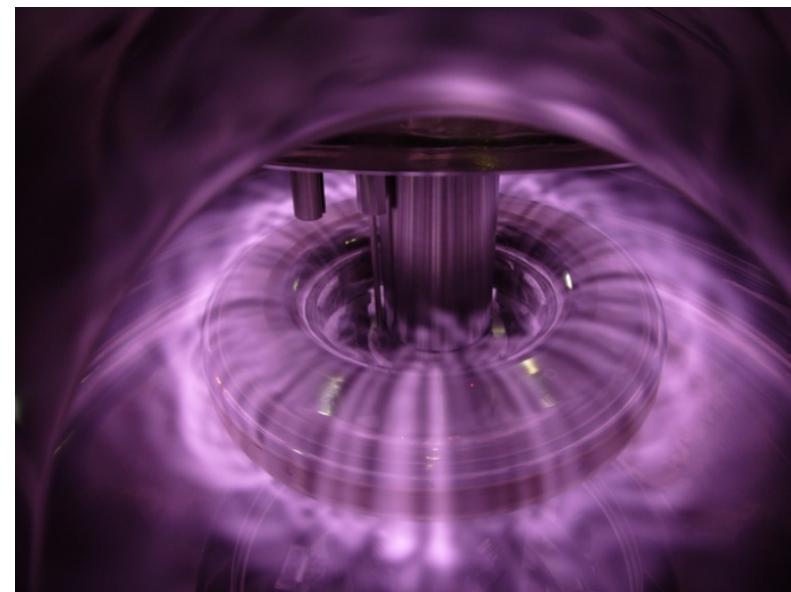


kurasse.jp/member/little-kinoko778/note/93577

vortex \rightarrow *self-organized confinement*

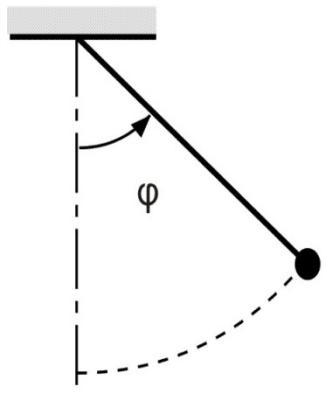


Jovian magnetosphere



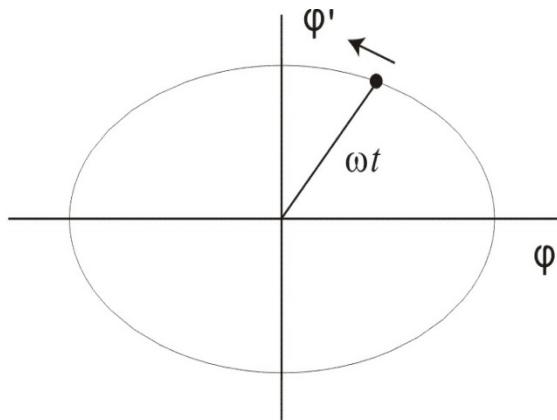
RT-1 magnetospheric plasma

“canonical vortex” = micro scale



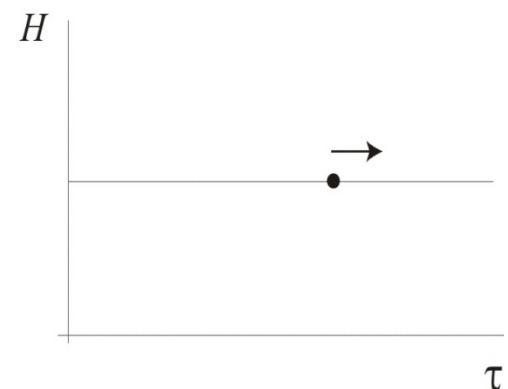
(a)

repetition



(b)

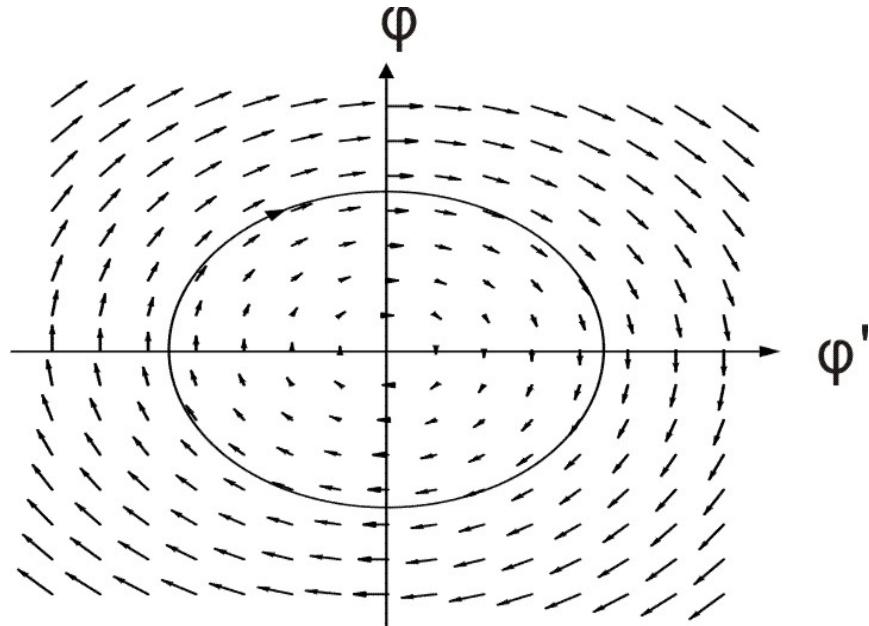
circulation



(c)

continuation

Hamiltonian formalism



$$J = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$$

Symplectic group
→ canonical vortex

matter = energy = Hamiltonian H

space-time = symplectic geometry J

$$\frac{d}{dt} F = \{H, F\}$$

$$\{H, F\} = \langle \partial_z H, J \partial_z F \rangle$$

What is *macroscopic* space-time ?

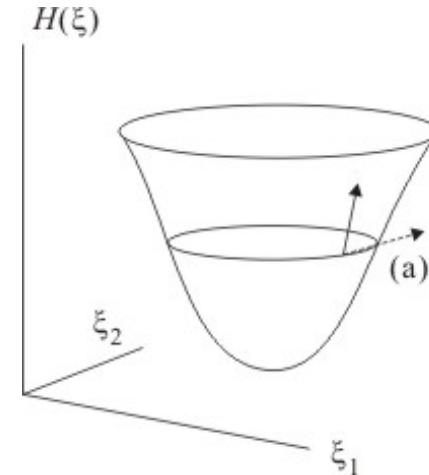
- How can magnetic field “confine” a plasma, despite the fact that $f = \exp(-\beta E)$?
→ This is possible because of “scale separation”.
- What is scale hierarchy?
- What is “state” and what is “space”?
.

micro = canonical / macro = non-canonical

General form of Hamiltonian systems

- Hamiltonian mechanics is dictated by J (Poisson operator) and H (Hamiltonian)

$$\frac{d}{dt} z = J \partial_z H(z)$$



Poisson bracket: $\{G, F\} = \langle \partial_z G(z), J \partial_z F(z) \rangle$

$$\frac{d}{dt} F(z) = \{H, F\}$$

$$\{\{F, G\}, H\} + \{\{G, H\}, F\} + \{\{H, F\}, G\} = 0$$

Examples of Hamiltonian systems

classical mechanics:

$$\frac{d}{dt} \mathbf{z} = \frac{d}{dt} \begin{pmatrix} q \\ p \end{pmatrix} = \begin{pmatrix} 0 & I \\ -I & 0 \end{pmatrix} \begin{pmatrix} \partial_q H \\ \partial_p H \end{pmatrix} = J \partial_z H$$

quantum mechanics:

$$\partial_t \psi = -i \partial_\psi \langle \mathcal{H} | \psi, \psi \rangle / 2 = J \partial_\psi H$$

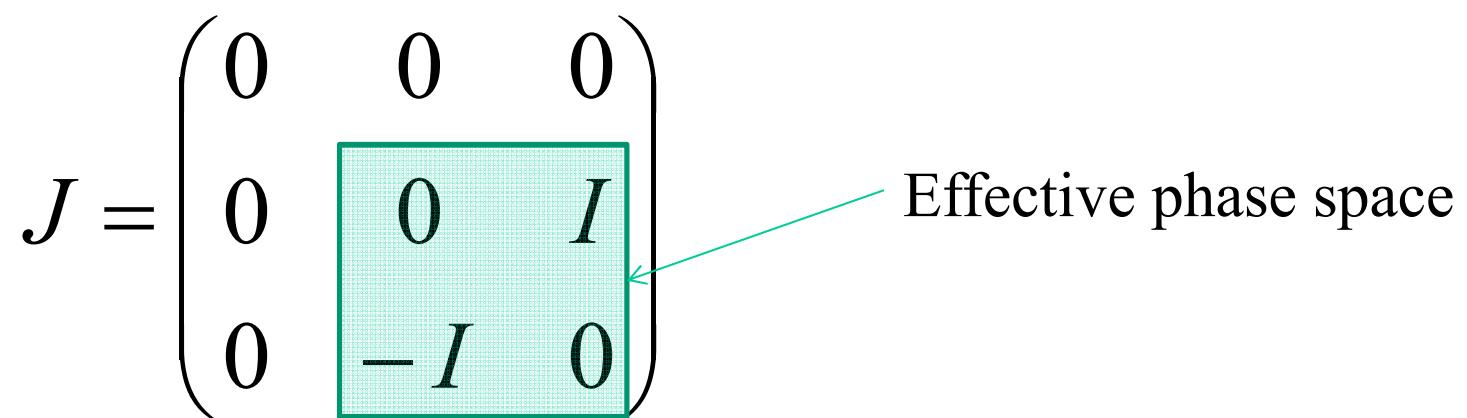
These examples are “canonical” because J are regular operators.

Non-canonical Hamiltonian mechanics

- Non-canonicality : $\text{Ker}(J)$

$$J = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & I \\ 0 & -I & 0 \end{pmatrix}$$

Effective phase space



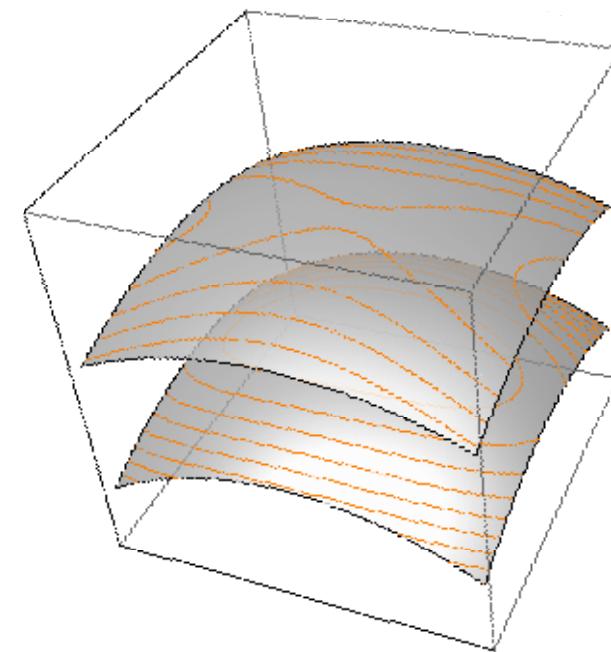
- $\text{Ker}(J) = \text{Coker}(J) \rightarrow \text{“topological defect”}$

non-canonical Hamiltonian mechanics

- $\text{Ker}(J) = \text{Coker}(J) \rightarrow$ “topological constraint”
- Foliation of $\text{Ker}(J) \rightarrow$ Casimir elements

$$\exists C \text{ s.t. } \{G, C\} = 0 \ (\forall G)$$

$$\text{i.e. } \partial_z C \in \text{Ker}(J)$$



Scale hierarchy of magnetized particles

$$H = \frac{m}{2} \left(V_c^2 + V_{\parallel}^2 + V_{\perp}^2 \right) + q\phi$$

$$\begin{aligned} H_c &= \mu\omega_c + \frac{m}{2} \left(V_{\parallel}^2 + V_{\perp}^2 \right) + q\phi \\ &= \mu\omega_c + \frac{(P_{\theta} - q\psi)^2}{2mr^2} + \frac{p_{\parallel}^2}{2m} + q\phi \end{aligned}$$

$$z = (\vartheta_c, \mu; \zeta, p_{\parallel}; \theta, P_{\theta})$$

Coarse graining → macro-hierarchy

$$z = (\vartheta_c, p_c; \zeta, p_{\parallel}; \theta, P_{\theta}) \rightarrow (\cancel{\vartheta}_c, \mu; \zeta, p_{\parallel}; \theta, P_{\theta})$$

Coarse-graining → non-canonicalization

$$J = \begin{pmatrix} J_c & 0 & 0 \\ 0 & J_c & 0 \\ 0 & 0 & J_c \end{pmatrix} \rightarrow J_{nc} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & J_c & 0 \\ 0 & 0 & J_c \end{pmatrix}$$

Boltzmann distribution on Casimir leaf

$$\delta(S - \alpha N - \beta E - \gamma M) = 0$$

$$f = c \exp(-\beta H_c - \gamma \mu)$$

Chemical potential

Quasi-particle number

$$S = - \int f \log f d^6 z$$

$$N = \int f d^6 z$$

$$E = \int H_c f d^6 z$$

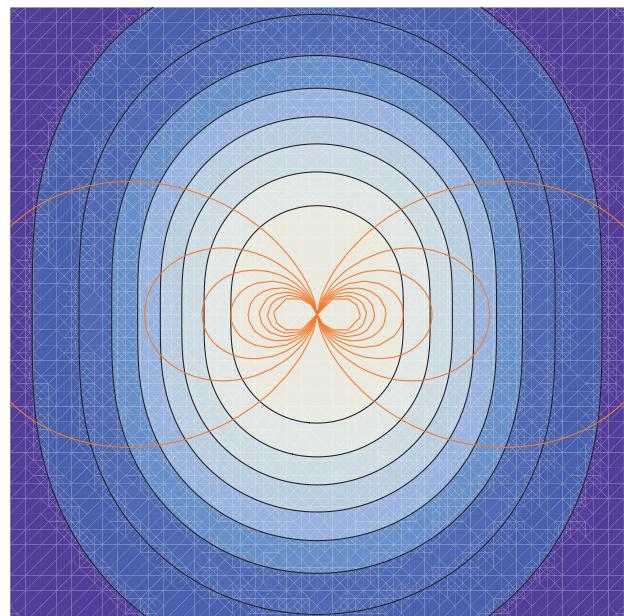
$$M = \int \mu f d^6 z$$

Embedding into the lab-frame space

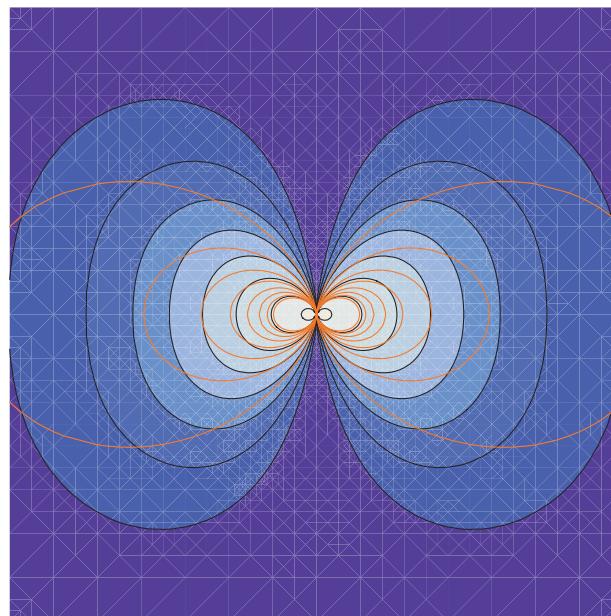
$$\begin{aligned}\rho(x) &= \int f d^3v = \int f(2\pi\omega_c / m) d\mu dv_{\parallel} dv_{\theta} \\ &= c \int \exp(-\beta H_c - \gamma \mu) \frac{2\pi\omega_c d\mu}{m} dv_{\parallel} dv_{\theta} \\ &= \frac{\omega_c(x)}{\omega_c(x) + \gamma}\end{aligned}$$

Density clump in lab-frame space

Adiabatic invariants = *number* of quasi-particles
→ thermodynamic distribution on a Casimir leaf



(A)



(B)

Topics in theoretical physics II

self-organization in foliated phasespace

Z. Yoshida (The University of Tokyo)

Collaborators:

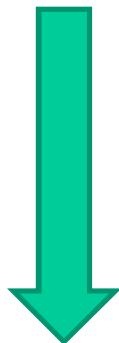
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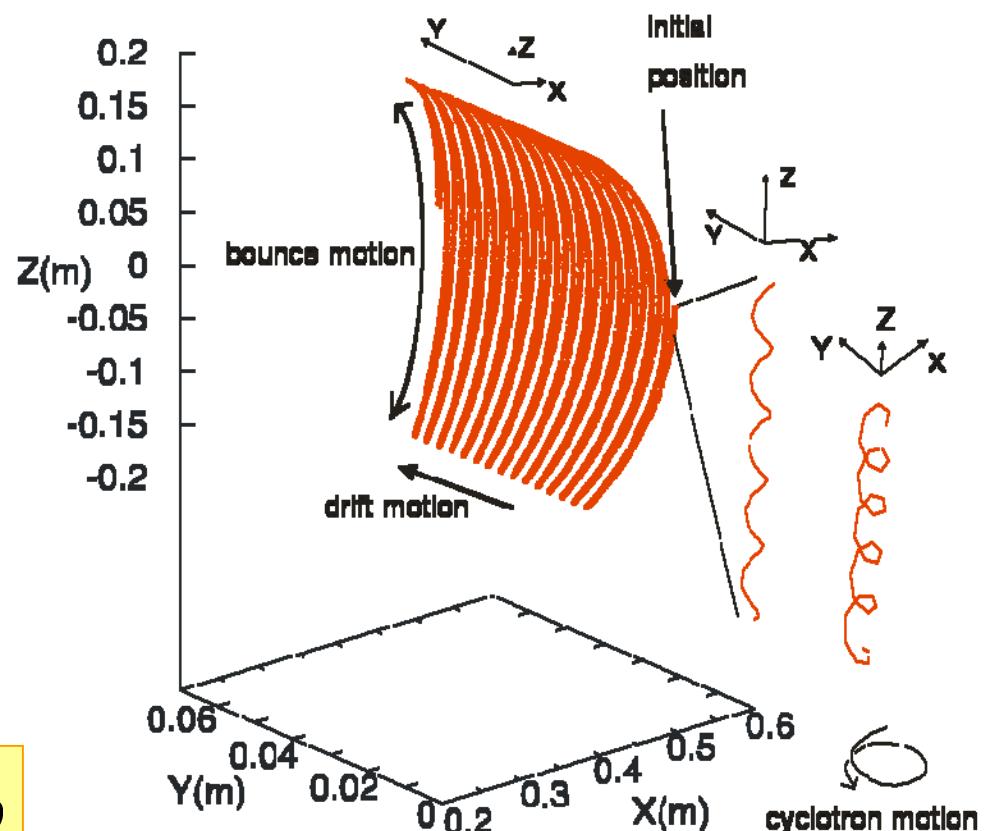
Scale hierarchy of magnetized particles

$$H = \frac{m}{2} (V_{\perp}^2 + V_{\parallel}^2) + q\phi$$



Guiding center = quasi-particle

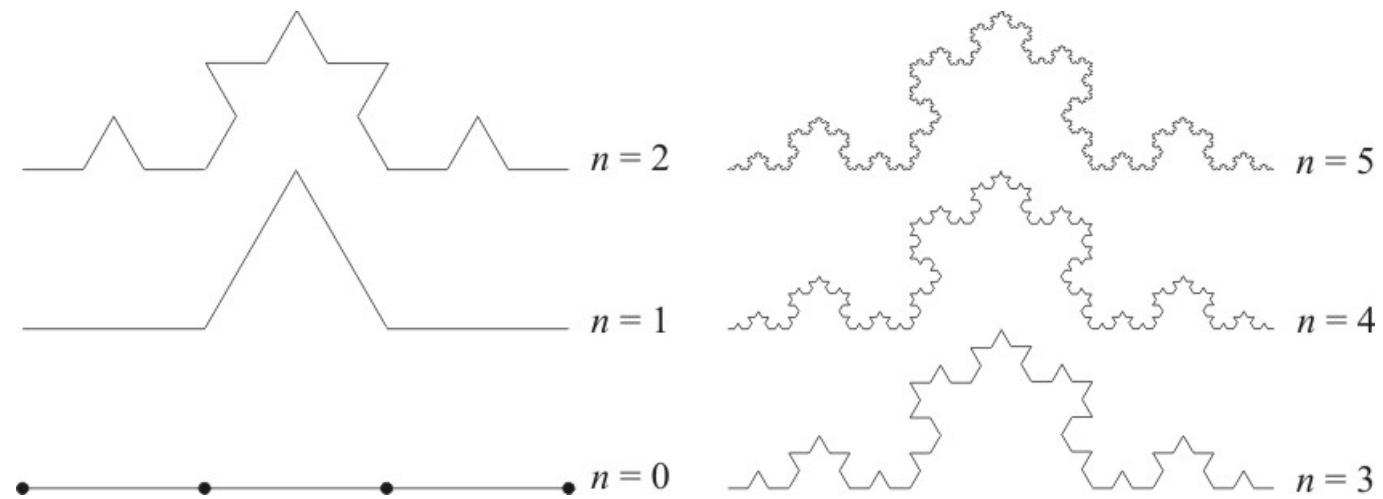
$$H_{gc} = \omega_c \mu + \omega_b J_{\parallel} + q\phi$$



subject defines a scale

$$z = (\vartheta_c, p_c; \zeta, p_{\parallel}; \theta, P_{\theta}) \rightarrow (\cancel{\vartheta}_c, \mu; \zeta, p_{\parallel}; \theta, P_{\theta})$$

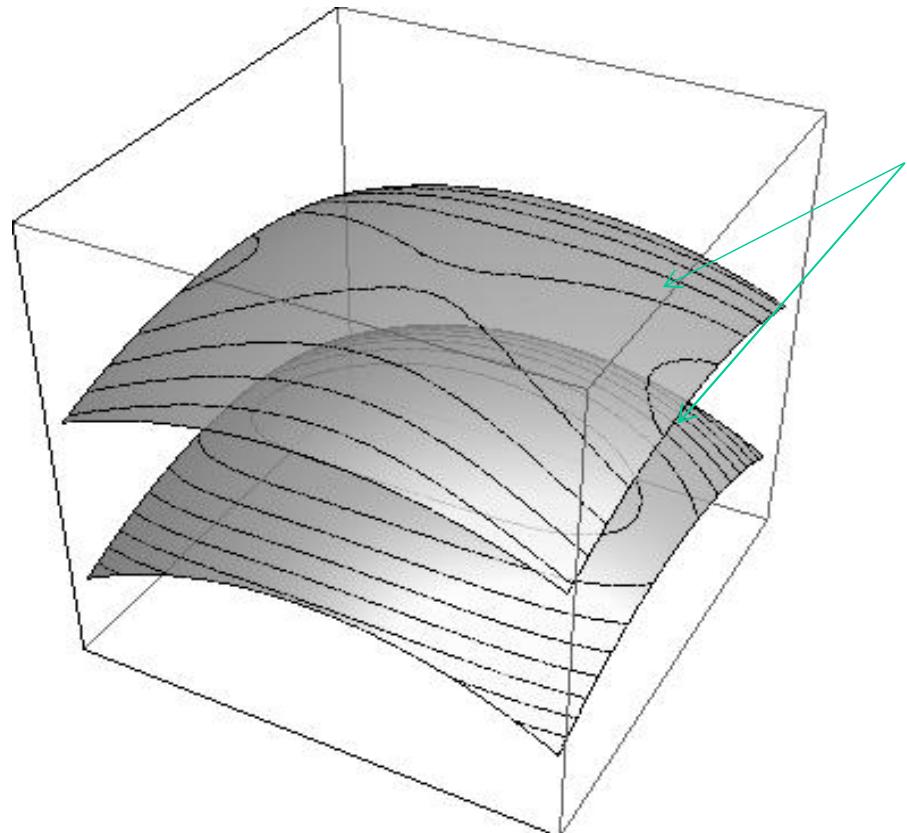
The definition (representation) of matter depends on the subject=scale.



Foliated phase space of *non-canonical* Hamiltonian system

- $\text{Ker}(J) = \{ \text{ grad } C(z) \}$

Casimir leaf



The “effective” phase space of constrained dynamics

- Equilibrium points?
- Stability?
- Thermal equilibrium?

Grand-canonical distribution on Casimir leaf

$$\delta(S - \alpha N - \beta E - \gamma M) = 0$$

$$f = c \exp[-\beta(H_c + \gamma' \mu)]$$

Chemical potential

Quasi-particle number

$$S = - \int f \log f d^6 z$$

$$N = \int f d^6 z$$

$$E = \int H_c f d^6 z$$

$$M = \int \mu f d^6 z$$

Embedding into the lab-frame space

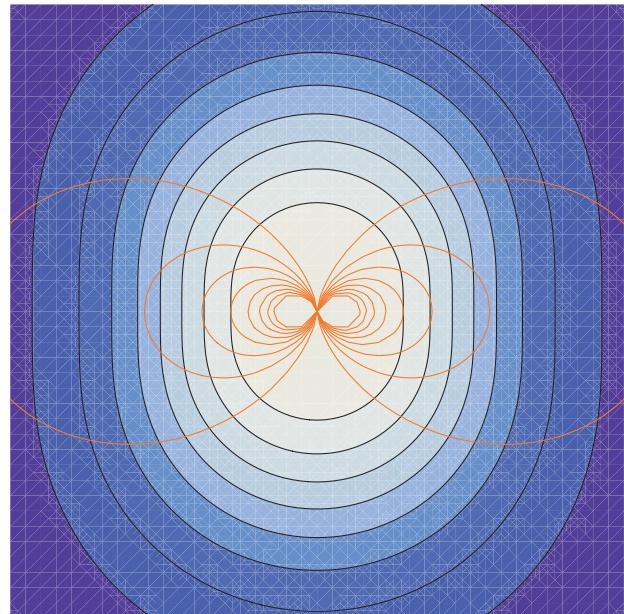
$$\begin{aligned}\rho(x) &= \int f d^3v = \int f \frac{2\pi\omega_c d\mu}{m} dv_{\parallel} \\ &= c \int \exp(-\beta H_c - \gamma \mu) \frac{2\pi\omega_c d\mu}{m} dv_{\parallel} \\ &= \frac{\omega_c(x)}{\omega_c(x) + \gamma}\end{aligned}$$

Density clump in lab-frame space

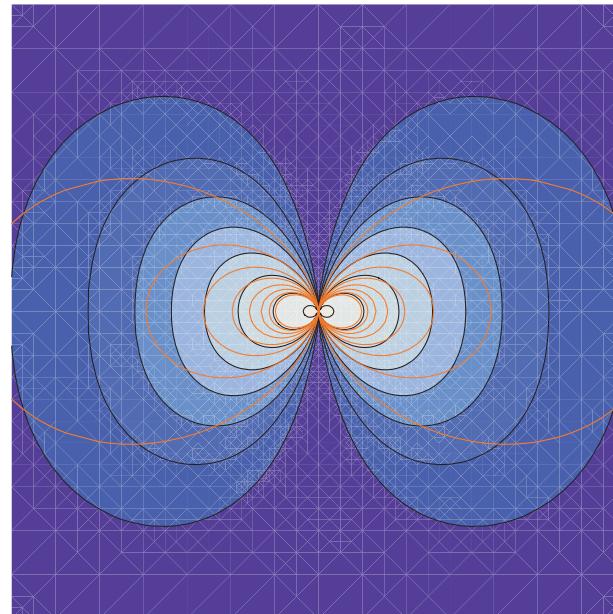
Adiabatic invariants

→ *phase space* of “quasi-particles”

→ thermodynamic distribution on a Casimir leaf



(A)



(B)

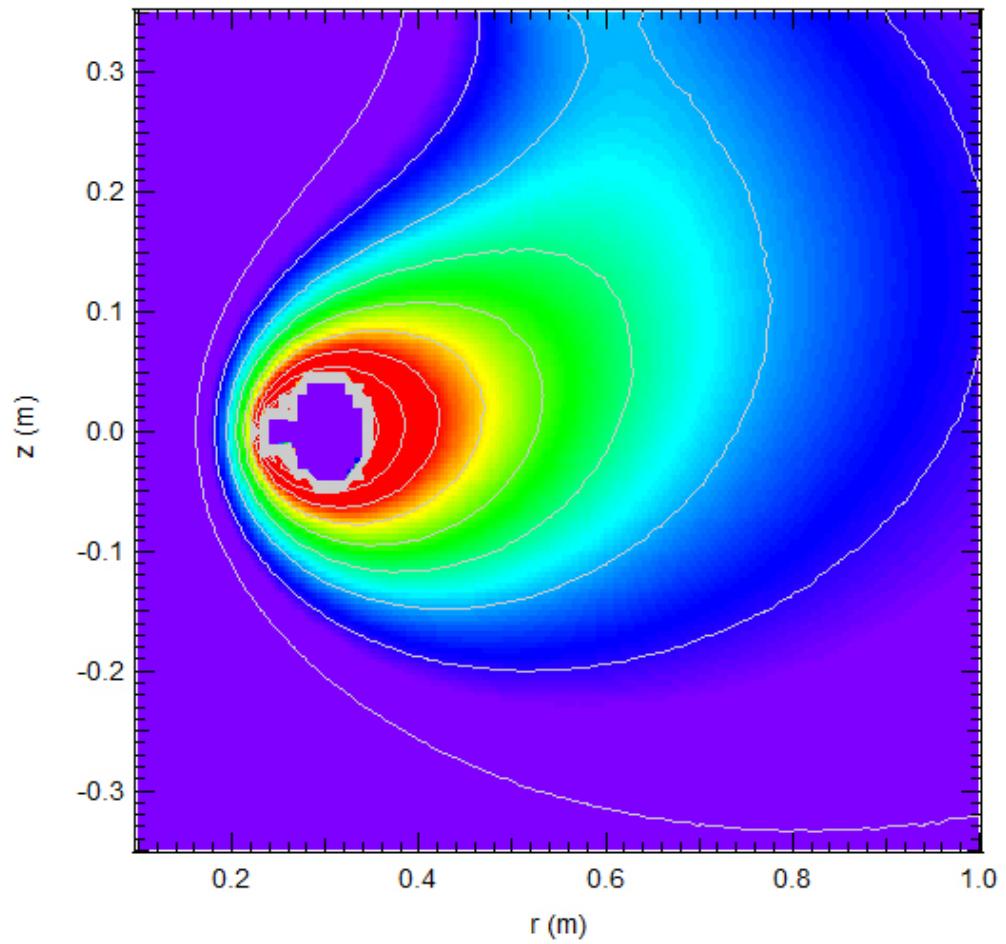
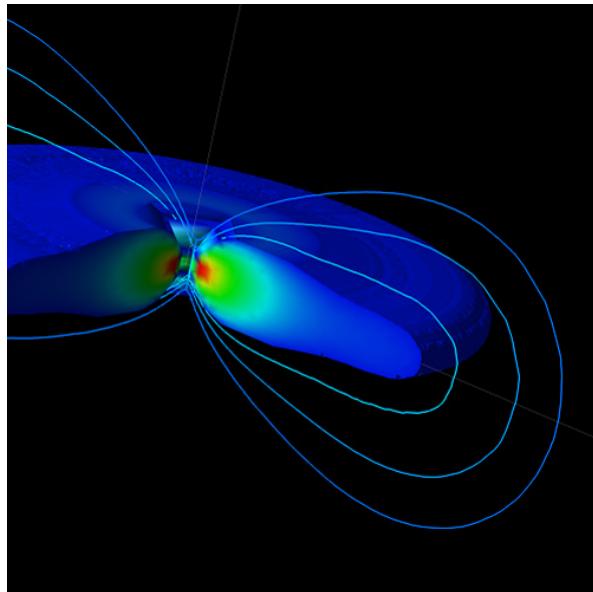
Conclusion (I)

- Scale hierarchy = foliated phase space
- Adiabatic invariant = *Casimir element*
→ foliation
- Distorted metric on a leaf → structure

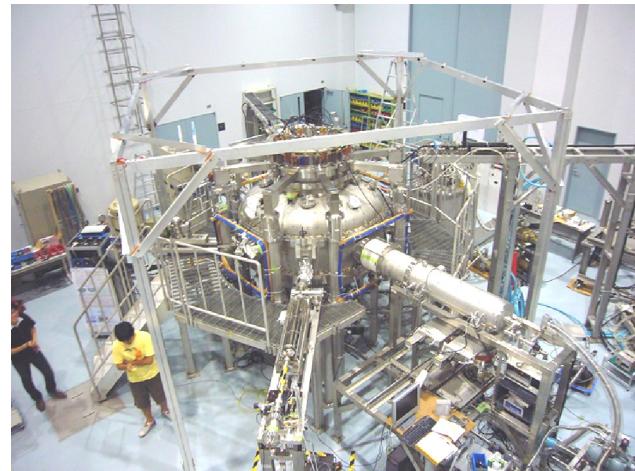
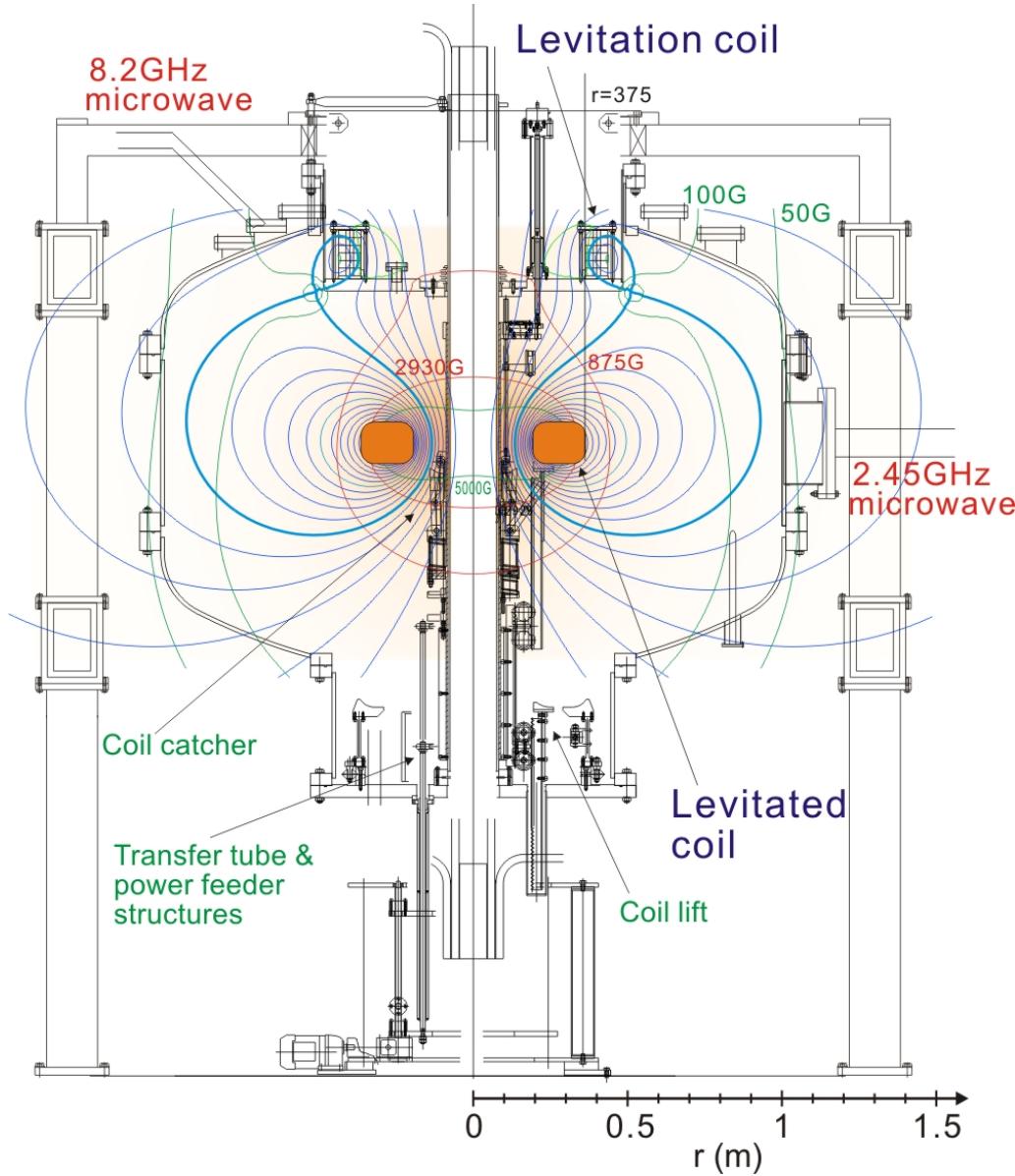
Z.Yoshida & S.M. Mahajan, *Self-organization in foliated phase space: construction of a scale hierarchy by adiabatic invariants of magnetized particles*, Prog. Theor. Exp. Phys. **2014** (2014), 073J01

A magnetosphere on the Earth

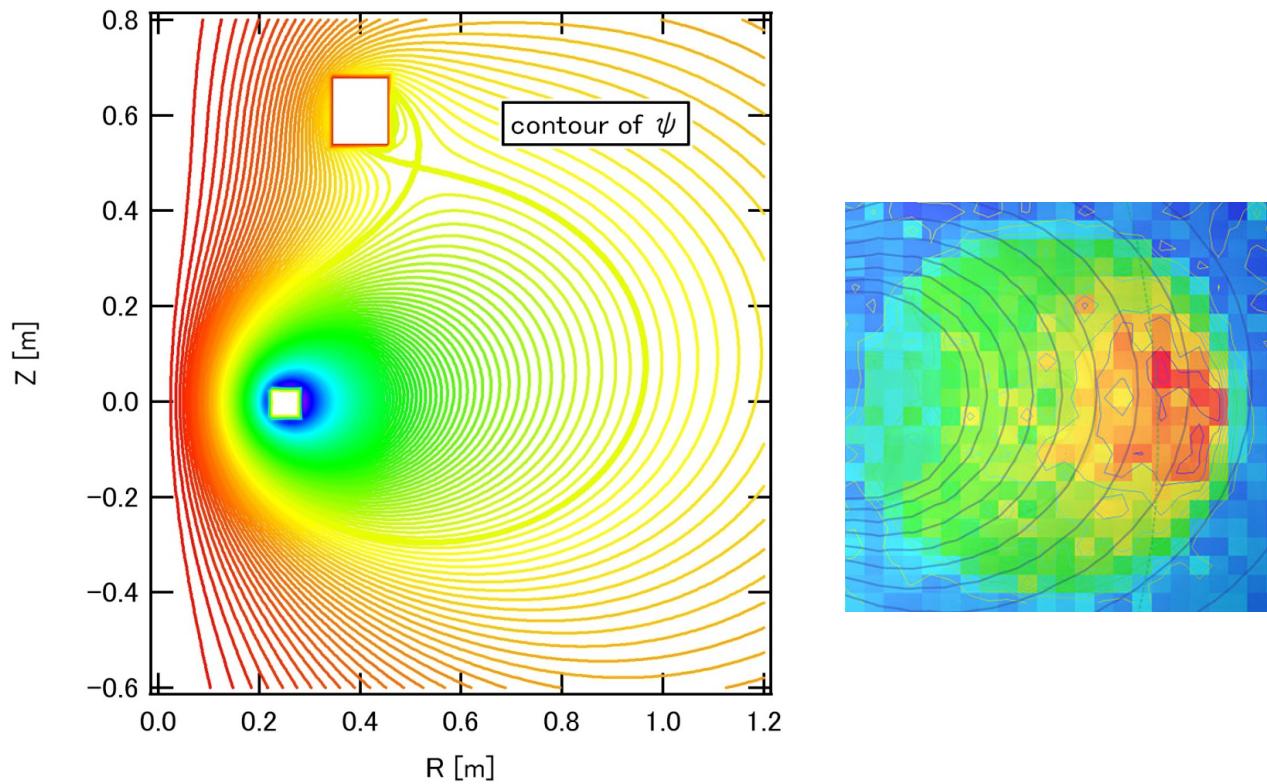
RT-1 project



Levitating HTC superconducting magnet system



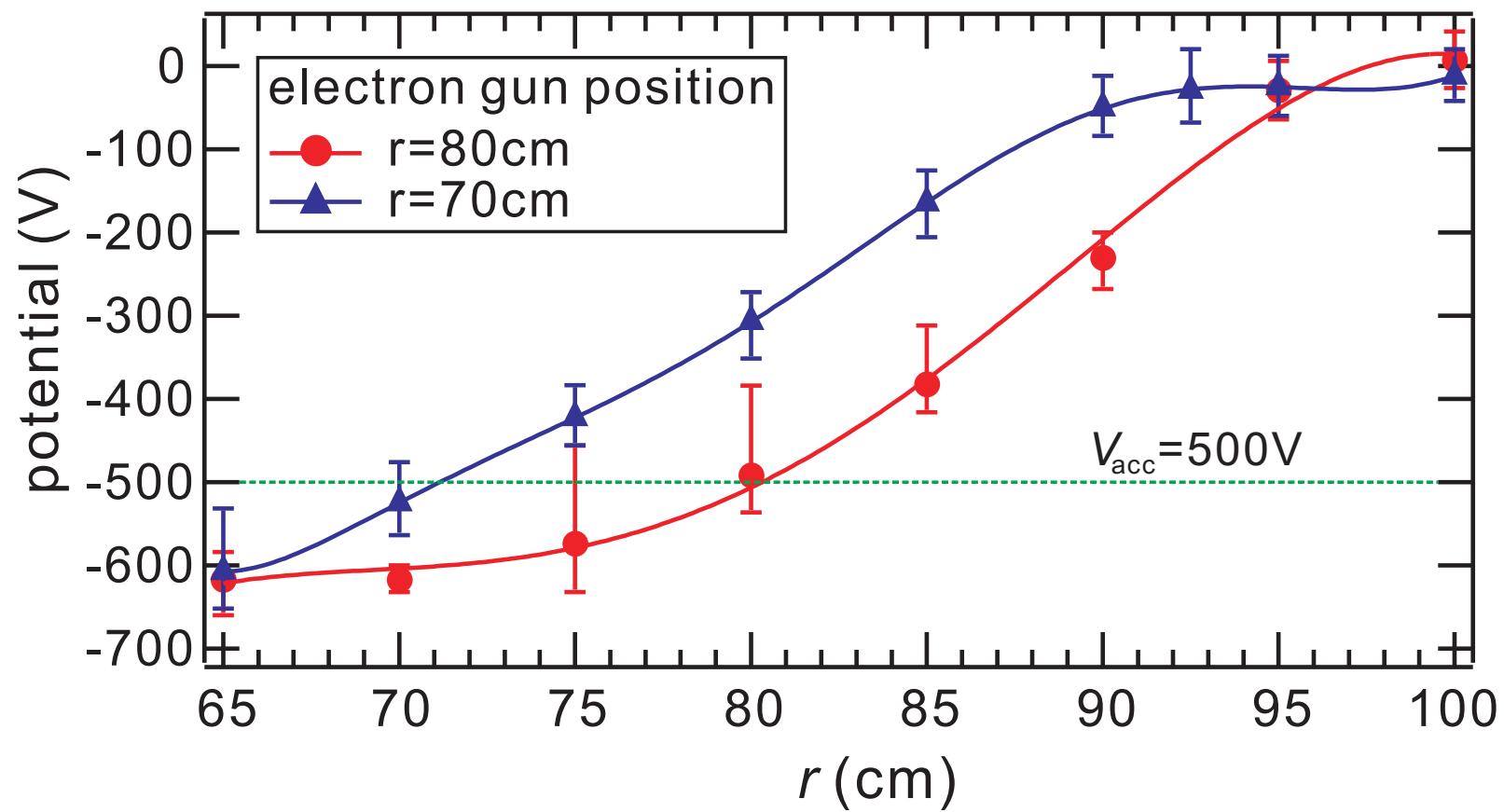
High-beta plasma confinement



$$T_e \approx 10 \sim 30 \text{ keV}, \quad n_e = 10^{16} \sim 10^{17} \text{ m}^{-3}$$

$$\beta \approx \beta_e \approx 0.7, \quad \tau_E \approx 0.5 \text{ sec}$$

Inward (up-hill) diffusion



ZY *et al.*, Phys. Rev. Lett. **104** (2010), 235004 .

space for “confinement”



kurasse.jp/member/little-kinoko778/note/93577

Conclusion (II)

- Experimental proof of *self-organized confinement*.
- Experimental evidence of *inward diffusion*.

Topics in theoretical physics III

self-organization in MHD plasma

Z. Yoshida (The University of Tokyo)

Collaborators:

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R.L. Dewar (ANU), F. Dobroff (Trieste)

N. Shatashvili (TSU)

MHD (incompressible)

- Naïve form:

$$\begin{aligned}\partial_t \mathbf{V} &= -P_\sigma(\nabla \times \mathbf{V}) \times \mathbf{V} + P_\sigma(\nabla \times \mathbf{B}) \times \mathbf{B}, \\ \partial_t \mathbf{B} &= \nabla \times (\mathbf{V} \times \mathbf{B}).\end{aligned}$$

- State vector:

$$\mathbf{u} = (\mathbf{V}, \mathbf{B}) \in L^2_\sigma(\Omega) \times L^2_\sigma(\Omega)$$

$$L^2_\sigma(\Omega) := \left\{ \mathbf{v} \in L^2_\sigma(\Omega); \nabla \cdot \mathbf{v} = 0, \mathbf{n} \cdot \mathbf{v} = 0 \right\}$$

$$P_\sigma : L^2(\Omega) \rightarrow L^2_\sigma(\Omega)$$

Hamiltonian form of MHD

- Hamiltonian

$$H(u) = \frac{1}{2} \int_{\Omega} (V^2 + B^2) dx = \frac{1}{2} \|u\|^2$$

- Poisson operator

$$J(u) = \begin{pmatrix} -P_\sigma(\nabla \times V) \times \circ & P_\sigma(\nabla \times \circ) \times \mathbf{B} \\ \nabla \times (\circ \times \mathbf{B}) & 0 \end{pmatrix}$$

- *Casimir elements:*

$$C_1(u) = \frac{1}{2} \langle A, \mathbf{B} \rangle, \quad C_2(u) = \langle V, \mathbf{B} \rangle$$

Beltrami equilibria on helicity leaves

- Beltrami equilibrium:

$$\partial_u H_\mu(u) = 0 \quad (H_\mu = H - \mu C_1).$$

$$\rightarrow \nabla \times \mathbf{B} = \mu \mathbf{B}, \quad \mathbf{B} \in L^2_\sigma(\Omega).$$

- Two classes of Beltrami eigenvalues:
 - (1) Self-adjoint curl operator S
 \rightarrow discrete real eigenvalues $\mu \in \{\lambda_1, \lambda_2, \dots\}$
 - (2) Non-self-adjoint extension T
(if Ω is multiply connected) $\rightarrow \forall \mu \in \mathbb{C}$

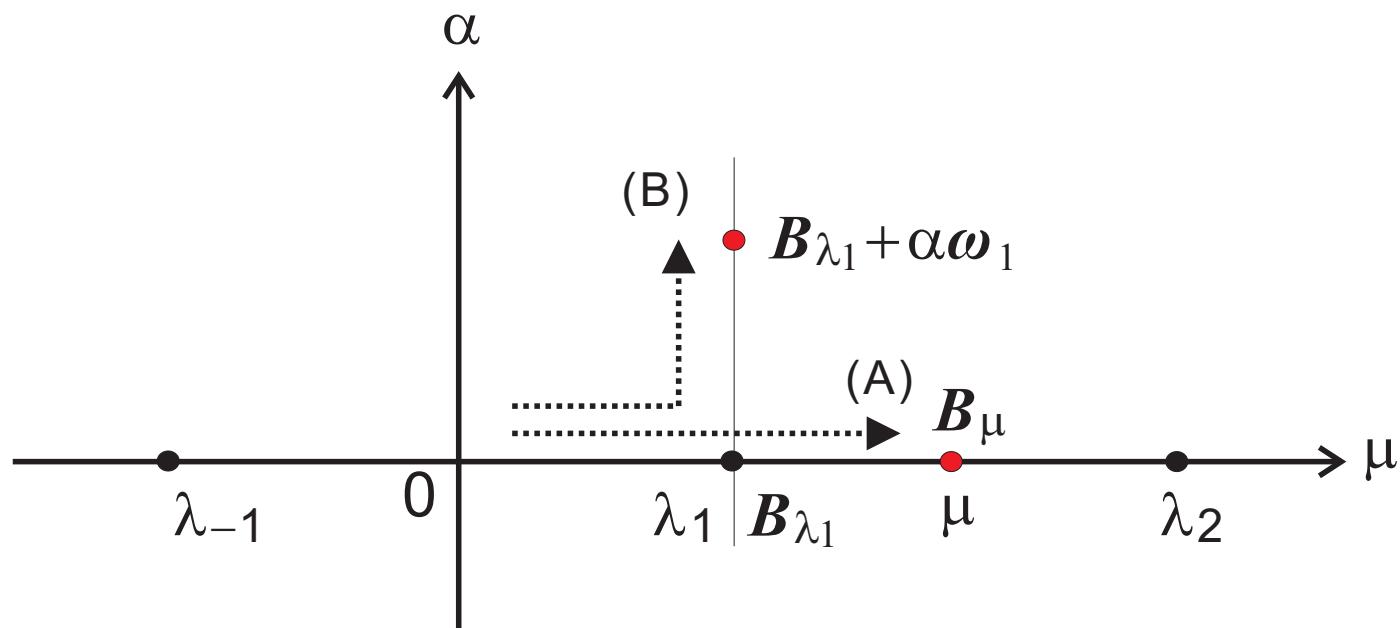
ZY & Y. Giga, Math. Z. **204** (1990) 235.

Bifurcation theorem

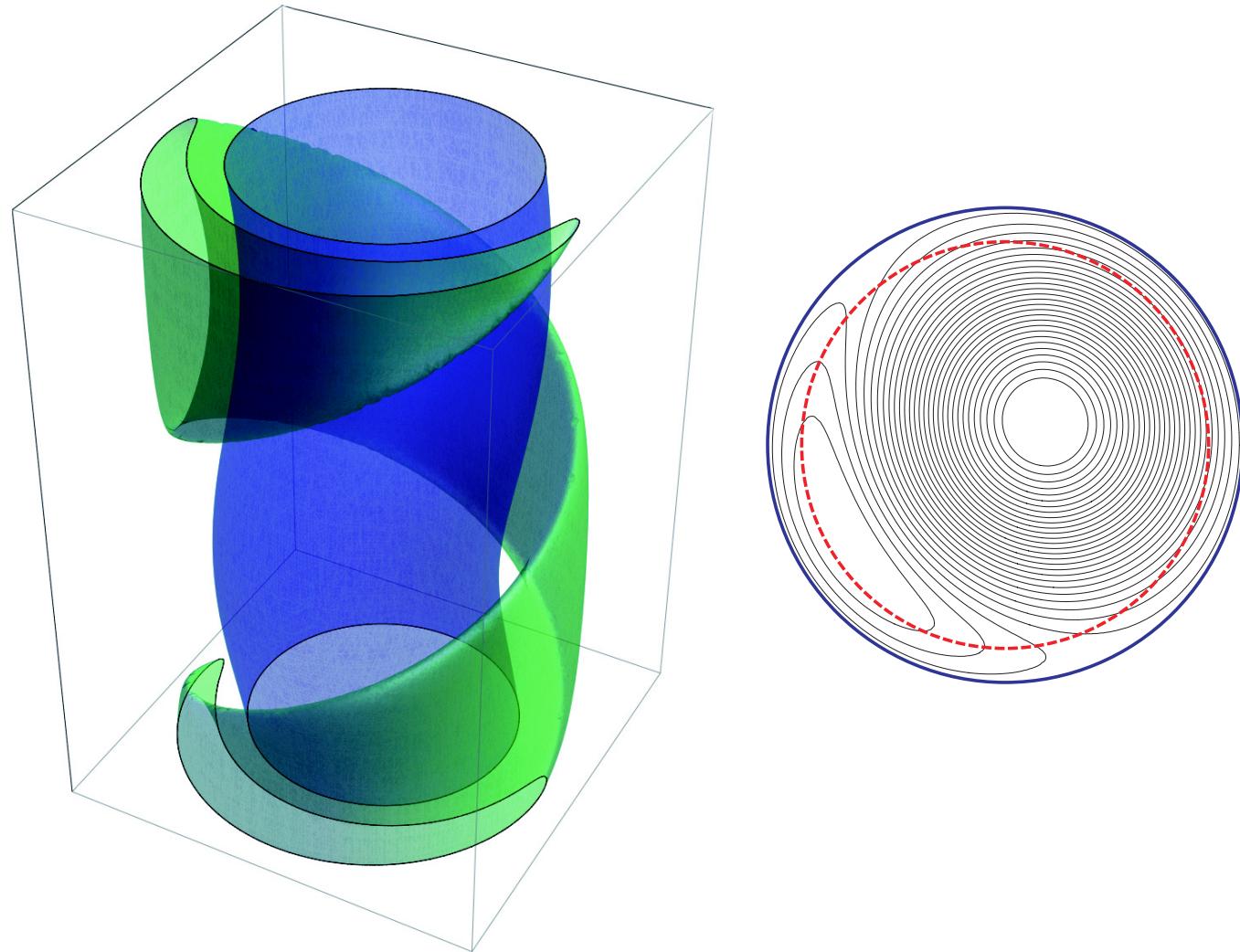
Theorem 1

Let A_H be the vector potential of B_H (cohomology).

If $\langle A_H, \omega_j \rangle = 0$, then branch-(B) bifurcates at $\mu = \lambda_j \in \sigma_p(S)$.



Bifurcated Beltrami equilibrium



More Casimirs: Linearized MHD

- Linearize near a Beltrami equilibrium \mathbf{B}_μ .
- Hamilton's operator

$$L_\mu = \begin{pmatrix} 1 & 0 \\ 0 & 1 - \mu S^{-1} \end{pmatrix}$$

- Poisson operator

$$J_\mu = \begin{pmatrix} 0 & P_\sigma(\nabla \times \circ) \times \mathbf{B}_\mu \\ \nabla \times (\circ \times \mathbf{B}_\mu) & 0 \end{pmatrix}$$

Resonance singularity

- $\text{Ker}(J_\mu)$ consists of singular eigenfunctions.

ν such that $\nabla \times (\mathbf{B}_\mu \times \nu) = 0$,

\mathbf{b} such that $\mathbf{B}_\mu \times (\nabla \times \mathbf{b}) = 0$

- In slab geometry: \mathbf{b} obeys

$$\mathbf{b} = (0, b_y(x), b_z(x)) e^{i(k_y y + k_z z)}$$

$$b_y(x) = i k_y \theta(x), \quad b_z(x) = i k_z \theta(x)$$

$$\Rightarrow [B_y(x)k_y + B_z(x)k_z] \partial_x \theta(x) = 0$$

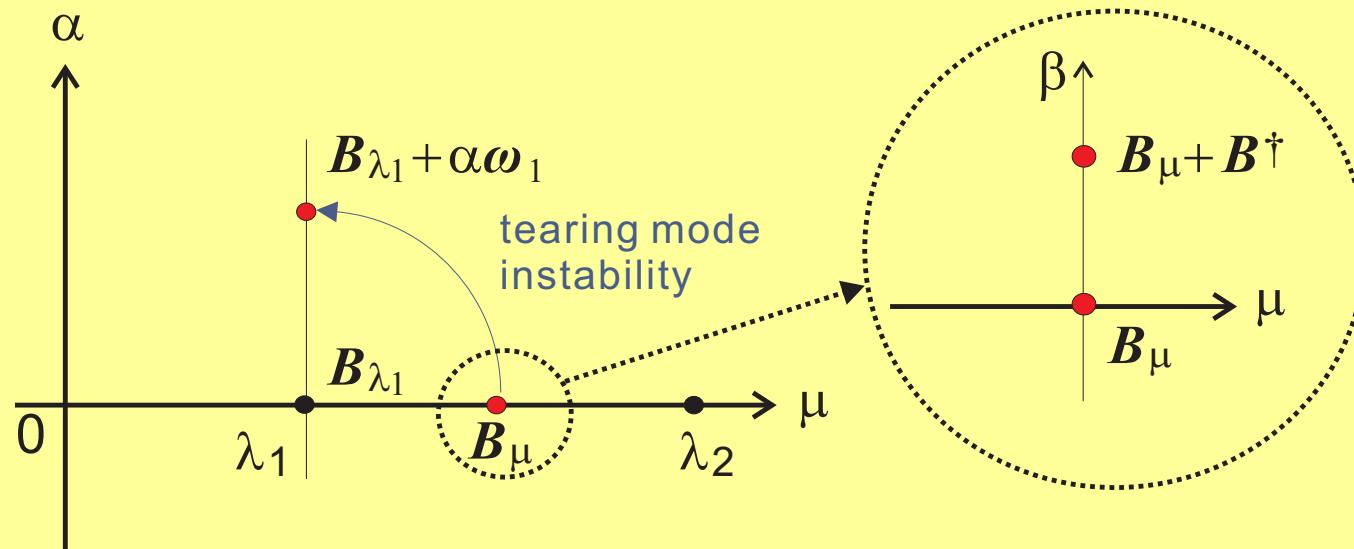
$$\Rightarrow \theta(x) = c Y(x - x^+)$$

$$\Rightarrow C_b = (\tilde{\mathbf{B}}, \mathbf{b})$$

Tearing mode (singular equilibria)

Theorem 2

When Ω has a symmetry, J_μ has helical flux Casimirs $C_b(\tilde{\mathbf{u}}) = (\tilde{\mathbf{B}}, \mathbf{b})$. The stationary point of the energy-Casimir functional gives a tearing-mode singular eigenfunction.



Excess energy of Beltrami equilibrium

- *We may estimate*

$$(\tilde{u}, L_\mu \tilde{u}) / 2 \geq (1 - \mu / \lambda_1) (\tilde{\mathbf{B}}, \boldsymbol{\omega}_1)^2 / 2$$

- *Under the constraint of the helical-flux Casimir,*

$$\min_{C_b(\tilde{u})=c_b} (\tilde{u}, L_\mu \tilde{u}) = (1 - \mu / \lambda_1) [c_b / (\mathbf{b}, \boldsymbol{\omega}_1)]^2 / 2$$

Conclusion (III)

- The *helicity* is a Casimir element that foliates the phase space.
- We find bifurcated equilibrium points on helicity leaves, because helicity leaves are distorted with respect to the energy norm.
- A tearing mode is an equilibrium on a leaf of singular (resonant) Casimir element

Z. Yoshida & R. L. Dewar; J. Phys. A **45** (2012), 365502 1-36.

Unfreezing Casimir \rightarrow Canonicalization

$$J_{nc} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \boxed{0 & I} \\ 0 & -I & 0 \end{pmatrix} \Rightarrow \tilde{J} = \begin{pmatrix} 0 & I & 0 & 0 \\ -I & 0 & 0 & 0 \\ 0 & 0 & 0 & I \\ 0 & 0 & -I & 0 \end{pmatrix}$$

Effective (macro) phase space

“recovered” angle variable

A diagram illustrating the transformation of matrix J_{nc} into matrix \tilde{J} . Matrix J_{nc} is a 3x3 matrix with a green box around its 2x2 submatrix $\begin{pmatrix} 0 & I \\ -I & 0 \end{pmatrix}$. Matrix \tilde{J} is a 4x4 matrix with a pink box around its top-left 2x2 submatrix $\begin{pmatrix} 0 & I \\ -I & 0 \end{pmatrix}$. A green arrow points from the green box to the text "Effective (macro) phase space". A pink arrow points from the pink box to the text "‘recovered’ angle variable".

Canonicalization of MHD

- Given a Casimir $C_b(\tilde{\mathbf{u}}) = (\tilde{\mathbf{B}}, \mathbf{b})$, we put

$$P_{\perp} \tilde{\mathbf{B}} = (\tilde{\mathbf{B}}, \mathbf{b}) \mathbf{b} \in \text{Ker}(J_{\mu}), \quad P_{\parallel} \tilde{\mathbf{B}} = \tilde{\mathbf{B}} - P_{\perp} \tilde{\mathbf{B}}$$

- Separating the kernel, we write

$$J_{\mu} = \begin{pmatrix} 0 & P_{\sigma}(\nabla \times P_{\parallel} \circ) \times \mathbf{B}_{\mu} & 0 \\ P_{\parallel} \nabla \times (\circ \times \mathbf{B}_{\mu}) & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

- Canonicalize

$$\tilde{J}_{\mu} = \left(\begin{array}{ccc|c} 0 & P_{\sigma}(\nabla \times P_{\parallel} \circ) \times \mathbf{B}_{\mu} & 0 & 0 \\ P_{\parallel} \nabla \times (\circ \times \mathbf{B}_{\mu}) & 0 & 0 & 0 \\ 0 & 0 & 0 & I \\ \hline 0 & 0 & -I & 0 \end{array} \right)$$

Singular perturbation

- Original Hamiltonian is independent of θ (Casimir \rightarrow Action); denoting $K=1-\mu S^{-1}$

$$L_\mu = \begin{pmatrix} 1 & 0 & 0 \\ 0 & P_{\parallel}K & P_{\parallel}K \\ 0 & P_{\perp}K & P_{\perp}K \end{pmatrix}$$

- perturbation

$$\tilde{L}_\mu = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & P_{\parallel}K & P_{\parallel}K & 0 \\ 0 & P_{\perp}K & P_{\perp}K & 0 \\ 0 & 0 & 0 & D \end{pmatrix}$$

Enlarged Hamiltonian system

- Denoting $p = C_b(\tilde{\boldsymbol{u}}) = (\tilde{\boldsymbol{B}}, \boldsymbol{b})$

$$\begin{cases} \dot{p} = -D\theta \\ \dot{\theta} = (K_\mu \tilde{\boldsymbol{B}}, \boldsymbol{b}) \end{cases}$$

- For $\tilde{\boldsymbol{B}} = p\boldsymbol{\omega}_1$,

$$(K_\mu \tilde{\boldsymbol{B}}, \boldsymbol{b}) = (1 - \mu / \lambda_1)(\boldsymbol{\omega}_1, \boldsymbol{b})p$$

- Beyond the bifurcation point, *negative mass*:

$$(1 - \mu / \lambda_1)(\boldsymbol{\omega}_1, \boldsymbol{b}) < 0$$

\rightarrow instability

Conclusion (IV)

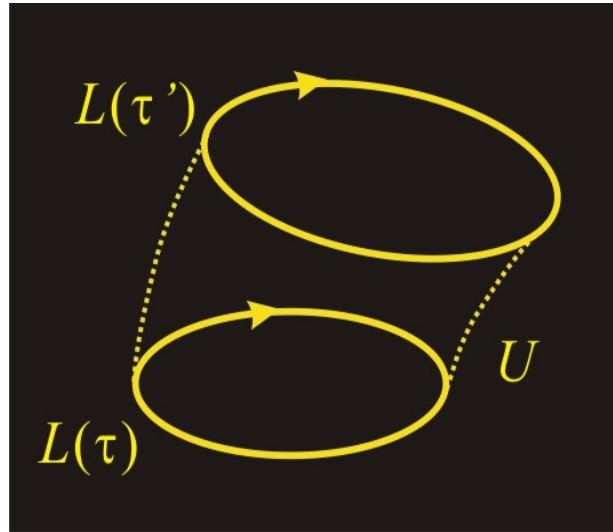
- Perturbing the Hamiltonian system by adding (recovering) an angle variable, we may **unfreeze** the Casimir invariant, allowing the tearing mode to emerge.

Z. Yoshida and P. J. Morrison, Unfreezing Casimir invariants: singular perturbations giving rise to forbidden instabilities, in *Nonlinear physical systems: spectral analysis, stability and bifurcation*, (ISTE and John Wiley and Sons, 2014), Chap. 18; arXiv:1303.0887

Remark

- Our Universe is not simply foliated:
singularities yield complex bifurcation of
leaves.
1. S.M. Mahajan and Z. Yoshida, PRL **105** (2010), 095005.
 2. Z. Yoshida, P. J. Morrison and F. Dobarro, J. Math. Fluid Mech. **16** (2014), 41—57; arXiv:1107.5118
 3. Z. Yoshida and P. J. Morrison, Fluid Dyn. Res. **46** (2014), 031412; arXiv:1401.7698
 4. Z. Yoshida, Y. Kawazura, and T. Yokoyama, J. Math. Phys. **55** (2014), 043101

“Loops” in a fluid / plasma



$$\frac{d}{dt} \left(\oint_{L(t)} \mathbf{P} \cdot d\mathbf{x} \right) = \oint_{L(t)} [\partial_t \mathbf{P} + (\nabla \times \mathbf{P}) \times \mathbf{V}] \cdot d\mathbf{x} = 0$$

if canonical momentum $\mathbf{P} = m\mathbf{V} + (q/c)\mathbf{A}$ obeys

$$\partial_t \mathbf{P} + (\nabla \times \mathbf{P}) \times \mathbf{V} = -\nabla \varphi \quad (\varphi = H + h).$$

Vortex and Entropy

$$-\oint \delta W$$

$$\begin{aligned} \frac{d}{dt} \left(\oint_{L(t)} \mathbf{P} \cdot d\mathbf{x} \right) &= \oint_{L(t)} [\partial_t \mathbf{P} + (\nabla \times \mathbf{P}) \times \mathbf{V}] \cdot d\mathbf{x} \\ &= \oint_{L(t)} -\nabla(H+h) \cdot d\mathbf{x} + \boxed{\int_{L(t)}^0 T \nabla S \cdot d\mathbf{x}} \end{aligned}$$

$$-\oint dE$$

$$\oint \delta Q$$

Relativistic circulation theorem

$$\frac{d}{ds} \left(\oint_{L(s)} P^\mu \cdot dx_\mu \right) = \oint_{L(s)} (\partial^\mu P^\nu - \partial^\nu P^\mu) U_\nu dx_\mu$$

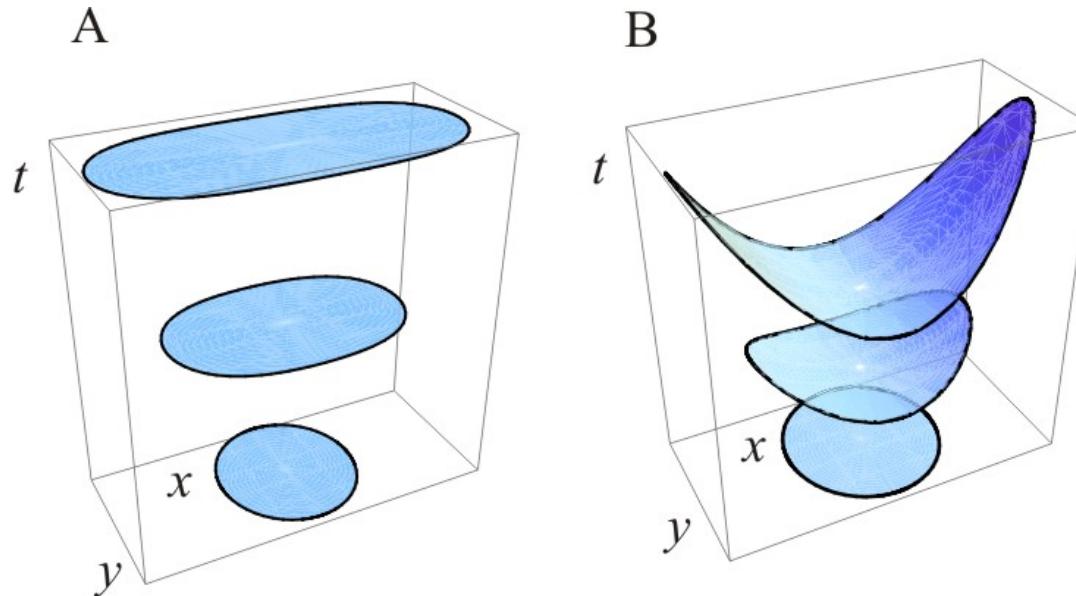
s : proper time, $U^\mu = (\gamma, \gamma V^j / c)$

relativistic equation of motion

$$(\partial^\mu P^\nu - \partial^\nu P^\mu) U_\nu = T \partial^\mu S$$

Relativistic circulation theorem applies on space-time:
relativistic distortion of space-time

Relativistic distortion of space-time



$$\frac{d}{dt}x_\mu = V_\mu$$

$$\frac{d}{ds}x_\mu = U_\mu$$

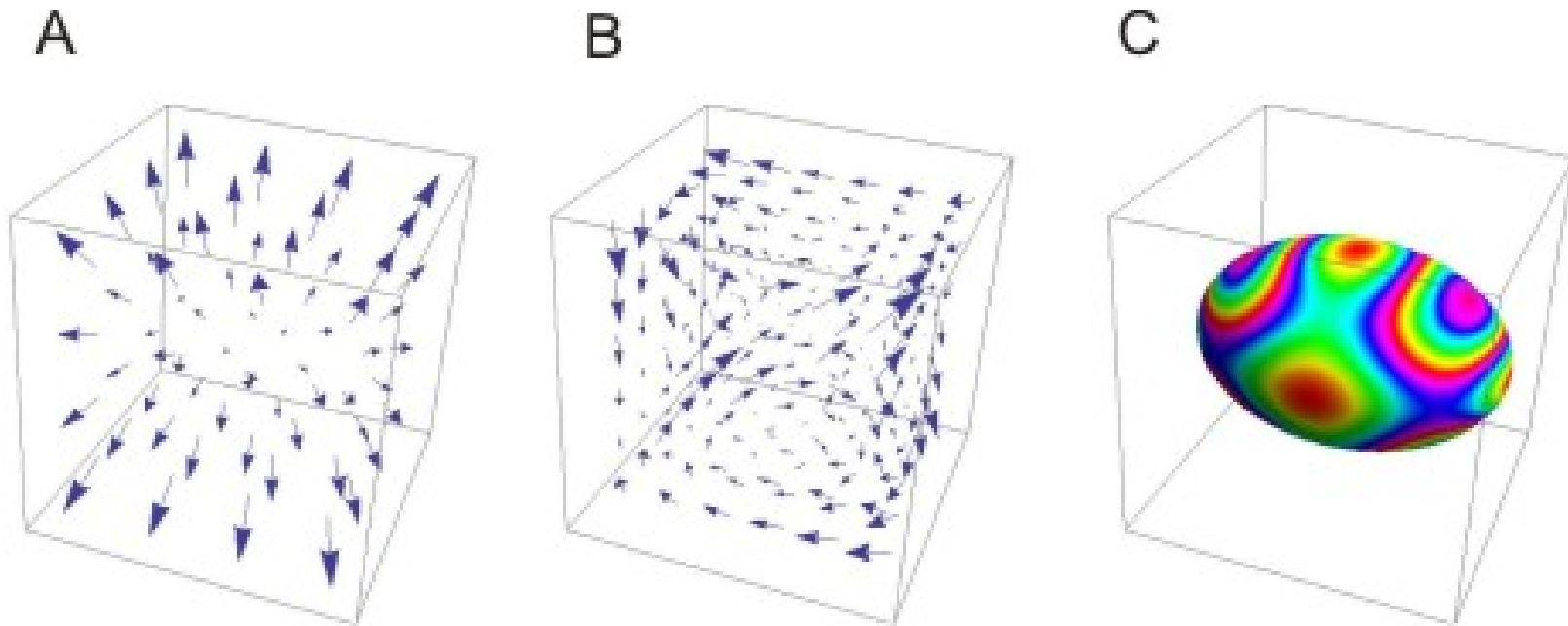
Relativity destroys the “synchrony” of a cycle.

→ “twists” space-time

→ generate circulation

$$\oint_{L(t)} \gamma^{-1} T dS$$

scalar \rightarrow axial vector (vortex/magnetic field)
cosmological origin of magnetic field



$$V \propto \nabla \varphi \quad \Rightarrow \quad \Omega \propto \nabla \gamma^{-1} \times \nabla \varphi$$

S. M. Mahajan & ZY, Phys. Rev. Lett. **105** (2010), 095005.

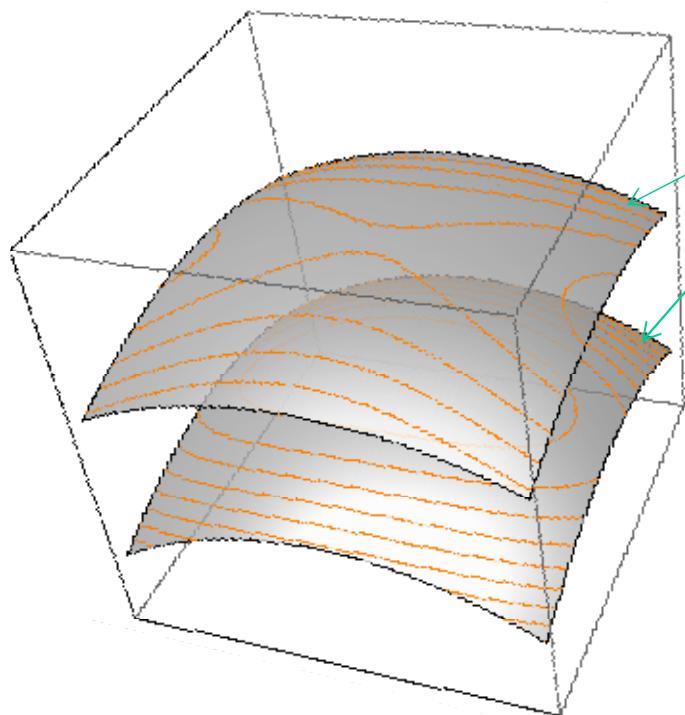
Conclusion (III)

- Unified (canonical) Vortex
 \rightarrow *flow-field (EM) coupling* .
- Circulation conserves as a quantum
(*helicity* is a Casimir)
- Relativity \rightarrow space-time distortion
 \rightarrow *creation* of vorticity (magnetic field)

Foliated phase space of *non-canonical* Hamiltonian system

- $\text{Ker}(J) = \text{Coker}(J) = \{ \text{ grad } C(\mathbf{u}) \}$

Casimir element



The “effective” phase space of constrained dynamics

- Equilibrium points?
- Stability?
- Thermal equilibrium?

Summary

- Macro = non-canonical
- Macro hierarchy = Casimir leaf
- Foliation → distorted space-time

Newton-Kant “space” → relativity, foliation
and beyond

The Euler eq. and vorticity eq.

- Euler equation of ideal fluid:

$$\begin{aligned}\partial_t u + (u \cdot \nabla) u &= -\nabla p, \quad \nabla \cdot u = 0, \\ n \cdot u &= 0\end{aligned}$$

- Vorticity formulation:

$$\begin{aligned}\partial_t \omega + \nabla \times (u \times \omega) &= (\omega := \nabla \times u) \\ &= \{\omega, \phi\} \quad (2D : u = \operatorname{curl} \phi, \omega = -\Delta \phi)\end{aligned}$$

- *Vorticity eq. in $H^{-1}(\Omega)$ is equivalent to Euler eq. in $L^2_\sigma(\Omega)$.*

Hamiltonian formalism of the Euler eq.

- Hamiltonian formalism

$$\partial_t \omega = J(\omega) \partial_\omega H(\omega) \quad [\text{in } H^{-1}(\Omega)]$$

- Hamiltonian

$$H(\omega) = \frac{1}{2} \int_{\Omega} (K\omega) \cdot \omega \, dx \quad [K := -\Delta^{-1} : H^{-1}(\Omega) \rightarrow H_0^1(\Omega)]$$

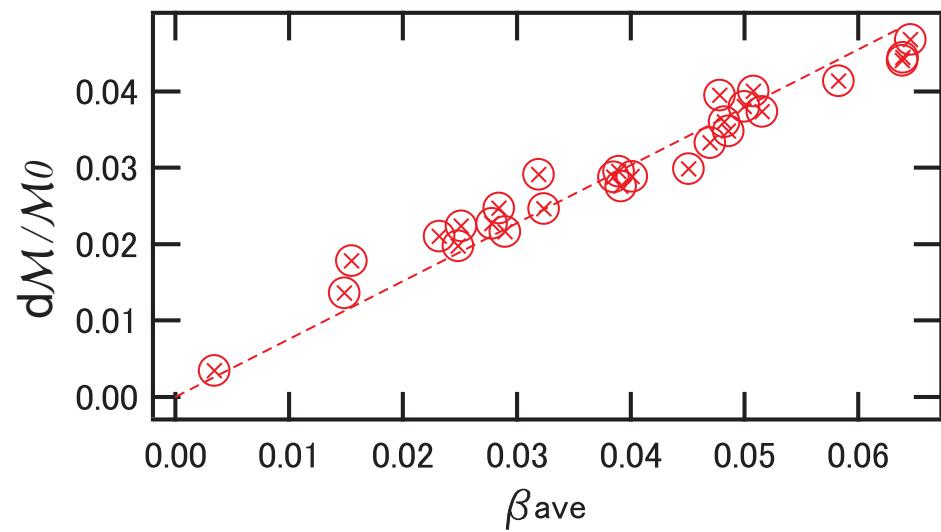
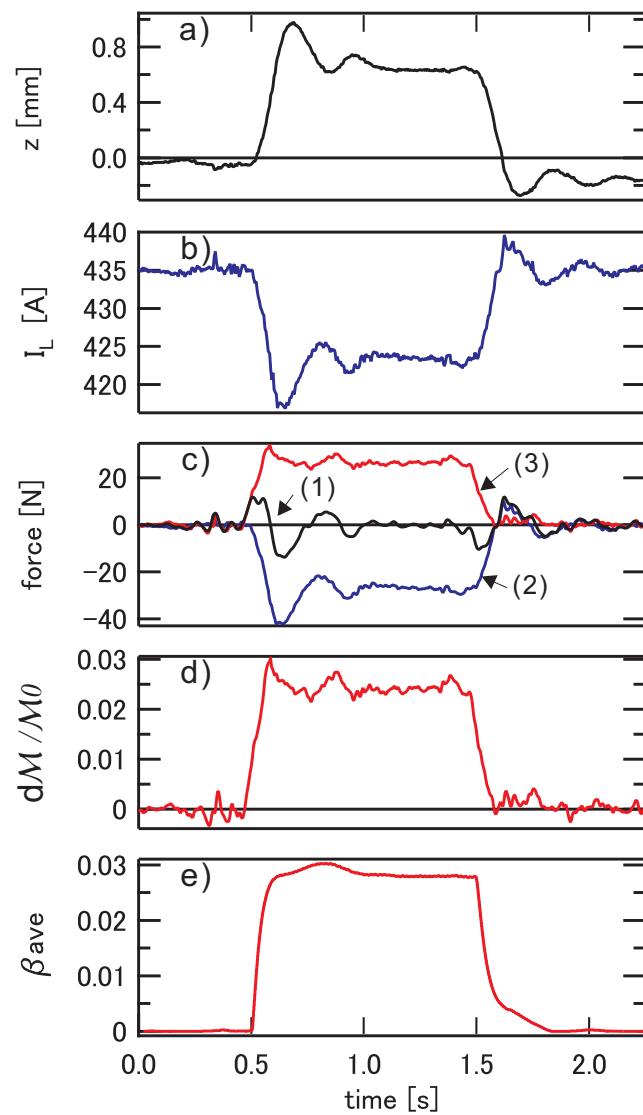
- Poisson operator

$$J(\omega) = \{\omega, \cdot\} : H_0^1(\Omega) \rightarrow H^{-1}(\Omega)$$

$$[\psi, \varphi] = \langle J(\omega)\psi, \varphi \rangle := \langle \omega, \{\psi, \varphi\} \rangle \quad [\omega \in C(\Omega)]$$

- *Known to be classically solvable for a Hölder continuous initial value: T. Kato (1967)*

Creation of magnetic moment by heating



$$\delta Q_{\text{ECH}} = \bar{B} dM = \bar{B} \sum_j d\mu_j$$