Lecture 2

ICTP 2015 Summer school

Neutrinos Selected topics



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Propagation of wave packets What happens?

Phase difference change

Due to different masses (dispersion relations) → phase velocities

Oscillations

Separation of wave packets

Due to different group velocities

Loss of coherence

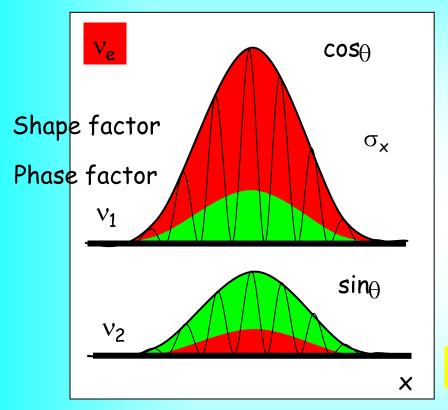
Spread of individual wave packets

Due to presence of waves with different momenta and energy in the packet

Loss of coherence within WP

Wave packet picture

2v- example



$$v_e = \cos\theta v_1 + \sin\theta v_2$$
 $v_{\mu} = -\sin\theta v_1 + \cos\theta v_2$
opposite phase

$$v_1 = \cos\theta v_e - \sin\theta v_\mu$$

$$v_2 = \cos\theta v_\mu + \sin\theta v_\epsilon$$

Interference of the same flavor parts

φ = 0

effective, frequency

$$|v(x,t)\rangle = \cos_{\theta} g_1(x - v_1 t)e^{i\phi_1}|v_1\rangle + \sin_{\theta} g_2(x - v_2 t)e^{i\phi_2}|v_2\rangle$$

$$\phi = \phi_2 - \phi_1$$

Oscillation phase

Changes with (x,t), for $\phi \neq 0$ components v_{μ} will not cancel \rightarrow appearance of v_{μ}

Oscillation phase

$$\phi = \phi_2 - \phi_1$$

$$\phi = \phi_2 - \phi_1 \qquad \phi_i = - E_i + p_i x$$

$$p_i = \sqrt{E_i^2 - m_i^2}$$

Dispersion relation

$$\phi = \Delta Et - \Delta px$$

These are averaged characteristics of WP

where

$$\Delta p = (dp/dE)\Delta E + (dp/dm^2)\Delta m^2 = 1/v_g \Delta E + (1/2p) \Delta m^2$$
insert

group velocity

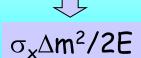
$$\phi = \Delta E/v_g (v_g t - x) + \frac{\Delta m^2}{2E} x$$
standard oscillation

 $\Delta E \sim \Delta m^2/2E$

 $<\sigma_{x}$

oscillation phase

Averaged energies

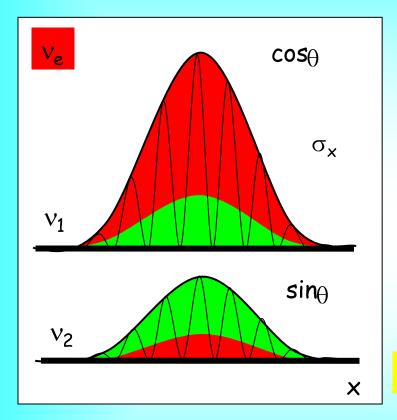


Oscillation effect over the size of WP usually-small



Phase difference along the wave packets is the same

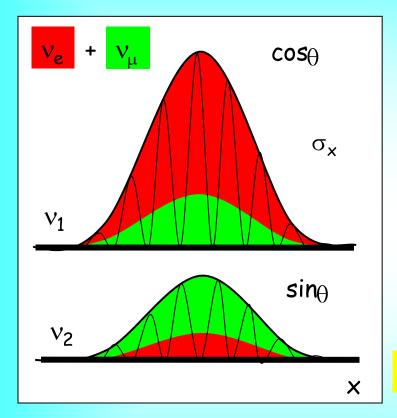
Oscillations



- Destructive interference of the muon parts
- Constructive interference of electron parts

 $\phi = 0$

Oscillations



- Destructive interference of the electron parts
- Constructive interference of muon parts

$$\phi = \pi$$

Detection:

As important as production should be considered symmetrically with production

Detection effect can be included in the generalized shape factors

$$g_k(x - v_k t) \rightarrow G_k(L - v_k t)$$

 $x \rightarrow L$ - distance between central points of the production and detection regions

HOMEWORK...

Oscillation probability

Amplitude of (survival) probability

$$A(v_e) = \langle v_e | v(x,t) \rangle = \cos^2\theta g_1(x - v_1 t) + \sin^2\theta g_2(x - v_2 t) e^{i\phi}$$

Probability in the moment of time t

$$P(v_e) = \int dx |\langle v_e | v(x,t) \rangle|^2 =$$

interference

=
$$\cos^4\theta + \sin^4\theta + 2\sin^2\theta \cos^2\theta \cos \phi \int dx g_1(x - v_1 t) g_2(x - v_2 t)$$

if $\int dx |g|^2 = 1$

If
$$g_1 = g_2$$

$$P(v_e) = 1 - 2 \sin^2\theta \cos^2\theta (1 - \cos\phi) = 1 - \sin^2 2\theta \sin^2\frac{1}{2}\phi$$

$$\phi = \frac{\Delta m^2 x}{2E} = \frac{2 \pi x}{L}$$



depth of oscillations

$$I_{v} = \frac{4 \pi E}{\Delta m^{2}}$$

Oscillation length

Physics summary

Appearance probability

$$P(v_{\mu}) = \sin^2 2\theta \sin^2 \frac{\pi x}{I_{\nu}}$$

All complications are "absorbed" in normalization or reduced to partial averaging of oscillations or lead to negligible corrections of order $\,\mathrm{m/E} << 1\,$

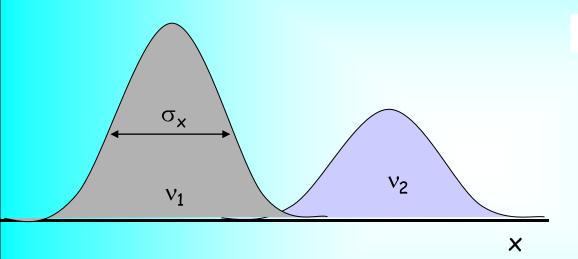
Oscillations - effect of the phase difference increase between mass eigenstates

Admixtures of the mass eigenstates v_i in a given neutrino state do not change during propagation

Flavors (flavor composition) of the eigenstates are fixed by the vacuum mixing angle

Coherence in propagation

In the configuration space: separation of the wave packets due to difference of group velocities



$$\Delta v_{qr} = \Delta m^2 / 2E^2$$

separation: $\Delta v_{qr} L = \Delta m^2 L/2E^2$

no overlap: $\Delta v_{gr} L > \sigma_x$

coherence length:

$$L_{coh} = \sigma_x E^2/\Delta m^2$$

In the energy space: averaging over oscillations

Oscillatory period in the energy space $E^T = 4\pi E^2/(\Delta m^2 L)$

Averaging (loss of coherence) if energy resolution $\sigma_{\rm F}$ is $E^{\rm T} < \sigma_{\rm E}$

→ leads to the same coherence length

If $E^T > \sigma_F$ - restoration of coherence even if the wave packets separated

Equivalence of considerations in p- and x- spaces

on blackboard...

L Stodolsky theorem

or in stationary state approximation

if there is no time tagging

Wave packets are unnecessary for computation of observable effects

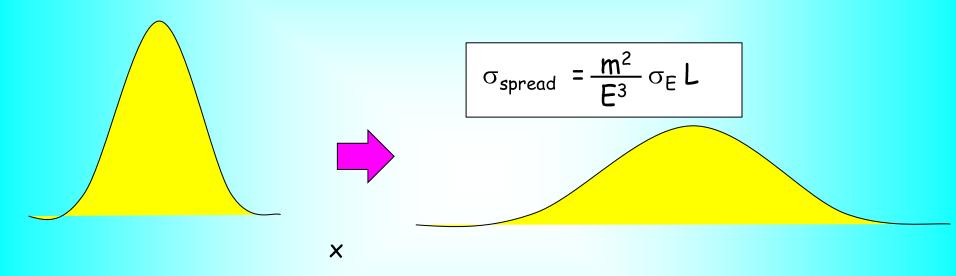
The only what is needed to make correct integration over

Production energy spectrum

Energy resolution of a detector

Spread of wave packets **

on blackboard...



Loss of coherence between different parts of the WP

Becomes classical describing that the highest energy neutrinos arrive first

No effect if considered in the p-space

Master equation

If loss of coherence and other complications related to WP picture are irrelevant -"point-like" picture

$$i \frac{d \Psi}{d t} = H \Psi$$

$$\Psi = \begin{bmatrix} \psi_e \\ \psi_\mu \\ \psi_\tau \end{bmatrix}$$

$$H = \frac{M M^+}{2E} + V(t)$$

generalization of

$$E \sim p + \frac{m^2}{2E}$$

V - matter effects
M is the mass matrix $V = diag(V_e, 0, 0) - effective potential$

Mixing matrix in vacuum

$$M M^{+} = U M_{diag}^{2} U^{+}$$

 $M_{diag}^{2} = diag (m_{1}^{2}, m_{2}^{2}, m_{3}^{2})$

Neutrino polarization vectors

$$\psi = \begin{pmatrix} v_e \\ v_{\tau} \end{pmatrix} \qquad Polarization vector: \\ P = \psi^+ \sigma/2 \psi \qquad (*)$$

$$P = \psi^+ \sigma/2 \psi \qquad (*)$$

$$P = \begin{pmatrix} Re \ v_e^+ v_{\tau}, \\ Im \ v_e^+ v_{\tau}, \\ v_e^+ v_e - 1/2 \end{pmatrix}$$

Evolution equation:

$$i \frac{d\Psi}{dt} = H\Psi$$

$$i \frac{d\Psi}{dt} = (B \sigma) \Psi$$
where $B = \frac{2\pi}{I_m} (\sin 2\theta_m, 0, \cos 2\theta_m)$

Differentiating P from (*) and using equation of motion for Ψ

$$\frac{dP}{dt} = (B \times P)$$

Coincides with equation for the electron spin precession in the magnetic field

Graphical representation

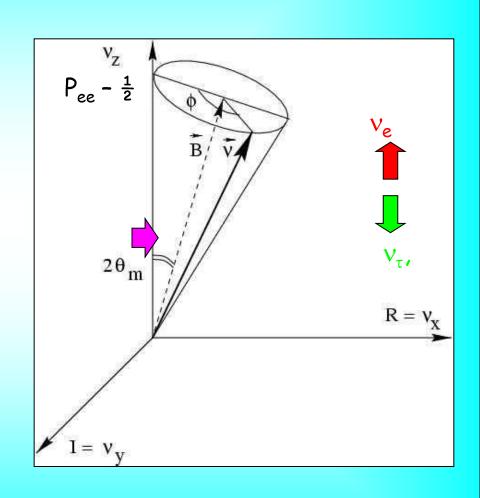
$$\overrightarrow{v} = P =$$
(Re $v_e^+ v_\tau$, Im $v_e^+ v_\tau$, $v_e^+ v_e - 1/2$)

$$\mathbf{B} = \frac{2\pi}{I_{\rm m}} (\sin 2\theta_{\rm m}, 0, \cos 2\theta_{\rm m})$$

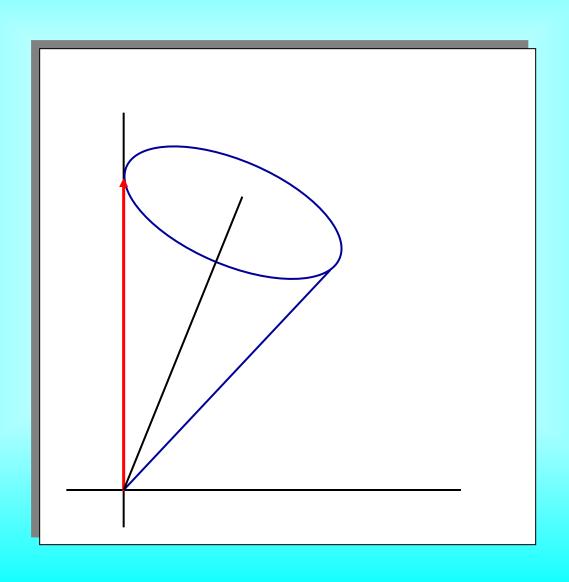
Evolution equation

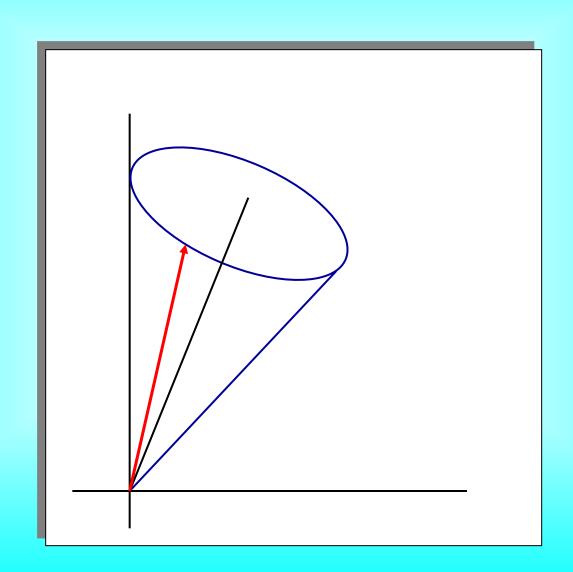
$$\frac{\overrightarrow{dv}}{dt} = (\overrightarrow{B} \times \overrightarrow{v})$$

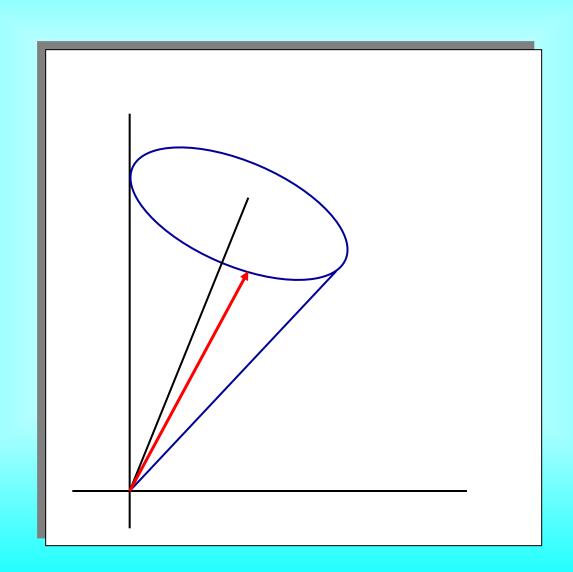
$$\phi = 2\pi t/I_m$$
 - phase of oscillations

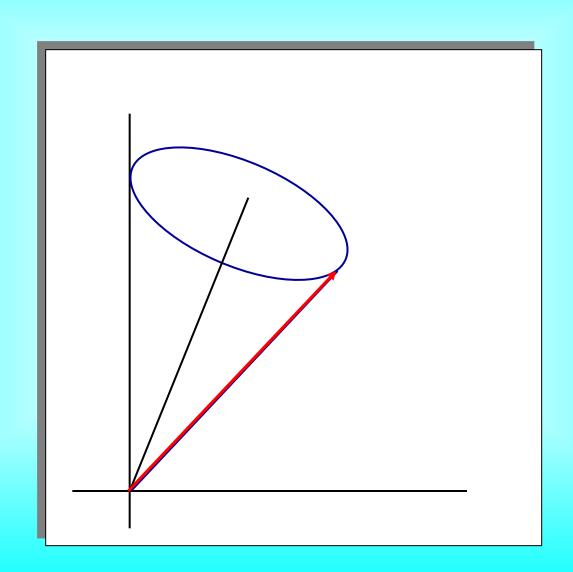


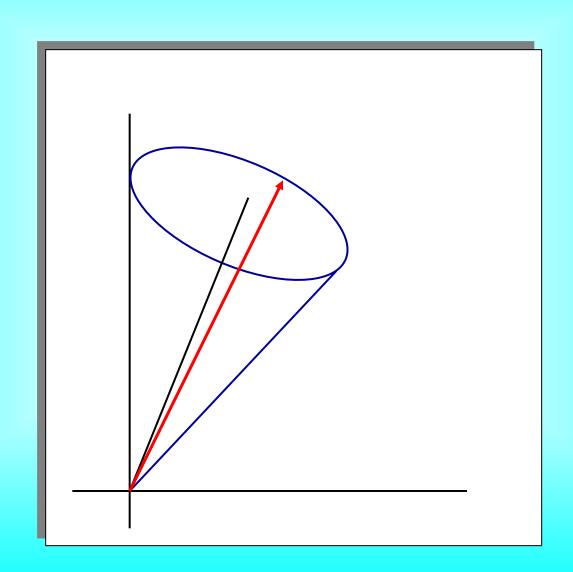
Oscillations

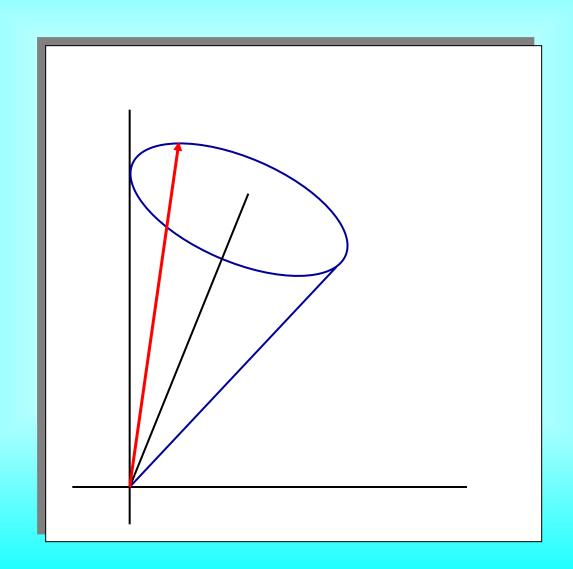


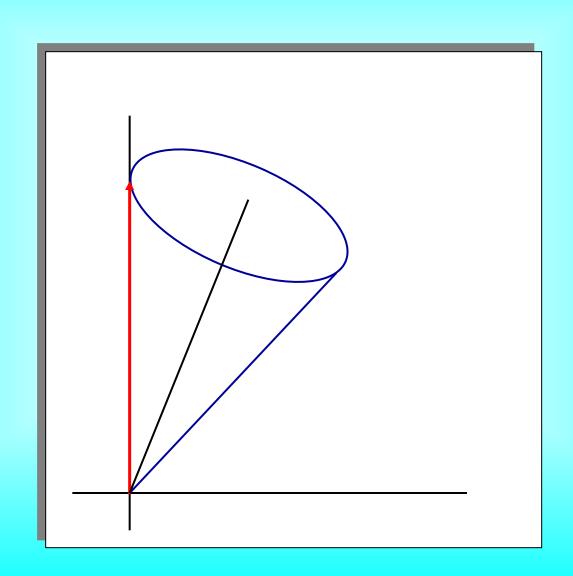




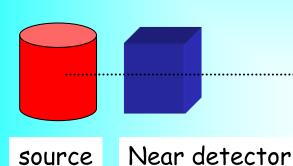




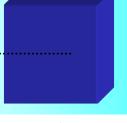




Experimental set-up



L - baseline



Far detector



 v_{α}

disappearance

 $\nu_{\!eta}$

appearance

$$\phi = \frac{\Delta m^2 x}{2E}$$

Oscillation probability - periodic function of

- distance L and
- inverse energy 1/E

Matter effects: Oscillations & flavor conversion

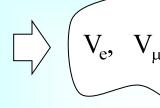
Matter potential

L. Wolfenstein, 1978

for
$$v_e v_\mu$$

at low energies Re A >> Im A inelastic interactions can be neglected

Elastic forward scattering



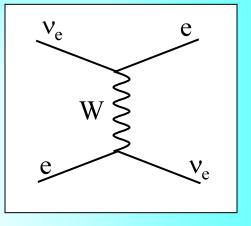
potentials

Refraction index:

$$n-1=V/p$$

for E = 10 MeV

$$n-1 = \begin{cases} \sim 10^{-20} & \text{inside the Earth} \\ < 10^{-18} & \text{inside the Sun} \end{cases}$$



difference of potentials

$$V = V_e - V_{\mu} = \sqrt{2} G_F n_e$$

 $V \sim 10^{-13}$ eV inside the Earth

Refraction length:

$$l_0 = \frac{2\pi}{V}$$

derivation

Matter potential

At low energies: neglect the inelastic scattering and absorption effect is reduced to the elastic forward scattering (refraction) described by the potential V:

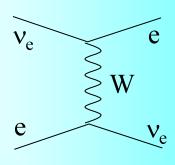
$$H_{int}(v) = \langle \psi \mid H_{int} \mid \psi \rangle = V \overline{\nu} v$$



CC interactions with electrons

$$H_{int} = \frac{G_F}{\sqrt{2}} \overline{\nu} \gamma^{\mu} (1 - \gamma_5) \nu \overline{e} \gamma_{\mu} (1 - \gamma_5) e$$

 ψ is the wave function of the medium



$$\begin{array}{ll} <\overline{e}\,\gamma_0\,(1-\gamma_5)\,e>=n_e & \text{- the electron number density} \\ <\overline{e}\,\stackrel{\rightarrow}{\gamma}\,e>=n_e\stackrel{\rightarrow}{v} \\ \end{array}$$

 $<\overline{e}\stackrel{\rightarrow}{\gamma}\gamma_5\,e>=\,n_e\stackrel{\rightarrow}{\lambda_e}\,$ - averaged polarization vector of $\,e\,$

For unpolarized medium at rest:

$$V = \sqrt{2} G_F n_e$$

Mixing in matter

in vacuumi

Effective Hamiltonian

Eigenstates

Eigenvalues

H₀

 v_1, v_2

 $m_1^2/2E$, $m_2^2/2E$

 \Rightarrow

 $H = H_0 + V$

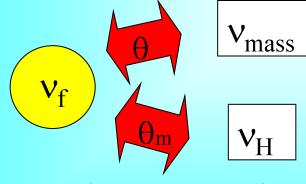
in matter:

 v_{1m} , v_{2m}

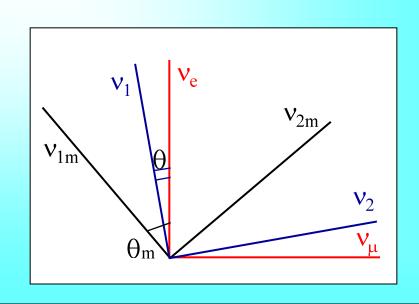
 H_{1m} , H_{2m}

depend on n_e, E

instantaneous



Mixing angle determines flavors (flavor composition) of eigenstates of propagation



ution equation

$$i \frac{dv_f}{dt} = H_{tot} v_f$$

$$oldsymbol{v}_{
m f} = \left(egin{array}{c} oldsymbol{v}_{
m e} \ oldsymbol{v}_{
m \mu} \end{array}
ight)$$

$$H_{tot} = H_{vac} + V$$
 is the total Hamiltonian

$$H_{\text{vac}} = \frac{M^2}{2E}$$
 is the vacuum (kinetic) part

$$\mathbf{V} = \begin{bmatrix} \mathbf{V}_{e} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} \end{bmatrix} \quad \text{matter part} \qquad \mathbf{V}_{e} = \sqrt{2} \,\mathbf{G}_{F} \mathbf{n}_{e}$$

$$V_e = \sqrt{2} G_F n_e$$

 H_{tot}

$$i \frac{d}{dt} \begin{bmatrix} v_e \\ v_{\mu} \end{bmatrix} = \begin{bmatrix} -\frac{\Delta m^2}{2E} \cos 2\theta + V_e & \frac{\Delta m^2}{4E} \sin 2\theta \\ \frac{\Delta m^2}{4E} \sin 2\theta & 0 \end{bmatrix} \begin{bmatrix} v_e \\ v_{\mu} \end{bmatrix}$$

Mixing in matter

Diagonalization of the Hamiltonian:

$$\sin^2 2\theta_{\rm m} = \frac{\sin^2 2\theta}{(\cos 2\theta - 2EV/\Delta m^2)^2 + \sin^2 2\theta}$$

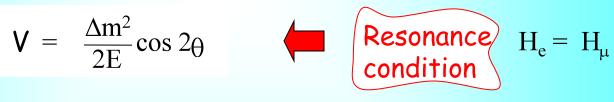
$$V = \sqrt{2} G_F n_e$$

Mixing is maximal if

$$V = \frac{\Delta m^2}{2E} \cos 2\theta$$

 $\sin^2 2\theta_m = 1$



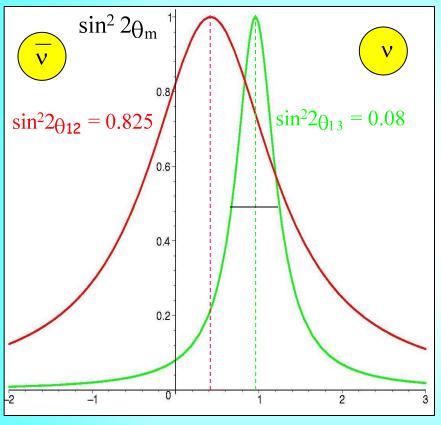


$$H_e = H_{\mu}$$

Difference of the eigenvalues

$$H_{2m} - H_{1m} = \frac{\Delta m^2}{2E} \sqrt{(\cos 2\theta - 2EV/\Delta m^2)^2 + \sin^2 2\theta}$$

Resonance



 $l_v / l_0 \sim n E$

Resonance width: Resonance layer:

$$\Delta n_R = 2n_R \tan 2\theta$$

 $n = n_R + / - \Delta n_R$

In resonance:

$$sin^2 2_{\theta m} = 1$$

Flavor mixing is maximal

$$1_{v} = 1_{0} \cos 2\theta$$



$$\approx$$

Refraction length

Level crossing

V. Rubakov, private comm.

N. Cabibbo, Savonlinna 1985

H. Bethe, PRL 57 (1986) 1271

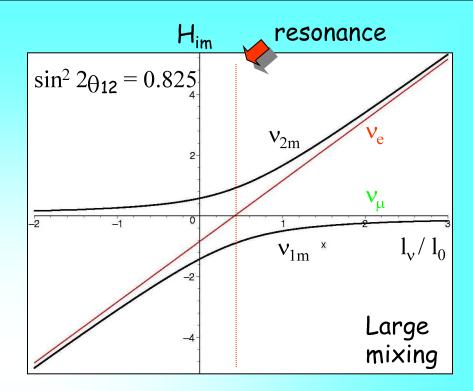
Dependence of the neutrino eigenvalues on the matter potential (density):

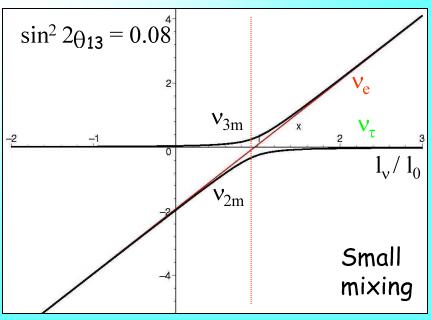
$$\frac{l_{\nu}}{l_0} = \frac{2E V}{\Delta m^2}$$

$$\frac{l_{v}}{l_{0}} = \cos 2\theta$$

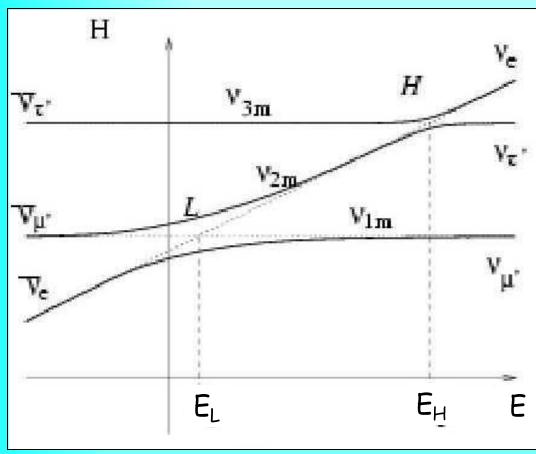
Crossing point - resonance

- the level split is minimal
- the oscillation length is maximal





Level crossings



Normal mass hierarchy

0.1 GeV

6 GeV

Resonance region

High energy range

Oscillations in matter

Oscillation probability constant density

$$P(v_e \rightarrow v_a) = \sin^2 2\theta_m \sin^2 \left(\frac{\pi L}{l_m}\right)$$
 half-phase ϕ Amplitude of oscillatory factor oscillations

$$\theta_m(E, n)$$
 - mixing angle in matter $l_m(E, n)$ - oscillation length in matter

$$l_{\rm m} = 2 \, \pi / (H_{\rm 2m} - H_{\rm 1m})$$

In vacuum:

$$\theta_{m} \to \theta$$

$$l_{m} \to l_{v}$$

Maximal effect:

$$\sin^2 2\theta_{\rm m} = 1$$



MSW resonance condition

$$\phi = \pi/2 + \pi k$$

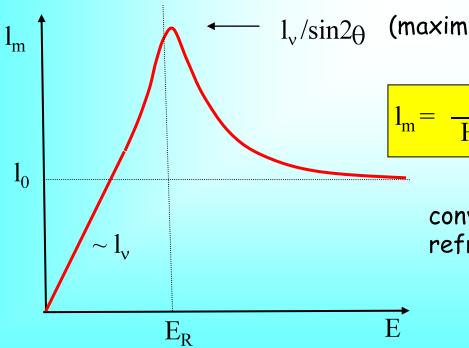
Oscillation length in matter

Oscillation length in vacuum

$$1_{v} = \frac{4\pi E}{\Delta m^2}$$

Refraction length

- determines the phase produced by interaction with matter



(maximum at $l_v = l_0 / \cos 2\theta$)

shifts with respect resonance energy:

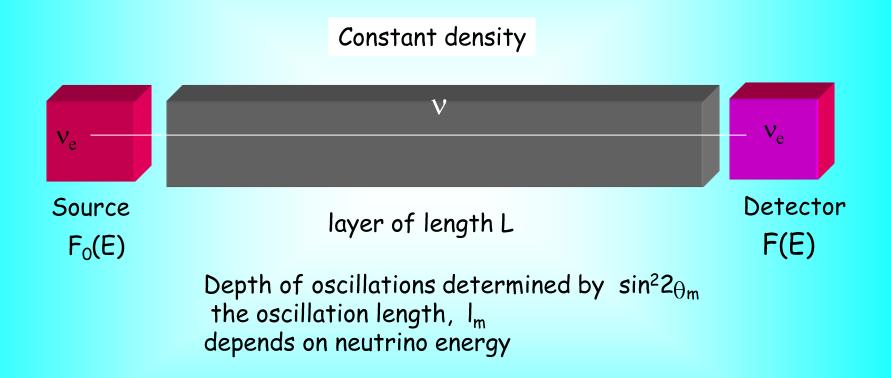
$$l_v(E_R) = l_0 \cos 2\theta$$

converges to the refraction length

Resonance enhancement of oscillations

Constant density

Resonance enhancement

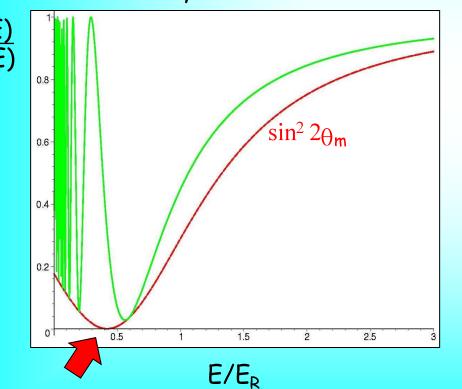


For neutrinos propagating in the mantle of the Earth

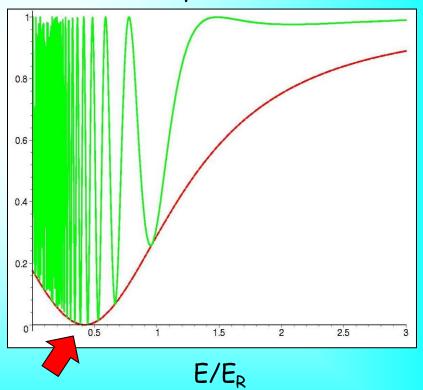
Large mixing $\sin^2 2\theta = 0.824$

Layer of length L $k = \pi L / l_0$

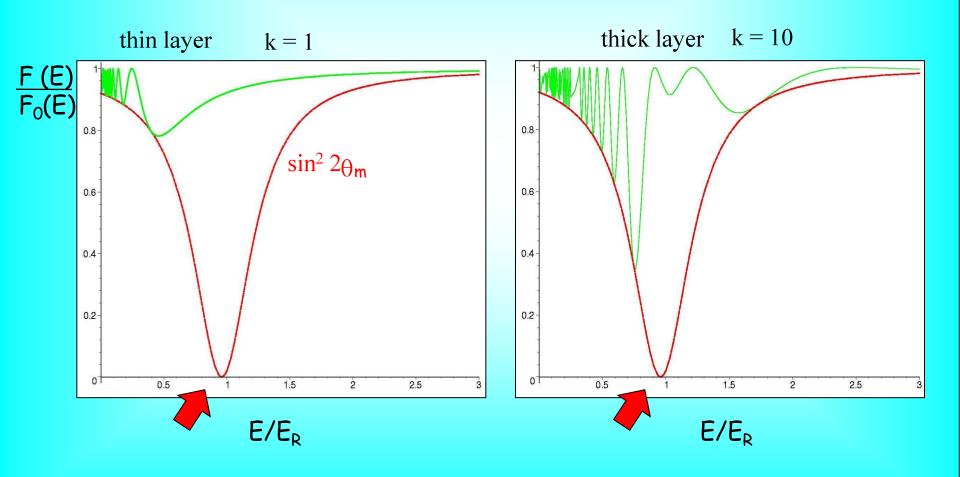




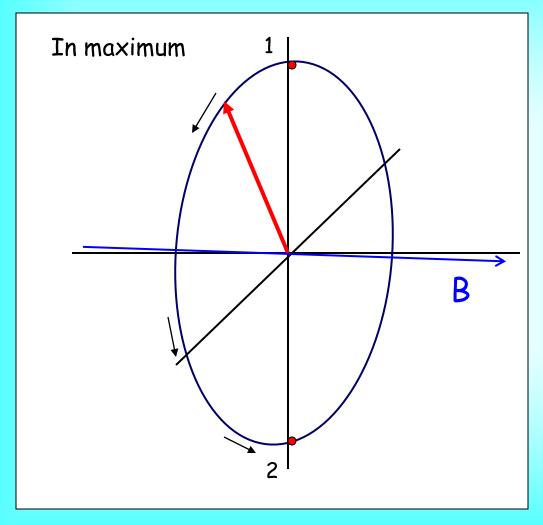
thick layer k = 10



Small mixing $\sin^2 2\theta = 0.08$

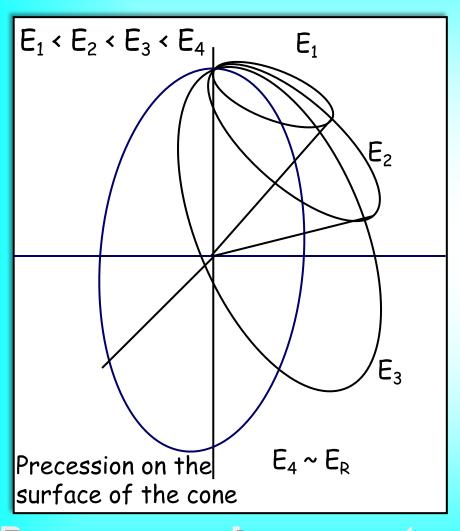


Resonance enhancement



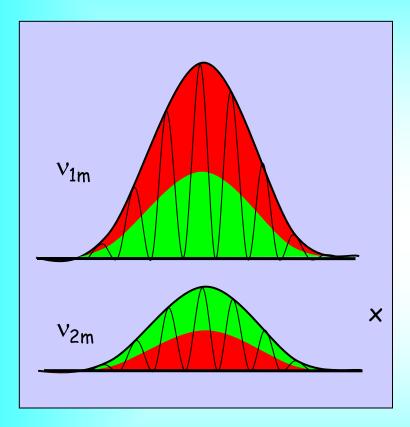


Graphic representation



Resonance enhancement

Oscillations in matter



Constant density medium: the same dynamics

Mixing changed phase difference changed

$$H_0 \rightarrow H = H_0 + V$$

$$v_k \rightarrow v_{mk}$$

eigenstates of H₀ eigenstates of H

$$\theta \rightarrow \theta_{m} (n)$$

Resonance - maximal mixing in matter - oscillations with maximal depth

$$\theta_{\rm m}$$
 = $\pi/4$

Resonance condition:

$$V = \cos 2\theta \frac{\Delta m^2}{2E}$$