## Challenges and future scenarios for High Energy Physics

International Centre for Theoretical Physics

17 June 2015

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## **The Standard Model of particle physics**



#### **Status of the Standard Model**

#### < 1973: theoretical foundations of the SM</p>

- renormalizability of SU(2)xU(1) with Higgs mechanism for EWSB
- asymptotic freedom, QCD as gauge theory of strong interactions
- KM description of CP violation

#### • Followed by 40 years of consolidation:

- **experimental** verification, via **discovery** of
  - **Fermions**: charm, tau, bottom, top (all discovered in the USA)
  - **Bosons**: gluon, W and Z, **Higgs** (all discovered in Europe)
- technical theoretical advances (higher-order calculations, lattice QCD, ...)
- experimental consolidation, via precision measurement of
  - EW radiative corrections
  - running of  $\alpha_s$
  - CKM parameters, ....

#### Remains to be verified:

mechanism at the origin of particles' masses: is the Higgs boson dynamics what prescribed by the SM, or are there other phenomena at work?

## On particles' masses

For a composite system the mass is obtained by solving the dynamics of the bound state  $\Rightarrow$  m=<E>/c<sup>2</sup> with <E>=<T+U>

Example: the proton mass. Dynamics of quarks and gluons inside the proton (they have negligible masses) ⇒ m<sub>D</sub> = 938 MeV

But what about elementary particles? Elementary  $\Rightarrow$  no internal dynamics



Need to develop a new framework within which to understand the origin and value of, for example, the electron mass

#### **However:**

- Why do we need a mechanism to accommodate the masses of elementary particles?
- How about just assigning mass values as parameters?

#### In other words:

WHY are particle physicists so obsessed with the problem of particles' masses?



+ 2 more "families" differing from the 1st one only in the mass of their elements

L-chirality R-chirality

## Parity asymmetry and mass for spin-1/2 particles

 $H \propto i \overline{\psi_L} \,\partial \cdot \gamma \,\psi_L + i \overline{\psi_R} \,\partial \cdot \gamma \,\psi_R + m \,\overline{\psi_L} \,\psi_R$ 



For a massive particle, chirality does not commute with the Hamiltonian, so it cannot be conserved

Chirality eigenstates cannot be Hamiltonian (physical) eigenstates

Nothing wrong with that in principle .... but chirality cannot be associated to a conserved charge!

The symmetry associated with the conservation of the weak charge must therefore be broken for leptons and quarks to have a mass

#### The SM solution ....

Time evolution of a massive particle:



The transition between L and R states, and the absorption of the changes in weak charge, are ensured by the interaction with a background scalar field,  $\mathbf{H}$ . Its "vacuum density" provides an infinite reservoir of weak charge.

The number "**v**" is the expectation value of the so-called **Higgs field**. The quantity " $\lambda$ " is characteristic of the particle interacting with the Higgs field. It can easily be shown that **this interaction leads to a mass m** ~  $\lambda$  **v** 



## Why is it difficult to study the Higgs ?

Like any other medium, the Higgs continuum background can be perturbed. Similarly to what happens if we bang on a table, creating sound waves, if we "bang" on the Higgs background we can stimulate "Higgs waves", i.e. what we call the Higgs boson ...

This requires not just energy (enough to create the H), but a large-mass probe (the H couples to mass, not to energy!)

Thus we typically need not just the energy required to produce the H, but also the energy required to produce the heavy particles that will stimulate its emission ...

☐ low rates, complex final states, large backgrounds, ....

\* Higgs particles are thus a bit like phonons ...

## Four main production mechanisms







#### Note on "the mass of the Universe"

- proton's mass arises from QCD dynamics, not from the mass of its constituent quarks. Half of it is kinetic energy of the tightly bound relativisitic quarks, the other half is binding energy ( $M=Ec^2$ , E=K+U, virial theorem....)

- the mass of particles composing Dark Matter does not need to arise from the coupling with the Higgs. E.g. in Supersymmetry models it could mostly come from the breaking of supersymmetry, nothing to do with the Higgs or EWSB

## What's to be learned from the Higgs, now that's been found?

The Higgs boson is directly connected to several key questions:

- What's the real origin of the Higgs potential, which breaks EW symmetry?
  - underlying strong dynamics? composite Higgs?
  - RG evolution from GUT scales, changing sign to quadratic term in V(H)?
  - Are there other Higgs-like states (e.g. H<sup>±</sup>, A<sup>0</sup>, H<sup>±±</sup>, ..., EW-singlets, ....) ?
- What happens at the EW phase transition (PT) during the Big Bang?
  - what's the order of the phase transition?
  - are the conditions realized to allow EW baryogenesis?
  - does the PT wash out possible pre-existing baryon asymmetry?
- Is there a relation between Higgs, EWSB, baryogenesis and Dark Matter?
- The hierarchy problem: what protects the smallness of  $m_H / m_{Plank,GUT,...}$ ?

## **Higgs selfcouplings**

The Higgs sector is defined in the SM by two parameters,  $\mu$  and  $\lambda$ : V(H)

$$\frac{\partial V_{SM}(H)}{\partial H}|_{H=v} = 0 \quad \text{and} \quad m_H^2 = \frac{\partial^2 V_{SM}(H)}{\partial H \partial H^*}|_{H=v} \quad \Rightarrow \quad \begin{array}{l} \mu = m_H \\ \lambda = \frac{m_H^2}{2v^2} \end{array}$$

These relations uniquely determine the strength of Higgs selfcouplings in terms of  $m_{\text{H}}$ 

$$\cdots \cdots \Rightarrow \mathbf{S}_{\mathbf{S}_{\mathbf{H}}} \Rightarrow 6\lambda v = \frac{3m_{H}^{2}}{v} \mathbf{\sim O(m_{top})} \qquad \mathbf{S}_{\mathbf{S}_{\mathbf{H}}} \Rightarrow 6\lambda = \frac{3m_{H}^{2}}{v^{2}} \mathbf{\sim O(I)}$$

Testing these relations is therefore an important test of the SM nature of the Higgs mechanism



Higgs selfcoupling and coupling to the top are the key elements to define the stability of the Higgs potential



0.10  $3\sigma$  bands in 0.08  $M_t = 173.1 \pm 0.6 \, \text{GeV} \, (\text{gray})$  $\alpha_3(M_Z) = 0.1184 \pm 0.0007$ (red) 0.06  $M_h = 125.7 \pm 0.3 \text{ GeV}$  (blue) Higgs quartic coupling  $\lambda$ 0.04 0.02  $M_t = 171.3 \, \text{GeV}$ 0.00  $\alpha_s(M_Z) = 0.1205$  $\alpha_s(M_Z) = 0.1163$ -0.02 $M_t = 174.9 \, \text{GeV}$ -0.0410<sup>10</sup> 1012  $10^{2}$  $10^{4}$  $10^{6}$  $10^{8}$  $10^{14}$   $10^{16}$   $10^{18}$ 1020

RGE scale  $\mu$  in GeV

Degrassi et al, http://arxiv.org/pdf/1205.6497

## The nature of the EW phase transition



Strong I<sup>st</sup> order phase transition  $\Rightarrow \langle \Phi_C \rangle > T_C$ 

In the SM this requires  $m_H \approx 80 \text{ GeV} \Rightarrow \text{new physics}$ , coupling to the Higgs and effective at scales O(TeV), must modify the Higgs potential to make this possible

Understanding the role of the EWPT in the evolution or generation of the baryon asymmetry of the Universe is a key target for future accelerators

- Experimental probes:
  - study of triple-Higgs couplings (... and quadruple, etc)
  - search for components of an extended Higgs sector (e.g. 2HDM, extra singlets, ...)
  - search for new sources of CP violation, originating from (or affecting) Higgs interactions

## H, the hierarchy problem, and physics beyond the SM

Calculating the radiative corrections to the Higgs mass in the SM poses an intriguing puzzle:

$$m_{H}^{2} = m_{0}^{2} - \frac{6G_{F}}{\sqrt{2}\pi^{2}} \left( m_{t}^{2} - \frac{1}{2}m_{W}^{2} - \frac{1}{4}m_{Z}^{2} - \frac{1}{4}m_{H}^{2} \right) \Lambda^{2} \sim m_{0}^{2} - (125 \,\text{GeV})^{2} \left( \frac{\Lambda}{400 \,\text{GeV}} \right)^{2}$$

$$\xrightarrow[\text{antitop}]{H^{-1}}_{\text{top}} + \frac{W}{H^{-1}} + \frac{W}{H^{-1}} + \dots \qquad \stackrel{\text{A= scale up to which the SM is valid}}{\text{walid}}$$

#### renormalizability =>

$$m_H^2(v) \sim m_H^2(\Lambda) - (\Lambda^2 - v^2)$$
,  $v = \langle H \rangle \sim 250 \text{GeV}$ 

Assuming  $\Lambda$  can extend up to the highest energy beyond which quantum gravity will enter the game, 10<sup>19</sup> GeV, keeping m<sub>H</sub> below 1 TeV requires a fine tuning among the different terms at a level of 10<sup>-34</sup>:

$$\frac{m_H^2(\Lambda) - \Lambda^2}{\Lambda^2} \sim \frac{v^2}{\Lambda^2} = O(10^{-34}) \text{ if } \Lambda \sim M_{Planck}$$

extremely **unnatural** if it is to be an accident !!

hierarchy, or fine tuning, problem

# **Higgs self-energy, Susy fix**



stability of the natural scale of the Higgs mass restored!

 $m_H \leq M_Z$  + radiative corrections ( $\propto \log(m_t/m_{stop}) \leq 135 \text{ GeV}$ 

## More in general ....

Tie the Higgs mass to some symmetry which protects it against quadratic divergencies

Supersymmetry

H (scalar) ↔ fermion

Gauge symmetry

H (scalar) ↔ 5th component of a gauge bosons in 5 dimensions or more

=> extra dimensional theories

Global symmetry

$$H \rightarrow H + a \Rightarrow L(H) = L(\partial H)$$

=> Little Higgs theories, Technicolor H=pseudo-goldstone boson

The manifestations of these new symmetries (e.g. new particles, new interactions) cannot be too far from the TeV scale, in order to solve the Higgs fine tuning issue in a **natural** way

## **Status of BSM**

- •Until few yrs ago, we had a benchmark model, MSSM, expected to deliver the following:
  - •low-mass Higgs  $h^0$ , no heavier than ~130 GeV
  - •~TeV scale squarks and gluinos, to be seen rapidly at the LHC
    - $\bullet \Rightarrow$  solution to the naturalness problem
  - extra Higgses ( $A^0 / H^0 / H^{\pm}$ ) observed at the LHC
  - candidate for DM, confirmed by direct detection
  - interesting flavour phenomenology
    - explanation of  $(g-2)_{\mu}$
    - sizable deviations from SM in  $B(B_S \rightarrow \mu^+ \mu^-)$
    - $\mu \! \rightarrow \! e \gamma$  observed at MEG, consistent with SUSY neutrino masses induced at the GUT scale
    - CPV in the Higgs or squark/gluino sector, to explain BAU
    - electric dipole moments (e, n) measured, consistent with previous point

- Given our knowledge 4-5 yrs back, all of this could have happened by now.
- Even models alternative to SUSY (extra dim, little Higgs, SILH, ...) had the potential of matching the "natural" predisposition of SUSY to solve problems and to provide rich phenomenological consequences across the fields (LHC, flavour, astro/cosmo)

## • None of the above happened.

- Thus a radical change in attitude in BSM model building is taking place, focusing on schemes that address individual issues or anomalies, leaving for later the understanding of the "grand picture"
- The above scenario may still happen, with a few-year delay, perhaps stretching a bit the "naturalness".
- This expectation is still high, and well justified

NATURALNESS, CHIRAL SYMMETRY, AND SPONTANEOUS

CHIRAL SYMMETRY BREAKING

G. 't Hooft

Institute for Theoretical Fysics

Utrecht, The Netherlands

## Naturalness is not a recent "fashion": it's an original sin of the SM itself ... See e.g.

Aug 1979. 23 pp. NATO Adv.Study Inst.Ser.B Phys. 59 (1980) 135

As we will see, naturalness will put the severest restriction on the occurrence of scalar particles in renormalizable theories. In fact we conjecture that this is the reason why light, weakly interacting scalar particles are not seen.

Pursuing naturalness beyond 1000 GeV will require theories that are immensely complex compared with some of the grand unified schemes.

A remarkable attempt towards a natural theory was made by Dimopoulos and Susskind <sup>2</sup>). These authors employ various kinds of confining gauge forces to obtain scalar bound states which may substitute the Higgs fields in the conventional schemes. In their model the observed fermions are still considered to be elementary.

Most likely a complete model of this kind has to be constructed step by step. One starts with the experimentally accessible aspects of the Glashow-Weinberg-Salam-Ward model. This model is natural if one restricts oneself to mass-energy scales below 1000 GeV. Beyond 1000 GeV one has to assume, as Dimopoulos and Susskind do, that the Higgs field is actually a fermion-antifermion composite field. Coupling this field to quarks and leptons in order to produce their mass, requires new scalar fields that cause naturalness to break down at 30 TeV or so. We're finally there, at I TeV, facing the fears about a light SM Higgs anticipated long ago

#### The observation of the Higgs where the SM predicted it would be, its SM-like properties, and the lack of BSM phenomena up to the TeV scale, make the *naturalness issue more puzzling than ever*

- Whether to keep believing in the MSSM or other specific BSM theories after LHC@8TeV is a matter of personal judgement. But the broad issue of *naturalness will ultimately require an understanding*.
- Naturalness remains a guiding principle to drive the search of new phenomena at the LHC

## **Possible reasons for the lack of signals ...**

- BSM particles are already being created at the LHC, but are hiding well:
  - compressed spectra: low MET, low ET, long lifetime heavy particles, ...
  - RPV
  - ....
- BSM is less "conventional", fine-tuning or direct search constraints less tight
  - NMSSM
  - non-degenerate squarks
  - ....
- The scale at which naturalness is restored is higher than the TeV: acceptable, but becoming less and less "natural" as the scale grows ....
- Naturalness is an ill guided principle  $\Rightarrow$  Anthropic principle

#### Example of ways out: explore less constrained SUSY models

Fraction of excluded models in the pMSSM (19 parameters MSSM)



Rizzo et al, arXiv:1211.1981

### Anomalies left over from run 1, some examples

$$Br[h \to \mu\tau] = (0.89^{+0.40}_{-0.37}) \%$$
CMS-PAS-HIG-14-005  

$$stat syst$$

$$R(K) = \frac{B \to K\mu^+\mu^-}{B \to Ke^+e^-} = 0.745^{+0.090}_{-0.074} \pm 0.036$$
LHCb, arXiv:1406.6482  

$$B \to K*\mu+\mu- \text{ anomaly}$$

$$B \to K*\mu+\mu- \text{ anomaly}$$

$$HCb, arXiv:1308.1707 \text{ and}$$

$$3fb^{-1} \text{ update LHCb-CONF-2015-002}$$

For possible interpretation within a single BSM model see e.g. Crivellin, D'Ambrosio, Heeck, arXiv:1501.00993 (2HDM w. gauged  $L_{\mu}-L_{\tau}$ )

## Anomalies left over from run 1, some examples

#### **Dileptons + jets + MET (SUSY searches)**



CMS, http://arxiv.org/abs/1502.06031

 $N_{jets (p_T>40 \text{ GeV})} \ge 2, E_T^{miss} > 150 \text{ GeV}$ or  $N_{jets (p_T>40 \text{ GeV})} \ge 3, E_T^{miss} > 100 \text{ GeV}$ 

low mass:  $m_{\parallel} = (20-70) \text{ GeV}$ On-Z:  $m_{\parallel} = (81-101) \text{ GeV}$ 

#### ATLAS, http://arxiv.org/abs/1503.03290

 $N_{jets (PT>35 GeV)} \ge 2$ ,  $E_T^{miss} > 225 GeV$  $H_T > 600 GeV$ 

On-Z: m<sub>∥</sub> = (81–101) GeV

## Anomalies left over from run 1, some examples

|                    | Low-              | mass            | On-Z             |                  |  |
|--------------------|-------------------|-----------------|------------------|------------------|--|
|                    | Central Forward   |                 | Central          | Forward          |  |
| Observed           | 860               | 163             | 487              | 170              |  |
| Flavor-symmetric   | $722\pm27\pm29$   | $155\pm13\pm10$ | $355\pm19\pm14$  | $131\pm12\pm8$   |  |
| Drell-Yan          | $8.2\pm2.6$       | $2.5\pm1.0$     | $116\pm21$       | $42\pm9$         |  |
| Total estimated    | $730 \pm 40$      | $158\pm16$      | $471\pm32$       | $173\pm17$       |  |
| Observed-estimated | $130^{+48}_{-49}$ | $5^{+20}_{-20}$ | $16^{+37}_{-38}$ | $-3^{+20}_{-21}$ |  |
| Significance       | 2.6 <i>σ</i>      | $0.3 \sigma$    | $0.4\sigma$      | <0.1 <i>o</i>    |  |

CMS, http://arxiv.org/abs/1502.06031

#### **⇒2.6** σ

... no signal on-peak

#### **σ(350 GeV) ratio | 3TeV/8TeV ~ 4.5**

#### ATLAS, http://arxiv.org/abs/1503.03290

| Channel                            | SR-Z ee             | SR-Ζ μμ                | SR-2           | Z same-flavour<br>combined |
|------------------------------------|---------------------|------------------------|----------------|----------------------------|
| Observed events                    | 16 <b>⇒3.0</b> C    | 13                     | ⇒ <b>I.6</b> σ | 29                         |
| Expected background events         | $4.2 \pm 1.6$       | $6.4 \pm 2.2$          |                | $10.6 \pm 3.2$             |
| Flavour-symmetric backgrounds      | $2.8 \pm 1.4$       | $3.3 \pm 1.6$          |                | $6.0 \pm 2.6$              |
| $Z/\gamma^*$ + jets (jet-smearing) | $0.05 \pm 0.04$     | $0.02^{+0.03}_{-0.02}$ |                | $0.07 \pm 0.05$            |
| Rare top                           | $0.18 \pm 0.06$     | $0.17 \pm 0.06$        |                | $0.35 \pm 0.12$            |
| WZ/ZZ diboson                      | $1.2 \pm 0.5$       | $1.7 \pm 0.6$          |                | $2.9 \pm 1.0$              |
| Fake leptons                       | $0.1^{+0.7}_{-0.1}$ | $1.2^{+1.3}_{-1.2}$    |                | $1.3^{+1.7}_{-1.3}$        |

... but no signal off-peak

#### **σ(800 GeV) ratio |3TeV/8TeV ~ 8.5**

#### Already more than 10 TH interpretation papers on arXiv ....





## **Dark Matter**

## Our thinking has shifted K. Zurek, Aspen 2014



From a single, stable weakly interacting particle ..... (WIMP, axion)

> Models: Supersymmetric light DM sectors, Secluded WIMPs, WIMPless DM, Asymmetric DM .. Production: freeze-in, freeze-out and decay, asymmetric abundance, non-thermal mechanicsms ..

 $M_p \sim 1 \text{ GeV}$ 

Standard Model

...to a hidden world with multiple states, new interactions

ASPEN 2014: https://indico.cern.ch/event/276476/

## Evidence building up for self-interacting DM





• A really large scattering cross section!  $\sigma \sim 1 \text{ cm}^2 (\text{m}_{\text{X}}/\text{g}) \sim 2 \times 10^{-24} \text{ cm}^2 (\text{m}_{\text{X}}/\text{GeV})$ For a WIMP:  $\sigma \sim 10^{-38} \text{ cm}^2 (\text{m}_{\text{X}}/100 \text{ GeV})$ 

SIDM indicates a new mass scale

Hai-BoYu, ASPEN 2014: https://indico.cern.ch/event/276476/

More in general, interest is growing in scenarios for EWSB with rich sectors of states only coupled to the SM particles via <u>weakly interacting</u> "portals"

How to move forward, towards finding the answer to key questions such as

- What's the origin of Dark matter ?
- What's the origin of matter/antimatter asymmetry in the universe?
- What's the origin of neutrino masses?
- What determines the number and interactions of different families of quarks and leptons?

## The "tools"

- Direct exploration of physics at the weak scale through highenergy colliders (linear/circular, ee/pp/ep/µµ)
- Quarks: flavour physics, EDM's
- Neutrinos: CP violation, mass hierarchy and absolute scale, majorana nature
- Charged leptons: flavour violation, g-2, EDMs
- Axions, axion-like's (ALPs), dark photons, ....

There is no experiment/facility, proposed or conceivable, in the lab or in space, accelerator or non-accelerator driven, which can guarantee to find an answer to any of the questions above

#### ⇒

- target broad and well justified scenarios
- consider the potential of given facilities to provide conclusive answers to relevant (and answerable!) questions
  - can we identify forms of no-lose theorems ?
- weigh the value of knowledge that will be acquired, no matter what, by a given facility (the value of "measurements")

Most of the "big questions" touch directly on weak scale physics.

There are relevant, well defined questions, whose answer can be found exploring the TeV scale, and which can help guide the evaluation of the future exptl facilities. E.g.

#### • Dark matter

• is TeV-scale dynamics (e.g. WIMPs) at the origin of Dark Matter ?

#### Baryogenesis

did it arise at the cosmological EW phase transition ?

### • EW Symmetry Breaking

what's the underlying dynamics? weakly interacting? strongly interacting? other interactions, players at the weak scale besides the SM Higgs ?

#### • Hierarchy problem

"natural" solution, at the TeV scale?

The exploration of the high-energy frontier can provide conclusive answers to several of these questions

- A complete study of the Higgs boson, of its interactions and of EWSB is a guaranteed deliverable of this programme ...
- ... accompanied by an ambitious discovery potential, sensitive to possible manifestations of new physics at the TeV scale

To address the scenarios raised by the question of "why don't we see new physics at the LHC" (i.e. (i) scale of new physics is too large, or (ii) signals are elusive), future facilities should guarantee

- precision
- sensitivity (to elusive signatures)
- extended energy/mass reach

The known faces at the energy frontier are the linear e<sup>+</sup>e<sup>-</sup> colliders, namely ILC and CLIC

The new kids in town: circular colliders

## **Dec 2011** Latest LHC data corner the Higgs boson to within a small mass window in the 115-130 GeV range

CERN-OPEN-2011-047 20 January 2012 Version 2.9 arXiv:1112.2518v1 [hep-ex]

#### A High Luminosity e<sup>+</sup>e<sup>-</sup> Collider in the LHC tunnel to study the Higgs Boson

Alain Blondel<sup>1</sup>, Frank Zimmermann<sup>2</sup> <sup>1</sup>DPNC, University of Geneva, Switzerland; <sup>2</sup>CERN, Geneva, Switzerland

**Abstract:** We consider the possibility of a 120x120 GeV e+e- ring collider in the LHC tunnel. A luminosity of  $10^{34}$ /cm<sup>2</sup>/s can be obtained with a luminosity life time of a few minutes. A high operation efficiency would require two machines: a low emittance collider storage ring and a separate accelerator injecting electrons and positrons into the storage ring to top up the beams every few minutes. A design inspired from the high luminosity b-factory design and from the LHeC design report is presented. Statistics of about 2x10<sup>4</sup> HZ events per year per experiment can be collected for a Standard Higgs Boson mass of 115-130 GeV.

#### Summer 2012. Higgs discovery => submissions to European Strategy Group Symposium

From the upgrade of the accelerator infrastructure in the LHC tunnel .....

| LEP3 – Higgs factory in the LHC tunnel<br>Prepared by Frank Zimmermann, CERN, 9 April 2012; revised on 3 August 2012 | CERN-ATS-2012-237   |
|--|---|
|  | High Energy LHC<br>Document prepared for the European HEP strategy update   |
|  | Oliver Brüning, Brennan Goddard, Michelangelo Mangano*, Steve Myers,<br>Lucio Rossi, Ezio Todesco and Frank Zimmerman |
|  | CERN, Accelerator & Technology Sector<br>* CERN, Physics Department   |

#### ..... to the development of more ambitious goals

| EDMS Nr: 1233485<br>Group reference: CERN/GS-SE          | 27 July 2012          |  |
|--|-----------------------|--|
| PRE-FEASIBILITY STUDY FOR AN 80KM T                      | UNNEL PROJECT AT CERN | LEP3 and TLEP:   |
| John Osborne (CERN), Caroline Waaijer (CERN), ARUP, GADZ |                       | High luminosity e <sup>+</sup> e <sup>-</sup> circular colliders for precise Higgs   |
|  |                       |  |
|  |                       | Alain Blondel (University of Geneva), John Ellis (King's College London),<br>Patrick Janot (CERN), Mike Koratzinos (University of Geneva), Marco Zanetti<br>(MIT), Frank Zimmermann (CERN) |

## Fall 2012 The idea caught up ...



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□ .....

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Sport ▼

Accelerators for a Higgs Factory: Linear vs. Circular (HF2012) (14-November 16, 2012)

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# ... and two efforts are formalized and develop into studies towards Conceptual Design Reports



~16 T  $\Rightarrow$  100 TeV *pp* in 100 km ~20 T  $\Rightarrow$  100 TeV *pp* in 80 km

- e<sup>+</sup>e<sup>-</sup> collider (FCC-ee) as potential intermediate step
- p-e (FCC-he) option
- 80-100 km infrastructure
   in Geneva area













23-29 March 2015 Marriott Georgetown Hotel US/Eastern timezone

See you in Rome next year! Note down April 11-15, 2016



The Standard Model (SM) of particle physics can describe the strong, weak and electromagnetic interactions under the framework of quantum gauge field theory. The theoretical predictions of SM are in excellent agreement with the past experimental measurements. Especially the 2013 Nobel Prize in physics was awarded to F. Englert and P. Higgs "for the theoretical discovery of a mechanism that contributes to our understanding of the origin of mass of subatomic particles, and which recently was confirmed through the discovery of the predicted fundamental particle, by the ATLAS and CMS experiments at CERN's Large Hadron Collider".

but fundamental questions remain unanswered, from an understanding of the origin of the electroweak s composition of the dark matter of the universe.

An extended high energy experimental program beyond the planned running of the LHC will be crucial to fully address these questions. The Center for Future High Energy Physics is dedicated to carrying out detailed studies on both the physics case and the design of possible future colliders. The immediate focus will be on circular colliders: an electron-positron collider as Z and Higgs factory, and a high-energy proton-proton collider.

| 512015 | China    |        |  |
|--------|----------|--------|--|
|        | Previous | worksh |  |
|        | Working  | Group  |  |



#### 1st CFHEP Symposium on circular collider physics

23-25 February 2014 IHEP Asia/Shanghai timezone Physics workshops spontaneously organized all over the world document better than anything else the physics results, and the interest of the community ....







from Monday, 26 January 2015 at 17:00 to Sunday, 1 February 2015 at 12:00 (America/Denver)

## **SLAC**

Workshop on Physics at a 100 TeV Collider April 23-25, 2014, SLAC



Organizing Committee Timothy Cohen (SLAC) Mike Hance (LBNL) Jay Wacker (SLAC) Michael Peskin (SLAC) Nima Arkani-Hamed (IAS)

www.slac.stanford.edu/th/100TeV.html

Hong Kong

## Key goals of a future circular collider complex

- Thorough measurements of the Higgs boson and its dynamics
- Significant extension, via direct and indirect probes, of the search for physics phenomena beyond the SM

Fulfilling these goals will also require dedicated attention to crucial ingredients, such as

- the progress of theoretical calculations for precision physics
- the experimental data needed to improve the knowledge of fundamental inputs such as SM parameters, PDFs and to assess/ reduce theoretical systematics
  - relevance of running  $e^+e^-$  at Z pole and tt threshold
  - relevance of ep programme
- Maximal exploitation of the facility, e.g.
  - physics with heavy ion collisions
  - physics with the injector complex

## Higgs couplings programme

- Precise measurement of main Higgs couplings:
  - W,Z bosons, 3rd generation fermions (⇒probe existence of BSM effective couplings, e.g. due to non-elementary nature of H, determine CP properties, etc.)
- Couplings to 2nd and 1st generation (⇒universality of Higgs mass-generation mechanism)
- Higgs selfcouplings (⇒probe Higgs potential, to test possible underlying structure of Higgs, deviations from "mexican hat", etc)
- Couplings to non-SM objects (e.g. invisible decays)
- non-SM couplings (e.g. forbidden decays)

## **Projections**



## Projections

| <b>g</b> hxy      | FCC-ee                             |
|-------------------|------------------------------------|
| ZZ                | 0.16%                              |
| WW                | 0.85%                              |
| ΥΥ                | I.7%                               |
| Zγ                |                                    |
| tt                |                                    |
| bb                | 0.88%                              |
| τт                | 0.94%                              |
| СС                | I.0%                               |
| SS                | H→Vγ, in progr.                    |
| μμ                | 6.4%                               |
| uu,dd             | H→Vγ, in progr.                    |
| ee                | $e^+e^- \rightarrow H$ , in progr. |
| HH                |                                    |
| BR <sub>exo</sub> | 0.48%                              |

model indep. fit of 240 GeV data

## **Projections**

N / 10ab<sup>-1</sup> σ

| 740 pb | 7.4 G  |
|--------|--|
| 82 pb  | 0.8 G  |
| I6 pb  | 160 M  |
| l I pb | 110 M  |
| 38 pb  | 380 M  |
| I.4 pb | 14 M   |
|        | 740 pb<br>82 pb<br>16 pb<br>11 pb<br>38 pb<br>1.4 pb |

 $\rightarrow$  extrapolation from HL-LHC estimates  $\rightarrow$  from ttH/ttZ

> FCC-hh ambitious but possible targets?

 $\rightarrow$  extrapolation from HL-LHC estimates

 $\rightarrow$  from HH  $\rightarrow$  bb  $\gamma\gamma$ 

 $\rightarrow$  for specific channels, like  $H \rightarrow e\mu$ , ...

| <b>g</b> hxy      | FCC-ee            | FCC-hh               |
|-------------------|-------------------|----------------------|
| ZZ                | 0.16%             |                      |
| WW                | 0.85%             |                      |
| γγ                | I.7%              |                      |
| Zγ                |                   | % ?                  |
| tt                |                   | % ?                  |
| bb                | 0.88%             |                      |
| ττ                | 0.94%             |                      |
| СС                | I.0%              |                      |
| SS                | H→Vγ, in progr.   |                      |
| μμ                | 6.4%              | 2% ?                 |
| uu,dd             | H→Vγ, in progr.   |                      |
| ee                | e⁺e⁻→H, in progr. |                      |
| HH                |                   | 5% ?                 |
| BR <sub>exo</sub> | 0.48%             | < 10 <sup>-6</sup> ? |

## @FCC-hh:

- ttH coupling:
  - 1% theoretical precision on  $y_{top}$ , from measurement of  $\sigma(ttH)/\sigma(ttZ)$  and using BR info from FCC-ee

## • H selfcoupling:

#### M.Son @ FCC week

| НН →<br>bЪγγ                    | Barr,Dolan,Englert,Lima,<br>Spannowsky<br>JHEP 1502 (2015) 016  | Contino, Azatov,<br>Panico, Son<br>arXiv:1502.00539   | He, Ren, Yao<br>(follow-up of Snowmass<br>study)  |
|---------------------------------|---|---|---|
| FCC <sub>@100TeV</sub><br>3/ab  | 30~40%  | 30%   | 15%   |
| FCC <sub>@100TeV</sub><br>30/ab | 10%   | 10%   | 5%  |
| $S/\sqrt{B}$                    | 8.4   | 15.2  | 16.5  |
| Details                         | <ul> <li>✓ <math>\lambda_{HHH}</math> modification only</li> <li>✓ <math>c \rightarrow b \&amp; j \rightarrow \gamma</math> included</li> <li>✓ Background systematics</li> <li>○ <math>b\bar{b}\gamma\gamma</math> not matched</li> <li>✓ <math>m_{\gamma\gamma} = 125 \pm 1 \text{ GeV}</math></li> </ul> | <ul> <li>✓ Full EFT approach</li> <li>No <math>c \to b \&amp; j \to \gamma</math></li> <li>✓ Marginalized</li> <li>✓ <math>b\bar{b}\gamma\gamma</math> matched</li> <li>✓ <math>m_{\gamma\gamma} = 125 \pm 5 \text{ GeV}</math></li> <li>✓ Jet /W<sub>had</sub> veto</li> </ul> | <ul> <li>✓ <math>\lambda_{HHH}</math> modification only</li> <li>✓ <math>c \rightarrow b \&amp; j \rightarrow \gamma</math> included</li> <li>○ No marginalization</li> <li>✓ <math>b\bar{b}\gamma\gamma</math> matched</li> <li>✓ <math>m_{\gamma\gamma} = 125 \pm 3 \text{ GeV}</math></li> </ul> |



# **BSM Higgs Sectors**

#### **Big Picture Motivations**

- Naturalness
  - SUSY
  - pGB
  - uncolored?
- Electroweak Phase Transition
  - Baryogenesis?
- Higgs Portal
  - Dark Matter?
  - Generic BSM

UV Completions & Rest of Theory

#### **IR Models**

D.Curtin @

FCC week

- SM+S (mixed/unmixed)
- SM+fermions
- 2HDM
- 2HDM+S
- SILH
- ....

#### **Observables at Current + Future Colliders**

- producing extra higgs states (incl. superpartners)
- Exotic Higgs Decays
- Electroweak Precision Observables
- Higgs coupling measurements
- Higgs portal direct production of new states
- Higgs self coupling measurements
- Zh cross section measurements



Interplay of EW precision tests (Tera-Z@FCC-ee), Higgs BR measurements (H@FCC-ee) and direct resonance searches (10-30 TeV, @ FCC-hh)

## Minimal stealthy model for a strong EWPT

 $V_0 = -\mu^2 |H|^2 + \lambda |H|^4 + \frac{1}{2}\mu_S^2 S^2 + \lambda_{HS} |H|^2 S^2 + \frac{1}{4}\lambda_S S^4$ FCC week

D.Curtin @

Unmixed SM+S. No exotic higgs decays, no higgs-singlet mixing, no EWPO, ....



 $\Rightarrow$  Appearance of first "**no-lose**" arguments for classes of compelling scenarios of new physics

## Scenarios for new physics

- Guidelines for the future
  - Search for all that's searchable!
  - Don't necessarily try to tie together under a single interpretation all TH issues and exptl puzzles ....
  - .... but still make reference to established conceptual frameworks as guiding principles to steer the exploration!

 Naturalness is one of the most compelling motivations for new physics near the weak scale.

- The LHC will eventually probe conventional "colorful" theories to (at best) the ~1% level.
- But it will leave kinematic regions in conventional theories

   and all regions of more novel theories essentially
   untested, and the status of naturalness truly unresolved.
- A Higgs factory & 100 TeV collider can uniformly probe natural symmetry-based theories to the ~1% level with powerful complementarity.

N.Craig @ FCC week

#### N.Craig



#### **Direct production of Dark Matter**

- The search for WIMP dark matter is largely out of the reach for the LHC. L.Wang @ FCC week LHC 14: reach to about a couple hundred GeV. - 100 TeV pp Collider significantly enhance the reach, a fact of 5–7 enhancement. - More detailed studies necessary. New ideas needed: more channels, detector design...
  - At the same time, it is clear that this should be one of the main motivations for going to a 100 TeV pp collider.

#### **Towards no-lose arguments for Dark Matter scenarios:**





## recall: A possible TLEP running programme

1. ZH threshold scan and 240 GeV running (200 GeV to 250 GeV) 5+ years @2 10^35 /cm2/s => 210^6 ZH events ++ returns at Z peak with TLEP-H configuration

for detector and beam energy calibration

**Higgs boson HZ studies** + WW, ZZ etc..

2. Top threshold scan and (350) GeV running 5+ years @5 10^34 /cm2/s → 10^6 ttbar pairs ++Zpeak

Top quark mass Hvv Higgs boson studies

- 3. Z peak scan and peak running , TLEP-Z configuration -> 10^12 Z decays → transverse polarization of 'single' bunches for precise E\_beam calibration 2 years  $Mz, \Gamma_z R_h$  etc... Precision tests and
- 4. WW threshold scan for W mass measurement and W pair studies 1-2 years  $\rightarrow$  10^8 W pairs ++Zpeak M<sub>w</sub>, and W properties
- 5. Polarized beams (spin rotators) at Z peak 1 year at BBTS=0.01/IP => 10<sup>11</sup> Z decays.

ALR, AFR<sup>pol</sup> etc

rare decays

etc...

6. more and upgrades....

# From the global programme, I–2 orders of magnitude more precise measurements of EW parameters

| x                              | Physics  | Present precision            |                                    | TLEP stat<br>Syst Precision       | TLEP key                   | Challenge                                |
|--------------------------------|--|------------------------------|------------------------------------|-----------------------------------|----------------------------|--|
| M <sub>Z</sub><br>MeV/c2       | Input  | 91187.5<br>±2.1              | Z Line shape<br>scan               | 0.005 MeV<br><±0.1 MeV            | E_cal                      | QED<br>corrections                       |
| $\Gamma_{z}$<br>MeV/c2         | Δρ (T)<br>(no Δα!)                                   | 2495.2<br>±2.3               | Z Line shape<br>scan               | 0.008 MeV<br><±0.1 MeV            | E_cal                      | QED<br>corrections                       |
| R <sub>I</sub>                 | $\alpha_{s,\delta_b}$                                | 20.767<br>± 0.025            | Z Peak                             | 0.0001<br>± 0.002<br>- 0.0002     | Statistics                 | QED<br>corrections                       |
| $N_{v}$                        | Unitarity of<br>PMNS,<br>sterile v's                 | 2.984<br>±0.008              | Z Peak<br>Z+γ(105/161)             | 0.00008<br>±0.004<br>0.0004-0.001 | ->lumi meast<br>Statistics | QED<br>corrections<br>to Bhabha<br>scat. |
| R <sub>b</sub>                 | δ <sub>b</sub>                                       | 0.21629<br>±0.00066          | Z Peak                             | 0.000003<br>±0.000020 - 60        | Statistics,<br>small IP    | Hemisphere correlations                  |
| <b>A</b> <sub>LR</sub>         | Δρ, ε <sub>3 ,</sub> Δα<br>(Τ, S )                   | 0.1514<br>±0.0022            | Z peak,<br>polarized               | ±0.000015                         | 4 bunch<br>scheme          | Design<br>experiment                     |
| M <sub>W</sub><br>MeV/c2       | Δρ, ε <sub>3 ,</sub> ε <sub>2,</sub> Δα<br>(T, S, U) | 80385<br>± <mark>15</mark>   | Threshold<br>(161 GeV)             | 0.3 MeV<br><1 MeV                 | E_cal &<br>Statistics      | QED<br>corections                        |
| m <sub>top</sub> 4/1<br>MeV/c2 | 2/ก็อิut   | 173200<br>± <mark>900</mark> | Threshold <sup>el FC</sup><br>Scan | Comer Circular<br>liders          | E_cal &<br>Statistics      | Theory limit at 100 MeV?                 |

## **Other aspects**

- The FCC will redefine the scope and role of the HEP laboratory that will host it, w.r.t. scope and role of previous HEP labs.
- For CERN, the scale of the project may require not just international participation, beyond the CERN member states, but also engagement of other science communities (low-energy nuclear physics, light sources, medical sciences, applied accelerator physics, advanced technology, ...)
- While the above has not entered our radars as yet, the least we can envisage today is maintaining at the FCC a rich and diverse HEP programme, fully exploiting the injector chain (fixed target experiments) and the beam options (heavy ions). The FCC study is mandated to explore these opportunities as well, and assess their impact on the whole project.

# High-density QCD in the final state: the Quark Gluon Plasma





 Lattice QCD predicts phase transition at T<sub>c</sub>~170 MeV

#### $\rightarrow$ Quark-Gluon Plasma

Confinement is removed



 Unique opportunity to study in the laboratory spatially-extended multiparticle QCD system



CC Kickoff WS, Geneva, 14.02.14





Andrea Daines

# Quark-Gluon Plasma studies at FCC



- QGP lifetime increases
- Collective phenomena enhanced (better tests of QGP transport)
- Initial temperature higher
- Equilibration times reduced





## **Questions to be addressed in future studies include:**

Larger number of degrees of freedom in QGP at FCC energy?  $\rightarrow$  g+u+d+s**+charm**? **Higher** Changes in the quarkonium spectra? does Y(1S) Temp. melt at FCC? How do studies of collective flow profit from higher multiplicity and stronger expansion? More stringent constraints on transport properties such as shear viscosity or other properties not accessible at the LHC Higher Hard probes are sensitive to medium properties. At energy FCC, longer in-medium path length and new, rarer **probes** become accessible. How can both features be exploited?

## The 5-year international FCC design study



- Goal of this effort: Conceptual design report (CDR) and first cost estimate ready for the next Strategy Group assessment (~2018)
- Likely next step: Commission a full technical design report (TDR), ready for the following Strategy Group assessment (~2024)
- Plausible next step at 2024 Strategy Review: Review TDR and updated cost estimate, in view of LHCI4@300fb<sup>-1</sup> results and more. Recommend CERN Council to approve, abort, or postpone.

==> we have ~10 years to articulate the physics case, focusing on the physics discussion and on the study of LHC results

## **Conclusions and final remarks**

- Major progress in the last year in the definition of the physics opportunities and challenges for future circular colliders
- ee and eh assessment of physics potential very mature, clear path outlined for the required theoretical efforts (precision!!) and well-defined detector requirements
- hh a bit behind, much work to be done, but concrete efforts to develop physics-driven performance benchmarks for detector design have started
- Rapidly increasing engagement of the theory community
- From the BSM perspective, the future circular collider facility is not just a quantitative upgrade of the LHC, but allows a deeper, and in some cases conclusive, exploration of fundamental theoretical issues
- For the Higgs, the future circular collider complex will be more than a *factory*. Rather a "Higgs valley<sup>\*</sup>": multiple independent, synergetic and complementary approaches to achieve precision (couplings), sensitivity (rare and forbidden decays) and perspective (role of Higgs dynamics in broad issues like EWSB and vacuum stability, baryogenesis, naturalness, etc)

#### \* in the sense of Silicon Valley ....

## The challenge: pulling (and holding) it together

- Civil engineering and technology:
  - caverns, magnets, cryogenics
- Costs: !!
- Sociology:
  - keep up the excitement and motivation over a 50 yr time window