

Edgeworth Streaming Model for redshift space distortions

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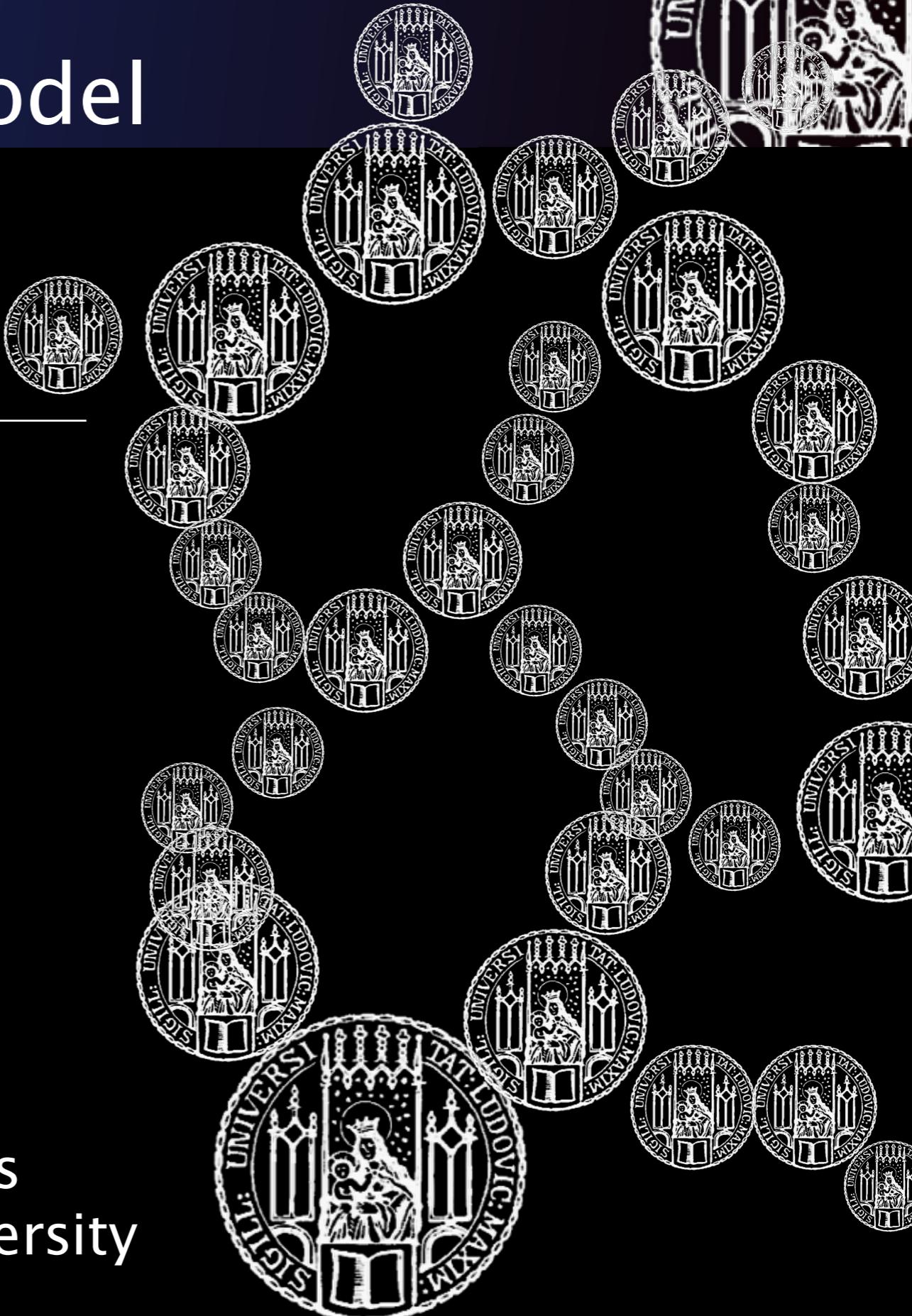
Advisor: Stefan Hofmann

in collaboration with

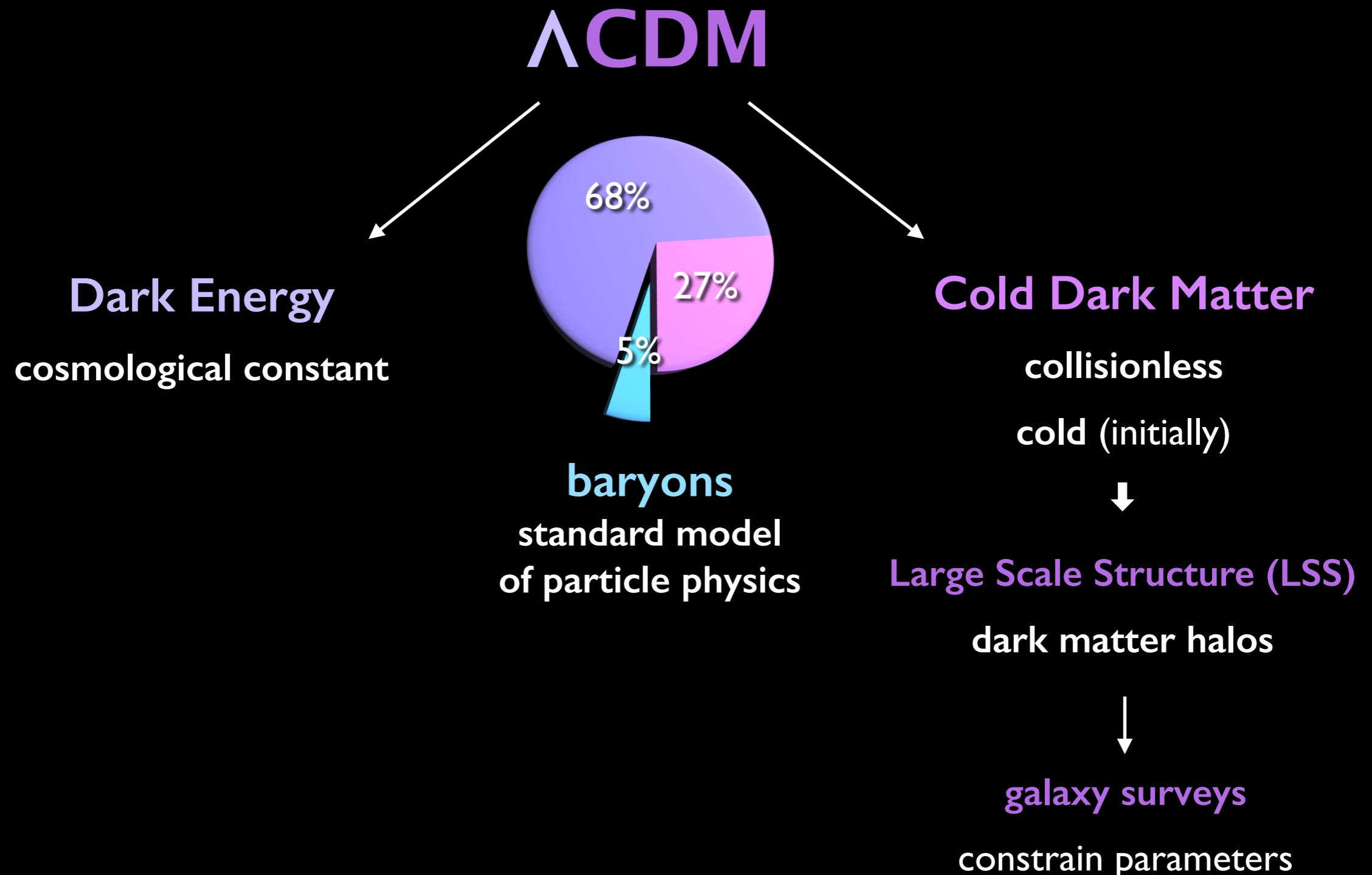
Jochen Weller and his group

co-authors

Michael Kopp, University of Cyprus
Ixandra Achitouv, Swinburne University
Thomas Haugg, LMU



Cosmological Standard Model



Redshift space correlation function

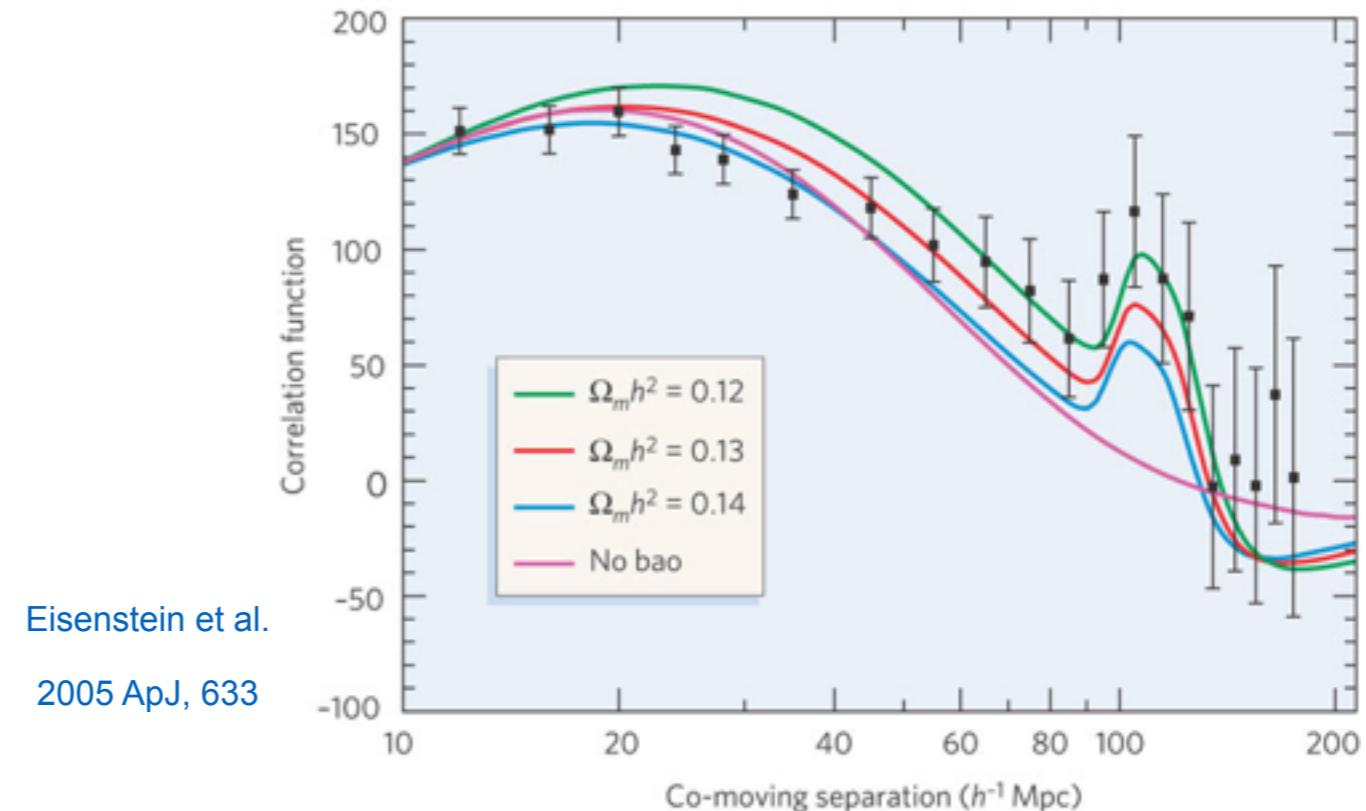


Correlation function

- measures excess probability

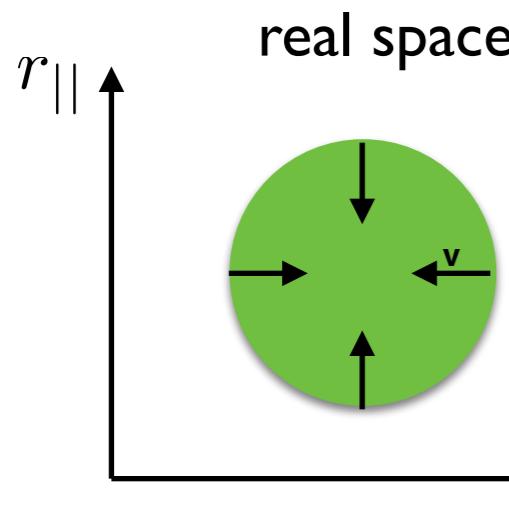
$$dP = n[1 + \xi(r)]dV$$

- biased CDM tracers: halos or galaxies
 - powerful probe for cosmology

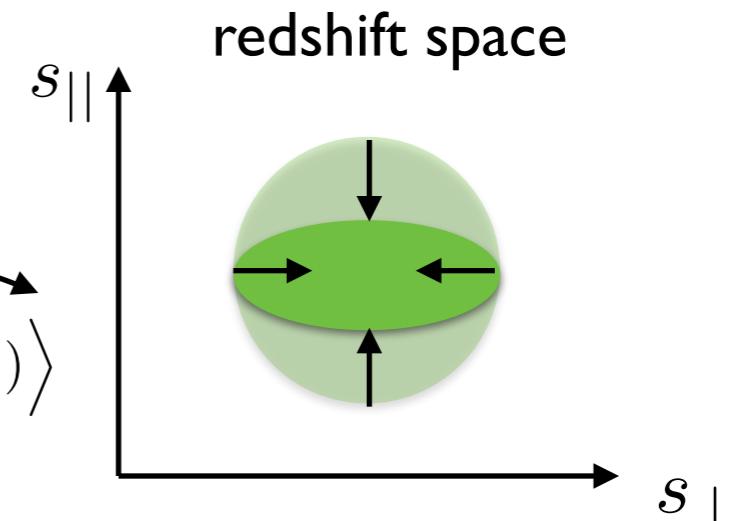


Redshift space distortions

- redshift observations affected by peculiar velocities $s_{||} = r_{||} + v/\mathcal{H}$ $s_{\perp} = r_{\perp}$



$$[1 + \delta_X(\mathbf{r})]d^3r = [1 + \delta_X(\mathbf{s})]d^3s$$
$$1 + \xi_X(s) = \left\langle (1 + \delta_X(s_1))(1 + \delta_X(s_2)) \right\rangle$$



Redshift space correlation function



Gaussian Streaming Model

$$1 + \xi_X(s_{||}, s_{\perp}, t) = \int_{-\infty}^{\infty} \frac{dr_{||}}{\sqrt{2\pi}\sigma_{12}} (1 + \xi_X(r, t)) \exp$$

real space
correlation

Gaussian velocity distribution
mean pairwise velocity & dispersion

$$\left[-\frac{(s_{||} - r_{||} - v_{12}(r, t)r_{||}/r)^2}{2\sigma_{12}^2(r, r_{||}, t)} \right]$$

Fisher (1995, *Astrophys.J.* 448,) & Reid & White (2011, *MNRAS* 417)

Redshift space correlation function



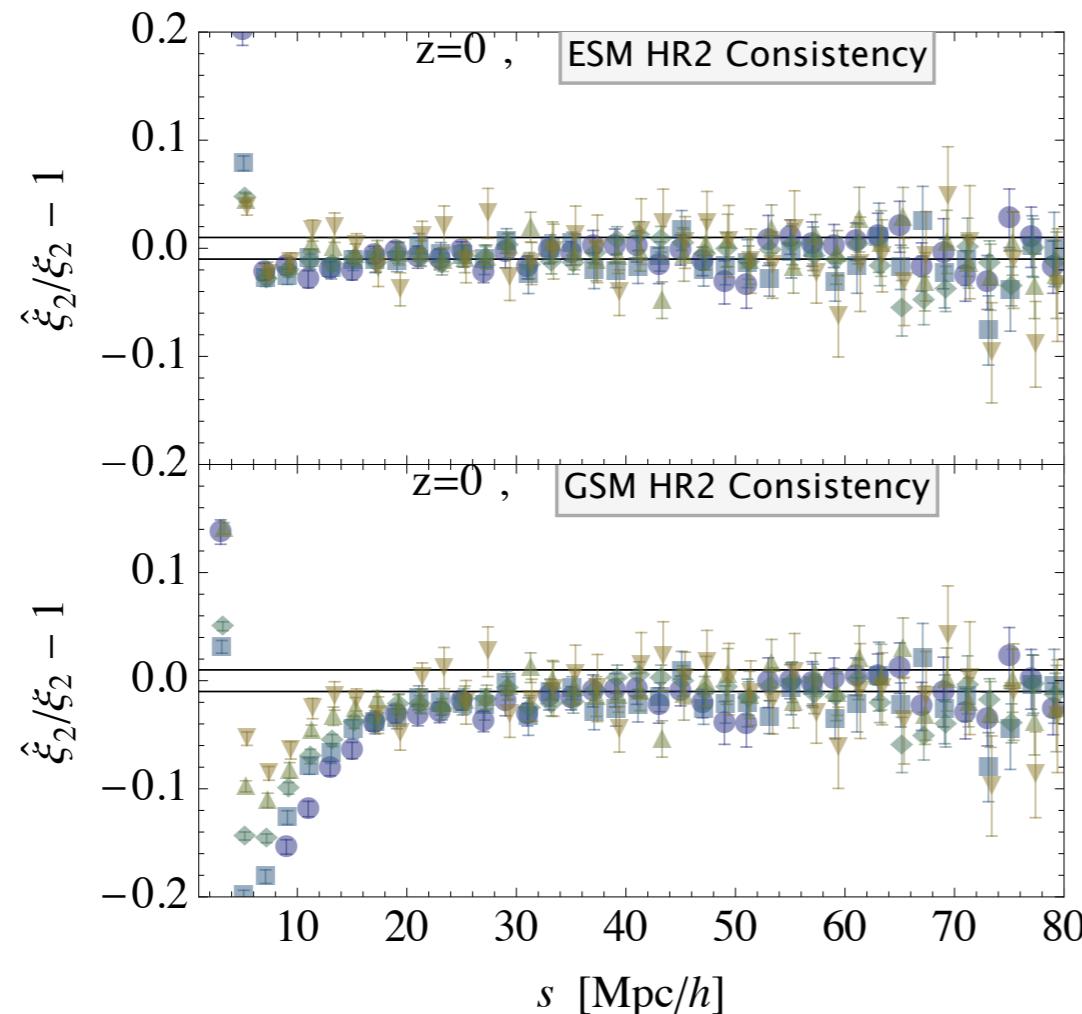
Edgeworth Streaming Model

$$1 + \xi_X(s_{||}, s_{\perp}, t) = \int_{-\infty}^{\infty} \frac{dr_{||}}{\sqrt{2\pi}\sigma_{12}} (1 + \xi_X(r, t)) \exp$$

**real space
correlation**

CU, Kopp and Haugg (2015, arXiv: 1503.08837)

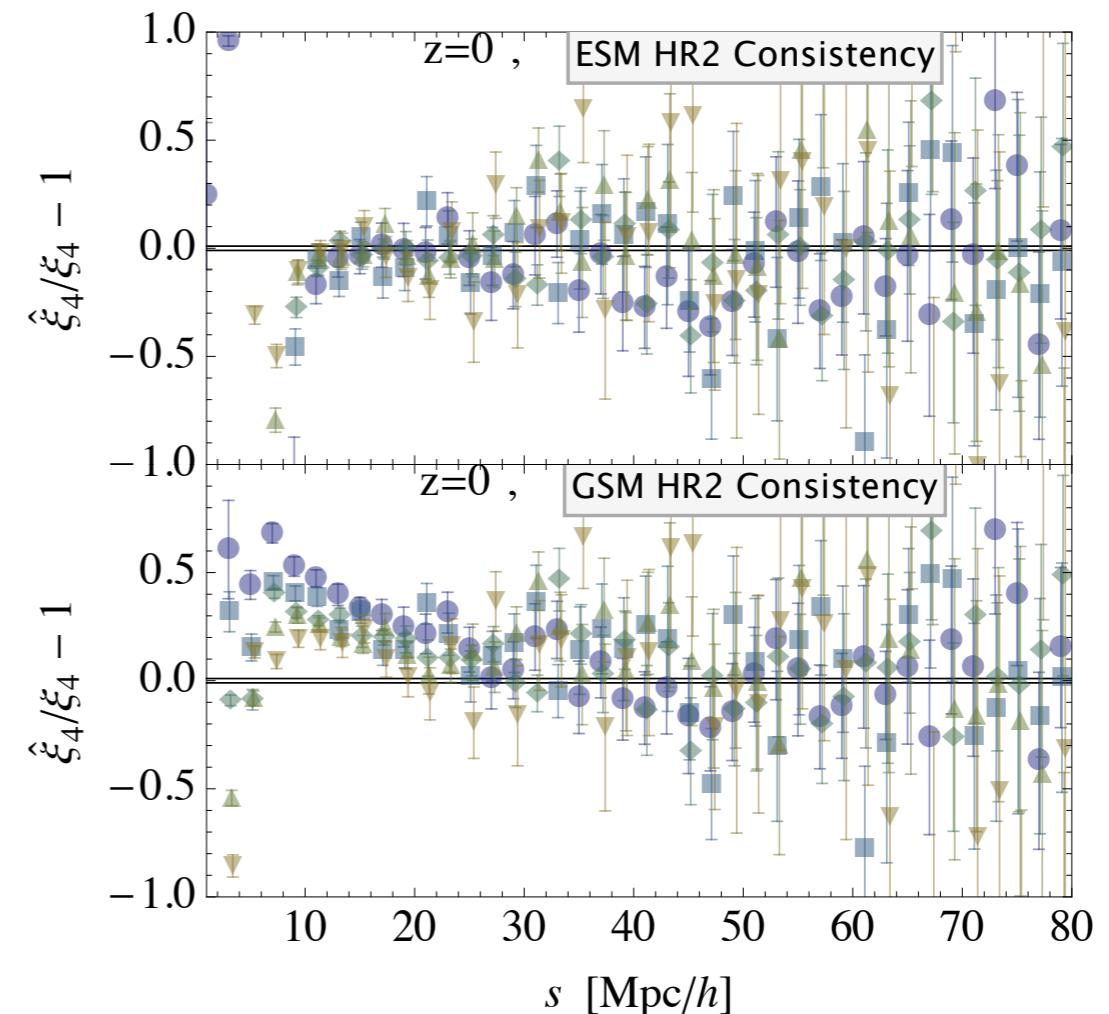
- improves accuracy on small scales
- 2% down to 10 Mpc/h (ESM) vs. 30 Mpc/h (GSM)



**Gaussian velocity distribution
mean pairwise velocity & dispersion**

$$\times \left(1 + \frac{\Lambda_{12}}{6\sigma_{12}^3} \left[\left(\frac{\Delta_{srv}}{\sigma_{12}} \right)^3 - 3 \frac{\Delta_{srv}}{\sigma_{12}} \right] \right)$$

pairwise velocity skewness



Streaming model ingredients



Gaussian Streaming Model

$$1 + \xi_X(s_{||}, s_{\perp}, t) = \int_{-\infty}^{\infty} \frac{dr_{||}}{\sqrt{2\pi}\sigma_{12}(r, r_{||}, t)} (1 + \xi_X(r, t)) \exp \left[-\frac{(s_{||} - r_{||} - v_{12}(r, t)r_{||}/r)^2}{2\sigma_{12}^2(r, r_{||}, t)} \right]$$

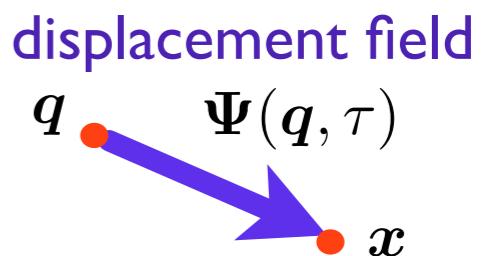
Wang, Reid & White (2014, MNRAS 437)

real space
correlation

Gaussian velocity distribution
mean pairwise velocity & dispersion

Lagrangian Perturbation Theory
+ local Lagrangian bias

- **Zel'dovich approximation** Zel'dovich (1970, A&A 5, 84)
 - 1st order Lagrangian PT
 - physically motivated resummation of SPT



- truncated Zel'dovich
 - Zel'dovich with smoothed input power spectrum
 - improves agreement with N-body

Coles, Melott, Shandarin (1993, MNRAS 260)

implementation
coarse-grained
dust model

CU & Kopp (arXiv: 1407.4810)
coarse-graining in Eulerian space
+ mapping to Lagrangian space

- Post Zel'dovich
 - higher order Lagrangian PT
 - partial resummation: Convolution LPT

Carlson et al. (2012, MNRAS 429)

Streaming model ingredients



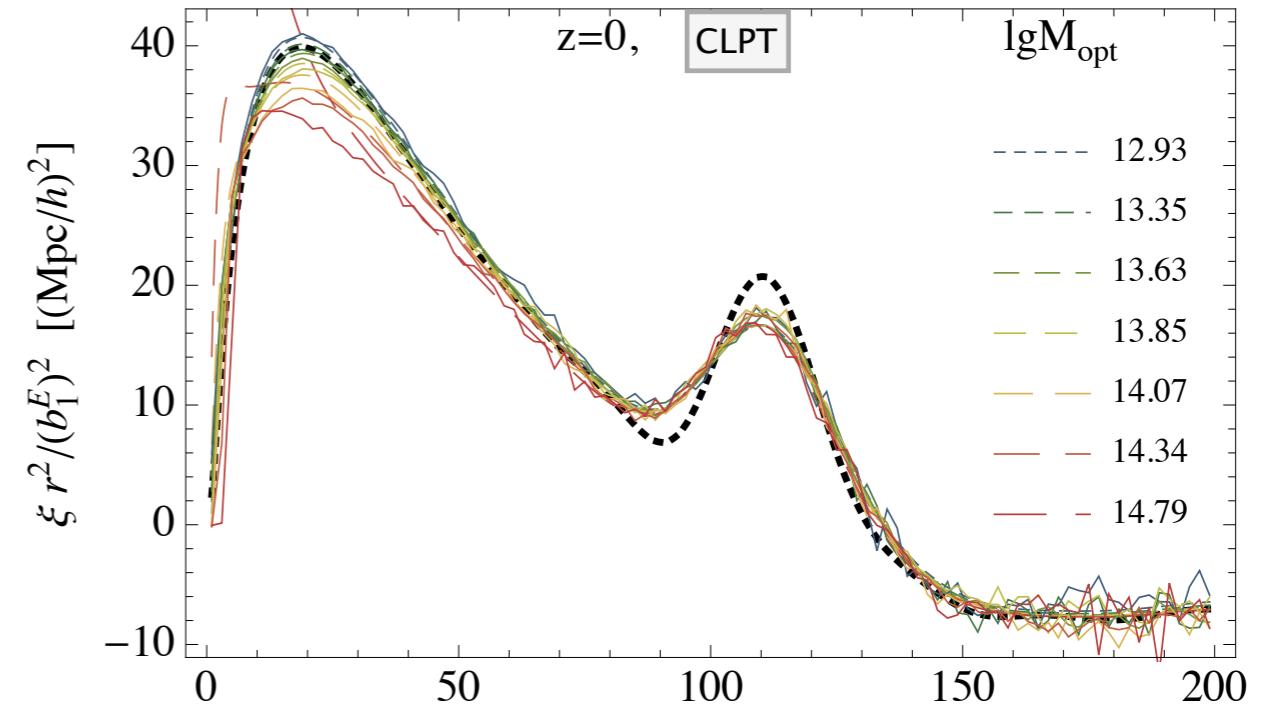
Truncated CLPT

Kopp, CU, Achitouv & Haugg (in preparation)

Real space halo correlation $\xi(r)$

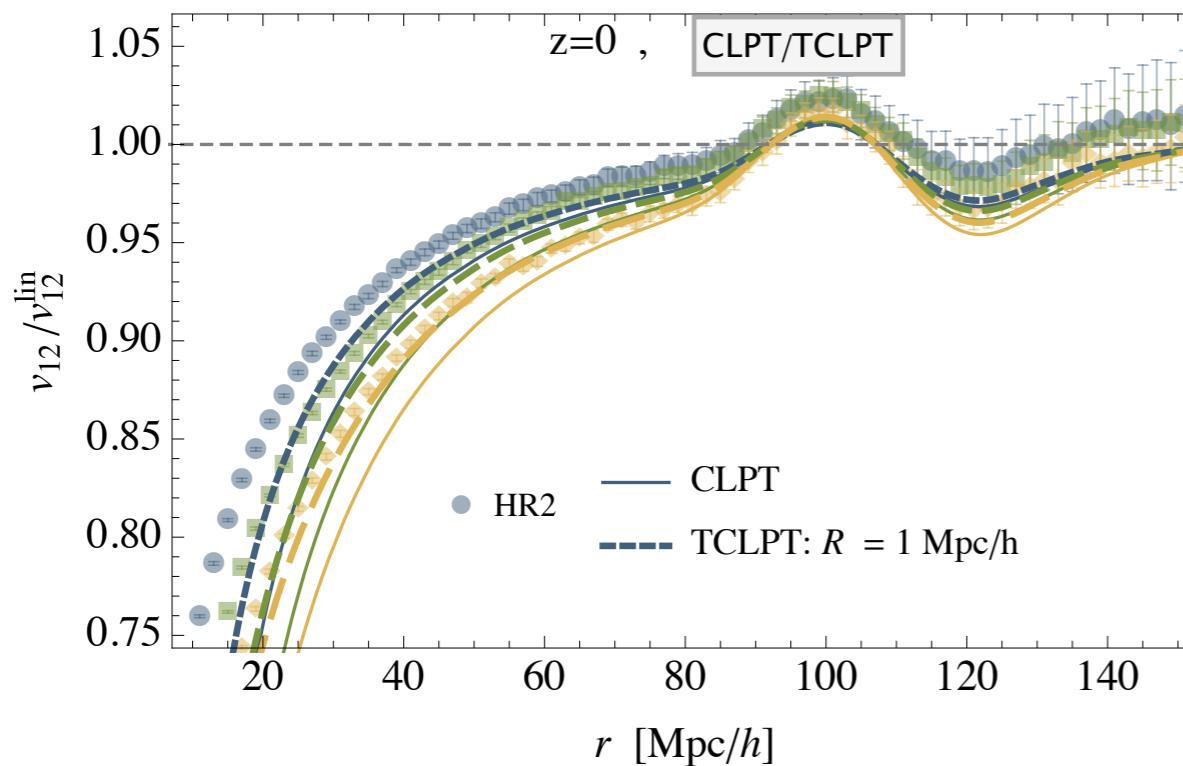
- best agreement for 1 Mpc/h
- smoothing in $R(M)$ worse
need to include peak bias

Baldauf, Desjacques & Seljak (arXiv: 1405.5885)



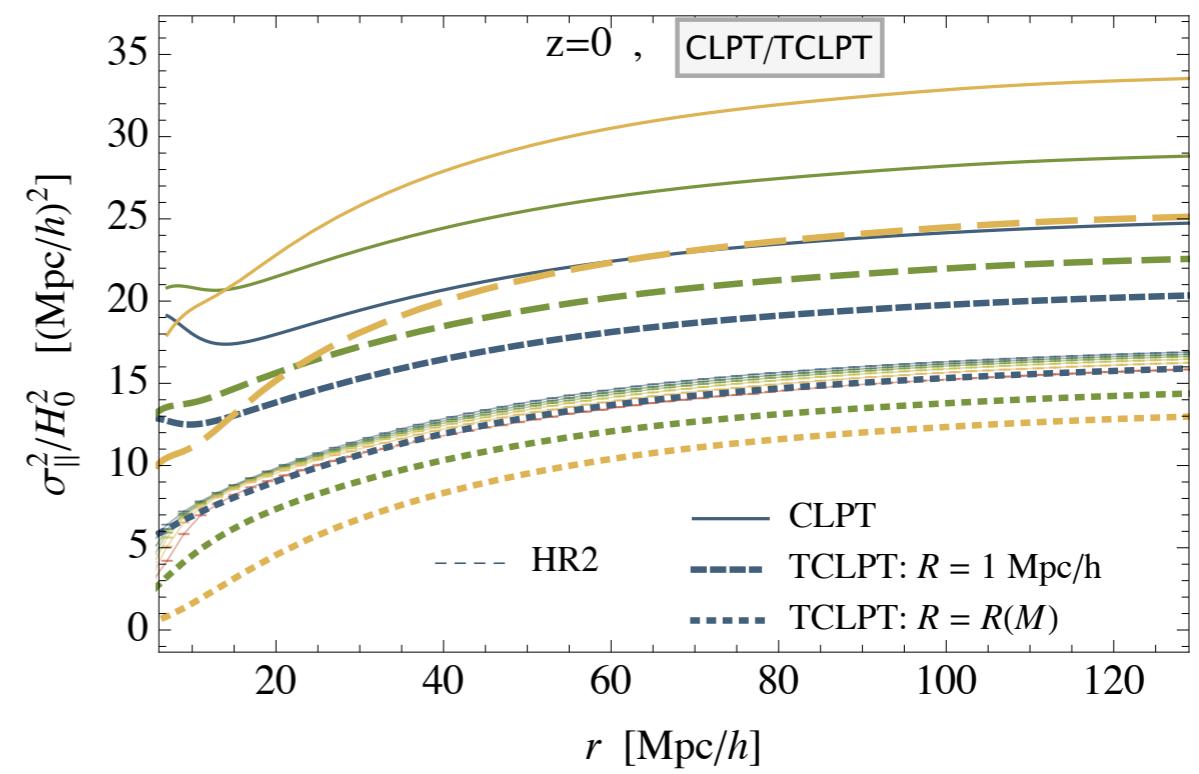
Pairwise velocity $v_{12}(r)$

- best agreement for 1 Mpc/h
similar as for DM CU & Kopp (arXiv: 1407.4810)



Pairwise velocity dispersion $\sigma_{12}(r)$

- best agreement for $R(M)$
- consider cumulant not moment



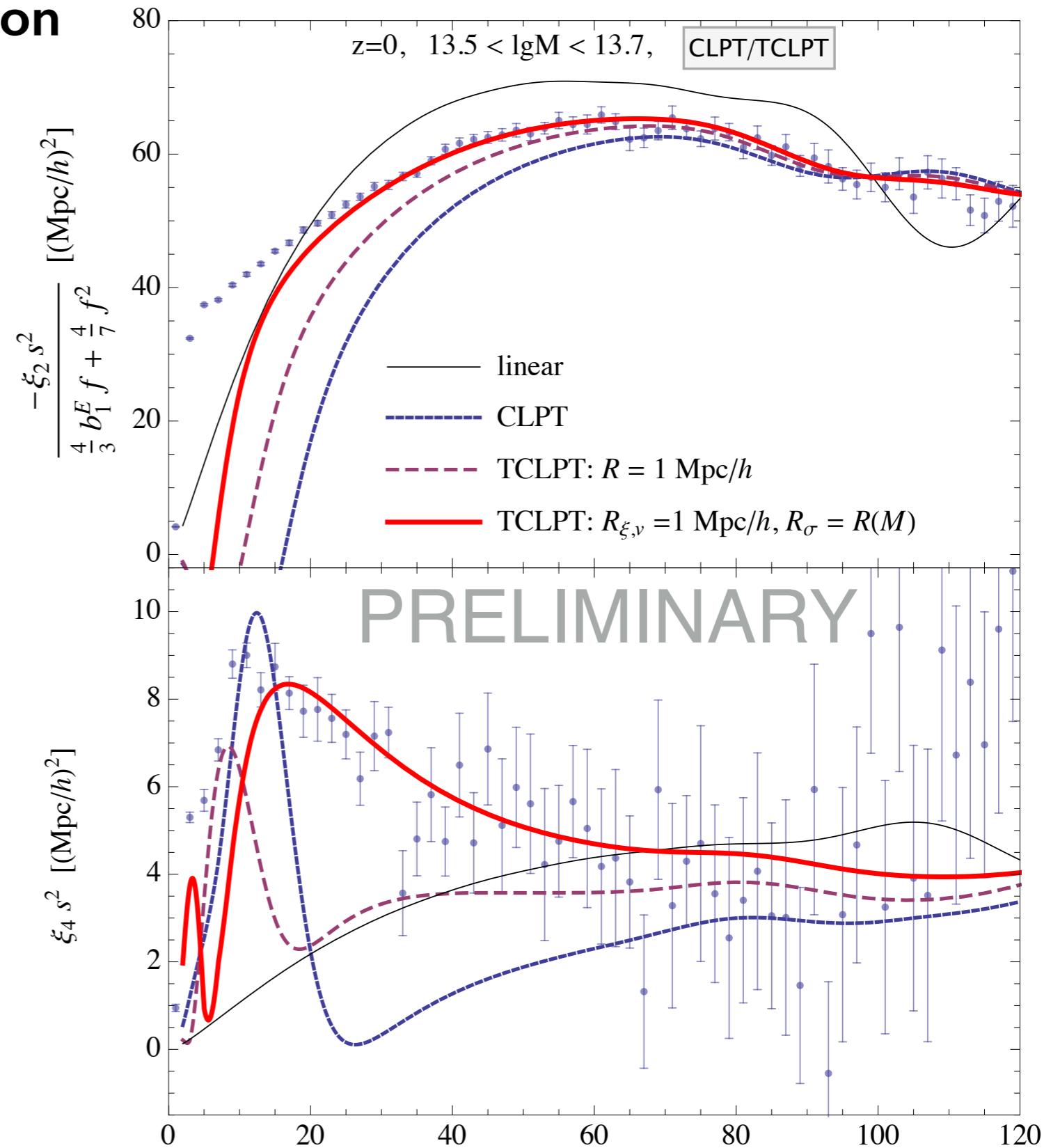
Streaming model predictions



Redshift space correlation function

Kopp, CU, Achitouv & Haugg (in preparation)

- plug streaming model ingredients
obtain redshift space multipoles
 - monopole $\xi_0(s)$
 - quadrupole $\xi_2(s)$
 - hexadecapole $\xi_4(s)$
- TCLPT outperforms CLPT
 - simultaneously improves all higher redshift-space multipoles
- best:TCLPT with hybrid smoothing
 - $\xi(r) \& v_{12}(r)$ on 1 Mpc/h
 - $\sigma_{12}(r)$ on Lagrangian scale $R(M)$



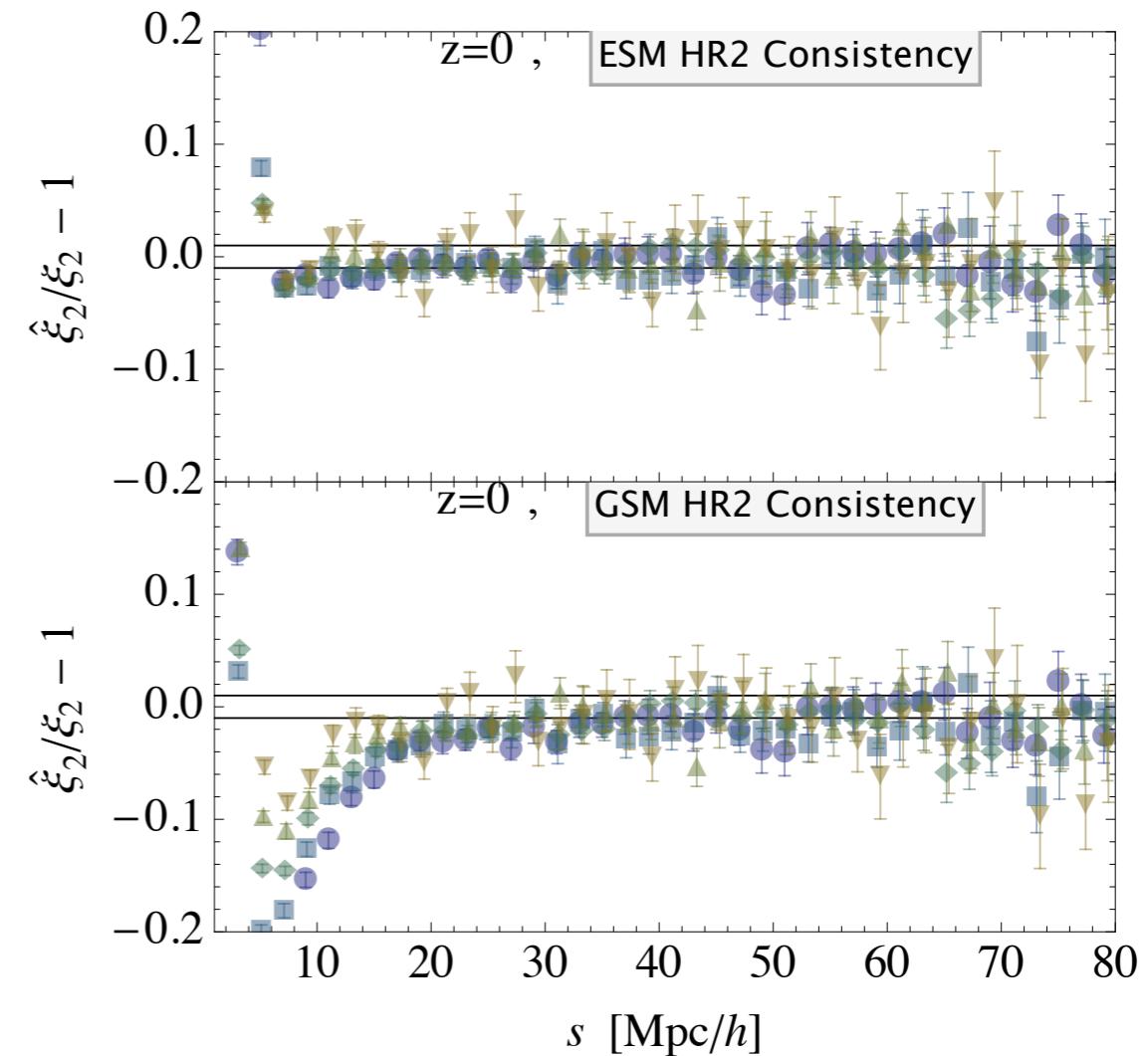
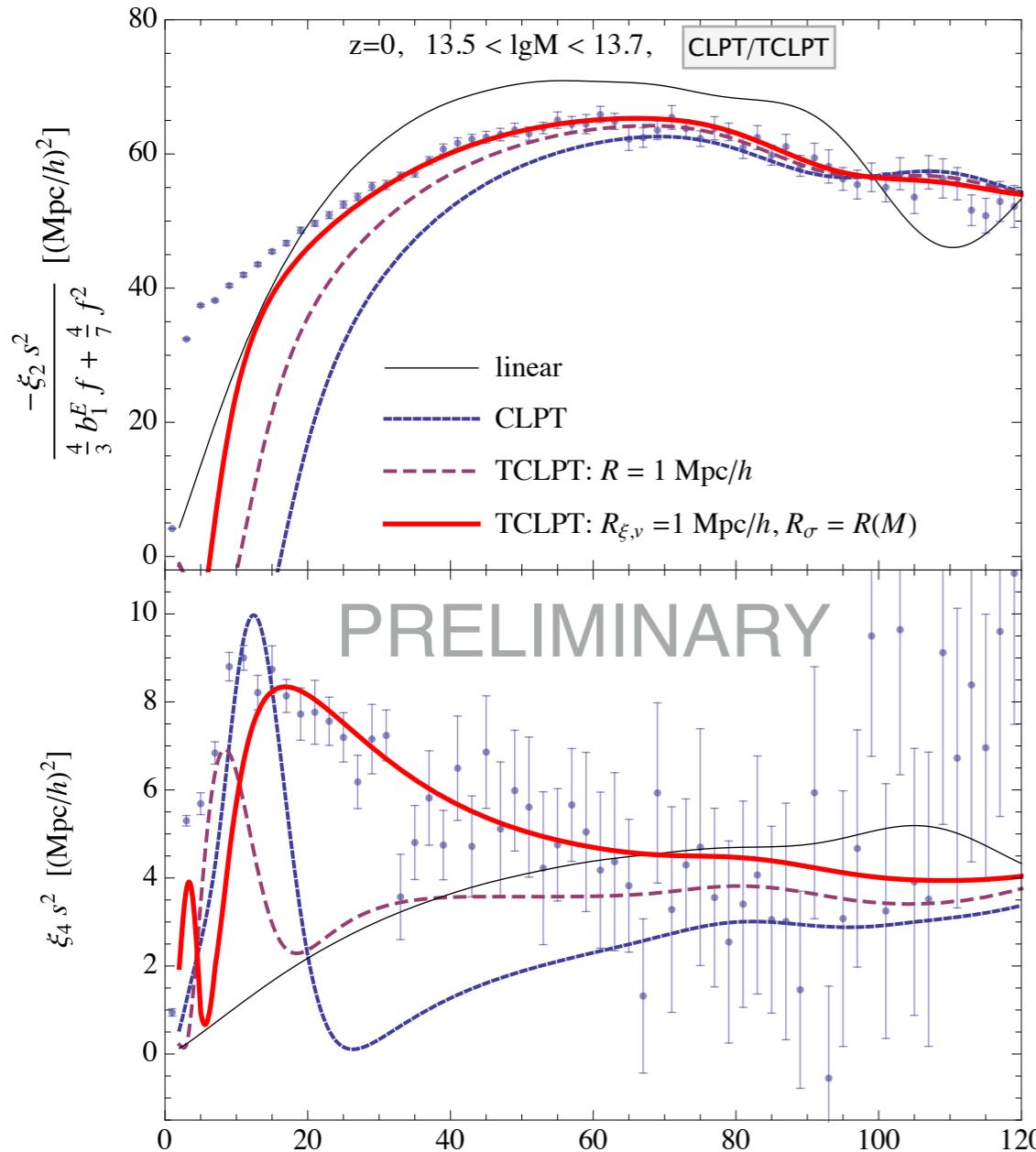


Summary

Edgeworth streaming model

- generalization of Gaussian streaming model
- pushed 2% accuracy from 30 down to 10 Mpc/h

CU, Kopp and Haugg (2015, arXiv: 1503.08837)



Redshift space halo correlation function

- coarse-graining in Eulerian vs Lagrangian space
CU, Kopp and Haugg (2015, arXiv: 1503.08837)
- truncated Post-Zel'dovich approximation (TCLPT)
 - use two smoothing scales: 1 Mpc/h & R(M)
Kopp, CU, Achitouv & Haugg (in preparation)