

Extracting Non-Gaussian Information from Large-scale Structure

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Advanced Workshop on Cosmological Structures from Reionization to Galaxies

ICTP, Trieste, Italy

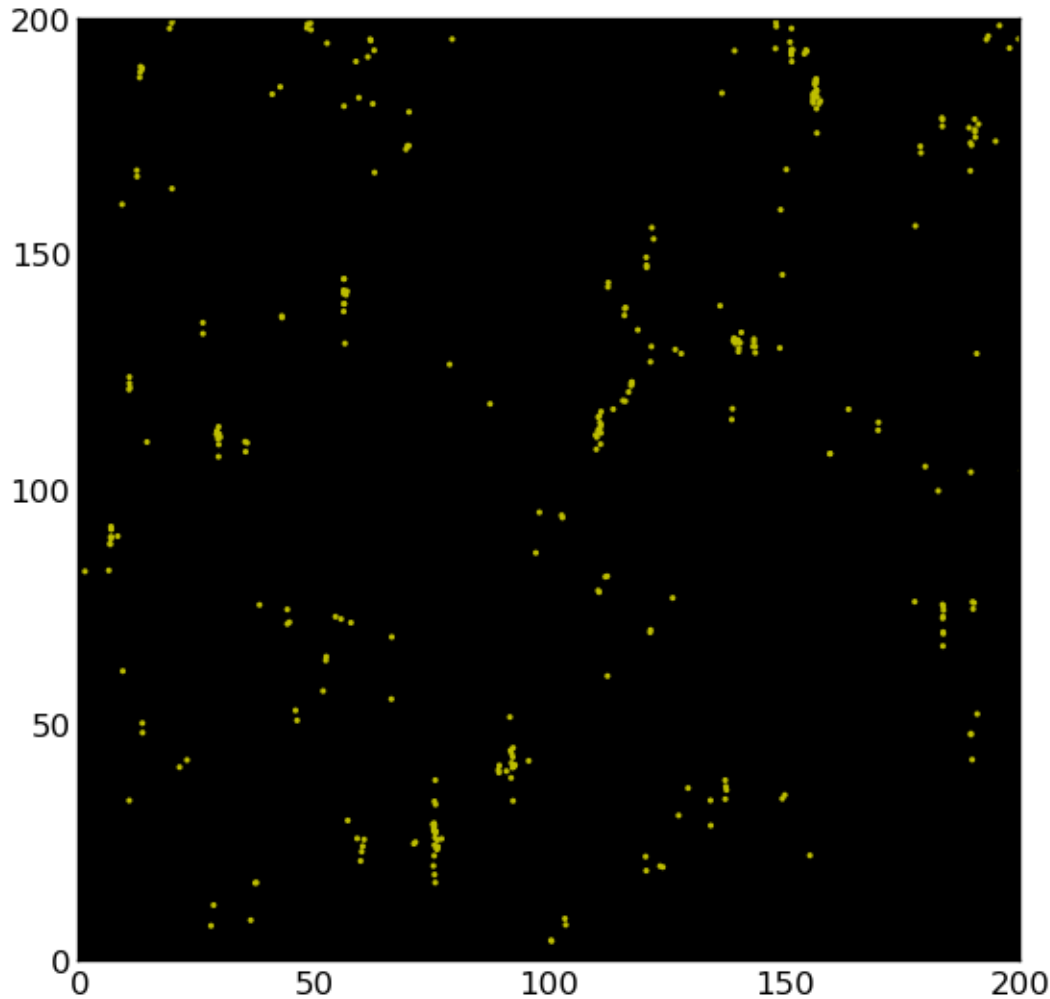
12 May, 2015

With: Mark Neyrinck, Alex Szalay, Shaun Cole, Peder Norberg

Outline

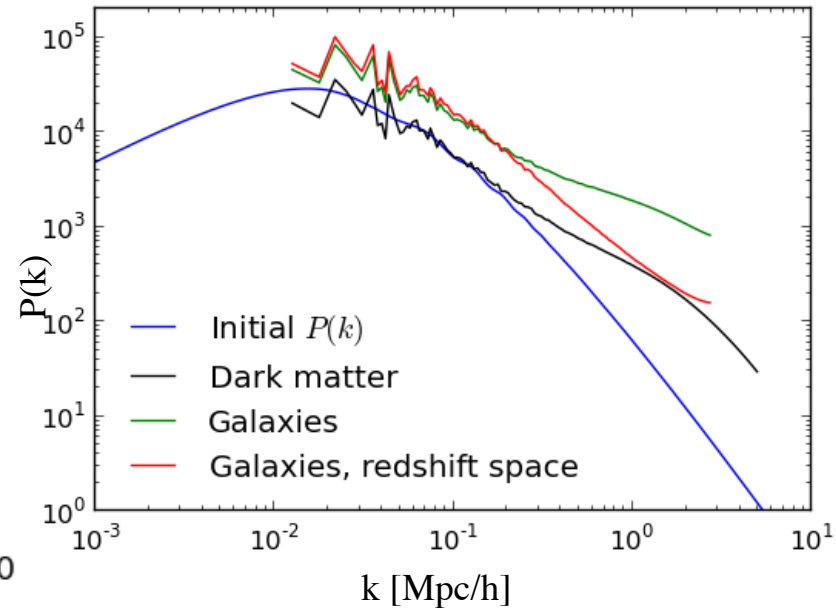
- Large-scale structure as a cosmological probe
- Beyond Gaussian statistics:
 - Log and Gaussianization transform
 - Higher-point statistics
- Summary and future work

Large-Scale Structure

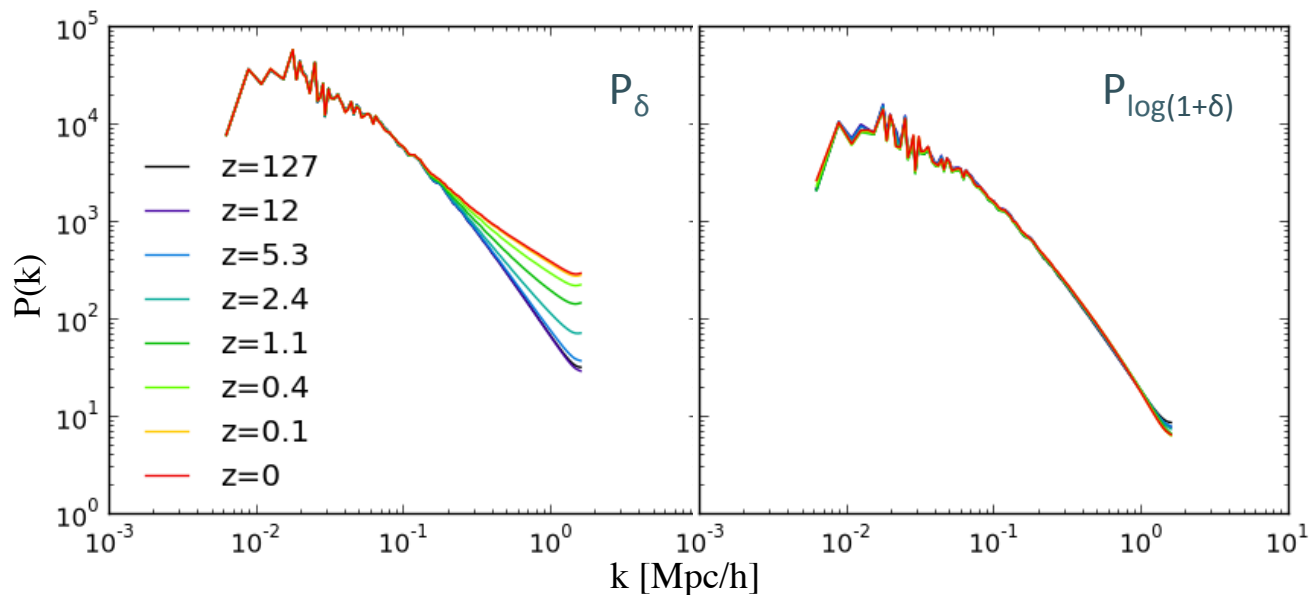
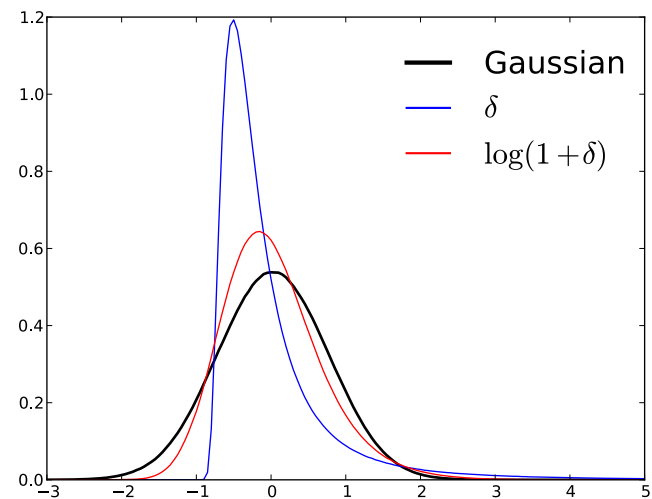
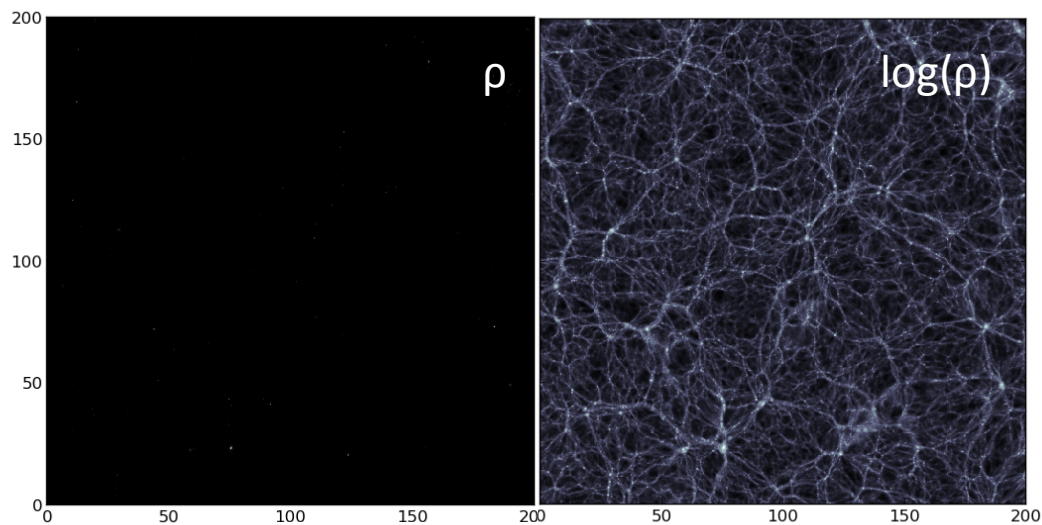


Systematics:

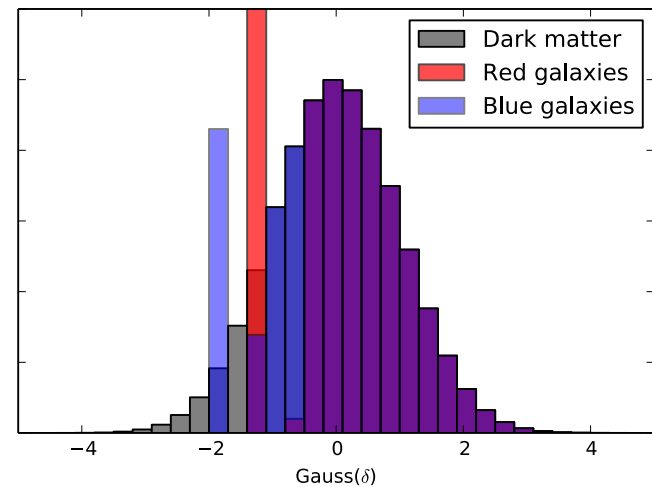
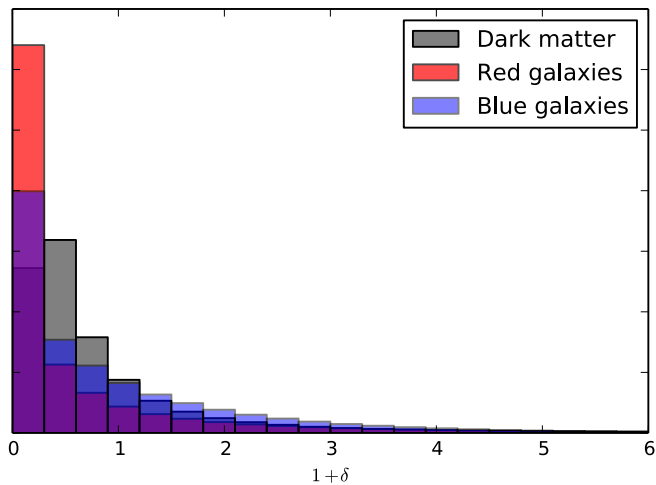
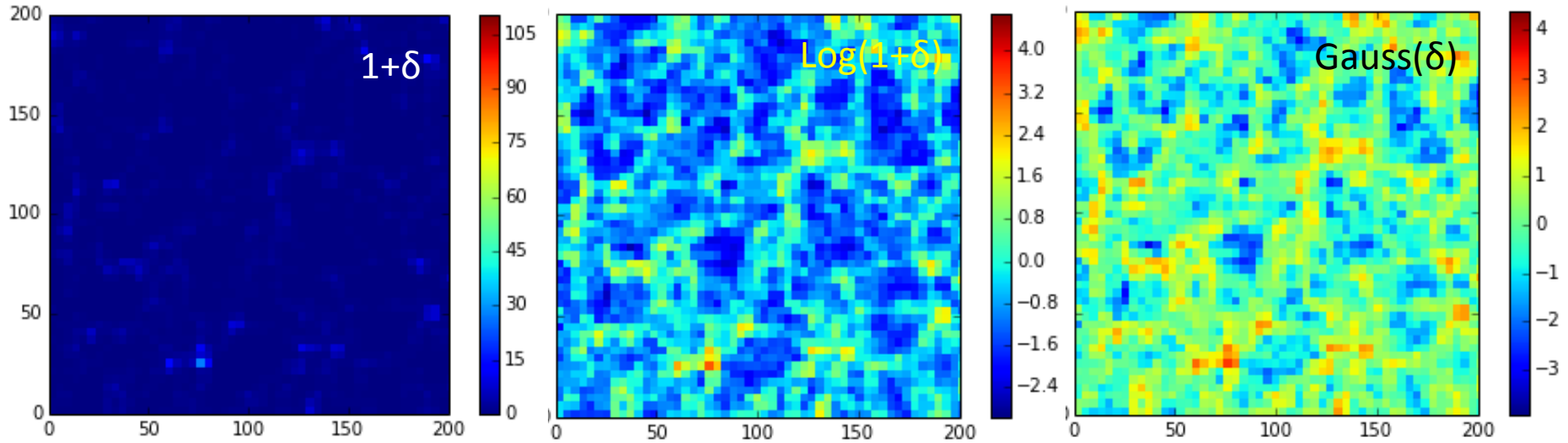
- Time evolution of matter distribution (nonlinearity)
- Galaxy bias
- Redshift-space distortions



Log transform



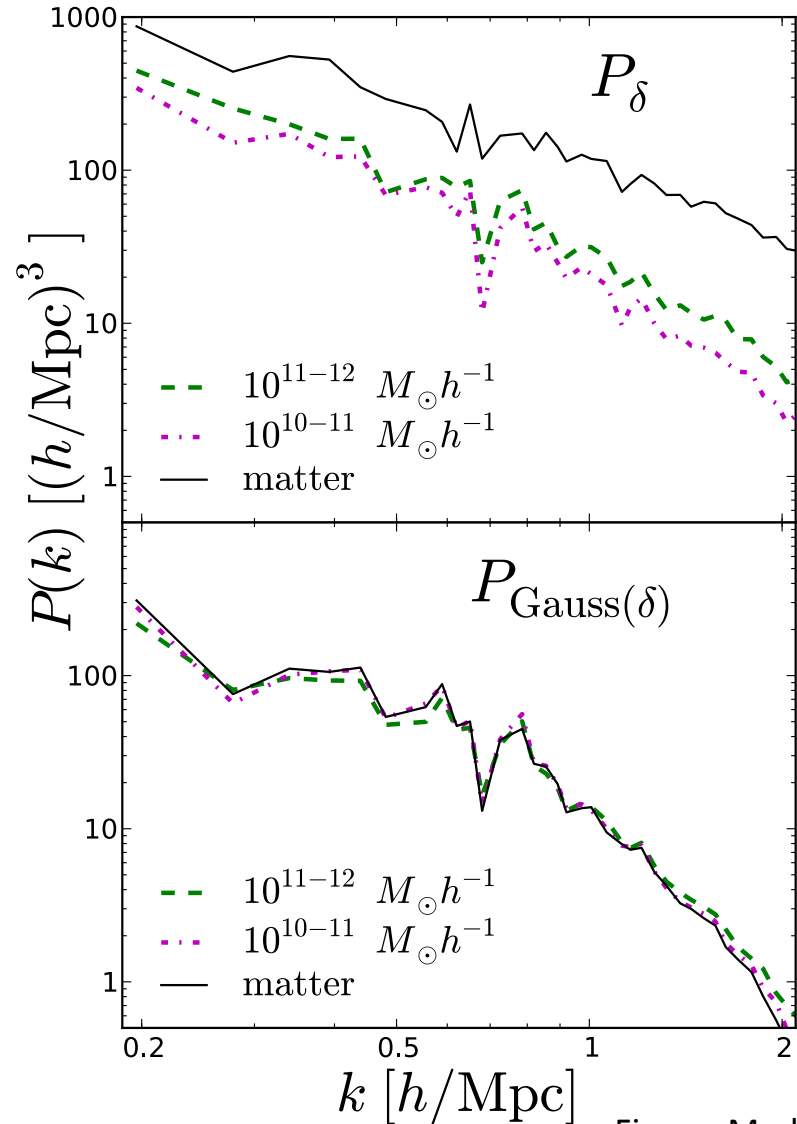
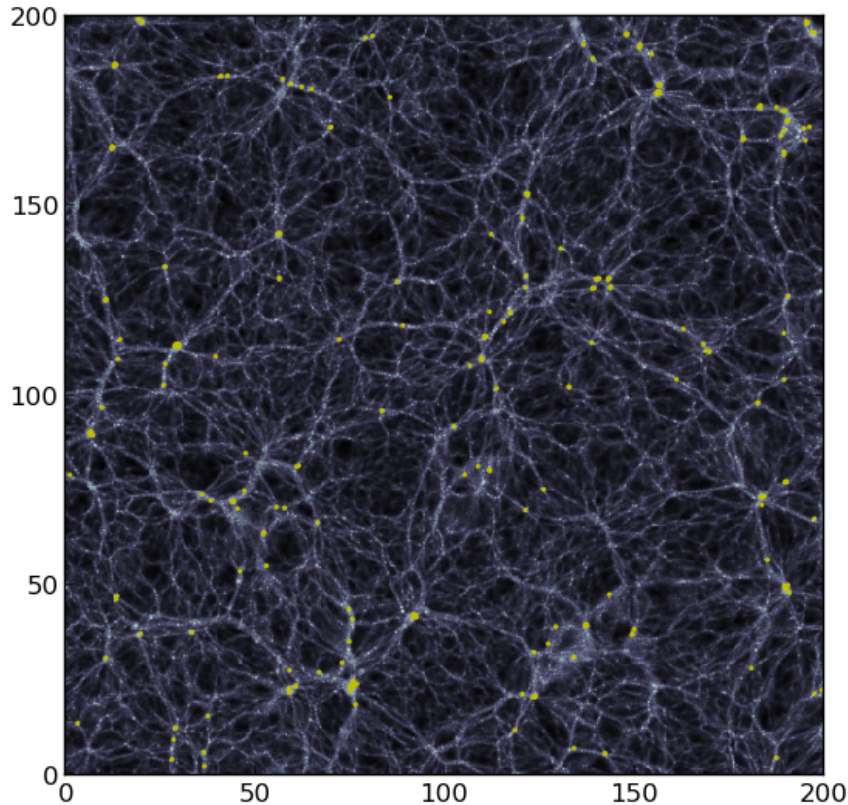
Gaussianization transform



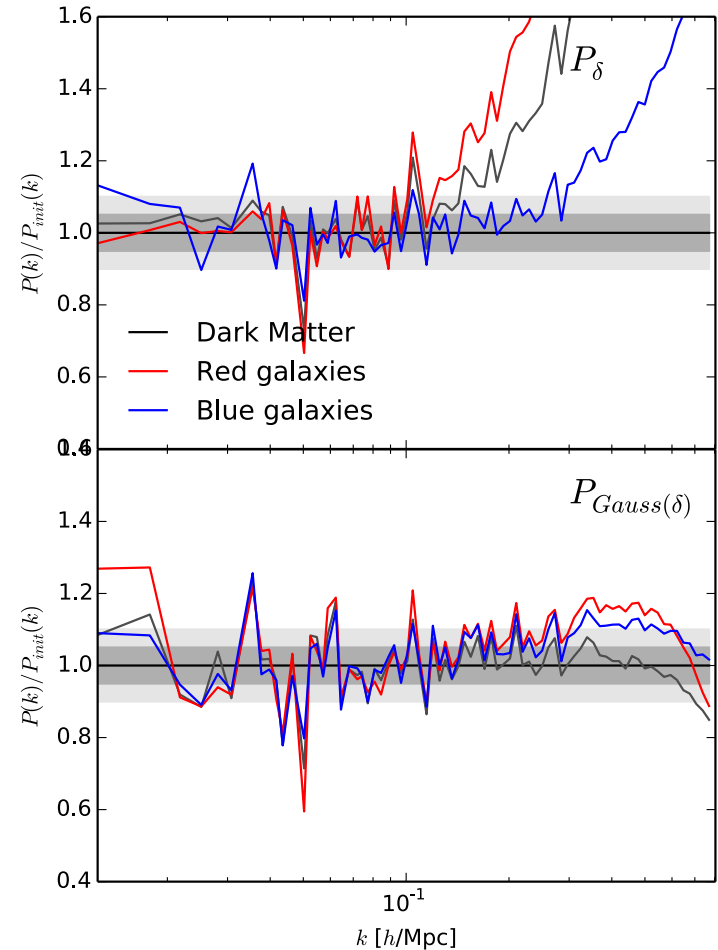
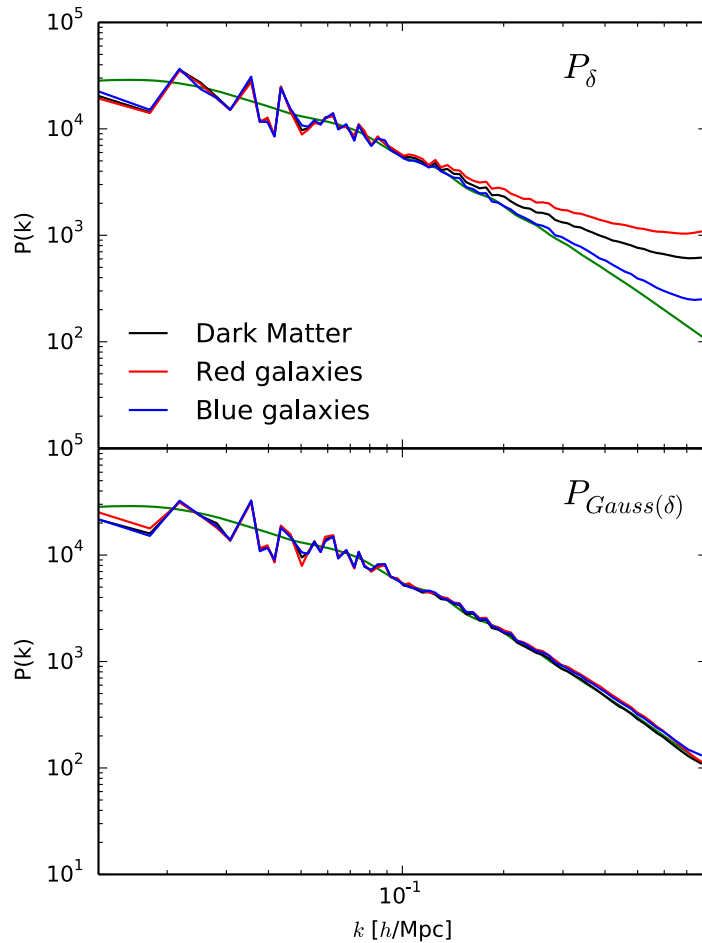
Decoupling Clustering and Tracer Bias

If $\delta_g = f(\delta_{DM})$

then $\text{Gauss}(\delta_g) = \text{Gauss}(\delta_{DM})$

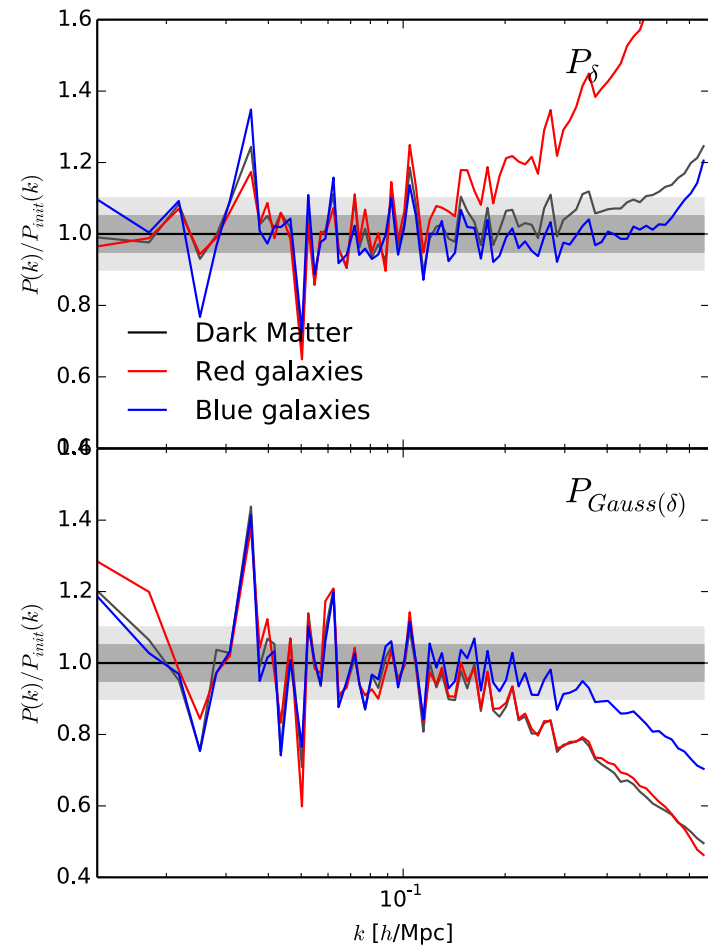
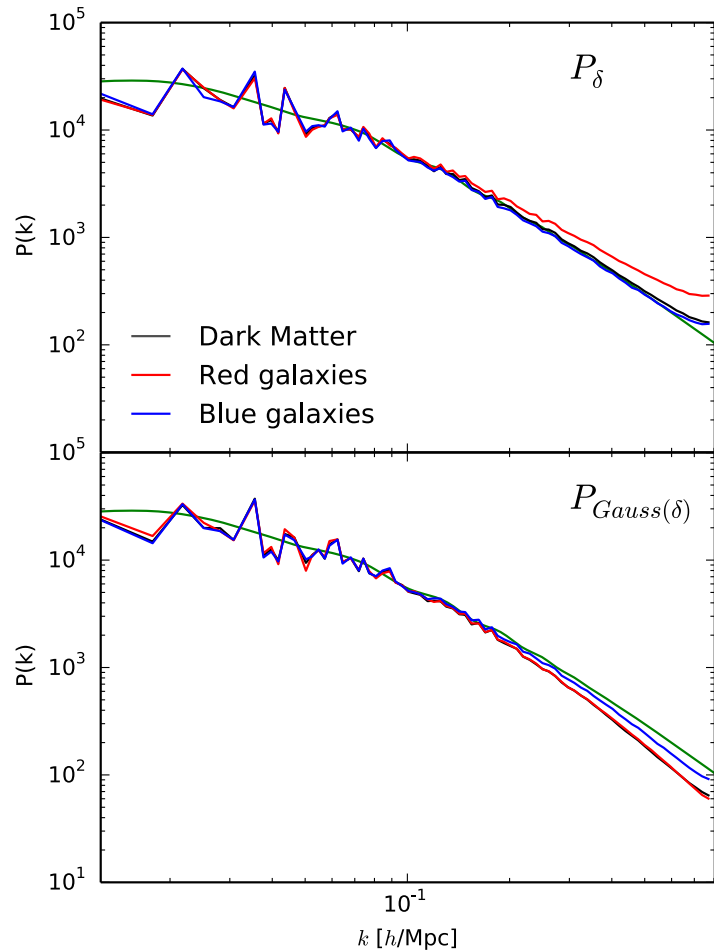


Decoupling Clustering and Tracer Bias



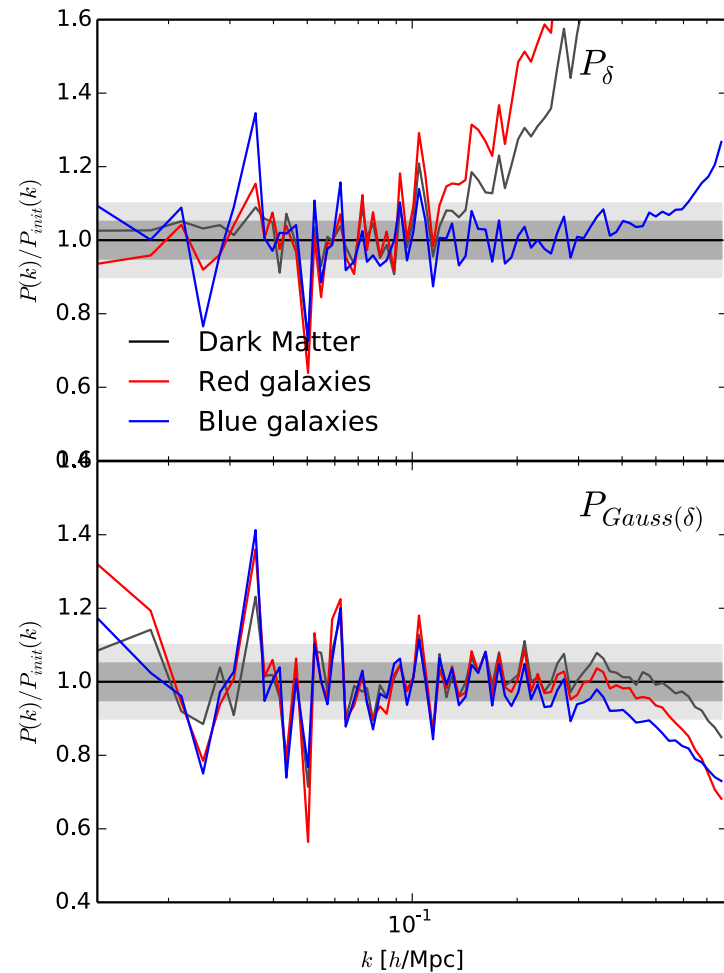
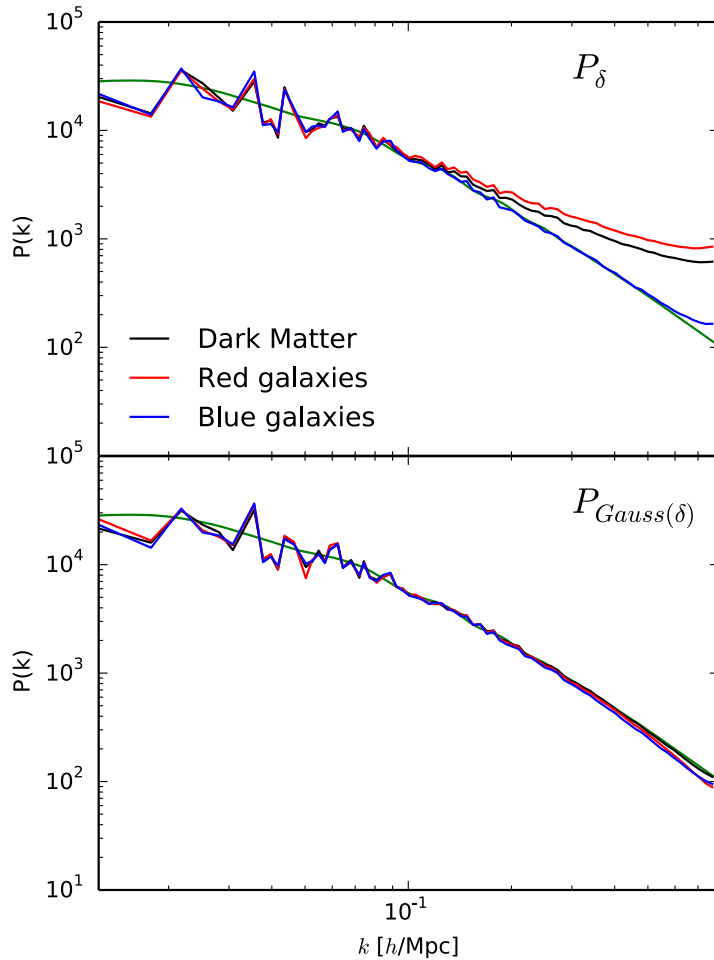
Real space

Decoupling Clustering and Tracer Bias



Redshift space

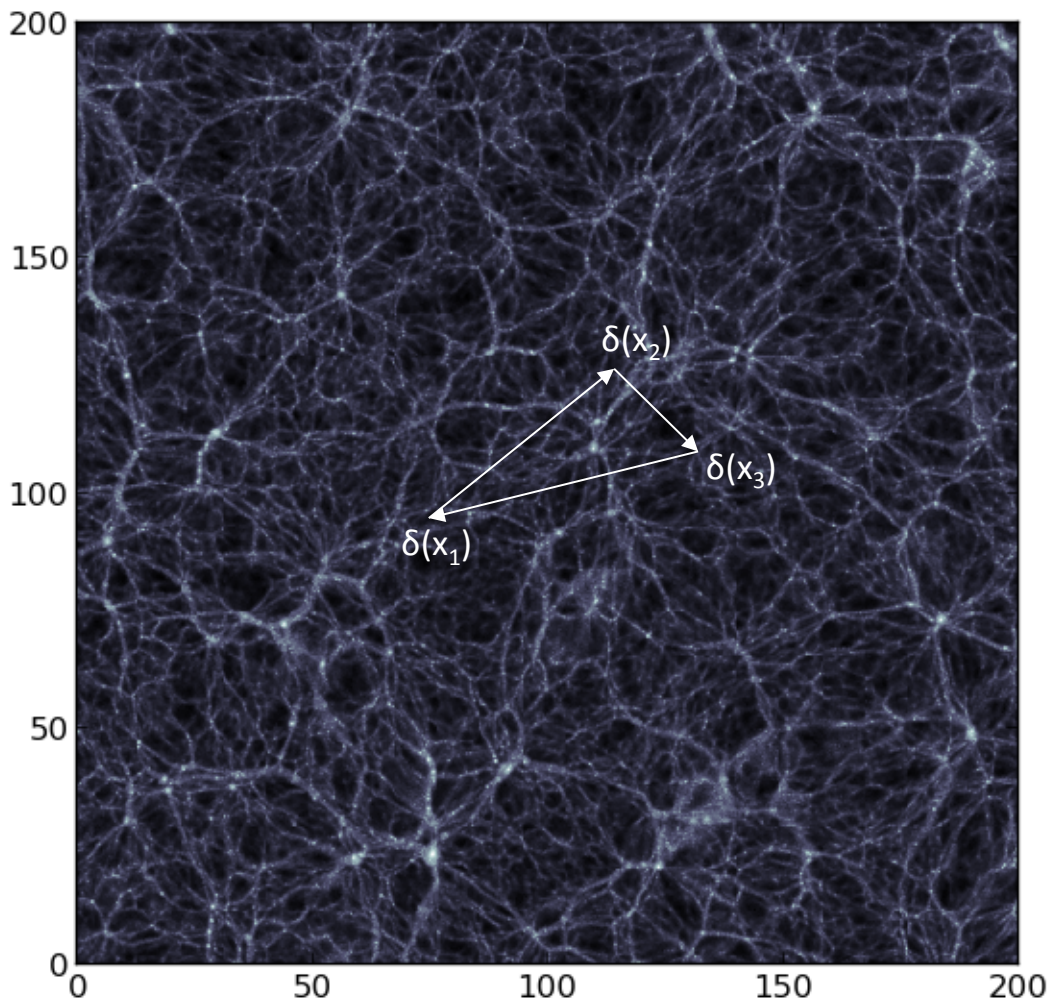
Decoupling Clustering and Tracer Bias



Higher-point statistics

$$\zeta(r_1, r_2, r_3) = \langle \delta(\mathbf{x}_1) \delta(\mathbf{x}_2) \delta(\mathbf{x}_3) \rangle$$

$$\langle \delta(\mathbf{k}_1) \delta(\mathbf{k}_2) \delta(\mathbf{k}_3) \rangle \equiv (2\pi)^3 B(k_1, k_2, k_3) \delta_D(\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3)$$



Galaxy 3-point correlation function/bispectrum contains information about:

- Galaxy bias
- Primordial non-Gaussianity/inflation
- Growth of structure/gravity

3-point Correlation Function

Lagrangian Perturbation Theory:

$$\mathbf{x}(\tau) = \mathbf{q} + \Psi(\mathbf{q}, \tau).$$

$$\Psi(\mathbf{q}, \tau) = \Psi^{(1)}(\mathbf{q}, \tau) + \Psi^{(2)}(\mathbf{q}, \tau) + \dots .$$

Real space result:

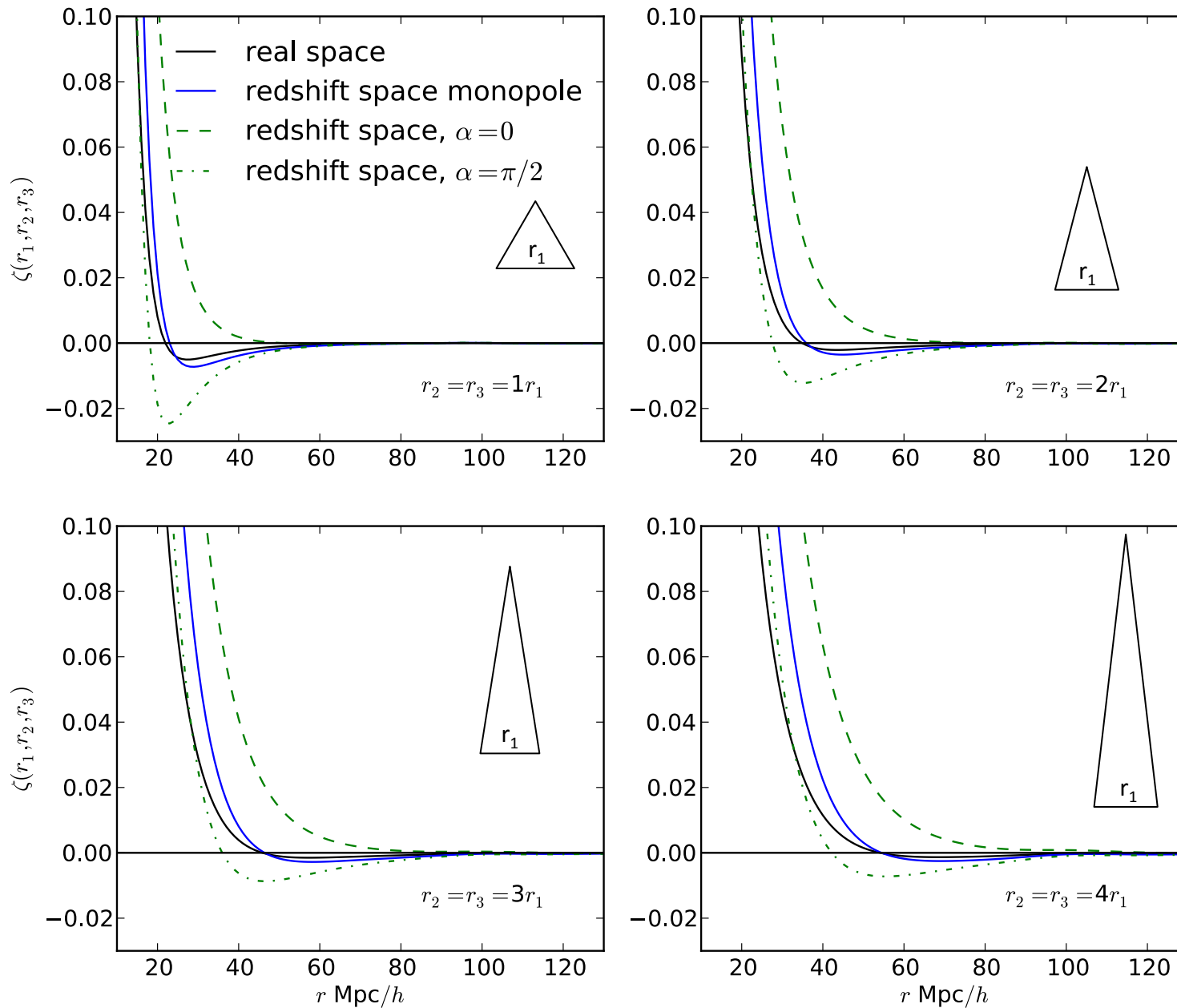
$$\zeta(r_1, r_2, r_3) = D^4 \left(\frac{34}{21} \xi_0^0(r_1) \xi_0^0(r_3) - \cos \theta_{31} (\xi_1^1(r_1) \xi_1^{-1}(r_3) + \xi_1^{-1}(r_1) \xi_1^1(r_3)) \right. \\ \left. + \frac{2}{21} (1 + 3 \cos 2\theta_{31}) \xi_2^0(r_1) \xi_2^0(r_3) + 2 \text{ cyclic} \right) .$$

RSD and (nonlocal) bias:

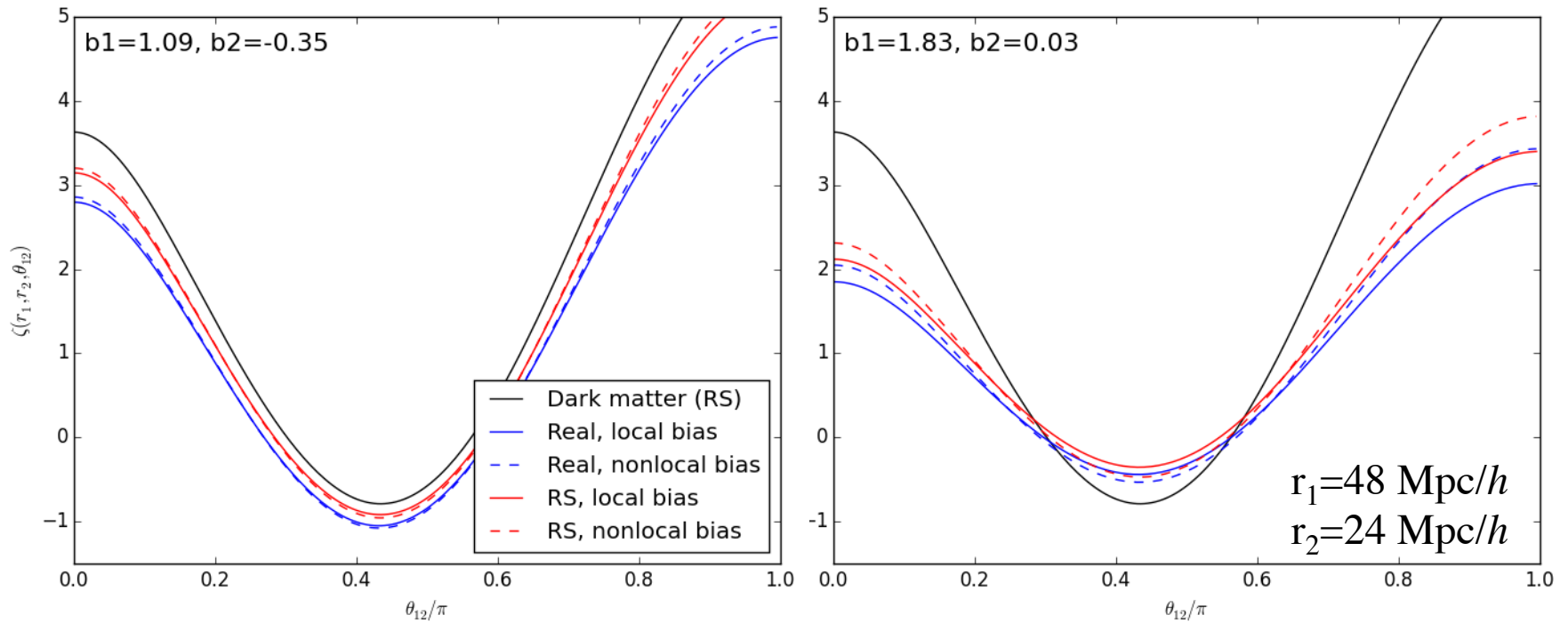
$$\mathbf{s} = \mathbf{x}(\mathbf{q}) + f(\Psi(\mathbf{q}) \cdot \hat{n}) \hat{n}$$

$$\delta_{x,g}(\mathbf{x}, t) = b_1 \delta(\mathbf{x}, t) + \frac{b_2}{2} (\delta^2(\mathbf{x}, t) - \sigma^2) \\ \left[+ \frac{b_{s^2}}{2} (s^2(\mathbf{x}, t) - \langle s^2 \rangle) \right]$$

Results: Dark Matter



Results: Biased tracers



Summary & Future Work

- Usual 2-point statistics of the matter density field do not capture the full cosmological information
- Log/gaussianization transform accesses non-Gaussian information in galaxy density fields:
 - Decouples clustering information and tracer bias
 - May not be as effective in redshift space
- Higher-point statistics also access non-Gaussian information
 - Bias and RSD must be included in analytic models
 - Will test configuration-space model against N-body simulations
 - Possibilities for extending model beyond tree-level PT, including Fingers of God, etc

Thank you!