

# Double scattering contribution to small-x processes: Mueller-Navelet Jets at the LHC

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in collaboration with

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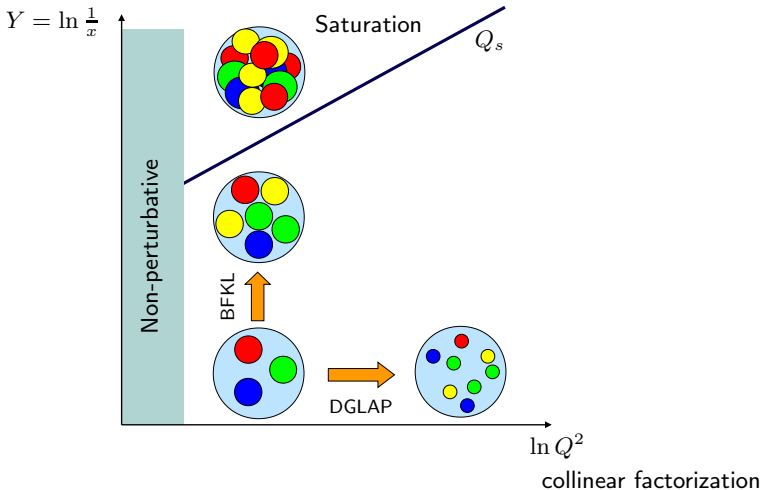
Based on:

D. Colferai, F. Schwennsen, L. Sz., S. Wallon: JHEP 1012 (2010) 026

B. Ducloué, L. Sz., S. Wallon: JHEP 1305 (2013) 096, Phys. Rev. Lett. 112 (2014) 082003  
Phys. Lett. B738 (2014) 311, Phys. Rev. D 92 (2015) 7, 076002

# The different regimes of QCD

$k_T$  factorization



# The specific case of QCD at large $s$

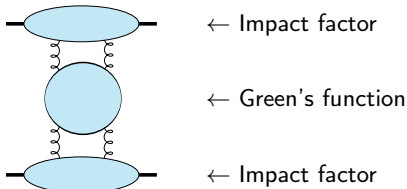
## QCD in the perturbative Regge limit

The amplitude can be written as:

$k_T$  factorization

$$\mathcal{A} = \begin{array}{c} \text{Diagram 1} \\ \sim s \end{array} + \left( \begin{array}{c} \text{Diagram 2} + \text{Diagram 3} + \dots \\ \sim s (\alpha_s \ln s) \end{array} \right) + \left( \begin{array}{c} \text{Diagram 4} + \dots \\ \sim s (\alpha_s \ln s)^2 \end{array} \right) + \dots$$

this can be put in the following form :



$$\sigma_{tot}^{h_1 h_2 \rightarrow anything} = \frac{1}{s} \text{Im} \mathcal{A} \sim s^{\alpha_{\mathbb{P}}(0)-1}$$

with  $\alpha_{\mathbb{P}}(0) - 1 = C \alpha_s + C' \alpha_s^2 + \dots$

$C > 0$  : Leading Log Pomeron

Balitsky, Fadin, Kuraev, Lipatov

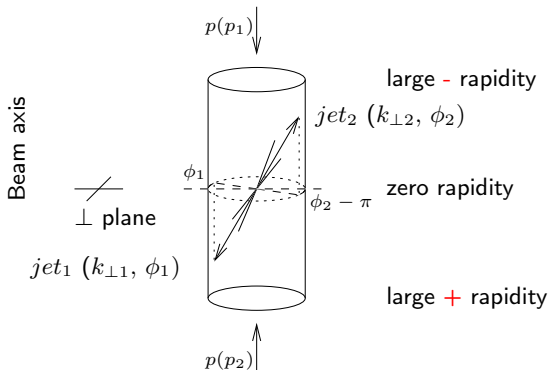
# Higher order corrections

- Higher order corrections to BFKL kernel are known at NLL order (Lipatov Fadin; Camici, Ciafaloni), now for arbitrary impact parameter  $\alpha_S \sum_n (\alpha_S \ln s)^n$  resummation
- impact factors are known in some cases at NLL
  - $\gamma^* \rightarrow \gamma^*$  at  $t = 0$  (Bartels, Colferai, Gieseke, Kyrieleis, Qiao; Balitski, Chirilli)
  - forward jet production (Bartels, Colferai, Vacca; Caporale, Ivanov, Murdaca, Papa, Perri; Chachamis, Hentschinski, Madrigal, Sabio Vera)
  - inclusive production of a pair of hadrons separated by a large interval of rapidity (Ivanov, Papa)
  - $\gamma_L^* \rightarrow \rho_L$  in the forward limit (Ivanov, Kotsky, Papa)

# Mueller-Navelet jets: Basics

## Mueller-Navelet jets

- Consider two jets (hadrons flying within a narrow cone) **separated by a large rapidity**, i.e. each of them almost fly in the direction of the hadron "close" to it, and with very similar transverse momenta
- Pure LO *collinear* treatment: these two jets should be emitted **back to back** at leading order:  $\Delta\phi - \pi = 0$  ( $\Delta\phi = \phi_1 - \phi_2 =$  relative azimuthal angle) and  $k_{\perp 1} = k_{\perp 2}$ . No phase space for (untagged) emission between them



# Master formulas

## $k_T$ -factorized differential cross section

$$\frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}} = \int d\phi_{J1} d\phi_{J2} \int d^2\mathbf{k}_1 d^2\mathbf{k}_2$$

$$\times \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1)$$

$$\times G(\mathbf{k}_1, \mathbf{k}_2, \hat{s})$$

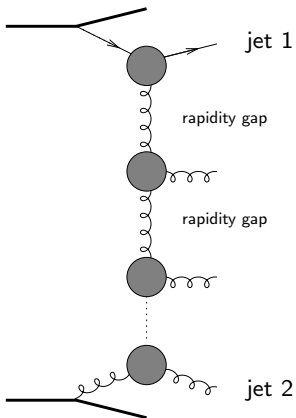
$$\times \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$

with  $\Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2) = \int dx_2 f(x_2) V(\mathbf{k}_2, x_2)$   $f \equiv \text{PDF}$   $x_J = \frac{|\mathbf{k}_J|}{\sqrt{s}} e^{y_J}$

# Mueller-Navelet jets: LL vs NLL

## LL BFKL

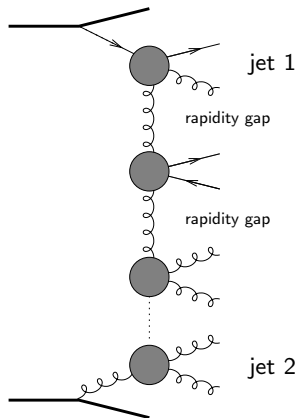
tree eff. vertex



$$\sum (\alpha_s \ln s)^n$$

## NLL BFKL

eff. vertex with 1-loop



$$\sum (\alpha_s \ln s)^n + \alpha_s \sum (\alpha_s \ln s)^n$$

# Results

## Results for a symmetric configuration

In the following we show results for

- $\sqrt{s} = 7 \text{ TeV}$
- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $0 < |y_1|, |y_2| < 4.7$

These cuts allow us to compare our predictions with the first experimental data on azimuthal correlations of **Mueller-Navelet** jets at the LHC presented by the **CMS** collaboration (CMS-PAS-FSQ-12-002)

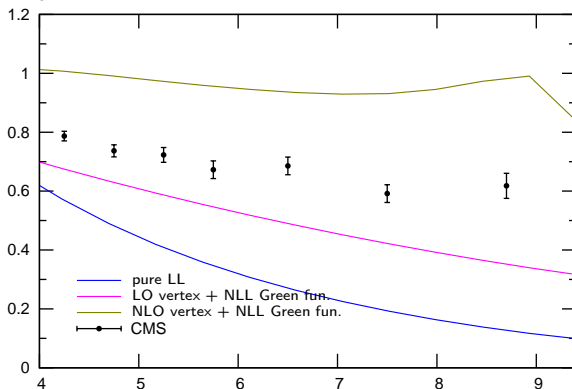
note: unlike experiments we have to set an upper cut on  $|\mathbf{k}_{J1}|$  and  $|\mathbf{k}_{J2}|$ . We have checked that our results do not depend on this cut significantly.



# Results: azimuthal correlations

## Azimuthal correlation $\langle \cos \varphi \rangle$

$$\frac{c_1}{c_0} = \langle \cos \varphi \rangle \equiv \langle \cos(\phi_{J1} - \phi_{J2} - \pi) \rangle$$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

$Y \equiv |y_1 - y_2|$

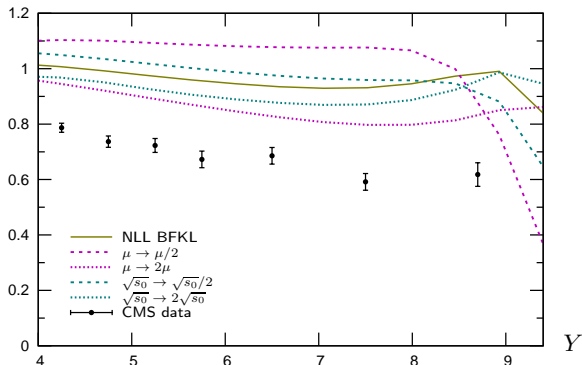
The NLO corrections to the jet vertex lead to a large increase of the correlation

Note: LO vertex + NLL Green done by F. Schwennsen, A. Sabio-Vera; C. Marquet, C. Royon

# Results: azimuthal correlations

## Azimuthal correlation $\langle \cos \varphi \rangle$

$$\langle \cos \varphi \rangle \equiv \langle \cos(\phi_{J1} - \phi_{J2} - \pi) \rangle$$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

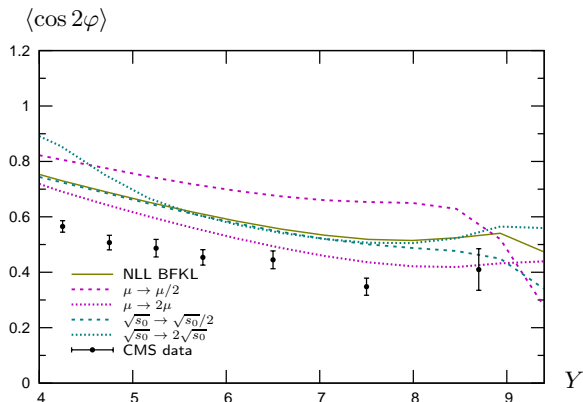
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

- NLL BFKL predicts a too small decorrelation
- The NLL BFKL calculation is still rather dependent on the scales, especially the renormalization / factorization scale

# Results: azimuthal correlations

## Azimuthal correlation $\langle \cos 2\varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

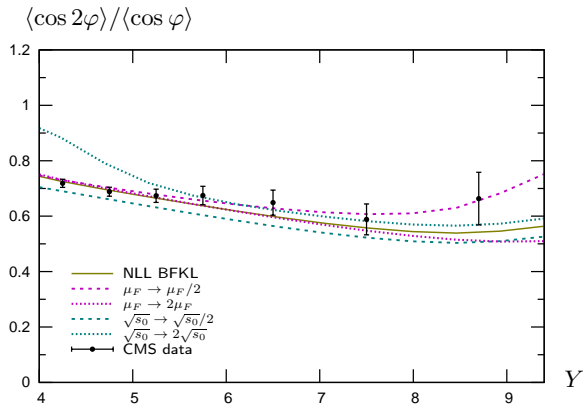
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

- The agreement with data is a little better for  $\langle \cos 2\varphi \rangle$  but still not very good
- This observable is also very sensitive to the scales

# Results: azimuthal correlations

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

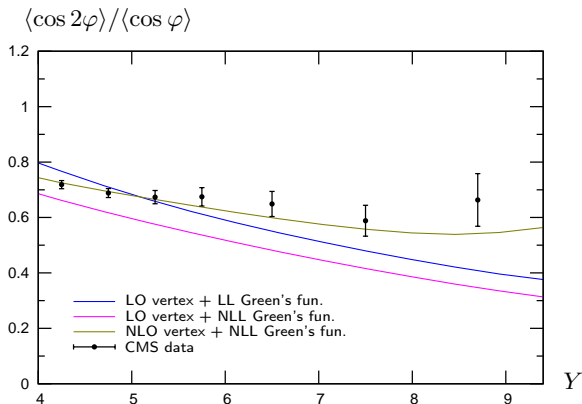
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

- This observable is more stable with respect to the scales than the previous ones
- The agreement with data is good across the whole  $Y$  range

# Results: azimuthal correlations

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

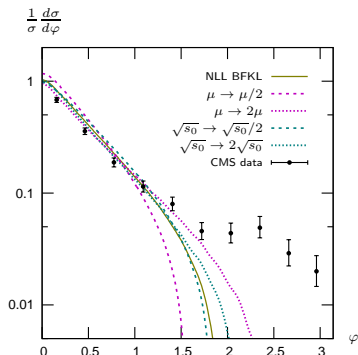
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

It is necessary to include the NLO corrections to the jet vertex to reproduce the behavior of the data at large  $Y$

# Results: azimuthal distribution

Azimuthal distribution (integrated over  $6 < Y < 9.4$ )



$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}.$$

- Our calculation predicts a too large value of  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  for  $\varphi \lesssim \frac{\pi}{2}$  and a too small value for  $\varphi \gtrsim \frac{\pi}{2}$
- It is not possible to describe the data even when varying the scales by a factor of 2

## Results: limitations

- The agreement of our calculation with the data for  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is good and quite stable with respect to the scales
- The agreement for  $\langle \cos n\varphi \rangle$  and  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  is not very good and very sensitive to the choice of the renormalization scale  $\mu_R$
- An all-order calculation would be independent of the choice of  $\mu_R$ . This feature is lost if we truncate the perturbative series  
 $\Rightarrow$  How to choose the renormalization scale?  
 'Natural scale': sometimes the typical momenta in a loop diagram are different from the natural scale of the process

We decided to use the **Brodsky-Lepage-Mackenzie (BLM)** procedure to fix the renormalization scale

# The BLM renormalization scale fixing procedure

The **Brodsky-Lepage-Mackenzie (BLM)** procedure resums the self-energy corrections to the gluon propagator at one loop into the running coupling.

First attempts to apply **BLM** scale fixing to **BFKL** processes lead to problematic results. **Brodsky, Fadin, Kim, Lipatov and Pivovarov** suggested that one should first go to a physical renormalization scheme like MOM and then apply the 'traditional' **BLM** procedure, i.e. identify the  $\beta_0$  dependent part and choose  $\mu_R$  such that it vanishes.

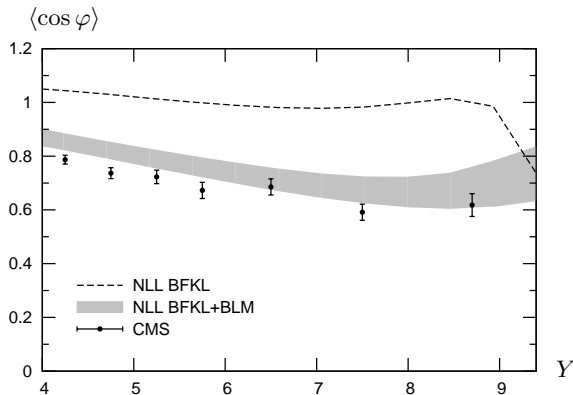
We followed this prescription for the full amplitude at NLL.

BLM procedure globally increases  $\mu$  of  $\alpha_s(\mu)$



# Results with BLM

## Azimuthal correlation $\langle \cos \varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

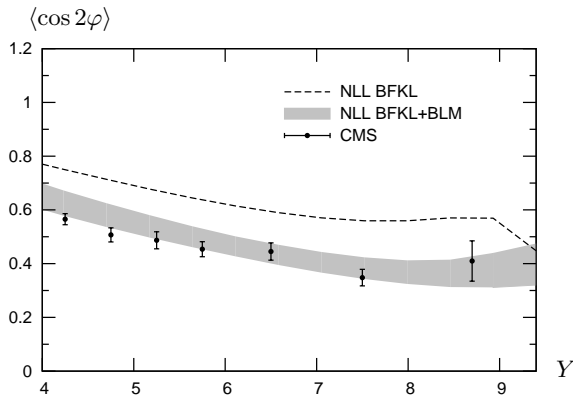
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

Using the BLM scale setting, the agreement with data becomes much better

# Results with BLM

## Azimuthal correlation $\langle \cos 2\varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

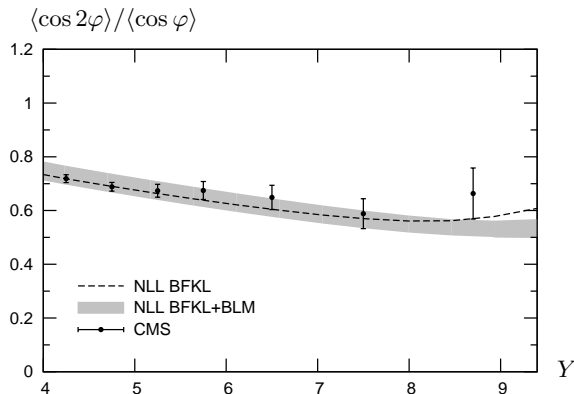
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

Using the BLM scale setting, the agreement with data becomes much better.

# Results with BLM

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$

$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$

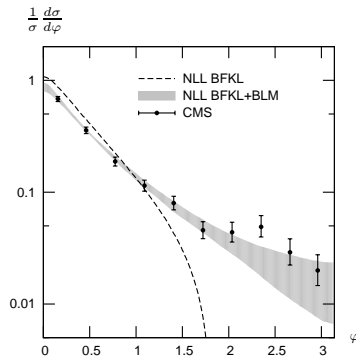
$0 < |y_1| < 4.7$

$0 < |y_2| < 4.7$

Because it is much less dependent on the scales, the observable  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is almost not affected by the BLM procedure and is still in good agreement with the data.

# Results with BLM

Azimuthal distribution (integrated over  $6 < Y < 9.4$ )



With the BLM scale setting the azimuthal distribution is in good agreement with the data across the full  $\varphi$  range.

## Comparison with fixed-order

Using the **BLM** scale setting:

- The agreement  $\langle \cos n\varphi \rangle$  with the data becomes much better
- The agreement for  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is still good and unchanged as this observable is weakly dependent on  $\mu_R$
- The azimuthal distribution is in much better agreement with the data

But the configuration chosen by **CMS** with  $\mathbf{k}_{J_{\min 1}} = \mathbf{k}_{J_{\min 2}}$  does not allow us to compare with a **fixed-order**  $\mathcal{O}(\alpha_s^3)$  treatment (i.e. without resummation)

- These calculations are unstable when  $\mathbf{k}_{J_{\min 1}} = \mathbf{k}_{J_{\min 2}}$  because the cancellation of some divergencies is difficult to obtain numerically
  - Presumably, resummation effects à la **Sudakov** could be important in the limit  $\mathbf{k}_{J_1} \simeq \mathbf{k}_{J_2}$  and require a special treatment
- Work in progress in collaboration with **A. H. Mueller, B-W. Xiao, F. Yuan**

## Comparison with fixed-order

### Results for an asymmetric configuration

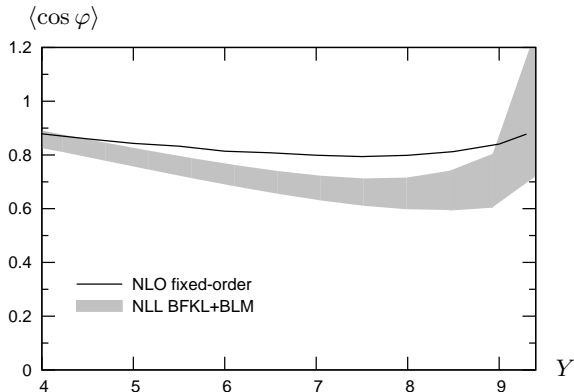
In this section we choose the cuts as

- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
- $0 < |y_1|, |y_2| < 4.7$

and we compare our results with the NLO fixed-order code Dijet (Aurenche, Basu, Fontannaz) in the same configuration

# Comparison with fixed-order

Azimuthal correlation  $\langle \cos \varphi \rangle$

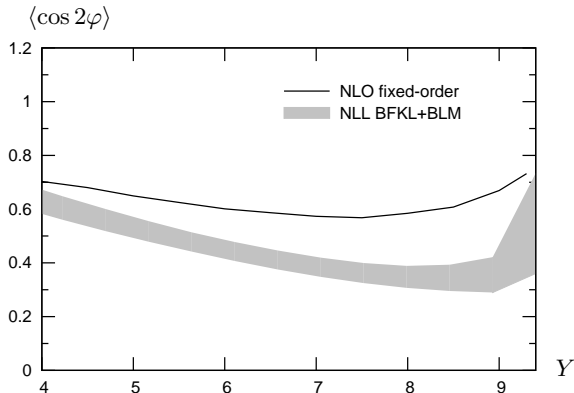


$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < |y_1| < 4.7$   
 $0 < |y_2| < 4.7$

The NLO fixed-order and NLL BFKL+BLM calculations are very close

# Comparison with fixed-order

## Azimuthal correlation $\langle \cos 2\varphi \rangle$



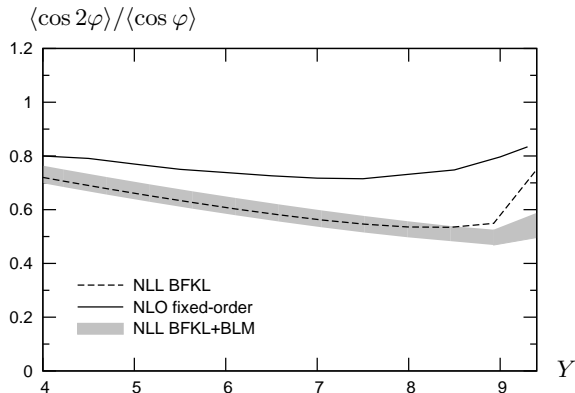
$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < |y_1| < 4.7$   
 $0 < |y_2| < 4.7$

The BLM procedure leads to a sizable difference between NLO fixed-order and NLL BFKL+BLM.



# Comparison with fixed-order

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$

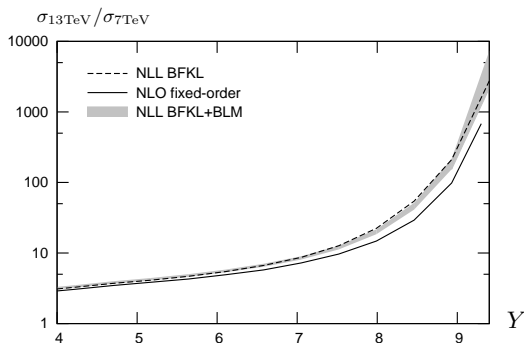


$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < |y_1| < 4.7$   
 $0 < |y_2| < 4.7$

Using **BLM** or not, there is a **sizable difference** between **BFKL** and fixed-order.

# Comparison with fixed-order

Cross section: 13 TeV vs. 7 TeV



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < |y_1| < 4.7$   
 $0 < |y_2| < 4.7$

- In a BFKL treatment, a strong rise of the cross section with increasing energy is expected.
- This rise is faster than in a fixed-order treatment

# Energy-momentum conservation

- It is necessary to have  $\mathbf{k}_{J_{\min 1}} \neq \mathbf{k}_{J_{\min 2}}$  for comparison with fixed order calculations but this can be problematic for **BFKL** because of energy-momentum conservation
- There is no strict energy-momentum conservation in **BFKL**
- This was studied at LO by **Del Duca and Schmidt**. They introduced an effective rapidity  $Y_{\text{eff}}$  defined as

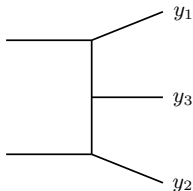
$$Y_{\text{eff}} \equiv Y \frac{\sigma^{2 \rightarrow 3}}{\sigma^{\text{BFKL}, \mathcal{O}(\alpha_s^3)}}$$

- When one replaces  $Y$  by  $Y_{\text{eff}}$  in the expression of  $\sigma^{\text{BFKL}}$  and truncates to  $\mathcal{O}(\alpha_s^3)$ , the exact  $2 \rightarrow 3$  result is obtained

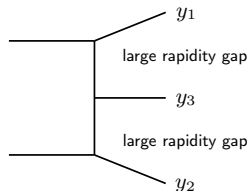
# Energy-momentum conservation

We follow the idea of [Del Duca and Schmidt](#), adding the NLO jet vertex contribution:

exact  $2 \rightarrow 3$

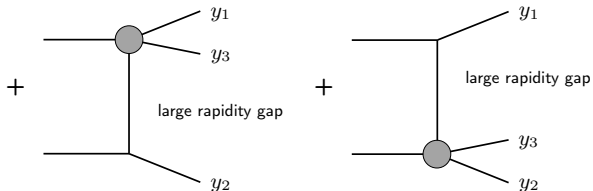


BFKL



one emission from the Green's function + LO jet vertex

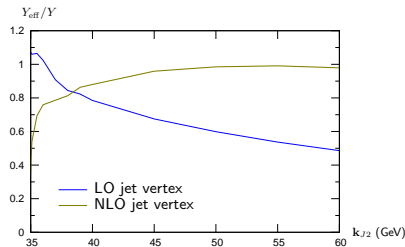
we have to take into account these additional  $\mathcal{O}(\alpha_s^3)$  contributions:



no emission from the Green's function + NLO jet vertex

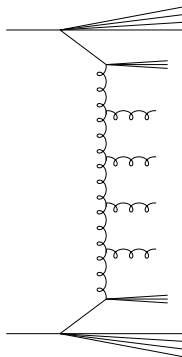
# Energy-momentum conservation

Variation of  $Y_{\text{eff}}/Y$  as a function of  $k_{J2}$  for fixed  $k_{J1} = 35$  GeV (with  $\sqrt{s} = 7$  TeV,  $Y = 8$ ):



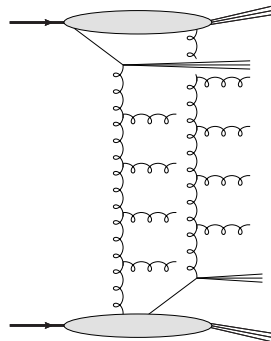
- With the **LO** jet vertex,  $Y_{\text{eff}}$  is much smaller than  $Y$  when  $k_{J1}$  and  $k_{J2}$  are significantly different
- This is the region important for comparison with fixed order calculations
- The improvement coming from the **NLO** jet vertex is very large in this region
- For  $k_{J1} = 35$  GeV and  $k_{J2} = 50$  GeV, typical of the values we used for comparison with fixed order, we get  $\frac{Y_{\text{eff}}}{Y} \simeq 0.98$  at NLO vs.  $\sim 0.6$  at LO

# Can Mueller-Navelet jets be a manifestation of multiparton interactions?



MN jets in the single partonic model

+

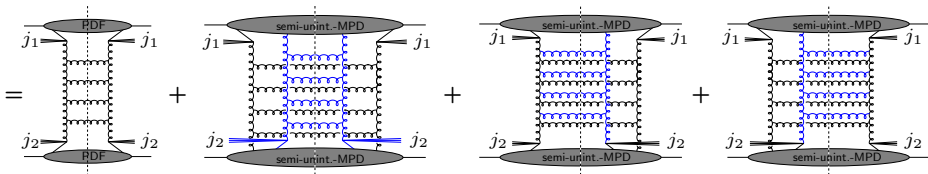
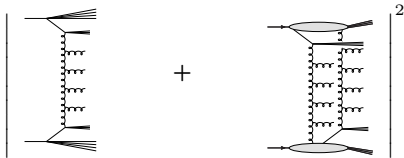


MN jets in MPI

here MPI = DPS (double parton scattering)

see also Maciula and Szczurek Phys. Rev. D90

# Can Mueller-Navelet jets be a manifestation of multiparton interactions?



single  $\mathbb{P}$  ladder

two  $\mathbb{P}$  ladders

interferences

scaling:  $s^{\alpha_{\mathbb{P}}}$

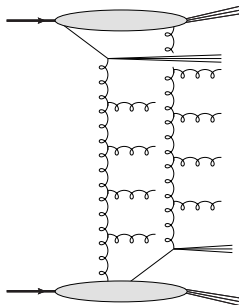
$(??) s^{2\alpha_{\mathbb{P}}}$

??

- The twist counting is not easy for MPI kinds of contributions at small  $x$
- $k_{\perp 1,2}$  are not integrated  $\Rightarrow$  MPI may be competitive, and enhanced by small- $x$  resummation
- Interference terms are not governed by **BJKP** (this is not a fully interacting 3-reggeons system) (for **BJKP**,  $\alpha_{\mathbb{P}} < 1 \Rightarrow$  suppressed)

## A phenomenological test: the problem

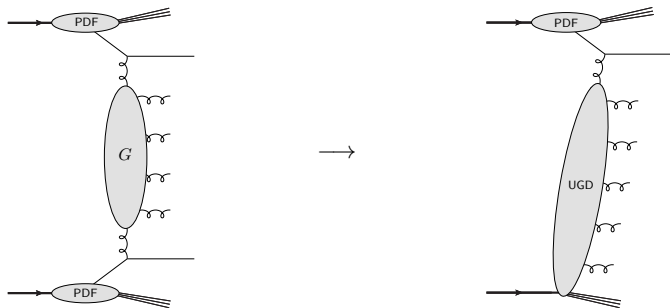
- Simplification: we neglect any interference contribution between the two mechanisms
- How to evaluate the DPS contribution?



- This would require some kind of "hybrid" double parton distributions, with
  - one collinear parton
  - one off-shell parton (with some  $k_{\perp}$ )
- Almost nothing is known on such distributions



## A phenomenological test: our ansatz



Mueller-Navelet jets production at LL accuracy

Inclusive forward jet production

Factorized ansatz for the DPS contribution:

$$\sigma_{\text{DPS}} = \frac{\sigma_{\text{fwd}} \sigma_{\text{bwd}}}{\sigma_{\text{eff}}}$$

Tevatron, LHC:  $\sigma_{\text{eff}} \simeq 15 \text{ mb}$

To account for some discrepancy between various measurements, we take

$$\sigma_{\text{eff}} \simeq 10 - 20 \text{ mb}$$

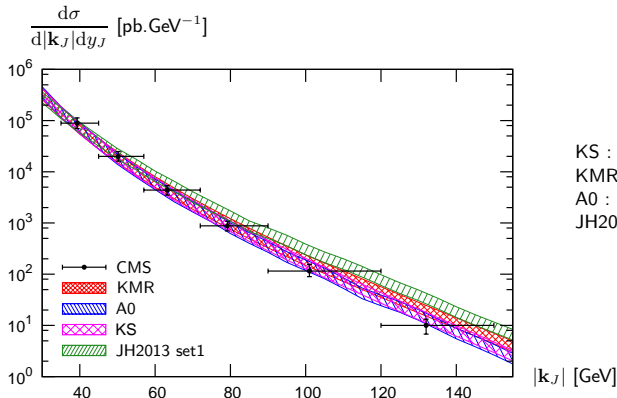
$$\mathcal{F}_g \left( \frac{\mathbf{k}_J^2}{s x_J}, |\mathbf{k}_J| \right)$$
$$\int d\mathbf{k}^2 \mathcal{F}_g(x, |\mathbf{k}|) = x f_g(x) \text{ (usual PDF)}$$

inclusive forward jet cross-section:

$$\frac{d\sigma}{d|\mathbf{k}_J|dy_J} = K \frac{\alpha_s}{|\mathbf{k}_J|} x_J (C_F f_q(x_J) + C_A f_g(x_J)) \mathcal{F}_g \left( \frac{\mathbf{k}_J^2}{s x_J}, |\mathbf{k}_J| \right)$$

## A phenomenological test

- We use **CMS** data at  $\sqrt{s} = 7$  TeV,  $3.2 < |y_J| < 4.7$
- We use various parametrization for the UGD
- For each parametrization we determine the range of  $K$  compatible with the **CMS** measurement in the lowest transverse momentum bin



	$K_{min}$	$K_{max}$
KS :	1.20	1.94
KMR :	1.05	1.69
A0 :	4.27	6.89
JH2013 :	2.44	3.94

$$\frac{d\sigma}{d|\mathbf{k}_J| dy_J} = K \frac{\alpha_s}{|\mathbf{k}_J|} x_J (C_F f_q(x_J) + C_A f_g(x_J)) \mathcal{F}_g \left( \frac{\mathbf{k}_J^2}{s x_J}, |\mathbf{k}_J| \right)$$

## SPS vs DPS: Results

## symmetric configuration

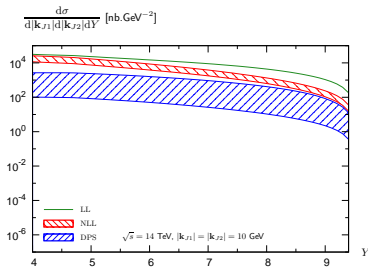
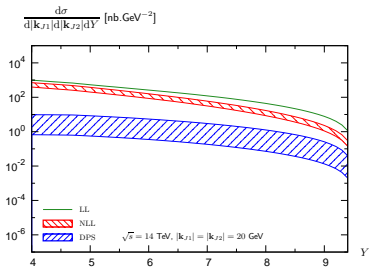
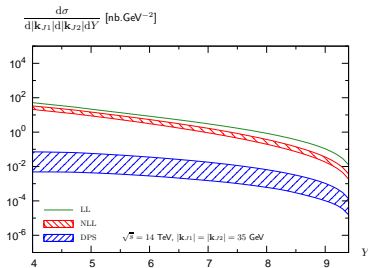
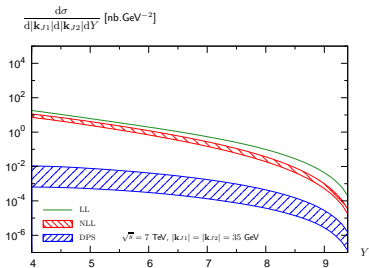
We will focus on four choices of kinematical cuts:

- $\sqrt{s} = 7 \text{ TeV}$ ,  $|\mathbf{k}_{J1}| = |\mathbf{k}_{J2}| = 35 \text{ GeV}$ ,  
(like in the CMS analysis for azimuthal correlations of MN jets)
- $\sqrt{s} = 14 \text{ TeV}$ ,  $|\mathbf{k}_{J1}| = |\mathbf{k}_{J2}| = 35 \text{ GeV}$ ,
- $\sqrt{s} = 14 \text{ TeV}$ ,  $|\mathbf{k}_{J1}| = |\mathbf{k}_{J2}| = 20 \text{ GeV}$ ,
- $\sqrt{s} = 14 \text{ TeV}$ ,  $|\mathbf{k}_{J1}| = |\mathbf{k}_{J2}| = 10 \text{ GeV}$  ← highest DPS effect expected

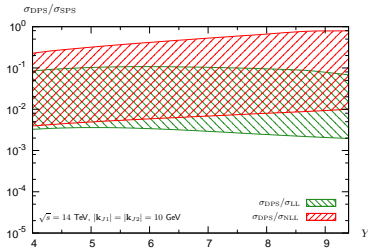
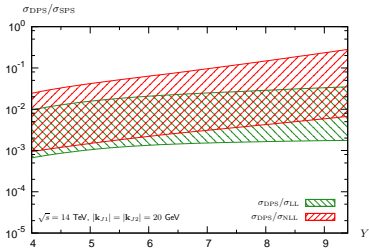
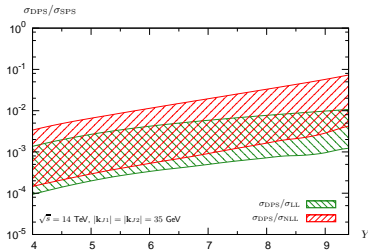
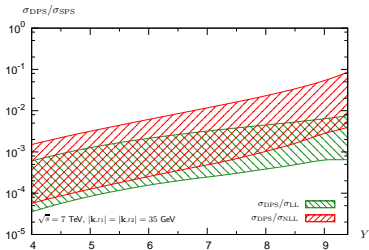
parameters:

- $0 < y_{J,1} < 4.7$  and  $-4.7 < y_{J,2} < 0$
- MSTW 2008 parametrization for PDFs
- In the case of the NLL BFKL calculation, anti- $k_t$  jet algorithm with  $R = 0.5$ .

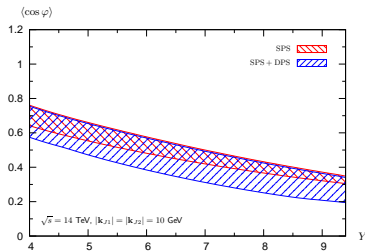
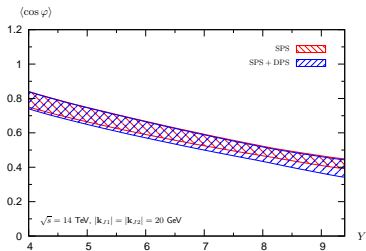
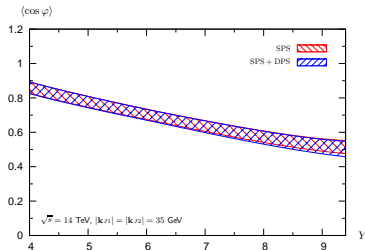
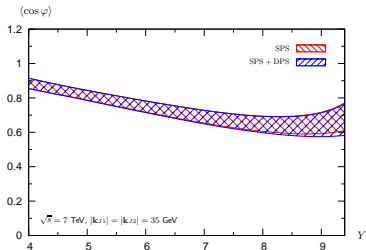
# SPS vs DPS: cross-sections



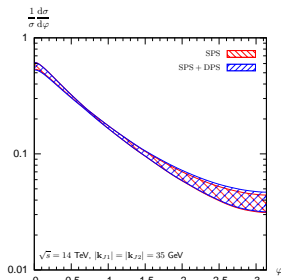
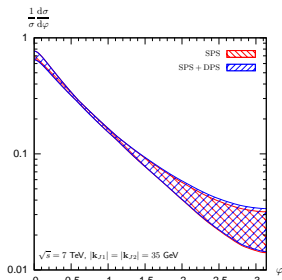
# SPS vs DPS: cross-sections (ratios)



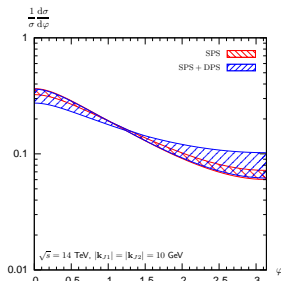
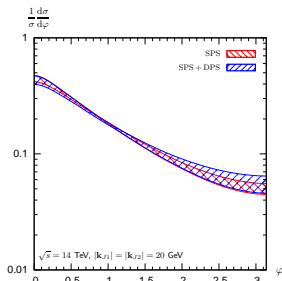
# SPS vs DPS: Azimuthal correlations



# SPS vs DPS: Azimuthal distributions

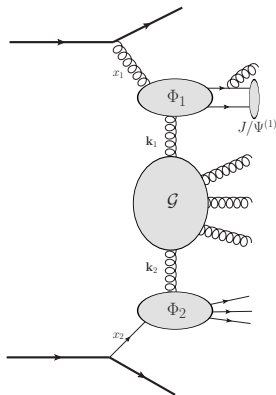


$$8 < Y < 9.4$$



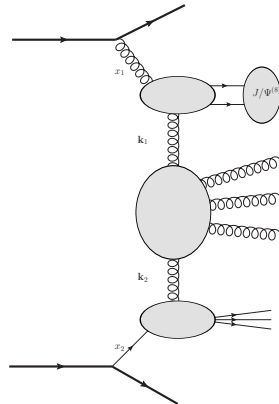


# Inclusive production of a forward $J/\psi$ + a backward jet



Color singlet mechanism

- Hard scales:  $k_J$  and  $M_{J/\psi}$
- Very promising at **ATLAS** (and **CMS**?)
- To be studied: cross-section study and azimuthal correlation



Color octet mechanism

Work in progress with LO vertex + NLO BFKL Green function

R. Boussarie, B. Ducloué, L. Sz., S. Wallon.

# Conclusions

- We studied **Mueller-Navelet** jets at full (vertex + Green's function) **NLL BFKL** accuracy and compared our results with the first data from the **LHC**
- The agreement with **CMS** data at 7 TeV is greatly improved by using the **BLM** scale fixing procedure
- $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is almost not affected by **BLM** and shows a clear difference between **NLO fixed-order** and **NLL BFKL** in an **asymmetric configuration**
- **Energy-momentum conservation** seems to be less severely violated with the NLO jet vertex
- We studied the effect of DPS contributions which could mimic the MN jet
  - For **cross-sections**: The uncertainty on DPS is very large. Still,  $\sigma_{DPS} < \sigma_{SPS}$  in the **LHC** kinematics
  - For **angular correlations**: including DPS **does not significantly modify our NLL BFKL prediction**
  - For low  $k_T$  and large  $Y$ , **the effect of DPS can become larger than the uncertainty on the NLL BFKL calculation.** One should focus on this region experimentally.

THANK YOU FOR YOUR ATTENTION

Backup

for  $k_{J1} = k_{J2} = 35\text{GeV}$

$$Y \sim 4 \quad \frac{G}{Q} \approx 5 - 10$$

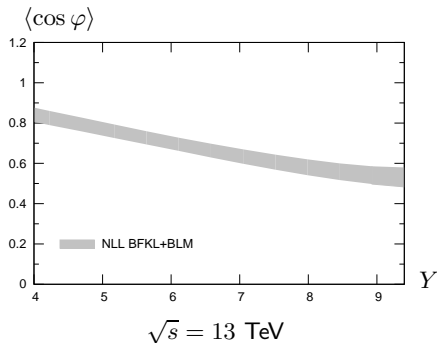
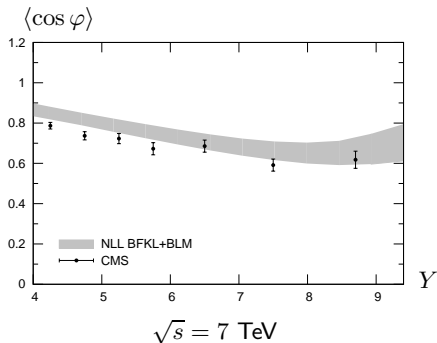
$$Y \sim 7 \quad \frac{G}{Q} \approx 1$$

$$Y \sim 9 \quad \frac{Q}{G} \approx 5 - 10$$

large  $Y$  large  $x$

## Comparison: 13 TeV vs. 7 TeV

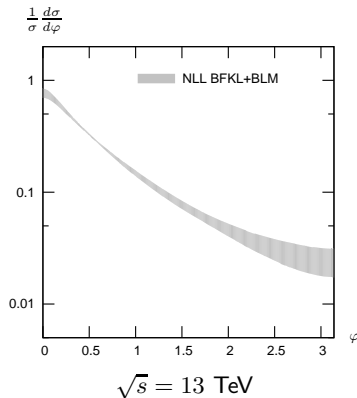
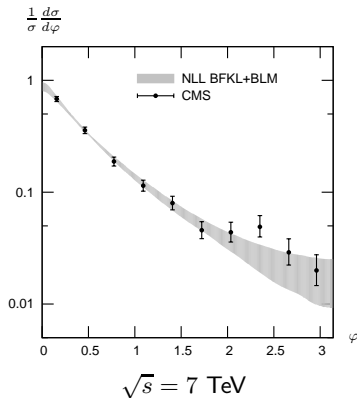
Azimuthal correlation  $\langle \cos \varphi \rangle$



The behavior is similar at 13 TeV and at 7 TeV

## Comparison: 13 TeV vs. 7 TeV

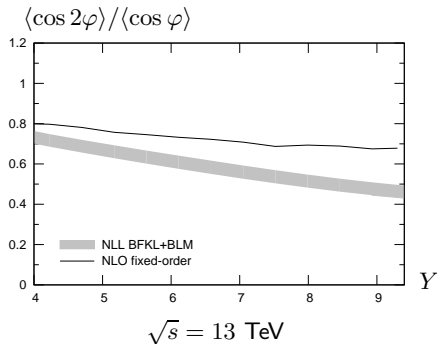
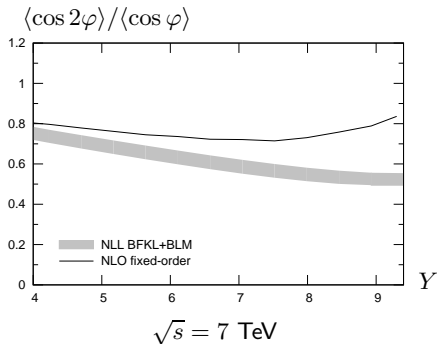
Azimuthal distribution (integrated over  $6 < Y < 9.4$ )



The behavior is similar at 13 TeV and at 7 TeV

## Comparison: 13 TeV vs. 7 TeV

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$   
(asymmetric configuration)

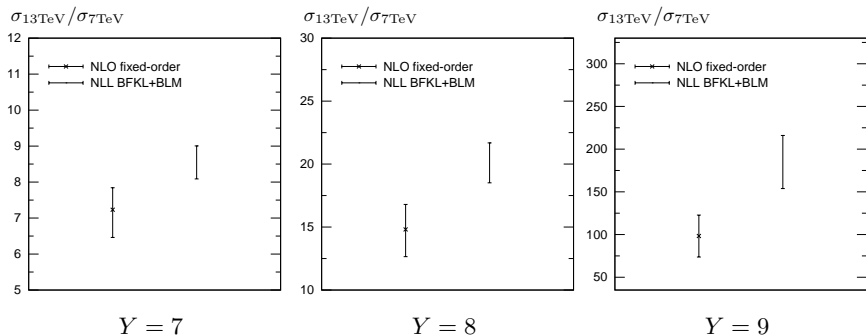


The difference between BFKL and fixed-order is smaller at 13 TeV than at 7 TeV



# Comparison: 13 TeV vs. 7 TeV

## Cross section



It is useful to define the coefficients  $\mathcal{C}_n$  as

$$\mathcal{C}_n \equiv \int d\phi_{J1} d\phi_{J2} \cos(n(\phi_{J1} - \phi_{J2} - \pi)) \\ \times \int d^2\mathbf{k}_1 d^2\mathbf{k}_2 \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1) G(\mathbf{k}_1, \mathbf{k}_2, \hat{s}) \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$

- $n = 0 \implies$  differential cross-section

$$\mathcal{C}_0 = \frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}}$$

- $n > 0 \implies$  azimuthal decorrelation

$$\frac{\mathcal{C}_n}{\mathcal{C}_0} = \langle \cos(n(\phi_{J,1} - \phi_{J,2} - \pi)) \rangle \equiv \langle \cos(n\varphi) \rangle$$

- sum over  $n \implies$  azimuthal distribution

$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}$$